

Paths, diffuse interreflections, caching and radiometry

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How we got here

- We want to render diffuse interreflections
 - strategy: compute approximation \hat{B} , then gather

$$B = E + (\rho\mathcal{K})E + (\rho\mathcal{K})(\hat{B} - E)$$

Exitance

Source term

mostly zero

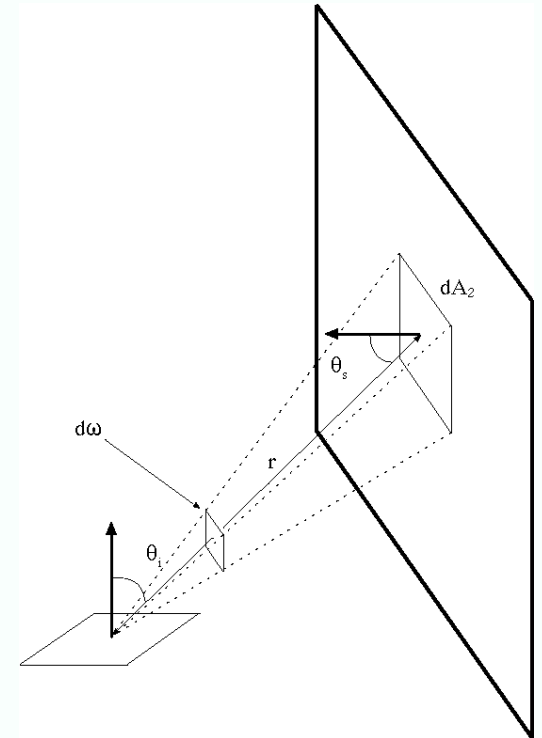
One or more bounces

Can change fast - shadows, etc.

Changes much more slowly, because \mathcal{K} smoothes,
so we should approximate this

Gathering

- We gather radiosity from B-hat
 - Here S is all the surfaces in the world



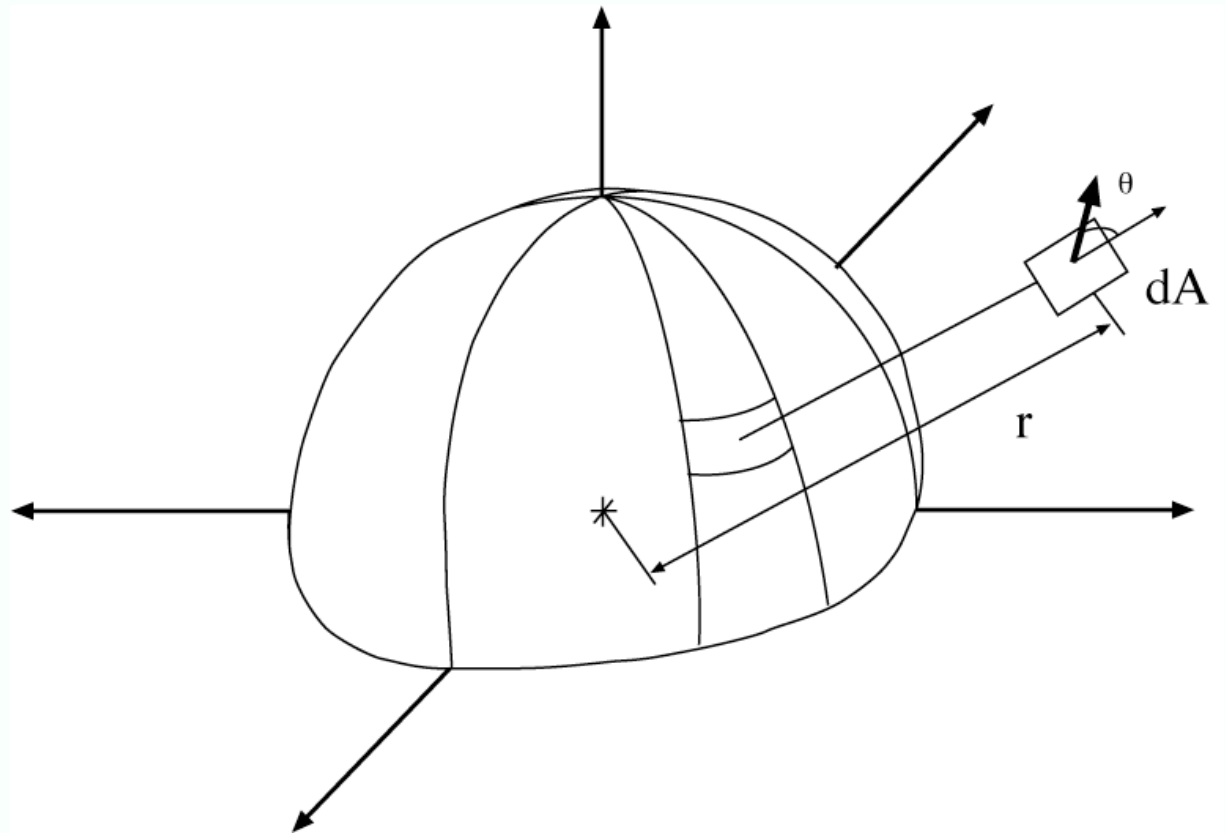
$$(\rho\mathcal{K})(\hat{B} - E) = \rho(\mathbf{x}) \int_S \frac{\cos \theta_i \cos \theta_s}{\pi r^2} Vis(\mathbf{x}, \mathbf{u})(\hat{B}(\mathbf{u}) - E(\mathbf{u}))dA_s$$

- Another integral
 - but not a good idea to integrate over dAs
 - too much area, too many samples
 - instead, integrate over hemisphere

Remember Solid Angle

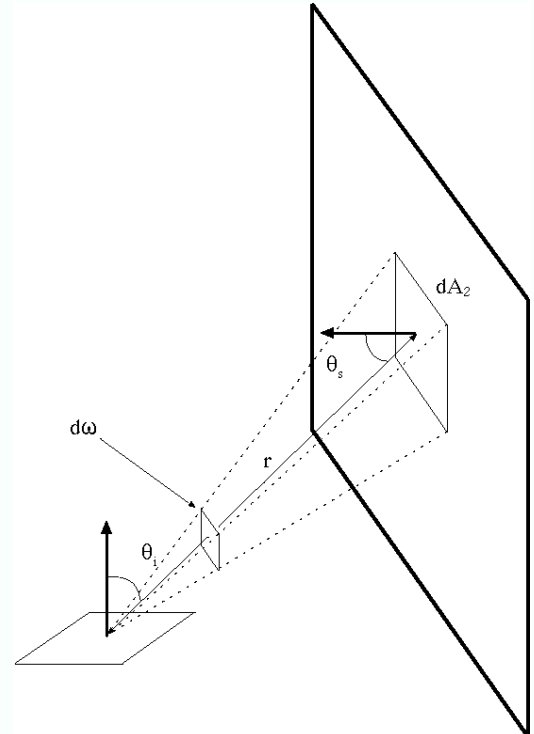
- By analogy with angle (in radians)
- The solid angle subtended by a patch area dA is given by

$$d\omega = \frac{dA \cos \vartheta}{r^2}$$



Changing variables

- Rather than integrate over all area, integrate over hemisphere
 - equivalently, integrate over solid angle



$$\frac{\cos \theta_s}{r^2} dA_s = d\omega_s$$

Changing variables

- Start with:

$$\rho(\mathbf{x}) \int_S \frac{\cos \theta_i \cos \theta_s}{\pi r^2} Vis(\mathbf{x}, \mathbf{u}) (\hat{B}(\mathbf{u}) - E(\mathbf{u})) dA_s$$

- Substitute:

$$\frac{\cos \theta_s}{r^2} dA_s = d\omega_s$$

- Get:

$$\rho(\mathbf{x}) \frac{1}{\pi} \int_{\Omega} \cos \theta_i (\hat{B}(\omega) - E(\omega)) d\omega$$

Value at far end of ray through angle

|

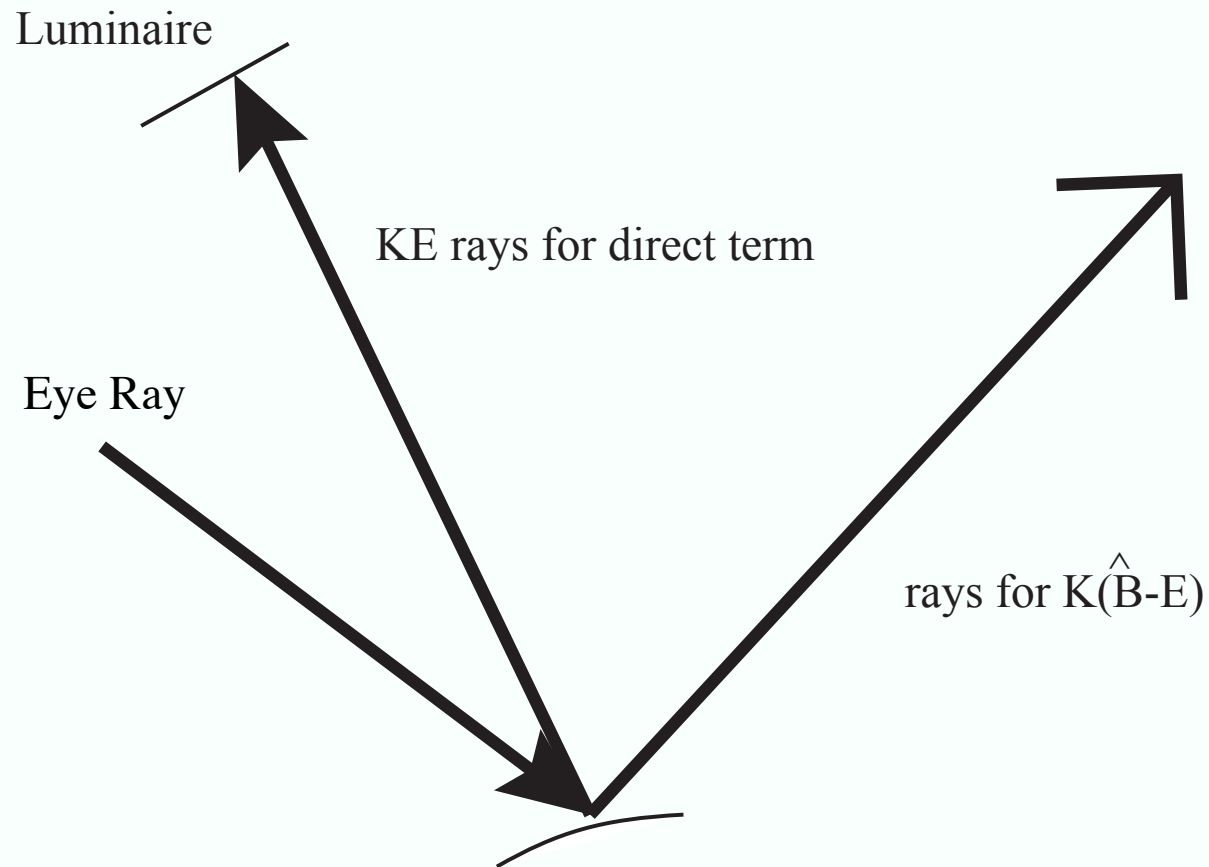
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Incoming hemisphere

Evaluating integral

- Procedure
 - Generate N uniform random samples on hemisphere
 - procedure described on whiteboard
 - Find $\hat{B-E}$ at far end of each ray
 - Average
- How big should N be?
 - Variance
 - estimate is a random variable, so must have variance
 - small N implies high variance, fast
 - large N implies low variance, slow
 - Variance will look like noise
 - but should be small, because the term is small
 - suggests small N is OK

Gathering from $\hat{B} - E$



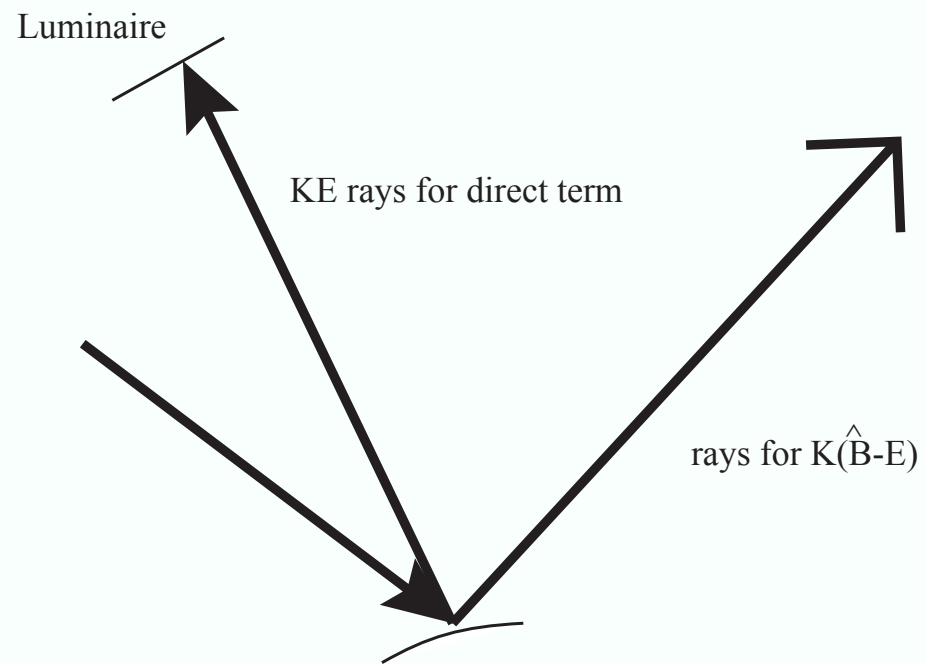
$$B = E + (\rho\mathcal{K})E + (\rho\mathcal{K})(\hat{B} - E)$$

Alternative: B-hat via random paths

- Notice that B-hat is also an integral
 - approximation to B
- Now from $B = E + (\rho\mathcal{K})B$
 - we expect $\hat{B} = E + (\rho\mathcal{K})\hat{B}$
 - so $\hat{B} - E = (\rho\mathcal{K})\hat{B}$
 - expand by substituting to get $\hat{B} = E + (\rho\mathcal{K})(E + (\rho\mathcal{K})\hat{B})$
 - ie $\hat{B} - E = (\rho\mathcal{K})(E + (\rho\mathcal{K})\hat{B})$
 - substitute from above to get $\hat{B} - E = (\rho\mathcal{K})E + (\rho\mathcal{K})(\hat{B} - E)$

Recursive evaluation

$$\text{shade}(x) = E(x) + \rho(x)\text{direct}(x) + \text{RKBME}(x)$$



Recursive evaluation: direct term

$$\text{direct}(x) = \sum_{l \in \text{luminares}} \text{directfromL}(x, l)$$

$\text{directfromL}(x, L)$

generate N uniform random samples u_i on luminaire L with area A_l
return $\frac{A_l}{N} \sum_i \frac{\cos \theta_x \cos \theta_u}{\pi r^2} E(u_i)$

We did this when we discussed area luminaires - no big mystery here

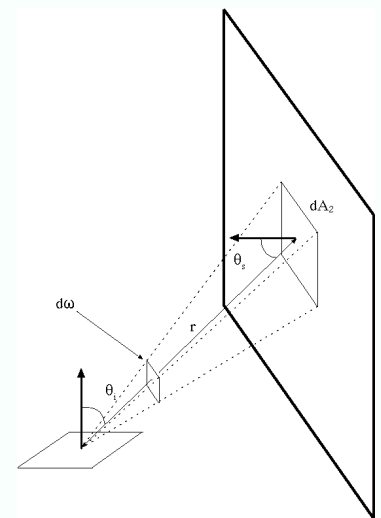
Recursive evaluation: Indirect term

This form isn't yet practical, because the recursion is infinite!

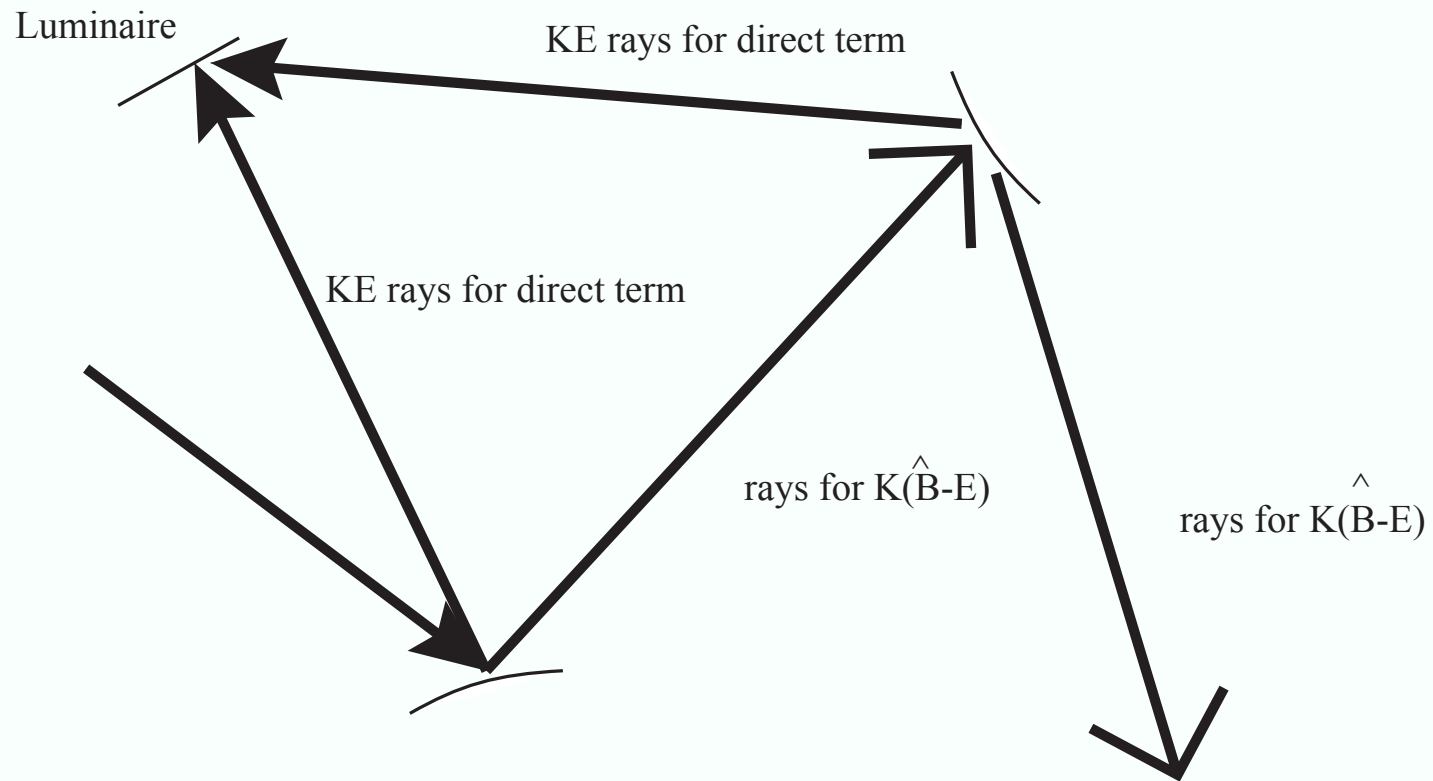
RKBME(x)

Generate M points p_i uniformly at random on unit hemisphere at x
For each point p_i , write u_i for the first hit on the ray from x to p_i
write $\cos \theta_{si}$ for the cosine at x of the i 'th direction

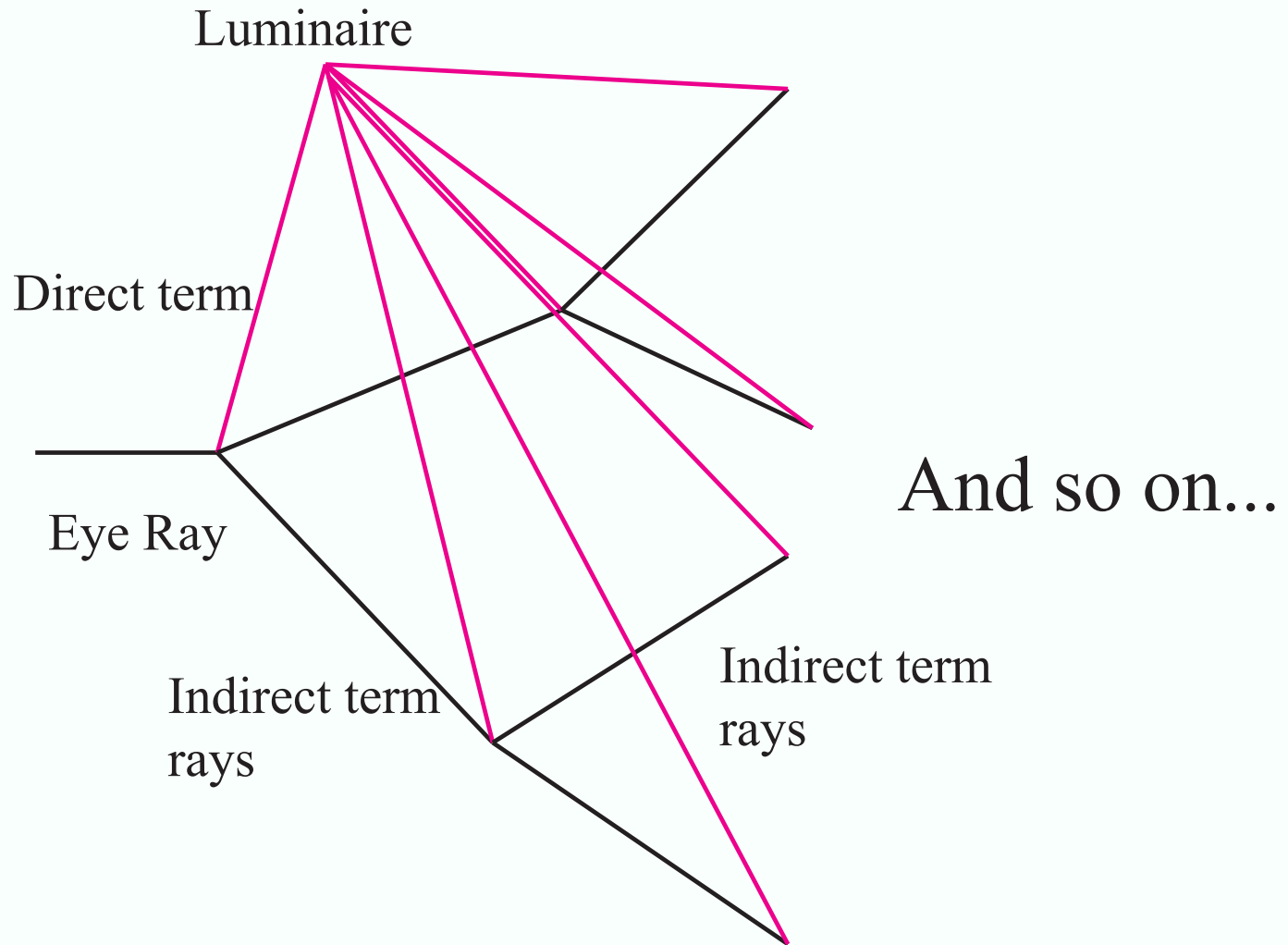
return $\rho(x)2\pi \frac{1}{\pi} \frac{1}{M} \sum_i (\rho(u_i)\text{direct}(u_i) + \text{RKBME}(u_i)) \cos \theta_{si}$



B-hat via random paths becomes a tree



B-hat via random paths becomes a tree



Recursive evaluation: Indirect term

Recursion no longer infinite, but estimate must be (very slightly) too small

RKBME(x , depth)

Generate M points p_i uniformly at random on unit hemisphere at x

For each point p_i , write u_i for the first hit on the ray from x to p_i

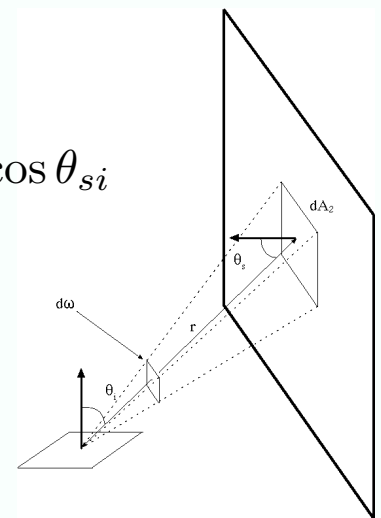
write $\cos \theta_{si}$ for the cosine at x of the i 'th direction

if depth==0

return 0

else

return $\rho(x)2\pi \frac{1}{\pi} \frac{1}{M} \sum_i (\rho(u_i)\text{direct}(u_i) + \text{RKBME}(u_i, \text{depth} - 1)) \cos \theta_{si}$



Recursive evaluation: Indirect term

Recursion no longer infinite, not as deep as previous,
but estimate must still be (very slightly) too small

RKBME(x , ρ_{acc})

Generate M points p_i uniformly at random on unit hemisphere at x

For each point p_i , write u_i for the first hit on the ray from x to p_i

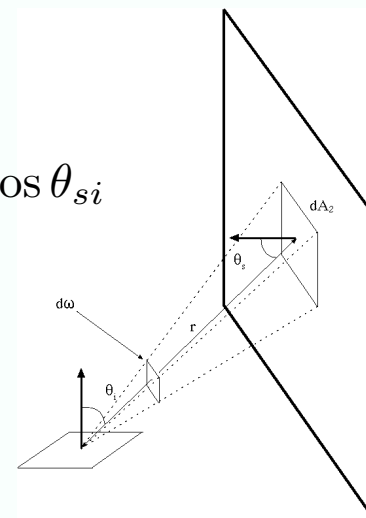
write $\cos \theta_{si}$ for the cosine at x of the i 'th direction

if $\rho_{acc} < \text{smallthresh}$

return 0

else

return $\rho(x) 2\pi \frac{1}{\pi} \frac{1}{M} \sum_i (\rho(u_i) \text{direct}(u_i) + \text{RKBME}(u_i, \rho(x) * \rho_{acc})) \cos \theta_{si}$



Russian roulette

- Consider a random process:
 - with probability p , return S
 - with probability $1-p$, return 0
- Expected value:
 - $p*S$
- We can use this to prune paths at random, mainly pruning when albedo is low

Russian roulette

Notice what's happened to the albedo term. When a path gets to low albedo surface, it has little chance of continuing. This is unbiased!

RKBME(x)

Generate v uniform random variable, $v \in [0, 1]$

if $v > \rho(x)$

return 0

else

Generate M points p_i uniformly at random on unit hemisphere at x

For each point p_i , write u_i for the first hit on the ray from x to p_i

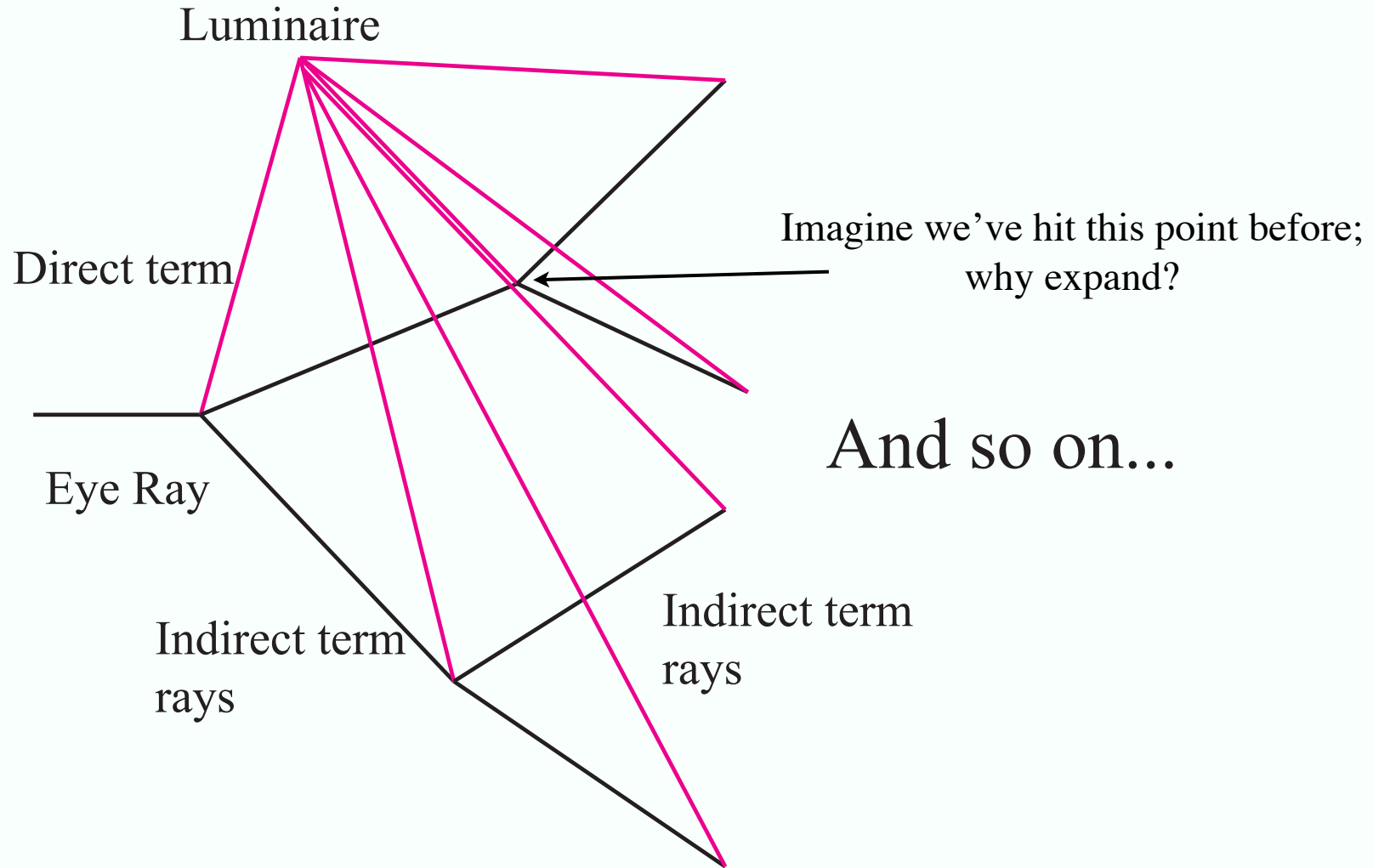
write $\cos \theta_{si}$ for the cosine at x of the i 'th direction

return $2\pi \frac{1}{\pi} \frac{1}{M} \sum_i (\rho(u_i) \text{direct}(u_i) + \text{RKBME}(u_i)) \cos \theta_{si}$

Light path analysis

- We've now done LD*E
 - russian roulette cleverly explores paths; if there's lots of albedo, paths tend to be long; else short.
 - russian roulette is a random process
 - random choice of directions; random choice to prune
 - unbiased
 - Expected value is the right answer
 - variance
 - because it's random
 - looks like image noise
 - seen this before in lenses, motion blur
 - control by
 - more rays (!)
 - caching
 - importance sampling (later)

Caching



Caching

RKBME(x)

Generate v uniform random variable, $v \in [0, 1]$

if $v > \rho(x)$

return 0

else

Interrogate cache - do we have an RKBME value close to x ?

if yes

return cache value

else

Generate M points p_i uniformly at random on unit hemisphere at x

For each point p_i , write u_i for the first hit on the ray from x to p_i

write $\cos \theta_{si}$ for the cosine at x of the i 'th direction

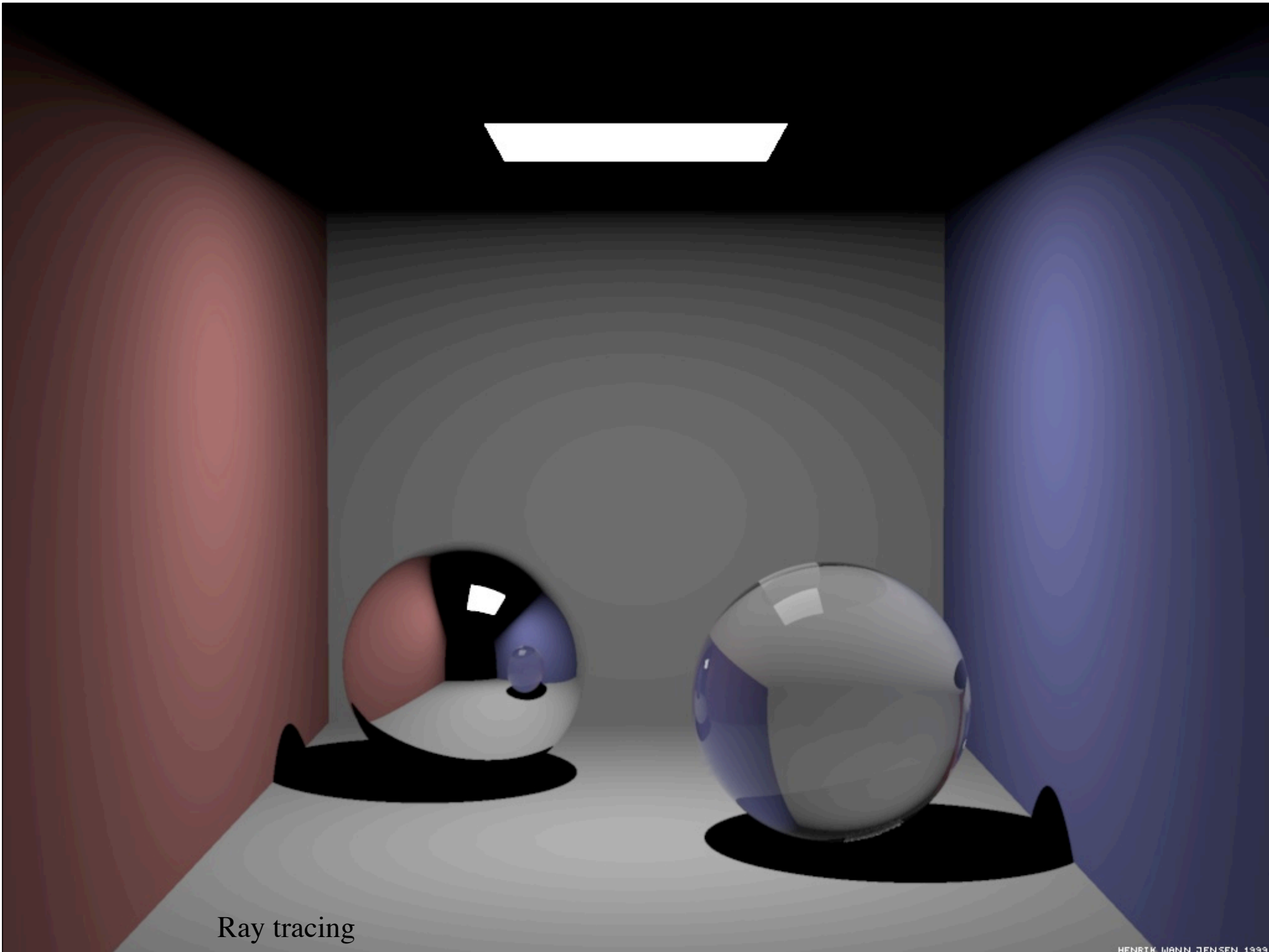
return $2\pi \frac{1}{\pi} \frac{1}{M} \sum_i (\rho(u_i) \text{direct}(u_i) + \text{RKBME}(u_i)) \cos \theta_{si}$



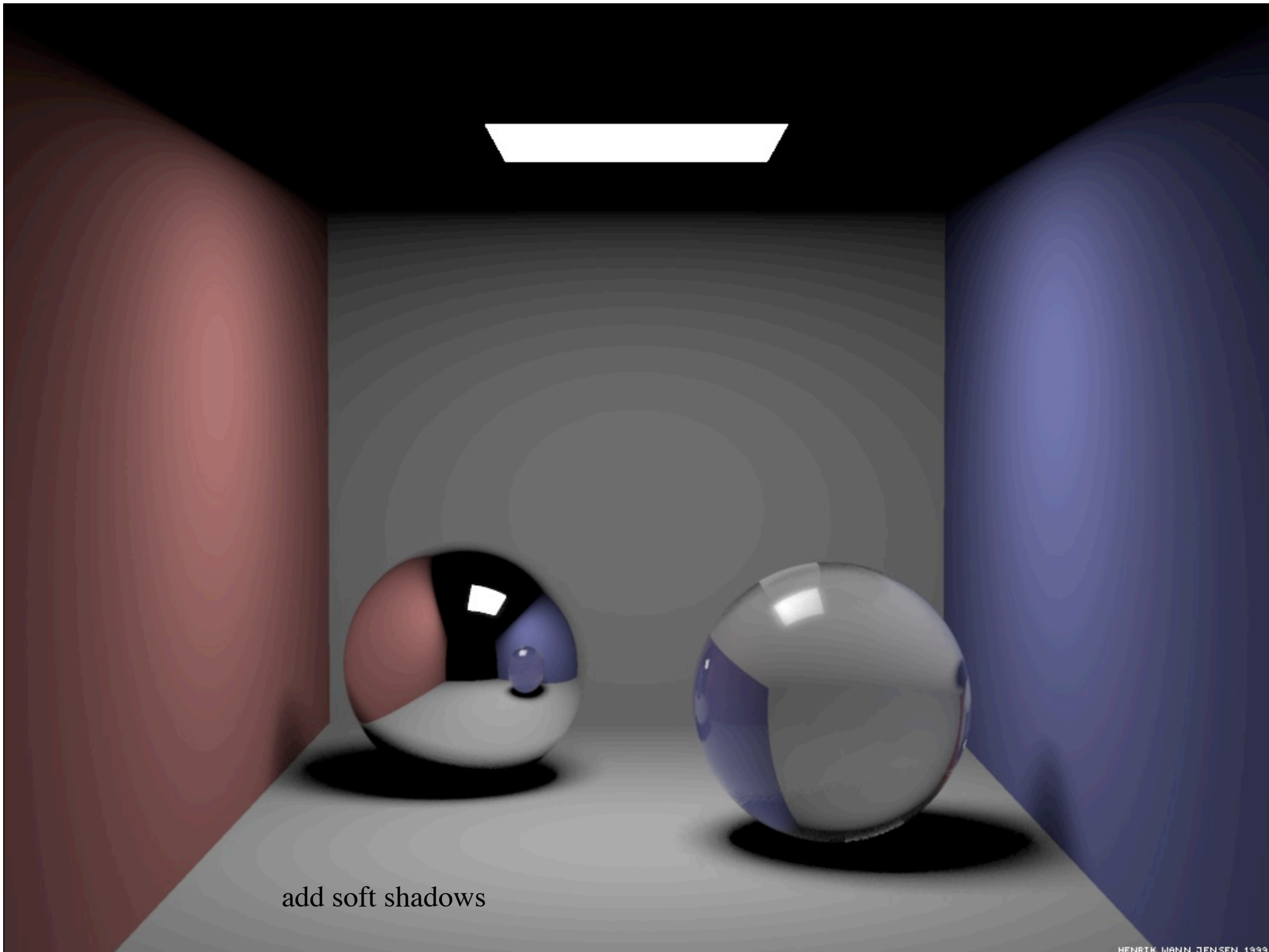
Irradiance cache vs path tracing, from Pharr + Humphreys, for
the same amount of cpu

Light path analysis

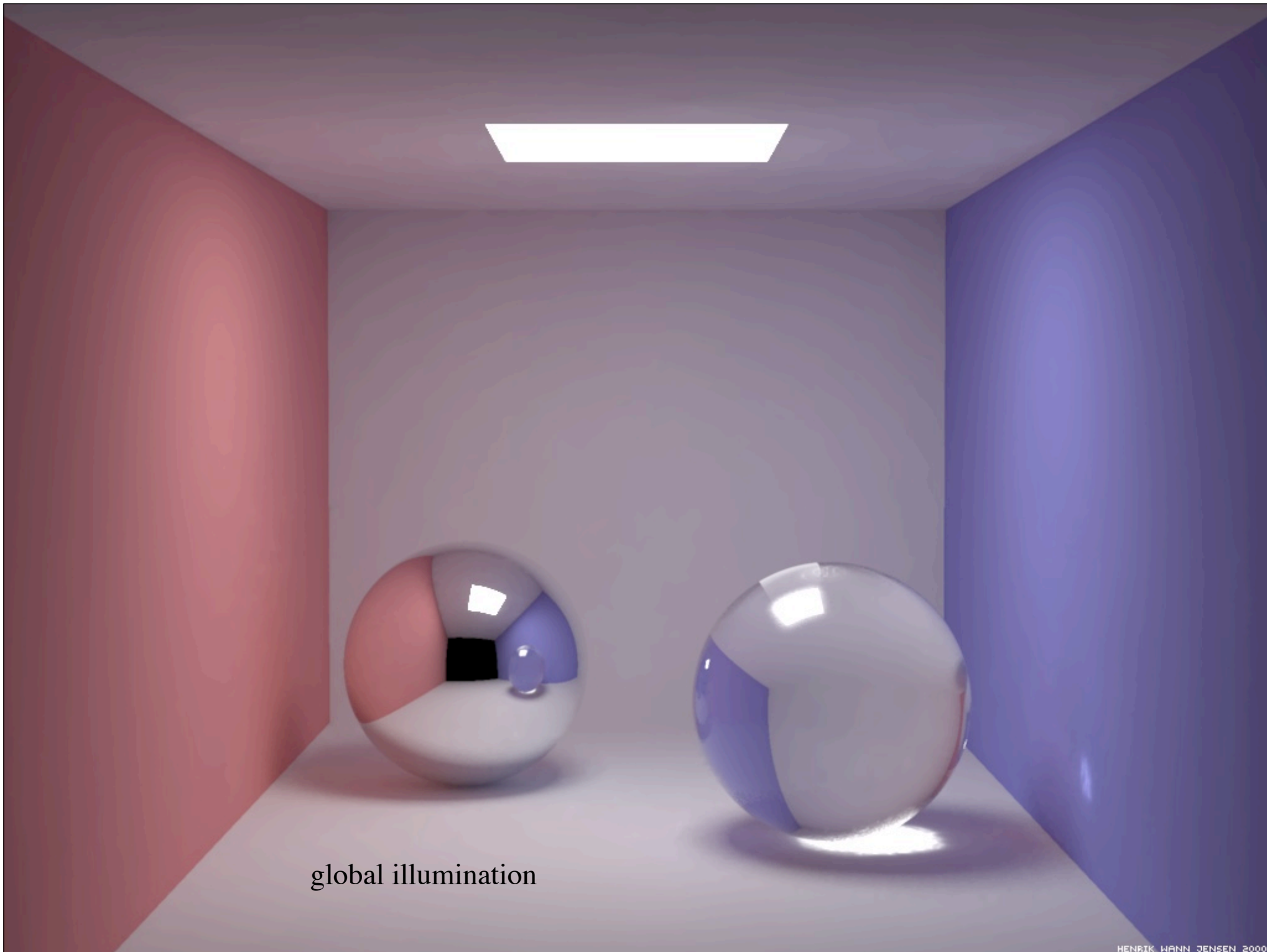
- Main strategy
 - build and evaluate light paths
- We can do other kinds of path like this, too
 - requires extra radiometry



Ray tracing



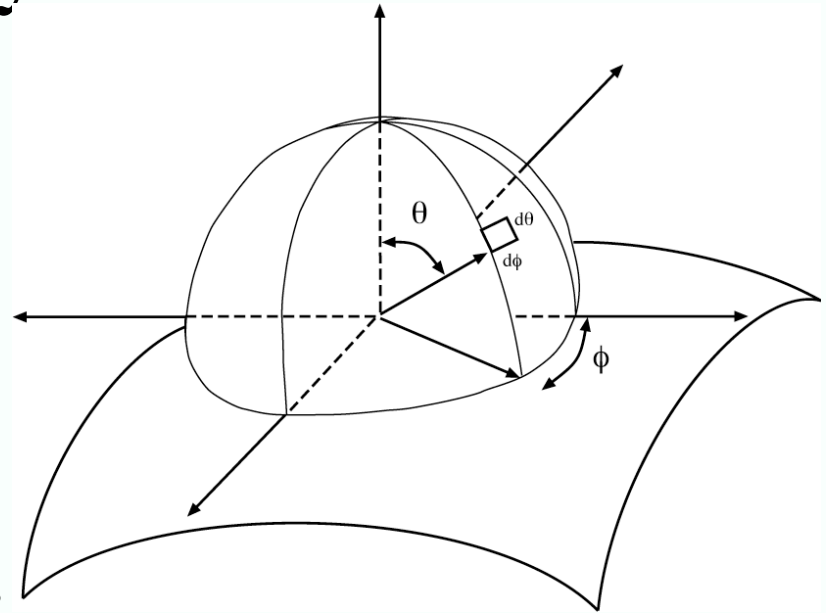
add soft shadows



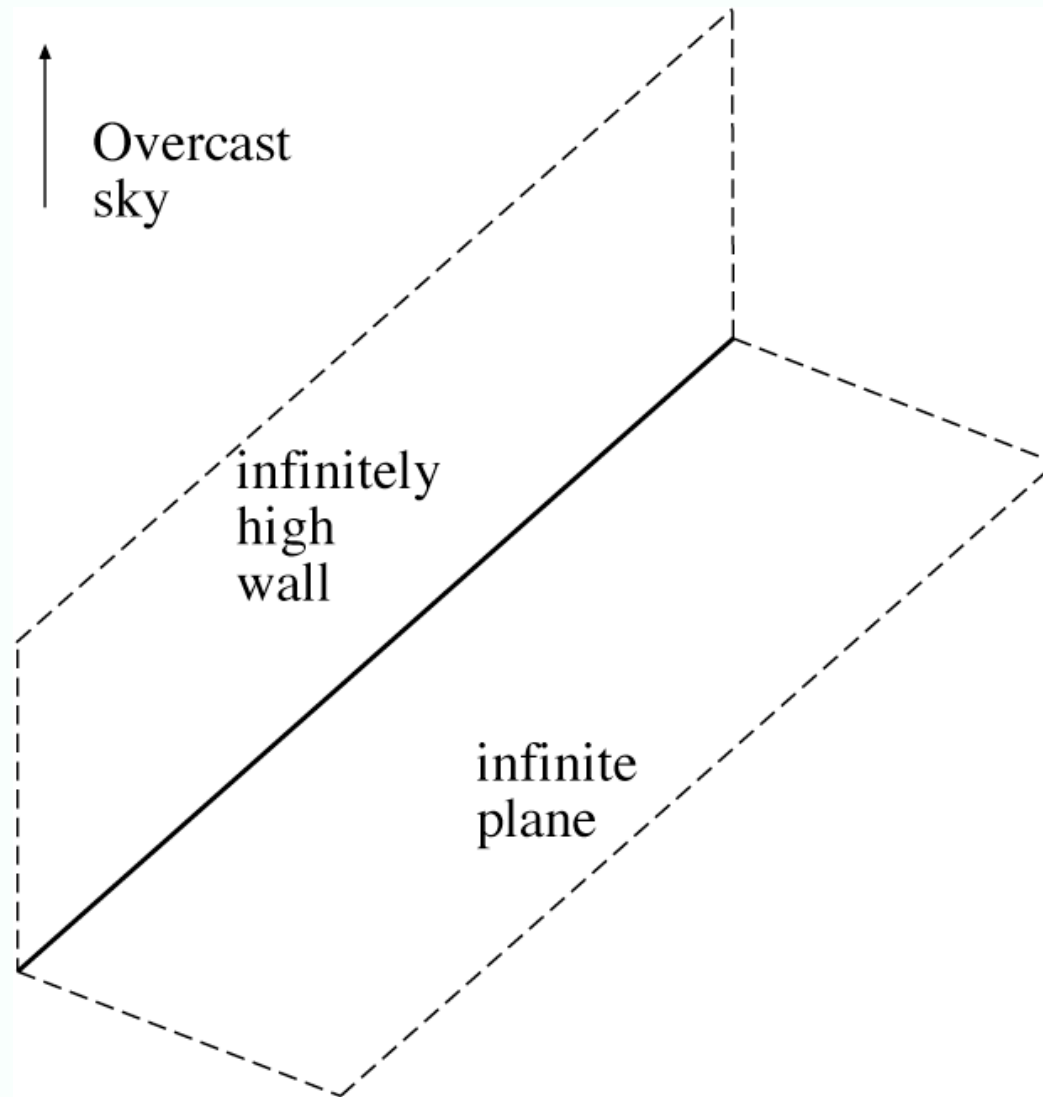
global illumination

Radiometry

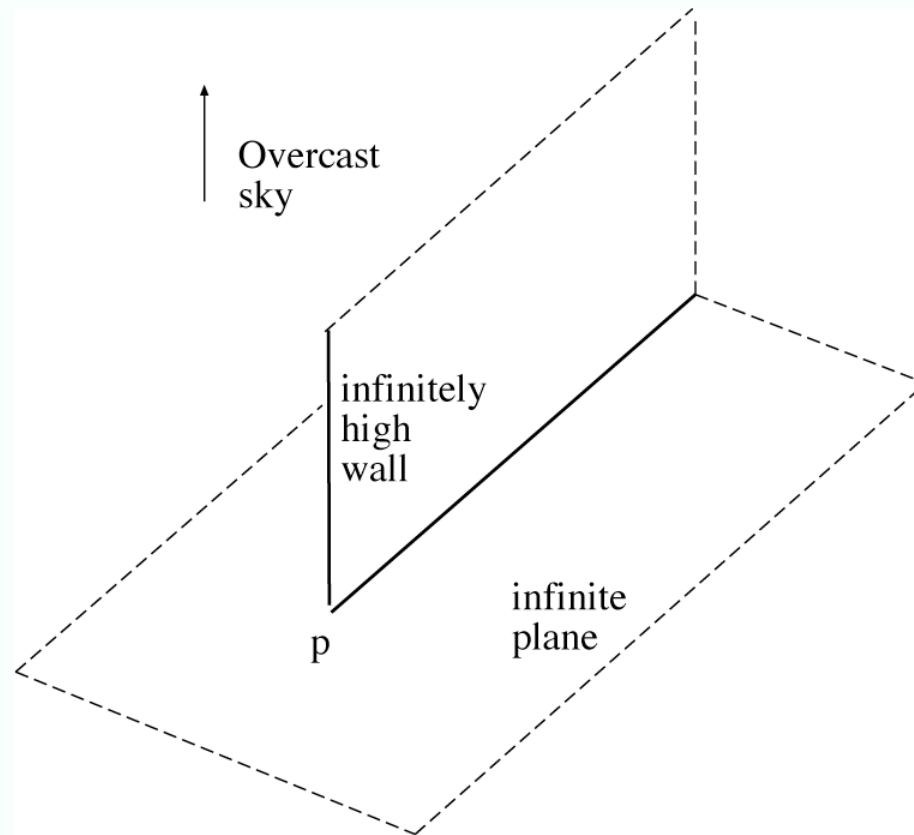
- Questions:
 - how “bright” will surfaces be?
 - what is “brightness”?
 - measuring light
 - interactions between light and surfaces
- Core idea - think about light arriving at a surface
- around any point is a hemisphere of directions
- Simplest problems can be dealt with by reasoning about this hemisphere



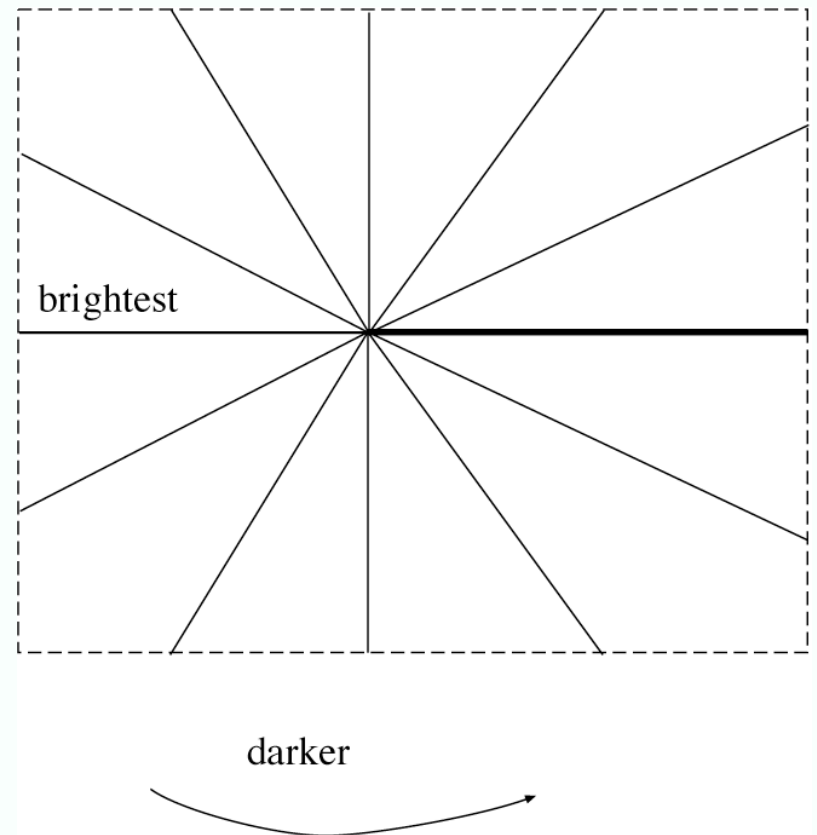
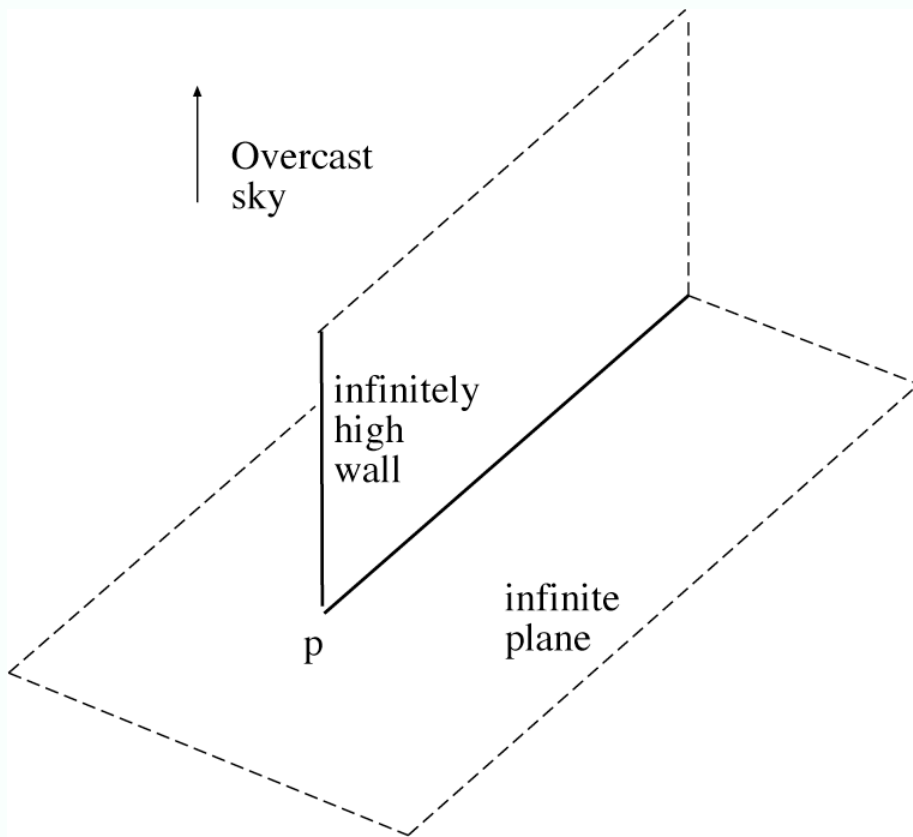
Lambert's wall



More complex wall



More complex wall



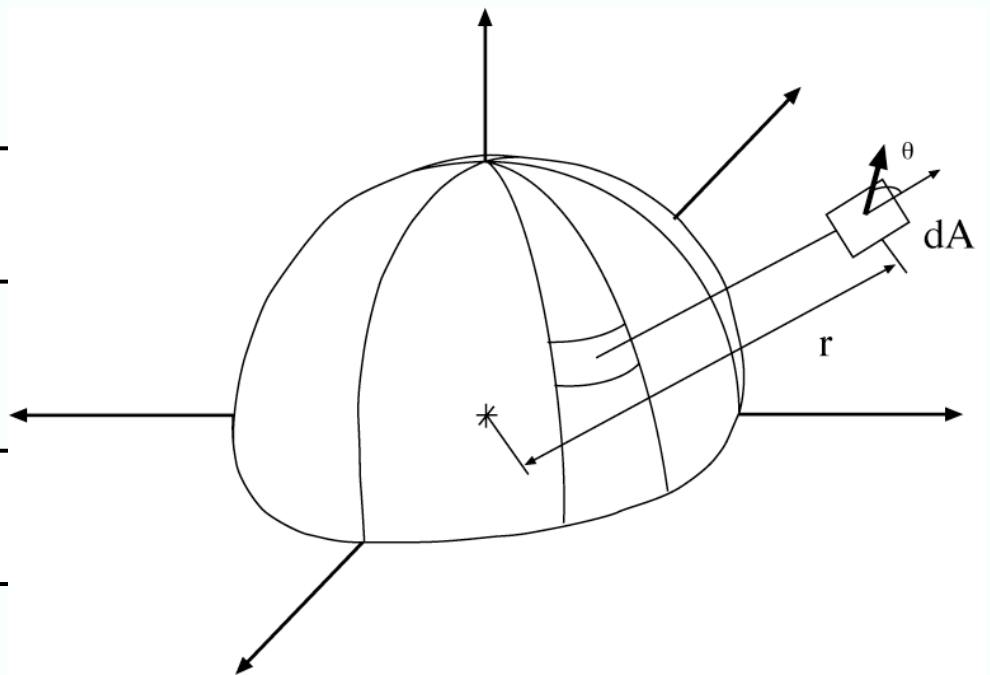
Solid Angle

- By analogy with angle (in radians)
- The solid angle subtended by a patch area dA is given by

$$d\omega = \frac{dA \cos \vartheta}{r^2}$$

- Another useful expression:

$$d\omega = \sin \vartheta (d\vartheta)(d\phi)$$



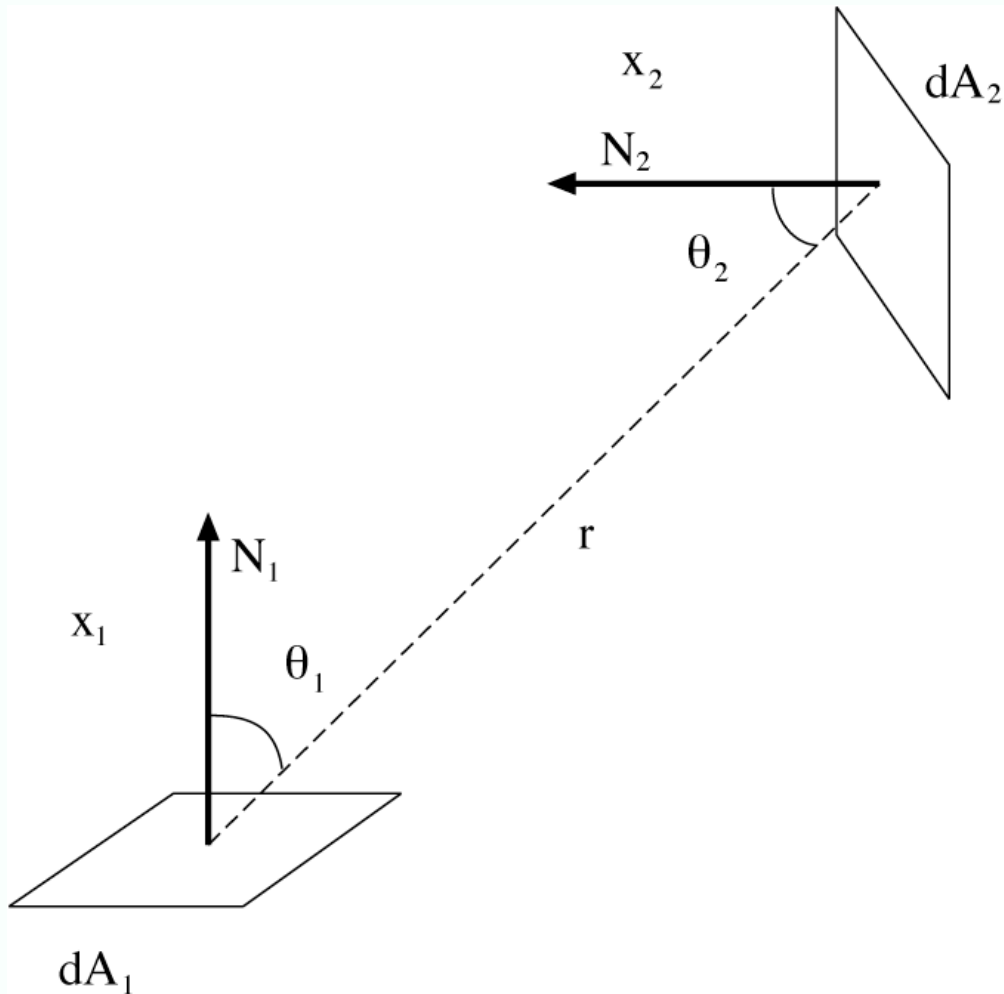
Radiance

- Measure the “amount of light” at a point, in a direction
- Property is:
Radiant power per unit foreshortened area per unit solid angle
- Units: watts per square meter per steradian ($\text{wm}^{-2}\text{sr}^{-1}$)
- Usually written as:

$$L(\underline{x}, \vartheta, \varphi)$$

- Crucial property:
In a vacuum, radiance leaving p in the direction of q is the same as radiance arriving at q from p – hence the units

Radiance is constant along straight lines



- Power 1->2, leaving 1:

$$L(\underline{x}_1, \vartheta, \varphi)(dA_1 \cos \vartheta_1) \left(\frac{dA_2 \cos \vartheta_2}{r^2} \right)$$

- Power 1->2, arriving at 2:

$$L(\underline{x}_2, \vartheta, \varphi)(dA_2 \cos \vartheta_2) \left(\frac{dA_1 \cos \vartheta_1}{r^2} \right)$$

Irradiance

- How much light is arriving at a surface?

- Sensible unit is Irradiance

$$L(\underline{x}, \vartheta, \varphi) \cos \vartheta d\omega$$

- Incident power per unit area not foreshortened

- This is a function of incoming angle.

- A surface experiencing radiance $L(x, \theta, \phi)$ coming in from $d\omega$ experiences irradiance

$$\int_{\Omega} L(\underline{x}, \vartheta, \varphi) \cos \vartheta \sin \vartheta d\vartheta d\varphi$$

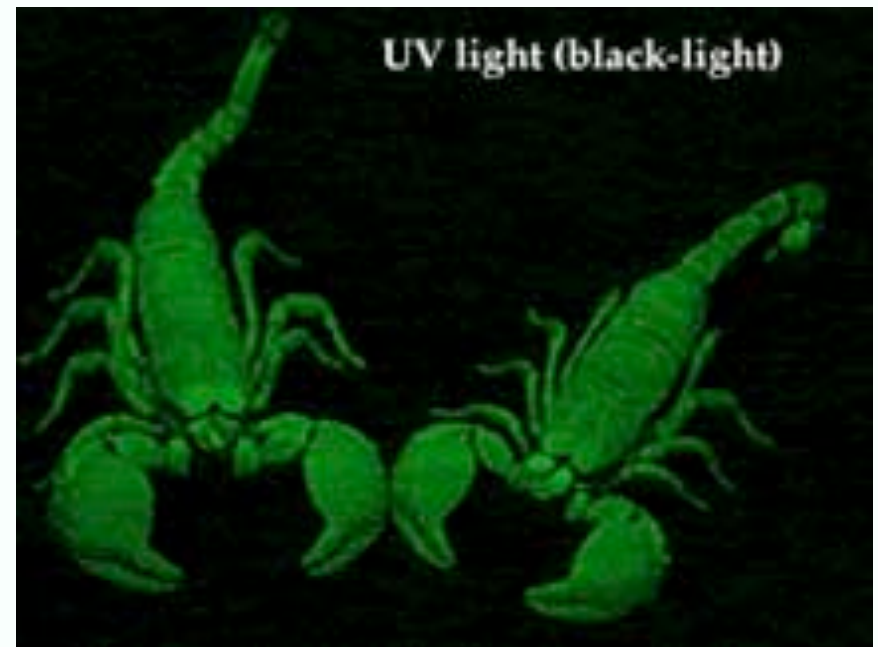
- Crucial property:

Total power arriving at the surface is given by adding irradiance over all incoming angles --- this is why it's a natural unit

Surfaces and the BRDF

- Many effects when light strikes a surface -- could be:
 - absorbed; transmitted. reflected; scattered
- Assume that
 - surfaces don't fluoresce
 - surfaces don't emit light (i.e. are cool)
 - all the light leaving a point is due to that arriving at that point
- Can model this situation with the Bidirectional Reflectance Distribution Function (BRDF)
- the ratio of the radiance in the outgoing direction to the incident irradiance

$$\rho_{bd}(\underline{x}, \vartheta_o, \varphi_o, \vartheta_i, \varphi_i) = \frac{L_o(\underline{x}, \vartheta_o, \varphi_o)}{\int L_i(\underline{x}, \vartheta_i, \varphi_i) \cos \vartheta_i d\omega}$$



BRDF

- Units: inverse steradians (sr-1)
- Symmetric in incoming and outgoing directions
- Radiance leaving in a particular direction:
 - add contributions from every incoming direction

$$\int_{\Omega} \rho_{bd}(\underline{x}, \vartheta_o, \varphi_o, \vartheta_i, \varphi_i) L_i(\underline{x}, \vartheta_i, \varphi_i) \cos \vartheta_i d\omega_i$$

Suppressing Angles - Radiosity

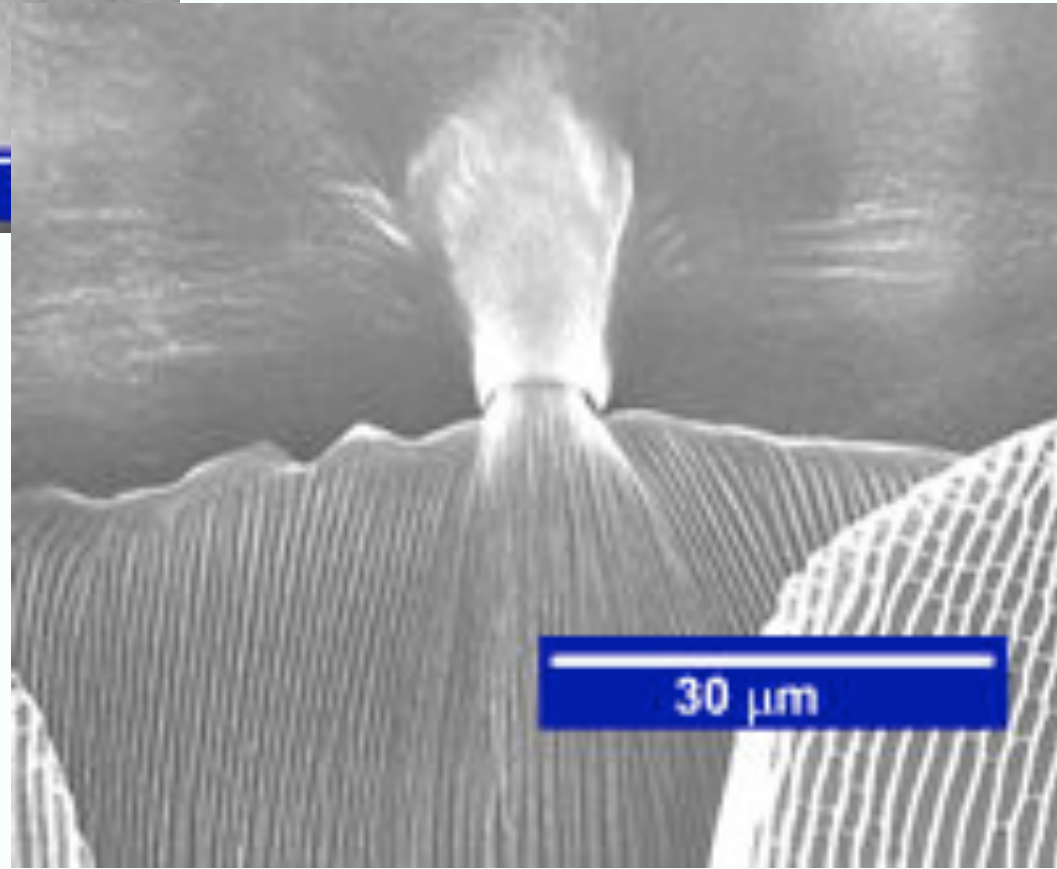
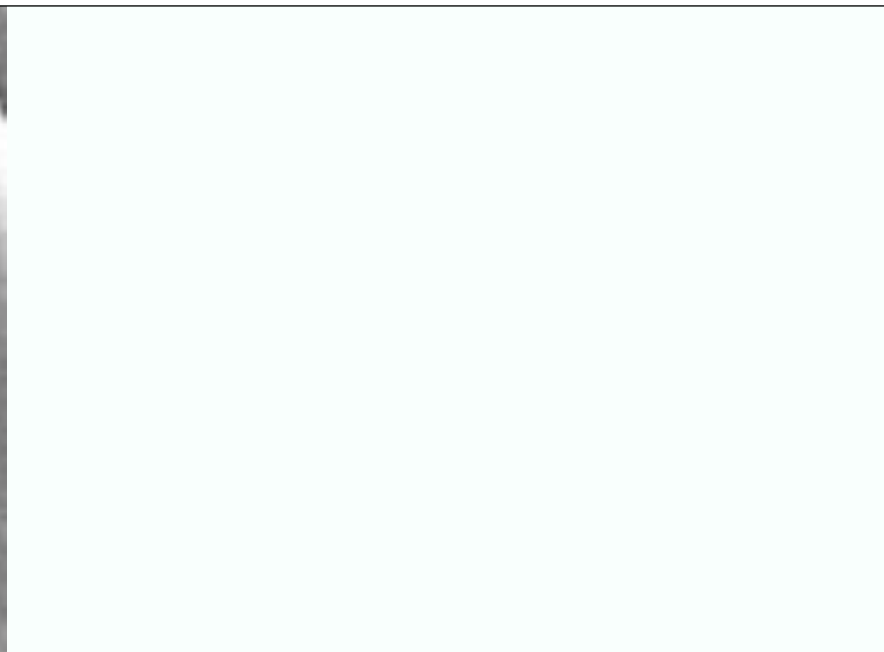
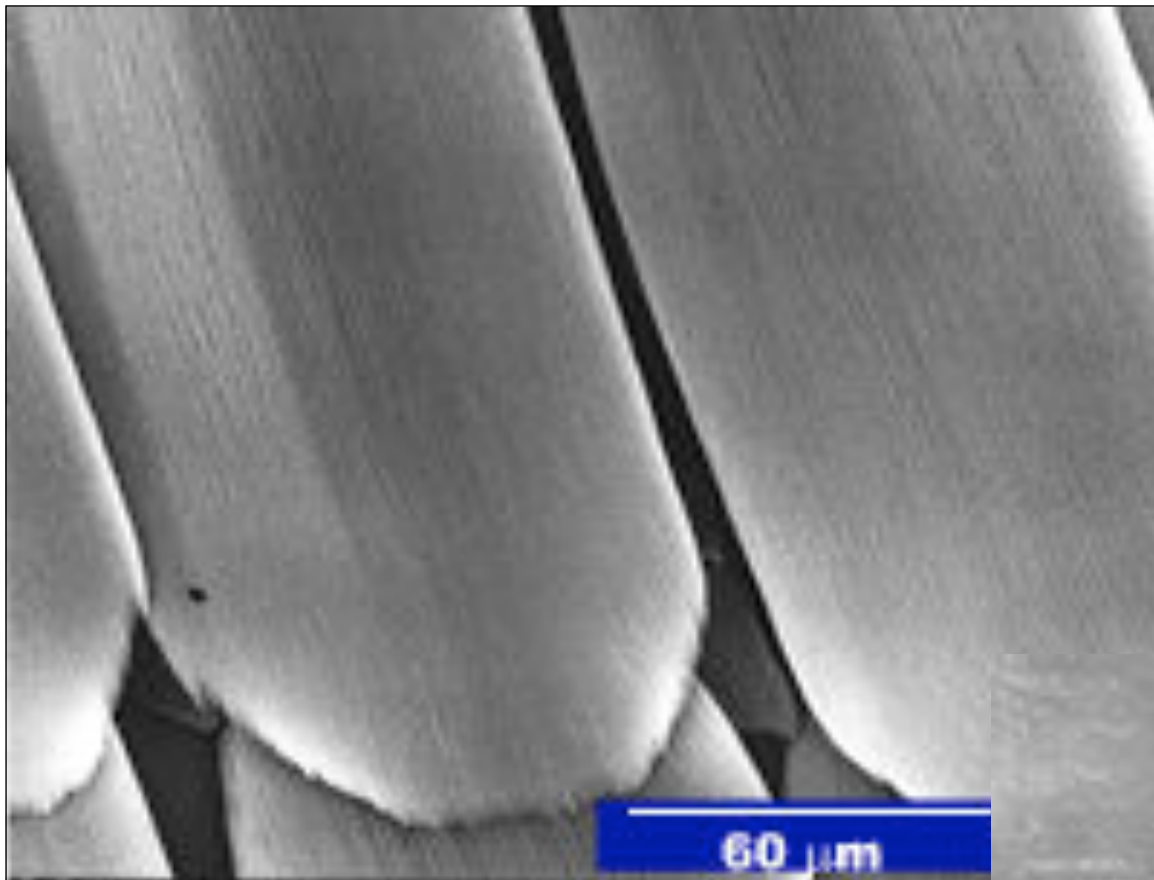
- In many situations, we do not really need angle coordinates
 - e.g. cotton cloth, where the reflected light is not dependent on angle
- Appropriate radiometric unit is radiosity
 - total power leaving a point on the surface, per unit area on the surface (Wm⁻²)
- Radiosity from radiance?
 - sum radiance leaving surface over all exit directions

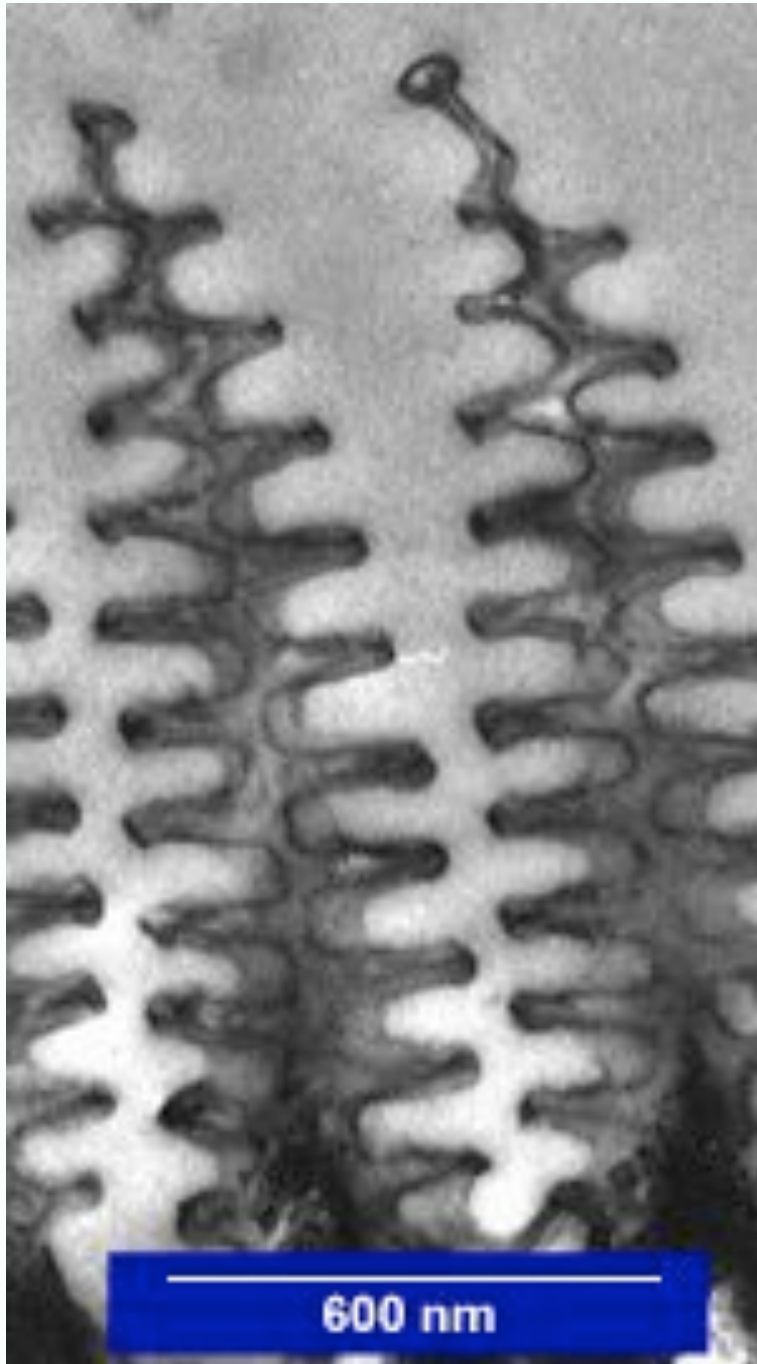
$$B(\underline{x}) = \int_{\Omega} L_o(\underline{x}, \vartheta, \varphi) \cos \vartheta d\omega$$

Radiosity

- Important relationship:
 - radiosity of a surface whose radiance is independent of angle (e.g. that cotton cloth)

$$\begin{aligned} B(\underline{x}) &= \int_{\Omega} L_o(\underline{x}, \vartheta, \varphi) \cos \vartheta d\omega \\ &= L_o(\underline{x}) \int_{\Omega} \cos \vartheta d\omega \\ &= L_o(\underline{x}) \int_0^{\pi/2} \int_0^{2\pi} \cos \vartheta \sin \vartheta d\varphi d\vartheta \\ &= \pi L_o(\underline{x}) \end{aligned}$$





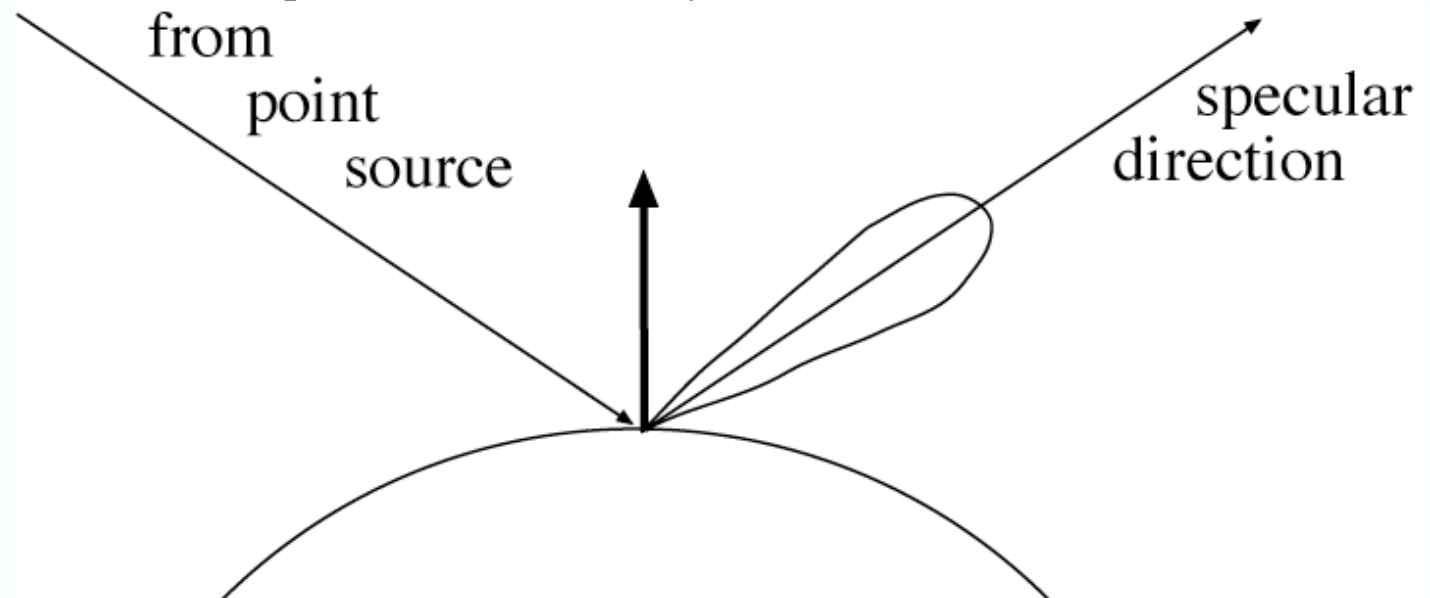
Lambertian surfaces and albedo

- For some surfaces, the BRDF is independent of direction
 - cotton cloth, carpets, matte paper, matte paints, etc.
 - radiance leaving the surface is independent of angle
 - Lambertian surfaces (same Lambert) or ideal diffuse surfaces
 - Use radiosity as a unit to describe light leaving the surface
 - percentage of incident light reflected is diffuse reflectance or albedo
- Useful fact:

$$\rho_{brdf} = \frac{\rho_d}{\pi}$$

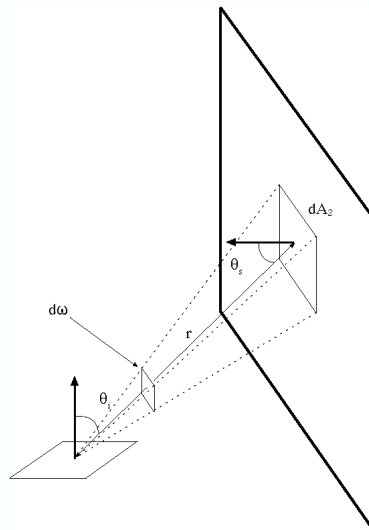
Specular surfaces

- Another important class of surfaces is specular, or mirror-like.
 - radiation arriving along a direction leaves along the specular direction
 - reflect about normal
 - some fraction is absorbed, some reflected
 - on real surfaces, energy usually goes into a lobe of directions
 - can write a BRDF, but requires the use of funny functions



Radiosity due to an area source

- rho is albedo
- E is exitance
- $r(x, u)$ is distance between points
- u is a coordinate on the source



$$\begin{aligned}
 B(x) &= \rho_d(x) \int_{\Omega} L_i(x, u \rightarrow x) \cos \theta_i d\omega \\
 &= \rho_d(x) \int_{\Omega} L_e(x, u \rightarrow x) \cos \theta_i d\omega \\
 &= \rho_d(x) \int_{\Omega} \left(\frac{E(u)}{\pi} \right) \cos \theta_i d\omega \\
 &= \rho_d(x) \int_{source} \left(\frac{E(u)}{\pi} \right) \cos \theta_i \left(\cos \theta_s \frac{dA_u}{r(x, u)^2} \right) \\
 &= \rho_d(x) \int_{source} E(u) \frac{\cos \theta_i \cos \theta_s}{\pi r(x, u)^2} dA_u
 \end{aligned}$$

The Rendering Equation- 1

- We can now write

Angle between normal
and incoming direction

$$L_o(\mathbf{x}, \omega_o) = L_e(\mathbf{x}, \omega_o) + \int_{\Omega} \rho_{bd}(\mathbf{x}, \omega_o, \omega_i) L_i(\mathbf{x}, \omega_i) \cos \theta_i d\omega_i$$

BRDF

Incoming irradiance

Average over hemisphere

Radiance emitted from surface at that point in that direction

Radiance leaving a point in a direction

The Rendering Equation - II

- This balance works for
 - each wavelength,
 - at any time, so
- So

$$L_o(\mathbf{x}, \omega_o, \lambda, t) = L_e(\mathbf{x}, \omega_o, \lambda, t) + \int_{\Omega} \rho_{bd}(\mathbf{x}, \omega_o, \omega_i, \lambda, t) L_i(\mathbf{x}, \omega_i, \lambda, t) \cos \theta_i d\omega_i$$