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CHAPTER 1

Notation and conventions

A dataset as a collection of d -tuples (a d -tuple is an ordered list of d elements). Tuples differ from vectors, because we can always add and subtract vectors, but we cannot necessarily add or subtract tuples. There are always N items in any dataset. There are always d elements in each tuple in a dataset. The number of elements will be the same for every tuple in any given tuple. Sometimes we may not know the value of some elements in some tuples.

We use the same notation for a tuple and for a vector. Most of our data will be vectors. We write a vector in bold, so \mathbf{x} could represent a vector or a tuple (the context will make it obvious which is intended).

The entire data set is $\{\mathbf{x}\}$. When we need to refer to the i 'th data item, we write \mathbf{x}_i . Assume we have N data items, and we wish to make a new dataset out of them; we write the dataset made out of these items as $\{\mathbf{x}_i\}$ (the i is to suggest you are taking a set of items and making a dataset out of them). If we need to refer to the j 'th component of a vector \mathbf{x}_i , we will write $x_i^{(j)}$ (notice this isn't in bold, because it is a component not a vector, and the j is in parentheses because it isn't a power). Vectors are always column vectors.

Terms:

- $\text{mean}(\{x\})$ is the mean of the dataset $\{x\}$ (definition 1, page 16).
- $\text{std}(x)$ is the standard deviation of the dataset $\{x\}$ (definition 2, page 17).
- $\text{var}(\{x\})$ is the standard deviation of the dataset $\{x\}$ (definition 3, page 21).
- $\text{median}(\{x\})$ is the standard deviation of the dataset $\{x\}$ (definition 4, page 22).
- $\text{percentile}(\{x\}, k)$ is the $k\%$ percentile of the dataset $\{x\}$ (definition 5, page 23).
- $\text{iqr}\{x\}$ is the interquartile range of the dataset $\{x\}$ (definition 7, page 23).
- $\{\hat{x}\}$ is the dataset $\{x\}$, transformed to standard coordinates (definition 8, page 28).
- Standard normal data is defined in definition 9, page 29).
- Normal data is defined in definition 10, page 30).
- $\text{corr}(\{(x, y)\})$ is the correlation between two components x and y of a dataset (definition 1, page 56).
- \emptyset is the empty set.
- Ω is the set of all possible outcomes of an experiment.
- Sets are written as \mathcal{A} .

- \mathcal{A}^c is the complement of the set \mathcal{A} (i.e. $\Omega - \mathcal{A}$).
- \mathcal{E} is an event (page 308).
- $P(\{\mathcal{E}\})$ is the probability of event \mathcal{E} (page 308).
- $P(\{\mathcal{E}\}|\{\mathcal{F}\})$ is the probability of event \mathcal{E} , conditioned on event \mathcal{F} (page 308).
- $p(x)$ is the probability that random variable X will take the value x ; also written $P(\{X = x\})$ (page 308).
- $p(x, y)$ is the probability that random variable X will take the value x and random variable Y will take the value y ; also written $P(\{X = x\} \cap \{Y = y\})$ (page 308).
- $\operatorname{argmax}_x f(x)$ means the value of x that maximises $f(x)$.
- $\hat{\theta}$ is an estimated value of a parameter θ .

Background information:

- *Cards:* A standard deck of playing cards contains 52 cards. These cards are divided into four suits. The suits are: spades and clubs (which are black); and hearts and diamonds (which are red). Each suit contains 13 cards: Ace, 2, 3, 4, 5, 6, 7, 8, 9, 10, Jack (sometimes called Knave), Queen and King. It is common to call Jack, Queen and King *court cards*.
- *Dice:* If you look hard enough, you can obtain dice with many different numbers of sides (though I've never seen a three sided die). We adopt the convention that the sides of an N sided die are labeled with the numbers $1 \dots N$, and that no number is used twice. Most dice are like this.
- *Fairness:* Each face of a fair coin or die has the same probability of landing upmost in a flip or roll.

1.1 SOME USEFUL MATHEMATICAL FACTS

The gamma function $\Gamma(x)$ is defined by a series of steps. First, we have that for n an integer,

$$\Gamma(n) = (n - 1)!$$

and then for z a complex number with positive real part (which includes positive real numbers), we have

$$\Gamma(z) = \int_0^\infty t^z \frac{e^{-t}}{t} dt.$$

By doing this, we get a function on positive real numbers that is a smooth interpolate of the factorial function. We won't do any real work with this function, so won't expand on this definition. In practice, we'll either look up a value in tables or require a software environment to produce it.

1.2 ACKNOWLEDGEMENTS

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— Your Name Here —

CHAPTER 2

First Tools for Looking at Data

The single most important question for a working scientist — perhaps the single most useful question anyone can ask — is: “what’s going on here?” Answering this question requires creative use of different ways to make pictures of datasets, to summarize them, and to expose whatever structure might be there. This is an activity that is sometimes known as “Descriptive Statistics”. There isn’t any fixed recipe for understanding a dataset, but there is a rich variety of tools we can use to get insights.

2.1 DATASETS

A dataset is a collection of descriptions of different instances of the same phenomenon. These descriptions could take a variety of forms, but it is important that they are descriptions of the same thing. For example, my grandfather collected the daily rainfall in his garden for many years; we could collect the height of each person in a room; or the number of children in each family on a block; or whether 10 classmates would prefer to be “rich” or “famous”. There could be more than one description recorded for each item. For example, when he recorded the contents of the rain gauge each morning, my grandfather could have recorded (say) the temperature and barometric pressure. As another example, one might record the height, weight, blood pressure and body temperature of every patient visiting a doctor’s office.

The descriptions in a dataset can take a variety of forms. A description could be **categorical**, meaning that each data item can take a small set of prescribed values. For example, we might record whether each of 100 passers-by preferred to be “Rich” or “Famous”. As another example, we could record whether the passers-by are “Male” or “Female”. Categorical data could be **ordinal**, meaning that we can tell whether one data item is larger than another. For example, a dataset giving the number of children in a family for some set of families is categorical, because it uses only non-negative integers, but it is also ordinal, because we can tell whether one family is larger than another.

Some ordinal categorical data appears not to be numerical, but can be assigned a number in a reasonably sensible fashion. For example, many readers will recall being asked by a doctor to rate their pain on a scale of 1 to 10 — a question that is usually relatively easy to answer, but is quite strange when you think about it carefully. As another example, we could ask a set of users to rate the usability of an interface in a range from “very bad” to “very good”, and then record that using -2 for “very bad”, -1 for “bad”, 0 for “neutral”, 1 for “good”, and 2 for “very good”.

Many interesting datasets involve **continuous** variables (like, for example, height or weight or body temperature) when you could reasonably expect to encounter any value in a particular range. For example, we might have the heights of

all people in a particular room; or the rainfall at a particular place for each day of the year; or the number of children in each family on a list.

You should think of a dataset as a collection of d -tuples (a d -tuple is an ordered list of d elements). Tuples differ from vectors, because we can always add and subtract vectors, but we cannot necessarily add or subtract tuples. We will always write N for the number of tuples in the dataset, and d for the number of elements in each tuple. The number of elements will be the same for every tuple, though sometimes we may not know the value of some elements in some tuples (which means we must figure out how to predict their values, which we will do much later).

Index	net worth	Index	Taste score	Index	Taste score
1	100, 360	1	12.3	11	34.9
2	109, 770	2	20.9	12	57.2
3	96, 860	3	39	13	0.7
4	97, 860	4	47.9	14	25.9
5	108, 930	5	5.6	15	54.9
6	124, 330	6	25.9	16	40.9
7	101, 300	7	37.3	17	15.9
8	112, 710	8	21.9	18	6.4
9	106, 740	9	18.1	19	18
10	120, 170	10	21	20	38.9

TABLE 2.1: *On the left, net worths of people you meet in a bar, in US \$; I made this data up, using some information from the US Census. The index column, which tells you which data item is being referred to, is usually not displayed in a table because you can usually assume that the first line is the first item, and so on. On the right, the taste score (I'm not making this up; higher is better) for 20 different cheeses. This data is real (i.e. not made up), and it comes from <http://lib.stat.cmu.edu/DASL/Datafiles/Cheese.html>.*

Each element of a tuple has its own type. Some elements might be categorical. For example, one dataset we shall see several times records entries for Gender; Grade; Age; Race; Urban/Rural; School; Goals; Grades; Sports; Looks; and Money for 478 children, so $d = 11$ and $N = 478$. In this dataset, each entry is categorical data. Clearly, these tuples are not vectors because one cannot add or subtract (say) Genders.

Most of our data will be vectors. We use the same notation for a tuple and for a vector. We write a vector in bold, so \mathbf{x} could represent a vector or a tuple (the context will make it obvious which is intended).

The entire data set is $\{\mathbf{x}\}$. When we need to refer to the i 'th data item, we write \mathbf{x}_i . Assume we have N data items, and we wish to make a new dataset out of them; we write the dataset made out of these items as $\{\mathbf{x}_i\}$ (the i is to suggest you are taking a set of items and making a dataset out of them).

In this chapter, we will work mainly with continuous data. We will see a variety of methods for plotting and summarizing 1-tuples. We can build these plots from a dataset of d -tuples by extracting the r 'th element of each d -tuple.

Mostly, we will deal with continuous data. All through the book, we will see many datasets downloaded from various web sources, because people are so generous about publishing interesting datasets on the web. In the next chapter, we will look at 2-dimensional data, and we look at high dimensional data in chapter 11.

2.2 WHAT'S HAPPENING? - PLOTTING DATA

The very simplest way to present or visualize a dataset is to produce a table. Tables can be helpful, but aren't much use for large datasets, because it is difficult to get any sense of what the data means from a table. As a continuous example, table 2.1 gives a table of the net worth of a set of people you might meet in a bar (I made this data up). You can scan the table and have a rough sense of what is going on; net worths are quite close to \$ 100, 000, and there aren't any very big or very small numbers. This sort of information might be useful, for example, in choosing a bar.

People would like to measure, record, and reason about an extraordinary variety of phenomena. Apparently, one can score the goodness of the flavor of cheese with a number (bigger is better); table 2.1 gives a score for each of thirty cheeses (I did not make up this data, but downloaded it from <http://lib.stat.cmu.edu/DASL/Datafiles/Cheese.html>). You should notice that a few cheeses have very high scores, and most have moderate scores. It's difficult to draw more significant conclusions from the table, though.

Gender	Goal	Gender	Goal
boy	Sports	girl	Sports
boy	Popular	girl	Grades
girl	Popular	boy	Popular
girl	Popular	boy	Popular
girl	Popular	boy	Popular
girl	Popular	girl	Grades
girl	Popular	girl	Sports
girl	Grades	girl	Popular
girl	Sports	girl	Grades
girl	Sports	girl	Sports

TABLE 2.2: *Chase and Dunner (?) collected data on what students thought made other students popular. As part of this effort, they collected information on (a) the gender and (b) the goal of students. This table gives the gender (“boy” or “girl”) and the goal (to make good grades — “Grades”; to be popular — “Popular”; or to be good at sports — “Sports”). The table gives this information for the first 20 of 478 students; the rest can be found at <http://lib.stat.cmu.edu/DASL/Datafiles/PopularKids.html>. This data is clearly categorical, and not ordinal.*

Table 2.2 shows a table for a set of categorical data. Psychologists collected data from students in grades 4-6 in three school districts to understand what factors students thought made other students popular. This fascinating data set can be found at <http://lib.stat.cmu.edu/DASL/Datafiles/PopularKids.html>, and was prepared by Chase and Dunner (?). Among other things, for each student

they asked whether the student's goal was to make good grades ("Grades", for short); to be popular ("Popular"); or to be good at sports ("Sports"). They have this information for 478 students, so a table would be very hard to read. Table 2.2 shows the gender and the goal for the first 20 students in this group. It's rather harder to draw any serious conclusion from this data, because the full table would be so big. We need a more effective tool than eyeballing the table.

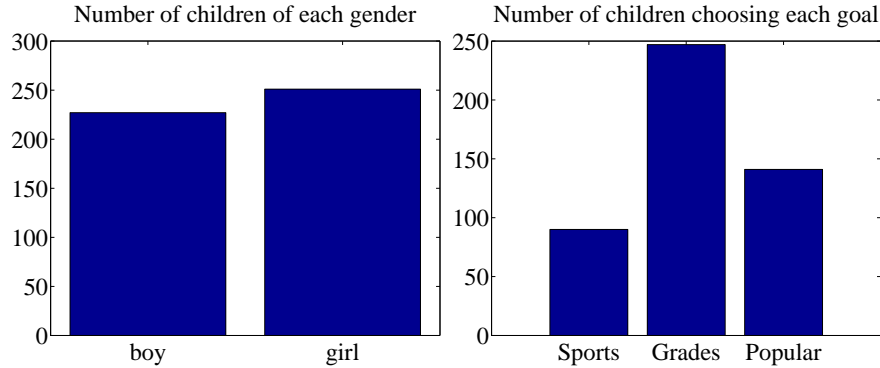


FIGURE 2.1: *On the left*, a bar chart of the number of children of each gender in the Chase and Dunner study (). Notice that there are about the same number of boys and girls (the bars are about the same height). *On the right*, a bar chart of the number of children selecting each of three goals. You can tell, at a glance, that different goals are more or less popular by looking at the height of the bars.

2.2.1 Bar Charts

A **bar chart** is a set of bars, one per category, where the height of each bar is proportional to the number of items in that category. A glance at a bar chart often exposes important structure in data, for example, which categories are common, and which are rare. Bar charts are particularly useful for categorical data. Figure 2.1 shows such bar charts for the genders and the goals in the student dataset of Chase and Dunner (). You can see at a glance that there are about as many boys as girls, and that there are more students who think grades are important than students who think sports or popularity is important. You couldn't draw either conclusion from Table 2.2, because I showed only the first 20 items; but a 478 item table is very difficult to read.

2.2.2 Histograms

Data is continuous when a data item could take any value in some range or set of ranges. In turn, this means that we can reasonably expect a continuous dataset contains few or no pairs of items that have *exactly* the same value. Drawing a bar chart in the obvious way — one bar per value — produces a mess of unit height bars, and seldom leads to a good plot. Instead, we would like to have fewer bars, each representing more data items. We need a procedure to decide which data items count in which bar.

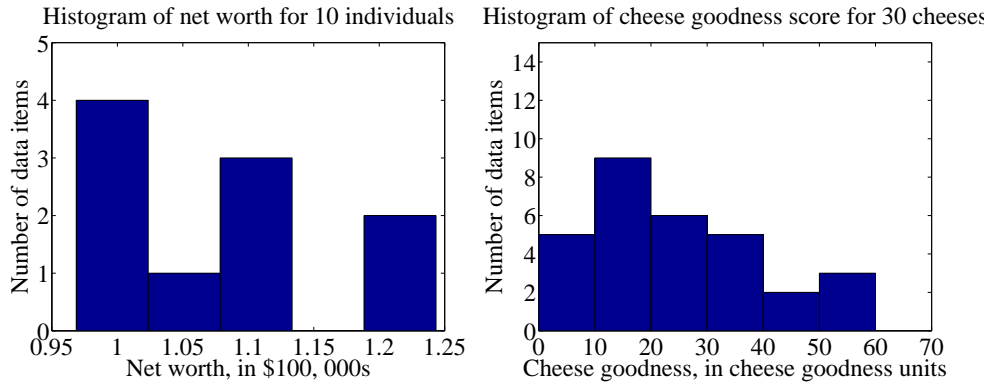


FIGURE 2.2: On the **left**, a histogram of net worths from the dataset described in the text and shown in table 2.1. On the **right**, a histogram of cheese goodness scores from the dataset described in the text and shown in table 2.1.

A simple generalization of a bar chart is a **histogram**. We divide the range of the data into intervals, which do not need to be equal in length. We think of each interval as having an associated pigeonhole, and choose one pigeonhole for each data item. We then build a set of boxes, one per interval. Each box sits on its interval on the horizontal axis, and its height is determined by the number of data items in the corresponding pigeonhole. In the simplest histogram, the intervals that form the bases of the boxes are equally sized. In this case, the height of the box is given by the number of data items in the box.

Figure 2.2 shows a histogram of the data in table 2.1. There are five bars — by my choice; I could have plotted ten bars — and the height of each bar gives the number of data items that fall into its interval. For example, there is one net worth in the range between \$102, 500 and \$107, 500. Notice that one bar is invisible, because there is no data in that range. This picture suggests conclusions consistent with the ones we had from eyeballing the table — the net worths tend to be quite similar, and around \$100, 000.

Figure 2.2 shows a histogram of the data in table 2.1. There are six bars (0-10, 10-20, and so on), and the height of each bar gives the number of data items that fall into its interval — so that, for example, there are 9 cheeses in this dataset whose score is greater than or equal to 10 and less than 20. You can also use the bars to estimate other properties. So, for example, there are 14 cheeses whose score is less than 20, and 3 cheeses with a score of 50 or greater. This picture is much more helpful than the table; you can see at a glance that quite a lot of cheeses have relatively low scores, and few have high scores.

2.2.3 How to Make Histograms

Usually, one makes a histogram by finding the appropriate command or routine in your programming environment. I use Matlab, and chapter 3 sketches some useful Matlab commands. However, it is useful to understand the procedures used.

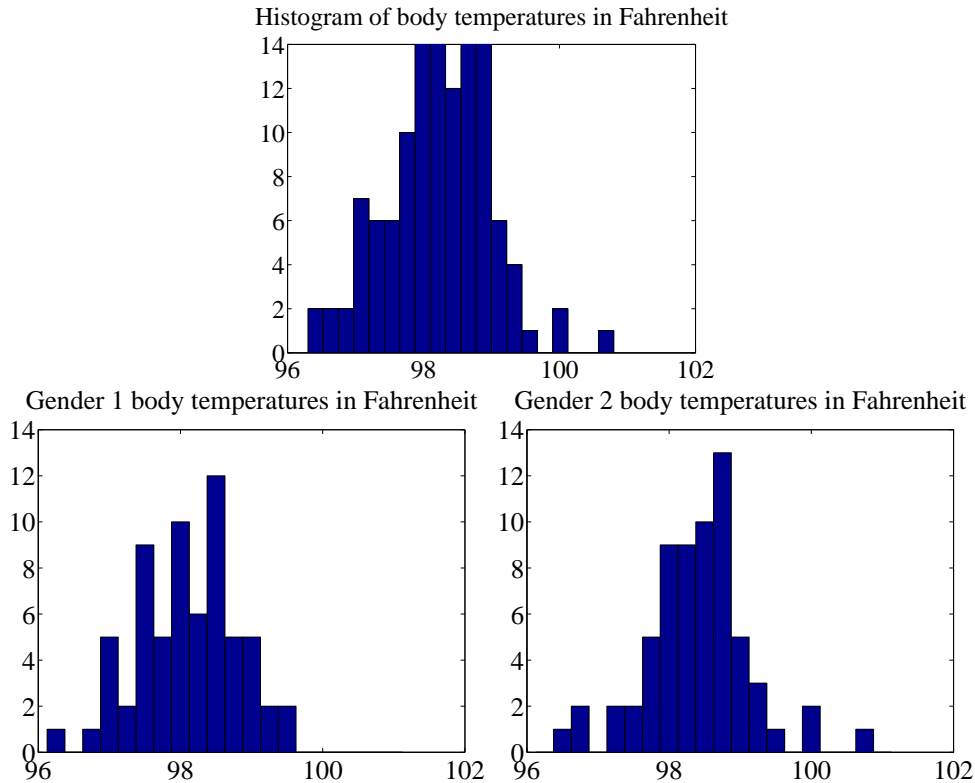


FIGURE 2.3: On **top**, a histogram of body temperatures, from the dataset published at <http://www2.stetson.edu/~jrasp/data.htm>. These seem to be clustered fairly tightly around one value. The **bottom row** shows histograms for each gender (I don't know which is which). It looks as though one gender runs slightly cooler than the other.

Histograms with Even Intervals: The easiest histogram to build uses equally sized intervals. Write x_i for the i 'th number in the dataset, x_{\min} for the smallest value, and x_{\max} for the largest value. We divide the range between the smallest and largest values into n intervals of even width $(x_{\max} - x_{\min})/n$. In this case, the height of each box is given by the number of items in that interval. We could represent the histogram with an n -dimensional vector of counts. Each entry represents the count of the number of data items that lie in that interval. Notice we need to be careful to ensure that each point in the range of values is claimed by exactly one interval. For example, we could have intervals of $[0 - 1)$ and $[1 - 2)$, or we could have intervals of $(0 - 1]$ and $(1 - 2]$. We could *not* have intervals of $[0 - 1]$ and $[1 - 2]$, because then a data item with the value 1 would appear in two boxes. Similarly, we could not have intervals of $(0 - 1)$ and $(1 - 2)$, because then a data item with the value 1 would not appear in any box.

Histograms with Uneven Intervals: For a histogram with even intervals, it is natural that the height of each box is the number of data items in that box.

But a histogram with even intervals can have empty boxes (see figure 2.2). In this case, it can be more informative to have some larger intervals to ensure that each interval has some data items in it. But how high should we plot the box? Imagine taking two consecutive intervals in a histogram with even intervals, and fusing them. It is natural that the height of the fused box should be the average height of the two boxes. This observation gives us a rule.

Write dx for the width of the intervals; n_1 for the height of the box over the first interval (which is the number of elements in the first box); and n_2 for the height of the box over the second interval. The height of the fused box will be $(n_1 + n_2)/2$. Now the *area* of the first box is $n_1 dx$; of the second box is $n_2 dx$; and of the fused box is $(n_1 + n_2) dx$. For each of these boxes, the *area* of the box is proportional to the number of elements in the box. This gives the correct rule: plot boxes such that the area of the box is proportional to the number of elements in the box.

2.2.4 Conditional Histograms

Most people believe that normal body temperature is 98.4° in Fahrenheit. If you take other people's temperatures often (for example, you might have children), you know that some individuals tend to run a little warmer or a little cooler than this number. I found data giving the body temperature of a set of individuals at <http://www2.stetson.edu/~jrasp/data.htm>. As you can see from the histogram (figure 2.3), the body temperatures cluster around a small set of numbers. But what causes the variation?

One possibility is gender. We can investigate this possibility by comparing a histogram of temperatures for males with histogram of temperatures for females. Such histograms are sometimes called **conditional histograms** or **class-conditional histograms**, because each histogram is conditioned on something (in this case, the histogram uses only data that comes from gender).

The dataset gives genders (as 1 or 2 - I don't know which is male and which female). Figure 2.3 gives the class conditional histograms. It does seem like individuals of one gender run a little cooler than individuals of the other, although we don't yet have mechanisms to test this possibility in detail (chapter 1).

2.3 SUMMARIZING 1D DATA

For the rest of this chapter, we will assume that data items take values that are continuous real numbers. Furthermore, we will assume that values can be added, subtracted, and multiplied by constants in a meaningful way. Human heights are one example of such data; you can add two heights, and interpret the result as a height (perhaps one person is standing on the head of the other). You can subtract one height from another, and the result is meaningful. You can multiply a height by a constant — say, $1/2$ — and interpret the result (A is half as high as B). Not all data is like this. Categorical data is often not like this. For example, you could not add “Grades” to “Popular” in any useful way.

2.3.1 The Mean

One simple and effective summary of a set of data is its **mean**. This is sometimes known as the **average** of the data.

Definition: 2.1 *Mean*

Assume we have a dataset $\{x\}$ of N data items, x_1, \dots, x_N . Their mean is

$$\text{mean}(\{x\}) = \frac{1}{N} \sum_{i=1}^{i=N} x_i.$$

For example, assume you're in a bar, in a group of ten people who like to talk about money. They're average people, and their net worth is given in table 2.1 (you can choose who you want to be in this story). The mean of this data is \$107,903.

An important interpretation of the mean is that it is the best guess of the value of a new data item, given no information at all. In the bar example, if a new person walked into this bar, and you had to guess that person's net worth, you should choose \$107,903.

Properties of the Mean The mean has several important properties you should remember:

- Scaling data scales the mean: or $\text{mean}(\{kx_i\}) = k\text{mean}(\{x_i\})$.
- Translating data translates the mean: or $\text{mean}(\{x_i + c\}) = \text{mean}(\{x_i\}) + c$.
- The sum of signed differences from the mean is zero. This means that

$$\sum_{i=1}^N (x_i - \text{mean}(\{x_i\})) = 0.$$

- Choose the number μ such that the sum of squared distances of data points to μ is minimized. That number is the mean. In notation

$$\arg \min_{\mu} \sum_i (x_i - \mu)^2 = \text{mean}(\{x_i\})$$

These properties are easy to prove (and so easy to remember). All but one proof is relegated to the exercises.

Proposition: $\arg \min_{\mu} \sum_i (x_i - \mu)^2 = \text{mean}(\{x\})$

Proof: Choose the number μ such that the sum of squared distances of data points to μ is minimized. That number is the mean. In notation:

$$\arg \min_{\mu} \sum_i (x_i - \mu)^2 = \text{mean}(\{x\})$$

We can show this by actually minimizing the expression. We must have that the derivative of the expression we are minimizing is zero at the value of μ we are seeking. So we have

$$\begin{aligned} \frac{d}{d\mu} \sum_{i=1}^N (x_i - \mu)^2 &= \sum_{i=1}^N 2(x_i - \mu) \\ &= 2 \sum_{i=1}^N (x_i - \mu) \\ &= 0 \end{aligned}$$

so that $2N\text{mean}(\{x\}) - 2N\mu = 0$, which means that $\mu = \text{mean}(\{x\})$.

Property 2.1: The Average Squared Distance to the Mean is Minimized

2.3.2 Standard Deviation and Variance

We would also like to know the extent to which data items are close to the mean. This information is given by the **standard deviation**, which is the root mean square of the offsets of data from the mean.

Definition: 2.2 *Standard deviation*

Assume we have a dataset $\{x\}$ of N data items, x_1, \dots, x_N . The standard deviation of this dataset is is:

$$\text{std}(x_i) = \sqrt{\frac{1}{N} \sum_{i=1}^N (x_i - \text{mean}(\{x\}))^2} = \sqrt{\text{mean}(\{(x_i - \text{mean}(\{x\}))^2\})}$$

You should think of the standard deviation as a scale. It measures the size of the average deviation from the mean for a dataset. When the standard deviation of a dataset is large, there are many items with values much larger than, or much smaller than, the mean. When the standard deviation is small, most data items have values close to the mean. This means it is helpful to talk about how many standard deviations away from the mean a particular data item is. Saying that data

item x_j is “within k standard deviations from the mean” means that

$$\text{abs}(x_j - \text{mean}(\{x\})) \leq k\text{std}(x_i).$$

Similarly, saying that data item x_j is “more than k standard deviations from the mean” means that

$$\text{abs}(x_i - \text{mean}(\{x\})) > k\text{std}(x).$$

As I will show below, there must be some data at least one standard deviation away from the mean, and there can be very few data items that are many standard deviations away from the mean.

Properties of the Standard Deviation Standard deviation has very important properties:

- Translating data does not change the standard deviation, i.e. $\text{std}(x_i + c) = \text{std}(x_i)$.
- Scaling data scales the standard deviation, i.e. $\text{std}(kx_i) = k\text{std}(x_i)$.
- For any dataset, there can be only a few items that are many standard deviations away from the mean. In particular, assume we have N data items, x_i , whose standard deviation is σ . Then there are at most $\frac{1}{k^2}$ data points lying k or more standard deviations away from the mean.
- For any dataset, there must be at least one data item that is at least one standard deviation away from the mean.

The first two properties are easy to prove, and are relegated to the exercises.

Proposition: Assume we have a dataset $\{x\}$ of N data items, x_1, \dots, x_N . Assume the standard deviation of this dataset is $\text{std}(x) = \sigma$. Then there are at most $\frac{1}{k^2}$ data points lying k or more standard deviations away from the mean.

Proof: Assume the mean is zero. There is no loss of generality here, because translating data translates the mean, but doesn't change the standard deviation. The way to prove this is to construct a dataset with the largest possible fraction r of data points lying k or more standard deviations from the mean. To achieve this, our data should have $N(1 - r)$ data points each with the value 0, because these contribute 0 to the standard deviation. It should have Nr data points with the value $k\sigma$; if they are further from zero than this, each will contribute more to the standard deviation, so the fraction of such points will be fewer. Because

$$\text{std}(x) = \sigma = \sqrt{\frac{\sum_i x_i^2}{N}}$$

we have that, for this rather specially constructed dataset,

$$\sigma = \sqrt{\frac{Nr k^2 \sigma^2}{N}}$$

so that

$$r = \frac{1}{k^2}.$$

We constructed the dataset so that r would be as large as possible, so

$$r \geq \frac{1}{k^2}$$

for any kind of data at all.

Property 2.2: For any dataset, it is hard for data items to get many standard deviations away from the mean.

The bound of box 2.3.2 is true for *any kind of data*. This bound implies that, for example, at most 100% of *any* dataset could be one standard deviation away from the mean, 25% of *any* dataset is 2 standard deviations away from the mean and at most 11% of *any* dataset could be 3 standard deviations away from the mean. But the configuration of data that achieves this bound is very unusual. This means the bound tends to wildly overstate how much data is far from the mean for most practical datasets. Most data has more random structure, meaning that we expect to see very much *less* data far from the mean than the bound predicts. For example, much data can reasonably be modelled as coming from a normal distribution (a topic we'll go into later). For such data, we expect that about 68% of the data is within one standard deviation of the mean, 95% is within two standard deviations of the mean, and 99.7% is within three standard deviations of the mean, and the percentage of data that is within ten standard deviations of the mean is essentially indistinguishable from 100%. This kind of behavior is quite common; the crucial point about the standard deviation is that you won't see much

data that lies many standard deviations from the mean, because you can't.

Proposition: $(\text{std}(x))^2 \leq \max_i (x_i - \text{mean}(\{x\}))^2$.

Proof: You can see this by looking at the expression for standard deviation. We have

$$\text{std}(x) = \sqrt{\frac{1}{N} \sum_{i=1}^{i=N} (x_i - \text{mean}(\{x\}))^2}.$$

Now, this means that

$$N(\text{std}(x))^2 = \sum_{i=1}^{i=N} (x_i - \text{mean}(\{x\}))^2.$$

But

$$\sum_{i=1}^{i=N} (x_i - \text{mean}(\{x\}))^2 \leq N \max_i (x_i - \text{mean}(\{x\}))^2$$

so

$$(\text{std}(x))^2 \leq \max_i (x_i - \text{mean}(\{x\}))^2.$$

Property 2.3: For any dataset, there must be at least one data item that is at least one standard deviation away from the mean.

Boxes 2.3.2 and 2.3.2 mean that the standard deviation is quite informative. Very little data is many standard deviations away from the mean; similarly, at least some of the data should be one or more standard deviations away from the mean. So the standard deviation tells us how data points are scattered about the mean.

Potential point of confusion: There is an ambiguity that comes up often here because two (very slightly) different numbers are called the standard deviation of a dataset. One — the one we use in this chapter — is an estimate of the scale of the data, as we describe it. The other differs from our expression very slightly; one computes

$$\sqrt{\frac{\sum_i (x_i - \text{mean}(\{x\}))^2}{N - 1}}$$

(notice the $N - 1$ for our N). If N is large, this number is basically the same as the number we compute, but for smaller N there is a difference that can be significant. Irritatingly, this number is also called the standard deviation; even more irritatingly, we will have to deal with it, but not yet. I mention it now because you may look up terms I have used, find this definition, and wonder. Don't worry - the N in our expressions is the right thing to use for what we're doing.

2.3.3 Variance

It turns out that thinking in terms of the square of the standard deviation, which is known as the **variance**, will allow us to generalize our summaries to apply to higher dimensional data.

Definition: 2.3 *Variance*

Assume we have a dataset $\{x\}$ of N data items, x_1, \dots, x_N . where $N > 1$. Their variance is:

$$\text{var}(\{x\}) = \frac{1}{N} \left(\sum_{i=1}^{i=N} (x_i - \text{mean}(\{x\}))^2 \right) = \text{mean}(\{(x_i - \text{mean}(\{x\}))^2\}).$$

One good way to think of the variance is as the mean-square error you would incur if you replaced each data item with the mean. Another is that it is the square of the standard deviation.

Properties of the Variance The properties of the variance follow from the fact that it is the square of the standard deviation. We have that:

- Translating data does not change the variance, i.e. $\text{var}(\{x + c\}) = \text{var}(\{x\})$.
- Scaling data scales the variance by a square of the scale, i.e. $\text{var}(\{kx\}) = k^2 \text{var}(\{x\})$.

While one could restate the other two properties of the standard deviation in terms of the variance, it isn't really natural to do so. The standard deviation is in the same units as the original data, and should be thought of as a scale. Because the variance is the square of the standard deviation, it isn't a natural scale (unless you take its square root!).

2.3.4 The Median

One problem with the mean is that it can be affected strongly by extreme values. Go back to the bar example, of section 2.3.1. Now Warren Buffett (or Bill Gates, or your favorite billionaire) walks in. What happened to the average net worth?

Assume your billionaire has net worth \$ 1, 000, 000, 000. Then the mean net worth suddenly has become

$$\frac{10 \times \$107,903 + \$1,000,000,000}{11} = \$91,007,184$$

But this mean isn't a very helpful summary of the people in the bar. It is probably more useful to think of the net worth data as ten people together with one billionaire. The billionaire is known as an **outlier**.

One way to get outliers is that a small number of data items are very different, due to minor effects you don't want to model. Another is that the data was misrecorded, or mistranscribed. Another possibility is that there is just too much variation in the data to summarize it well. For example, a small number of extremely wealthy people could change the average net worth of US residents

dramatically, as the example shows. An alternative to using a mean is to use a **median**.

Definition: 2.4 *Median*

The median of a set of data points is obtained by sorting the data points, and finding the point halfway along the list. If the list is of even length, it's usual to average the two numbers on either side of the middle. We write

$$\text{median}(\{x_i\})$$

for the operator that returns the median.

For example,

$$\text{median}(\{3, 5, 7\}) = 5,$$

$$\text{median}(\{3, 4, 5, 6, 7\}) = 5,$$

and

$$\text{median}(\{3, 4, 5, 6\}) = 4.5.$$

For much, but not all, data, you can expect that roughly half the data is smaller than the median, and roughly half is larger than the median. Sometimes this property fails. For example,

$$\text{median}(\{1, 2, 2, 2, 2, 2, 2, 3\}) = 2.$$

With this definition, the median of our list of net worths is \$107,835. If we insert the billionaire, the median becomes \$108,930. Notice by how little the number has changed — it remains an effective summary of the data.

Properties of the median You can think of the median of a dataset as giving the “middle” or “center” value. This means it is rather like the mean, which also gives a (slightly differently defined) “middle” or “center” value. The mean has the important properties that if you translate the dataset, the mean translates, and if you scale the dataset, the mean scales. The median has these properties, too:

- Translating data translates the median, i.e. $\text{median}(\{x + c\}) = \text{median}(\{x\}) + c$.
- Scaling data scales the median by the same scale, i.e. $\text{median}(\{kx\}) = k\text{median}(\{x\})$.

Each is easily proved, and proofs are relegated to the exercises.

2.3.5 Interquartile Range

Outliers can affect standard deviations severely, too. For our net worth data, the standard deviation without the billionaire is \$9265, but if we put the billionaire in there, it is $\$3.014 \times 10^8$. When the billionaire is in the dataset, all but one of

the data items lie about a third of a standard deviation away from the mean; the other one (the billionaire) is many standard deviations away from the mean. In this case, the standard deviation has done its work of informing us that there are huge changes in the data, but isn't really helpful.

The problem is this: describing the net worth data with billionaire as a having a mean of $\$9.101 \times 10^7$ with a standard deviation of $\$3.014 \times 10^8$ really isn't terribly helpful. Instead, the data really should be seen as a clump of values that are near \$100,000 and moderately close to one another, and one massive number (the billionaire outlier).

One thing we could do is simply remove the billionaire and compute mean and standard deviation. This isn't always easy to do, because it's often less obvious which points are outliers. An alternative is to follow the strategy we did when we used the median. Find a summary that describes scale, but is less affected by outliers than the standard deviation. This is the **interquartile range**; to define it, we need to define percentiles and quartiles, which are useful anyway.

Definition: 2.5 *Percentile*

The k 'th percentile is the value such that $k\%$ of the data is less than or equal to that value. We write $\text{percentile}(\{x\}, k)$ for the k 'th percentile of dataset $\{x\}$.

Definition: 2.6 *Quartiles*

The first quartile of the data is the value such that 25% of the data is less than or equal to that value (i.e. $\text{percentile}(\{x\}, 25)$). The second quartile of the data is the value such that 50% of the data is less than or equal to that value, which is usually the median (i.e. $\text{percentile}(\{x\}, 50)$). The third quartile of the data is the value such that 75% of the data is less than or equal to that value (i.e. $\text{percentile}(\{x\}, 75)$).

Definition: 2.7 *Interquartile Range*

The interquartile range of a dataset $\{x\}$ is $\text{iqr}\{x\} = \text{percentile}(\{x\}, 75) - \text{percentile}(\{x\}, 25)$.

Like the standard deviation, the interquartile range gives an estimate of how widely the data is spread out. But it is quite well-behaved in the presence of outliers. For our net worth data without the billionaire, the interquartile range is \$12350; with the billionaire, it is \$17710.

Properties of the interquartile range You can think of the interquartile range of a dataset as giving an estimate of the scale of the difference from the mean. This means it is rather like the standard deviation, which also gives a (slightly differently defined) scale. The standard deviation has the important properties that if you translate the dataset, the standard deviation translates, and if you scale the dataset, the standard deviation scales. The interquartile range has these properties, too:

- Translating data does not change the interquartile range, i.e. $\text{iqr}\{x + c\} = \text{iqr}\{x\}$.
- Scaling data scales the interquartile range by the same scale, i.e. $\text{iqr}\{kx\} = k^2\text{iqr}\{x\}$.

Each is easily proved, and proofs are relegated to the exercises.

2.3.6 Using Summaries Sensibly

One should be careful how one summarizes data. For example, the statement that “the average US family has 2.6 children” invites mockery (the example is from Andrew Vickers’ book *What is a p-value anyway?*), because you can’t have fractions of a child — no family has 2.6 children. A more accurate way to say things might be “the average of the number of children in a US family is 2.6”, but this is clumsy. What is going wrong here is the 2.6 is a mean, but the number of children in a family is a categorical variable. Reporting the mean of a categorical variable is often a bad idea, because you may never encounter this value (the 2.6 children). For a categorical variable, giving the median value and perhaps the interquartile range often makes much more sense than reporting the mean.

For continuous variables, reporting the mean is reasonable because you could expect to encounter a data item with this value, even if you haven’t seen one in the particular data set you have. It is sensible to look at both mean and median; if they’re significantly different, then there is probably something going on that is worth understanding. You’d want to plot the data using the methods of the next section before you decided what to report.

You should also be careful about how precisely numbers are reported (equivalently, the number of significant figures). Numerical and statistical software will produce very large numbers of digits freely, but not all are always useful. This is a particular nuisance in the case of the mean, because you might add many numbers, then divide by a large number; in this case, you will get many digits, but some might not be meaningful. For example, Vickers (ibid) describes a paper reporting the mean length of pregnancy as 32.833 weeks. That fifth digit suggests we know the mean length of pregnancy to about 0.001 weeks, or roughly 10 minutes. Neither medical interviewing nor people’s memory for past events is that detailed. Furthermore, when you interview them about embarrassing topics, people quite often lie. There is no prospect of knowing this number with this precision.

People regularly report silly numbers of digits because it is easy to miss the harm caused by doing so. But the harm is there: you are implying to other people, and to yourself, that you know something more accurately than you do. At some point, someone will suffer for it.

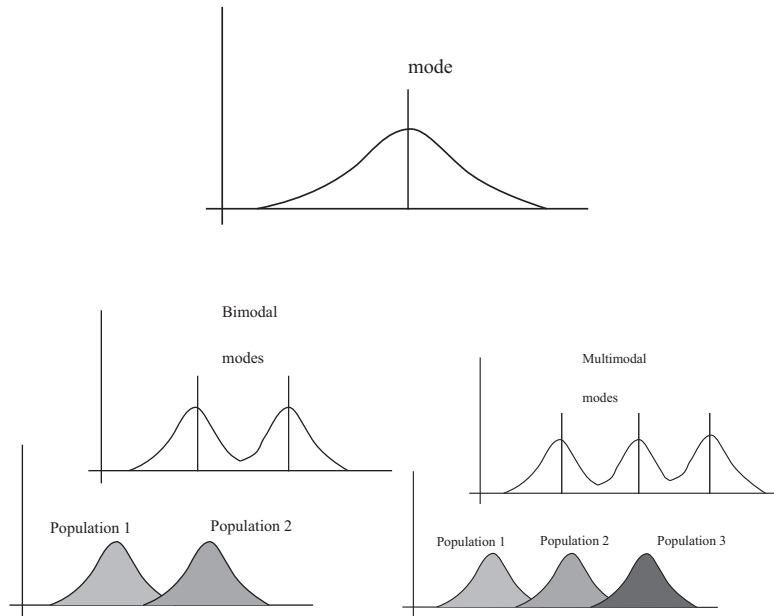


FIGURE 2.4: Many histograms are unimodal, like the example on the **top**; there is one peak, or mode. Some are bimodal (two peaks; **bottom left**) or even multimodal (two or more peaks; **bottom right**). One common reason (but not the only reason) is that there are actually two populations being conflated in the histograms. For example, measuring adult heights might result in a bimodal histogram, if male and female heights were slightly different. As another example, measuring the weight of dogs might result in a multimodal histogram if you did not distinguish between breeds (eg chihuahua, terrier, german shepherd, pyrenean mountain dog, etc.).

2.4 PLOTS AND SUMMARIES

Knowing the mean, standard deviation, median and interquartile range of a dataset gives us some information about what its histogram might look like. In fact, the summaries give us a language in which to describe a variety of characteristic properties of histograms that are worth knowing about (Section 2.4.1). Quite remarkably, many different datasets have the same shape of histogram (Section 2.4.2). For such data, we know roughly what percentage of data items are how far from the mean.

Complex datasets can be difficult to interpret with histograms alone, because it is hard to compare many histograms by eye. Section 2.4.3 describes a clever plot of various summaries of datasets that makes it easier to compare many cases.

2.4.1 Some Properties of Histograms

The **tails** of a histogram are the relatively uncommon values that are significantly larger (resp. smaller) than the value at the peak (which is sometimes called the **mode**). A histogram is **unimodal** if there is only one peak; if there are more than one, it is **multimodal**, with the special term **bimodal** sometimes being used for

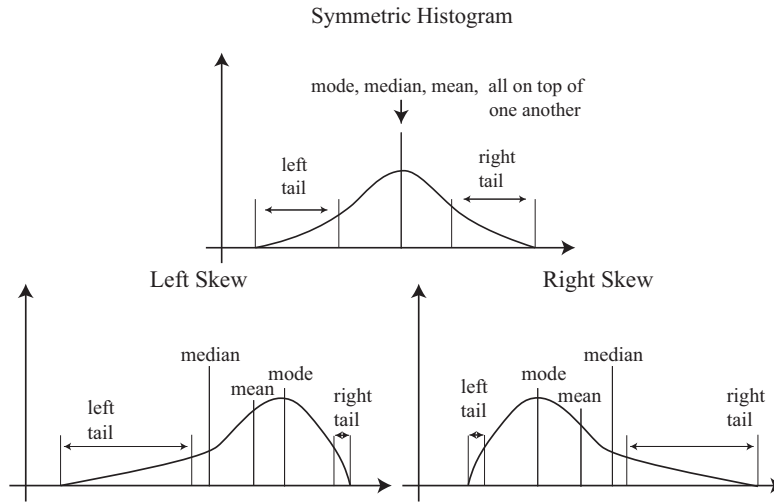


FIGURE 2.5: On the **top**, an example of a symmetric histogram, showing its tails (relatively uncommon values that are significantly larger or smaller than the peak or mode). **Lower left**, a sketch of a left-skewed histogram. Here there are few large values, but some very small values that occur with significant frequency. We say the left tail is “long”, and that the histogram is left skewed (confusingly, this means the main bump is to the right). **Lower right**, a sketch of a right-skewed histogram. Here there are few small values, but some very large values that occur with significant frequency. We say the right tail is “long”, and that the histogram is right skewed (confusingly, this means the main bump is to the left).

the case where there are two peaks (Figure 2.4). The histograms we have seen have been relatively symmetric, where the left and right tails are about as long as one another. Another way to think about this is that values a lot larger than the mean are about as common as values a lot smaller than the mean. Not all data is symmetric. In some datasets, one or another tail is longer (figure 2.5). This effect is called **skew**.

Skew appears often in real data. SOCR (the Statistics Online Computational Resource) publishes a number of datasets. Here we discuss a dataset of citations to faculty publications. For each of five UCLA faculty members, SOCR collected the number of times each of the papers they had authored had been cited by other authors (data at http://wiki.stat.ucla.edu/socr/index.php/SOCR_Data_Dinov_072108_H_Index_Pubs). Generally, a small number of papers get many citations, and many papers get few citations. We see this pattern in the histograms of citation numbers (figure 2.6). These are very different from (say) the body temperature pictures. In the citation histograms, there are many data items that have very few citations, and few that have many citations. This means that the right tail of the histogram is longer, so the histogram is skewed to the right.

One way to check for skewness is to look at the histogram; another is to compare mean and median (though this is not foolproof). For the first citation

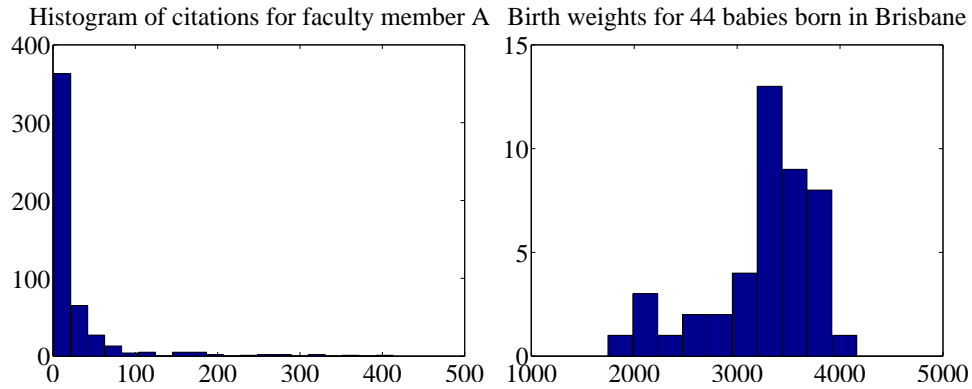


FIGURE 2.6: *On the left*, a histogram of citations for a faculty member, from data at http://wiki.stat.ucla.edu/socr/index.php/SOCR_Data_Dinov_072108_H_Index_Pubs. Very few publications have many citations, and many publications have few. This means the histogram is strongly right-skewed. *On the right*, a histogram of birth weights for 44 babies borne in Brisbane in 1997. This histogram looks slightly left-skewed.

histogram, the mean is 24.7 and the median is 7.5; for the second, the mean is 24.4, and the median is 11. In each case, the mean is a lot bigger than the median. Recall the definition of the median (form a ranked list of the data points, and find the point halfway along the list). For much data, the result is larger than about half of the data set and smaller than about half the dataset. So if the median is quite small compared to the mean, then there are many small data items and a small number of data items that are large — the right tail is longer, so the histogram is skewed to the right.

Left-skewed data also occurs; figure 2.6 shows a histogram of the birth weights of 44 babies born in Brisbane, in 1997 (from http://www.amstat.org/publications/jse/jse_data_archive.htm). This data appears to be somewhat left-skewed, as birth weights can be a lot smaller than the mean, but tend not to be much larger than the mean.

Skewed data is often, but not always, the result of constraints. For example, good obstetrical practice tries to ensure that very large birth weights are rare (birth is typically induced before the baby gets too heavy), but it may be quite hard to avoid some small birth weights. This could skew birth weights to the left (because large babies will get born, but will not be as heavy as they could be if obstetricians had not interfered). Similarly, income data can be skewed to the right by the fact that income is always positive. Test mark data is often skewed — whether to right or left depends on the circumstances — by the fact that there is a largest possible mark and a smallest possible mark.

2.4.2 Standard Coordinates and Normal Data

It is useful to look at lots of histograms, because it is often possible to get some useful insights about data. However, in their current form, histograms are hard to

compare. This is because each is in a different set of units. A histogram for length data will consist of boxes whose horizontal units are, say, metres; a histogram for mass data will consist of boxes whose horizontal units are in, say, kilograms. Furthermore, these histograms typically span different ranges.

We can make histograms comparable by (a) estimating the “location” of the plot on the horizontal axis and (b) estimating the “scale” of the plot. The location is given by the mean, and the scale by the standard deviation. We could then normalize the data by subtracting the location (mean) and dividing by the standard deviation (scale). The resulting values are unitless, and have zero mean. They are often known as **standard coordinates**.

Definition: 2.8 *Standard coordinates*

Assume we have a dataset $\{x\}$ of N data items, x_1, \dots, x_N . We represent these data items in standard coordinates by computing

$$\hat{x}_i = \frac{(x_i - \text{mean}(\{x\}))}{\text{std}(x)}.$$

We write $\{\hat{x}\}$ for a dataset that happens to be in standard coordinates.

Standard coordinates have some important properties. Assume we have N data items. Write x_i for the i 'th data item, and \hat{x}_i for the i 'th data item in standard coordinates (I sometimes refer to these as “normalized data items”). Then we have

$$\text{mean}(\{\hat{x}\}) = 0.$$

We also have that

$$\text{std}(\hat{x}) = 1.$$

An extremely important fact about data is that, for many kinds of data, histograms of these standard coordinates look the same. Many completely different datasets produce a histogram that, in standard coordinates, has a very specific appearance. It is symmetric, unimodal; it looks like a narrow bump. If there were enough data points and the histogram boxes were small enough, the curve would look like the curve in figure 2.7. This phenomenon is so important that data of this form has a special name.

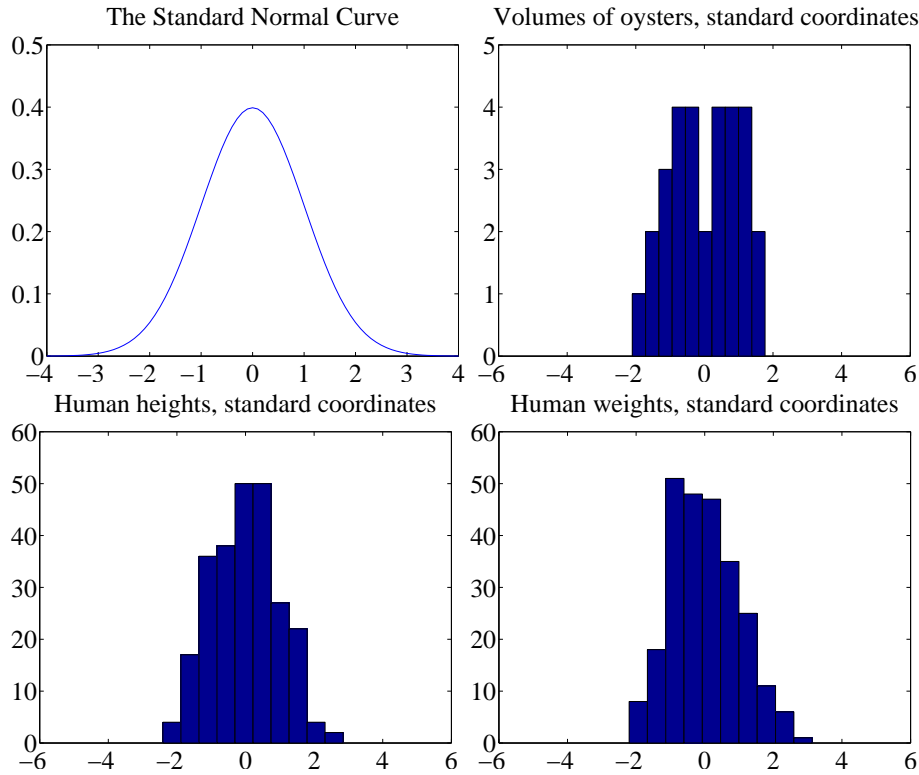


FIGURE 2.7: Data is standard normal data when its histogram takes a stylized, bell-shaped form, plotted above. One usually requires a lot of data and very small histogram boxes for this form to be reproduced closely. Nonetheless, the histogram for normal data is unimodal (has a single bump) and is symmetric; the tails fall off fairly fast, and there are few data items that are many standard deviations from the mean. Many quite different data sets have histograms that are similar to the normal curve; I show three such datasets here.

Definition: 2.9 *Standard normal data*

Data is **standard normal data** if, when we have a great deal of data, the histogram is a close approximation to the **standard normal curve**. This curve is given by

$$y(x) = \frac{1}{\sqrt{2\pi}} e^{(-x^2/2)}$$

(which is shown in figure 2.7).

Definition: 2.10 *Normal data*

Data is **normal data** if, when we subtract the mean and divide by the standard deviation (i.e. compute standard coordinates), it becomes standard normal data.

It is not always easy to tell whether data is normal or not, and there are a variety of tests one can use, which we discuss later. However, there are many examples of normal data. Figure 2.7 shows a diverse variety of data sets, plotted as histograms in standard coordinates. These include: the volumes of 30 oysters (from http://www.amstat.org/publications/jse/jse_data_archive.htm; look for 30oysters.dat.txt); human heights (from <http://www2.stetson.edu/~jrasp/data.htm>; look for bodyfat.xls, with two outliers removed); and human weights (from <http://www2.stetson.edu/~jrasp/data.htm>; look for bodyfat.xls, with two outliers removed).

Properties of normal data For the moment, assume we know that a dataset is normal. Then we expect it to have the following properties:

- If we normalize it, its histogram will be close to the standard normal curve. This means, among other things, that the data is not significantly skewed.
- About 68% of the data lie within one standard deviation of the mean. We will prove this later.
- About 95% of the data lie within two standard deviations of the mean. We will prove this later.
- About 99% of the data lie within three standard deviations of the mean. We will prove this later.

In turn, these properties imply that data that contains outliers (points many standard deviations away from the mean) is not normal. This is usually a very safe assumption. It is quite common to model a dataset by excluding a small number of outliers, then modelling the remaining data as normal. For example, if I exclude two outliers from the height and weight data from <http://www2.stetson.edu/~jrasp/data.htm>, the data looks pretty close to normal.

2.4.3 Boxplots

It is usually hard to compare multiple histograms by eye. One problem with comparing histograms is the amount of space they take up on a plot, because each histogram involves multiple vertical bars. This means it is hard to plot multiple overlapping histograms cleanly. If you plot each one on a separate figure, you have to handle a large number of separate figures; either you print them too small to see enough detail, or you have to keep flipping over pages.

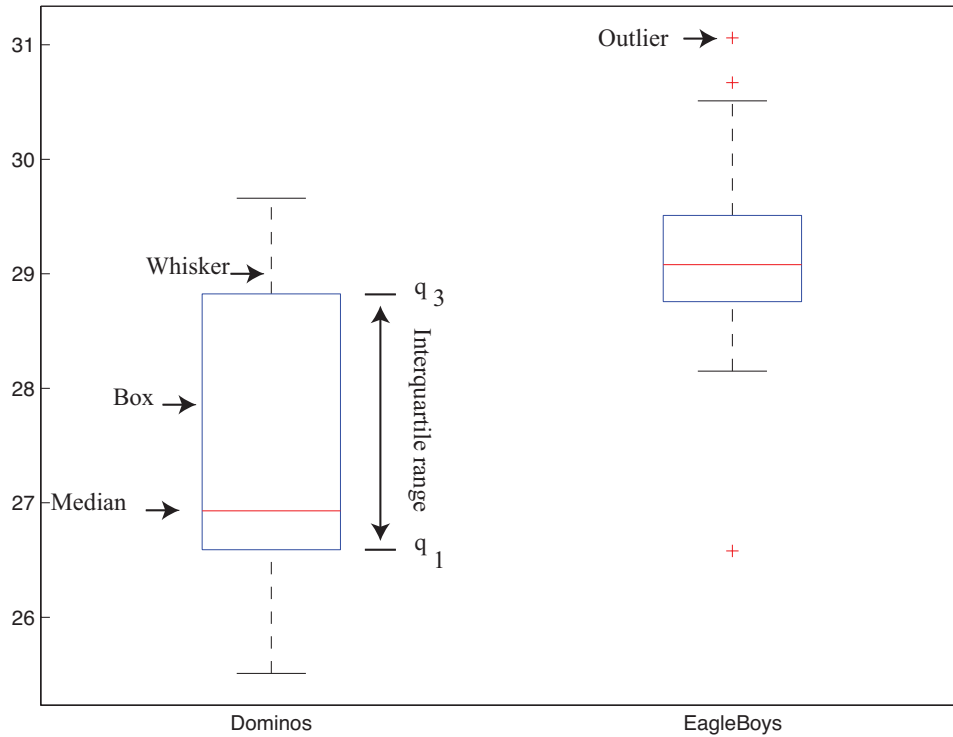


FIGURE 2.8: A boxplot showing the box, the median, the whiskers and two outliers. Notice that we can compare the two datasets rather easily; the next section explains the comparison.

A **boxplot** is a way to plot data that simplifies comparison. A boxplot displays a dataset as a vertical picture. There is a vertical box whose height corresponds to the interquartile range of the data (the width is just to make the figure easy to interpret). Then there is a horizontal line for the median; and the behavior of the rest of the data is indicated with whiskers and/or outlier markers. This means that each dataset makes is represented by a vertical structure, making it easy to show multiple datasets on one plot *and interpret the plot* (Figure 2.8).

To build a boxplot, we first plot a box that runs from the first to the third quartile. We then show the median with a horizontal line. We then decide which data items should be outliers. A variety of rules are possible; for the plots I show, I used the rule that data items that are larger than $q_3 + 1.5(q_3 - q_1)$ or smaller than $q_1 - 1.5(q_3 - q_1)$, are outliers. This criterion looks for data items that are more than one and a half interquartile ranges above the third quartile, or more than one and a half interquartile ranges below the first quartile.

Once we have identified outliers, we plot these with a special symbol (crosses in the plots I show). We then plot whiskers, which show the range of non-outlier data. We draw a whisker from q_1 to the smallest data item that is not an outlier, and from q_3 to the largest data item that is not an outlier. While all this sounds

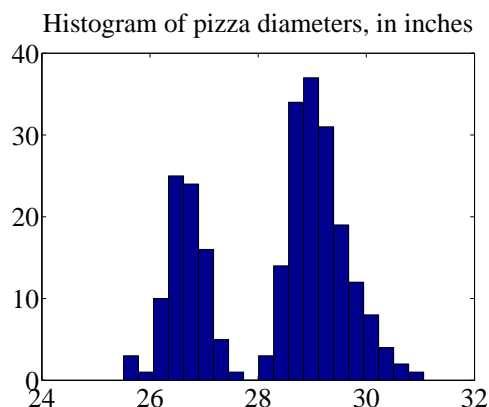


FIGURE 2.9: A histogram of pizza diameters from the dataset described in the text. Notice that there seem to be two populations.

complicated, any reasonable programming environment will have a function that will do it for you. Figure 2.8 shows an example boxplot. Notice that the rich graphical structure means it is quite straightforward to compare two histograms.

2.5 WHOSE IS BIGGER? INVESTIGATING AUSTRALIAN PIZZAS

At http://www.amstat.org/publications/jse/jse_data_archive.htm, you will find a dataset giving the diameter of pizzas, measured in Australia (search for the word “pizza”). This website also gives the backstory for this dataset. Apparently, EagleBoys pizza claims that their pizzas are always bigger than Dominos pizzas, and published a set of measurements to support this claim (the measurements were available at <http://www.eagleboys.com.au/realsizepizza> as of Feb 2012, but seem not to be there anymore).

Whose pizzas are bigger? and why? A histogram of all the pizza sizes appears in figure 2.9. We would not expect every pizza produced by a restaurant to have exactly the same diameter, but the diameters are probably pretty close to one another, and pretty close to some standard value. This would suggest that we’d expect to see a histogram which looks like a single, rather narrow, bump about a mean. This is not what we see in figure 2.9 — instead, there are two bumps, which suggests two populations of pizzas. This isn’t particularly surprising, because we know that some pizzas come from EagleBoys and some from Dominos.

If you look more closely at the data in the dataset, you will notice that each data item is tagged with the company it comes from. We can now easily plot conditional histograms, conditioning on the company that the pizza came from. These appear in figure 2.10. Notice that EagleBoys pizzas seem to follow the pattern we expect — the diameters are clustered tightly around one value — but Dominos pizzas do not seem to be like that. This is reflected in a boxplot (figure 2.11), which shows the range of Dominos pizza sizes is surprisingly large, and that EagleBoys pizza sizes have several large outliers. There is more to understand about this data. The dataset contains labels for the type of crust and the type of topping — perhaps

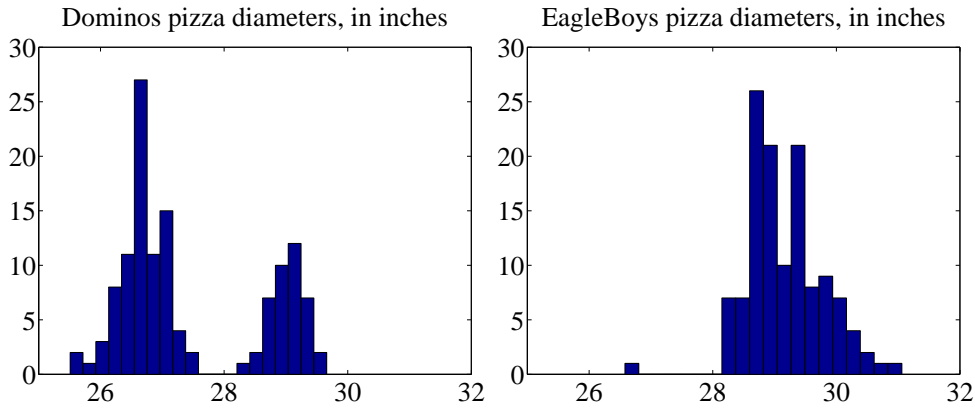


FIGURE 2.10: *On the left, the class-conditional histogram of Dominos pizza diameters from the pizza data set; on the right, the class-conditional histogram of EagleBoys pizza diameters. Notice that EagleBoys pizzas seem to follow the pattern we expect — the diameters are clustered tightly around a mean, and there is a small standard deviation — but Dominos pizzas do not seem to be like that. There is more to understand about this data.*

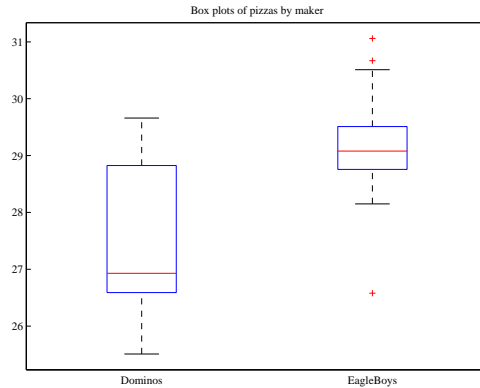


FIGURE 2.11: *Boxplots of the pizza data, comparing EagleBoys and Dominos pizza. There are several curiosities here: why is the range for Dominos so large (25.5-29)? EagleBoys has a smaller range, but has several substantial outliers; why? One would expect pizza manufacturers to try and control diameter fairly closely, because pizzas that are too small present risks (annoying customers; publicity; hostile advertising) and pizzas that are too large should affect profits.*

these properties affect the size of the pizza?

EagleBoys produces DeepPan, MidCrust and ThinCrust pizzas, and Dominos produces DeepPan, ClassicCrust and ThinNCrispy pizzas. This may have something to do with the observed patterns, but comparing six histograms by eye is unattractive. A boxplot is the right way to compare these cases (figure 2.12). The

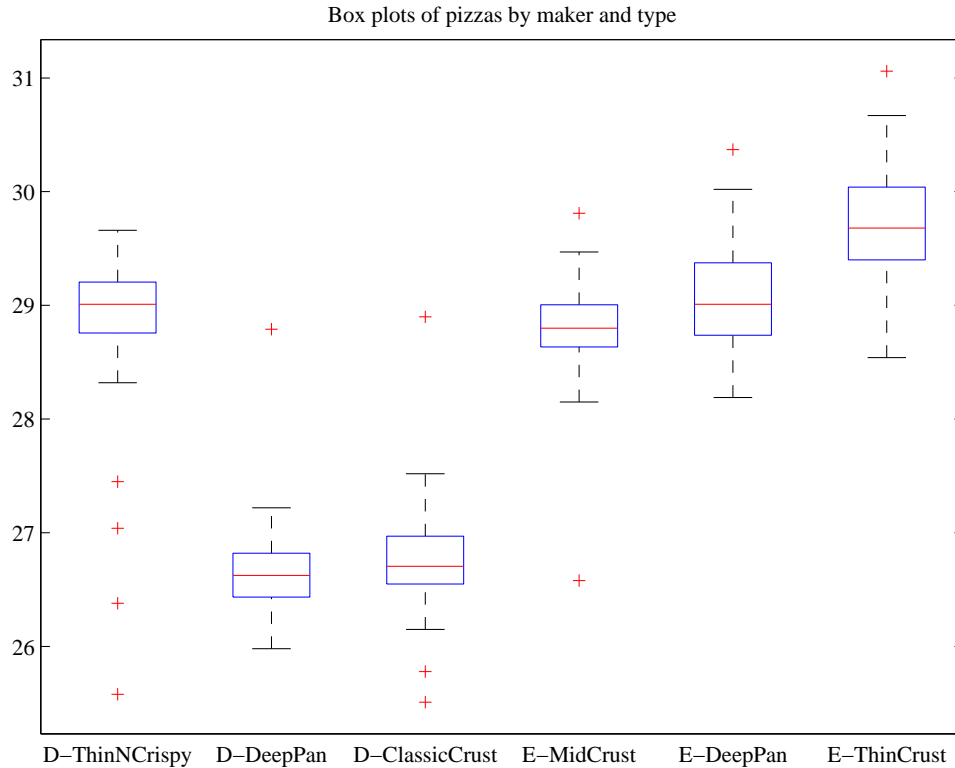


FIGURE 2.12: *Boxplots for the pizza data, broken out by type (thin crust, etc.).*

boxplot gives some more insight into the data. Dominos thin crust appear to have a narrow range of diameters (with several outliers), where the median pizza is rather larger than either the deep pan or the classic crust pizza. EagleBoys pizzas all have a range of diameters that is (a) rather similar across the types and (b) rather a lot like the Dominos thin crust. There are outliers, but few for each type.

Another possibility is that the variation in size is explained by the topping. We can compare types and toppings by producing a set of conditional boxplots (i.e. the diameters for each type and each topping). This leads to rather a lot of boxes (figure 2.13), but they're still easy to compare by eye. The main difficulty is that the labels on the plot have to be shortened. I made labels using the first letter from the manufacturer ("D" or "E"); the first letter from the crust type (previous paragraph); and the first and last letter of the topping. Toppings for Dominos are: Hawaiian; Supreme; BBQMeatlovers. For EagleBoys, toppings are: Hawaiian; SuperSupremo; and BBQMeatlovers. This gives the labels: 'DCBs'; (Dominos; ClassicCrust; BBQMeatlovers); 'DCHn'; 'DCSe'; 'DDBs'; 'DDHn'; 'DDSe'; 'DTBs'; 'DTHn'; 'DTSe'; 'EDBs'; 'EDHn'; 'EDSo'; 'EMBs'; 'EMHn'; 'EMSo'; 'ETBs'; 'ETHn'; 'ETSo'. Figure 2.13 suggests that the topping isn't what is important, but the crust (group the boxplots by eye).

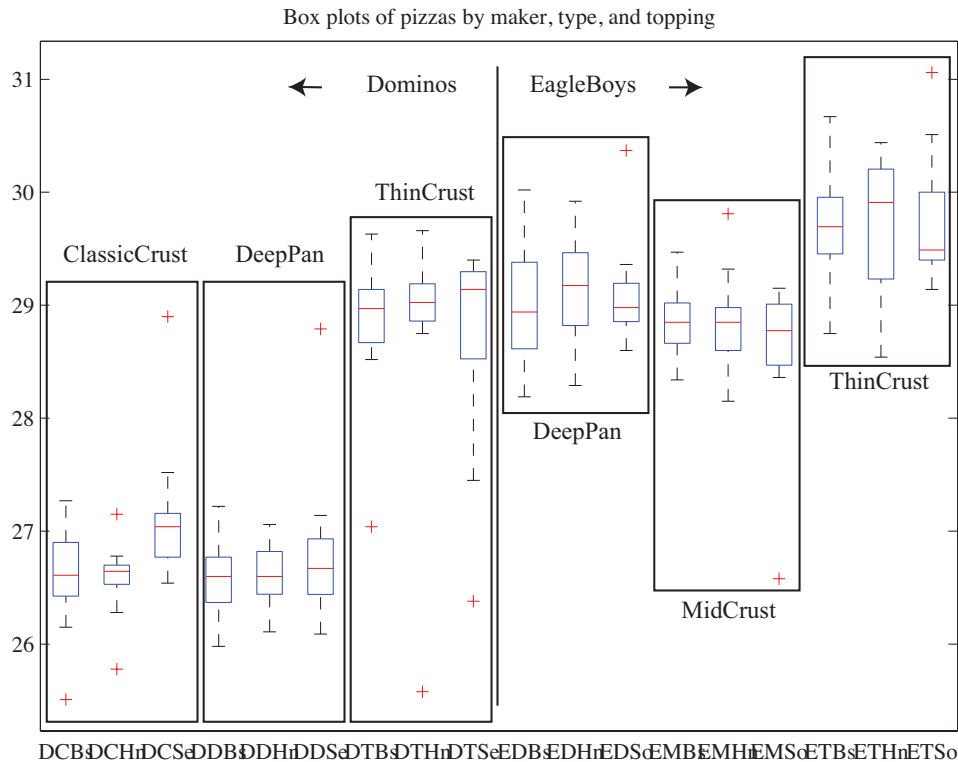


FIGURE 2.13: The pizzas are now broken up by topping as well as crust type (look at the source for the meaning of the names). I have separated Dominos from EagleBoys with a vertical line, and grouped each crust type with a box. It looks as though the issue is not the type of topping, but the crust. EagleBoys seems to have tighter control over the size of the final pizza.

What could be going on here? One possible explanation is that EagleBoys have tighter control over the size of the final pizza. One way this could happen is that all EagleBoys pizzas start the same size and shrink the same amount in baking, whereas all Dominos pizzas start a standard diameter, but different Dominos crusts shrink differently in baking. Another way is that Dominos makes different size crusts for different types, but that the cooks sometimes get confused. Yet another possibility is that Dominos controls portions by the mass of dough (so thin crust diameters tend to be larger), but EagleBoys controls by the diameter of the crust.

You should notice that this is more than just a fun story. If you were a manager at a pizza firm, you'd need to make choices about how to control costs. Labor costs, rent, and portion control (i.e. how much pizza, topping, etc. a customer gets for their money) are the main thing to worry about. If the same kind of pizza has a wide range of diameters, you have a problem, because some customers are getting too much (which affects your profit) or too little (which means they might call someone else). But making more regular pizzas might require more skilled (and so

more expensive) labor. The fact that Dominos and EagleBoys seem to be following different strategies successfully suggests that more than one strategy might work. But you can't choose if you don't know what's happening. As I said at the start, "what's going on here?" is perhaps the single most useful question anyone can ask.

2.6 WHAT YOU MUST REMEMBER

You should be able to:

- Plot a bar chart for a dataset.
- Plot a histogram for a dataset.
- Tell whether the histogram is skewed or not, and in which direction.
- Plot a box plot for one or several datasets.
- Interpret a box plot.

You should remember:

- The definition and properties of the mean.
- The definition and properties of the standard deviation.
- The definition and properties of the variance.
- The definition and properties of the median.
- The definition and properties of the interquartile range.

PROBLEMS

- 2.1. Show that $\text{mean}(\{kx\}) = k\text{mean}(\{x\})$ by substituting into the definition.
- 2.2. Show that $\text{mean}(\{x + c\}) = \text{mean}(\{x\}) + c$ by substituting into the definition.
- 2.3. Show that $\sum_{i=1}^N (x_i - \text{mean}(\{x\})) = 0$ by substituting into the definition.
- 2.4. Show that $\text{std}(x + c) = \text{std}(x)$ by substituting into the definition (you'll need to recall the properties of the mean to do this).
- 2.5. Show that $\text{std}(kx) = k\text{std}(x)$ by substituting into the definition (you'll need to recall the properties of the mean to do this).
- 2.6. Show that $\text{median}(\{x + c\}) = \text{median}(\{x\}) + c$ by substituting into the definition.
- 2.7. Show that $\text{median}(\{kx\}) = k\text{median}(\{x\})$ by substituting into the definition.
- 2.8. Show that $\text{iqr}\{x + c\} = \text{iqr}\{x\}$ by substituting into the definition.
- 2.9. Show that $\text{iqr}\{kx\} = k\text{iqr}\{x\}$ by substituting into the definition.

PROGRAMMING EXERCISES

- 2.10. You can find a data set showing the number of barrels of oil produced, per year for the years 1880-1984 at <http://lib.stat.cmu.edu/DASL/Datafiles/Oilproduction.html>. Is a mean a useful summary of this dataset? Why?

- 2.11.** You can find a dataset giving the cost (in 1976 US dollars), number of megawatts, and year of construction of a set of nuclear power plants at <http://lib.stat.cmu.edu/DASL/Datafiles/NuclearPlants.html>.
- (a) Are there outliers in this data?
 - (b) What is the mean cost of a power plant? What is the standard deviation?
 - (c) What is the mean cost per megawatt? What is the standard deviation?
 - (d) Plot a histogram of the cost per megawatt. Is it skewed? Why?
- 2.12.** You can find a dataset giving the sodium content and calorie content of three types of hot dog at <http://lib.stat.cmu.edu/DASL/Datafiles/Hotdogs.html>. The types are Beef, Poultry, and Meat (a rather disturbingly vague label). Use class-conditional histograms to compare these three types of hot dog with respect to sodium content and calories.
- 2.13.** You will find a dataset giving (among other things) the number of 3 or more syllable words in advertising copy appearing in magazines at <http://lib.stat.cmu.edu/DASL/Datafiles/magadsdat.html>. The magazines are grouped by the education level of their readers; the groups are 1, 2, and 3 (the variable is called GRP in the data).
- (a) Use a boxplot to compare the number of three or more syllable words for the ads in magazines in these three groups. What do you see?
 - (b) Use a boxplot to compare the number of sentences appearing in the ads in magazines in these three groups. What do you see?

CHAPTER 3

Intermezzo - Programming Tools

Matlab is a programming environment widely used in numerical analysis and computer vision circles, among other. I will use Matlab in examples here, because I'm fairly fluent in Matlab, and give some guidelines. Many universities, including UIUC, have free student licenses for Matlab, and there are numerous reference books and tutorials you can look at for details of syntax, etc. Matlab's focuses on representations of vectors and matrices. The syntax is rather like Fortran, except that there are some special matrix and vector operations. Fluent Matlab programmers replace iterations with these operations to get faster code; some people are very good at this sort of trick.

An alternative widely used in statistical circles is R, which is an open-source language rather like S (a paid license statistics package). I didn't know much R at start of writing, and haven't managed to pick up the kind of fluency required to prepare figures in it, so I won't show R code here. But you should be aware of it as an alternative.

If you want to know what a Matlab command does, you can use `help` or `doc`. For example, you could type `doc sum`. In this chapter, I'll show code I used to make some of the figures and examples in chapter one, to give you some examples for Matlab code.

Listing 3.1 shows how I made the bar chart of genders in Figure 2.1. There are some points to notice here. I cut-and-pasted the data from the web page (<http://lib.stat.cmu.edu/DASL/Datafiles/PopularKids.html>), put the result into a spreadsheet to make a .xls file, saved that, then read it into Matlab using `xlsread`. I did this because it meant I didn't have to write any kind of file reader. You can learn more about `xlsread` by doing `doc xlsread`. It isn't perfect — or, at least, the one on my Mac isn't — because it doesn't like some kinds of format, and it seems not to like very long files. It was enough for this purpose. It produces a cell array (`doc cell`) for the text fields, and an ordinary array for the numeric fields. The cell array is an array whose entries can contain objects like strings, etc. You will find the question of how to get a dataset from the form in which you find it into a form you want in your programming environment to be a persistent nuisance. It's usually approached with quick-and-dirty solutions like this one.

You should have noticed one nice thing. Once I've actually gotten my data into a reasonable form, making the bar chart is easy (`bar` does the trick). It's more trouble to get the title on the figure and the fonts right, etc. than to make the figure. This is true for histograms, as well (listing 2.2). In this case, the data was just a vector, so I cut it out from the web page and stuck it into the code. Look at `doc hist`, and notice you can change the number of bars and also the centers of the bars if you wish.

I have listed useful commands for computing summaries below. To give you an example of normalizing data, there is listing 3.3.

Matlab will do boxplots, too. In listing 3.4, I show the Matlab code I used to make Figure 2.11.

Boxplots can get pretty elaborate, and putting the right marker in the right place gets interesting. I show the code I used to make Figure 2.13, which required a certain amount of fooling around with strings to get the right marker and put it in the right place.

Some useful commands:

- `bar` makes a bar chart.
- `hist` makes a histogram.
- `xlsread` will read many spreadsheet files.
- `mean` computes the mean of a dataset represented as a column vector. If you give it an array, it will compute the mean of each column.
- `std` computes the standard deviation of a dataset represented as a column vector. If you give it an array, it will compute the standard deviation of each column.
- `var` computes the variance of a dataset represented as a column vector. If you give it an array, it will compute the variance of each column.
- `median` computes the median of a dataset represented as a column vector. If you give it an array, it will compute the median of each column.
- `prctile` can compute a set of percentiles for a dataset represented as a column vector. If you give it an array, it will compute those percentiles for each column. Note it takes two argument; you should do `doc prctile`.
- `boxplot` will produce a boxplot; there are many arguments and tricks, so you should do `doc boxplot`.

Listing 3.1: Matlab code used to make the gender bar chart in figure 2.1

```

cd( '~/Current/courses/ProbCourse/SomeData/MatlabCode' );
[num, txt]=xlsread( '../Datasets/schooldata.xls' );
%
set(0, 'DefaultAxesFontName', 'TimesₒNewₒRoman')
set(0, 'DefaultAxesFontSize', 28)
% Change default text fonts for figures
% so they print well
set(0, 'DefaultTextFontname', 'TimesₒNewₒRoman')
set(0, 'DefaultTextFontSize', 28)
wv=zeros( size(num, 1), 1);
wv2=zeros( size(num, 1), 1);
for i=1:size(num, 1)
    if strcmp(txt(i, 7), 'Sports')
        wv(i)=1;
    elseif strcmp(txt(i, 7), 'Grades')
        wv(i)=2;
    elseif strcmp(txt(i, 7), 'Popular')
        wv(i)=3;
    end
    if strcmp(txt(i, 1), 'boy')
        wv2(i)=1;
    elseif strcmp(txt(i, 1), 'girl')
        wv2(i)=2;
    end
end
% really stupid parser
genders=[sum(wv2==1), sum(wv2==2)];
% sum knows what to do with a vector
figure(1); % sets the window in which the figure
% will be put
clf % clears the figure
bar(genders)
set(gca, 'XTick', [1, 2]);
set(gca, 'XTickLabel', {'boy', 'girl'});
title('Numberₒofₒchildrenₒofₒeachₒgender');
figure(1);
% brings it to the front
%print -depsc2 ../Figures/childgenderbar.eps
% this makes an eps figure and puts in the right
% place

```


Listing 3.2: Matlab code used to make the cheese histogram in figure 2.2

```

cheeses=[12.3, 20.9, 39, 47.9, 5.6, 25.9, ...
        37.3, 21.9, 18.1, 21, 34.9, 57.2, 0.7, ...
        25.9, 54.9, 40.9, 15.9, 6.4, 18, 38.9, ...
        14, 15.2, 32, 56.7, 16.8, 11.6, 26.5, ...
        0.7, 13.4, 5.5];
figure(2);
hist(cheeses, [5, 15, 25, 35, 45, 55])
figure(2);
axis([0, 70, 0, 15])
ylabel('Number_of_data_items');
xlabel('Cheese_goodness_in_cheese_goodness_units');
title('Histogram_of_cheese_goodness_score_for_30_cheeses');
set(gca, 'XTick', [0:7]*10);
set(gca, 'YTick', [0:2:15]);
figure(2);

```

Listing 3.3: Matlab code used to make the normalized oyster volume histogram in figure 2.7

```

set(0, 'DefaultAxesFontName', 'Times_New_Roman')
set(0, 'DefaultAxesFontSize', 28)
% Change default text fonts.
set(0, 'DefaultTextFontname', 'Times_New_Roman')
set(0, 'DefaultTextFontSize', 28)
cd('~\Current\Courses\Probcourse\SomeData\DataSets\');
% this is where I keep the file, but you may have it somewhere else
[num2, txt, raw]=xlsread('30oystersdata.xls');
% index weight volume
wdat=num2(:, 4); % this is the volume field
wm=mean(wdat);
wsd=std(wdat);
figure(1);
hist((wdat-wm)/wsd); figure(1)
axis([-6 6 0 5])
% sets the axes of the plot
title('Volumes_of_oysters_standard_coordinates');
print -depsc2 noysters.eps

```

Listing 3.4: Matlab code used to make the boxplot in figure 2.11

```

% pizza size data
set(0, 'DefaultAxesFontName', 'Times_New_Roman')
set(0, 'DefaultAxesFontSize', 11)
% Change default text fonts.
set(0, 'DefaultTextFontname', 'Times_New_Roman')
set(0, 'DefaultTextFontSize', 11)
cd('~ / Current / Courses / Probcourse / SomeData / DataSets / ');
[num, txt, raw]=xlsread('cleanpizzasize.xls');
ndat=size(num, 1);
figure(1); boxplot(num(:, 5), txt(:, 2)); figure(1)
title('Box_plots_of_pizzas_by_maker');
print -depsc2 pmakerboxes.eps

```

Listing 3.5: Matlab code used to make the boxplot in figure 2.13

```

% pizza size data
set(0, 'DefaultAxesFontName', 'Times_New_Roman')
set(0, 'DefaultAxesFontSize', 11)
% Change default text fonts.
set(0, 'DefaultTextFontname', 'Times_New_Roman')
set(0, 'DefaultTextFontSize', 11)
cd('~ / Current / Courses / Probcourse / SomeData / DataSets / ');
[num, txt, raw]=xlsread('cleanpizzasize.xls');
ndat=size(num, 1);
t2=txt(:, 1);
for i=1:ndat
    foo=txt(i, 2); bar=foo{1}; c1=bar(1);
    % this gets a letter for the maker
    foo=txt(i, 3); bar=foo{1}; c2=bar(1);
    % this gets a letter for the crust
    foo=txt(i, 4); bar=foo{1}; c3=bar(1); c4=bar(end);
    % this gets first, last letter for the topping
    t2(i)={strcat(c1, c2, c3, c4)};
end
figure(2); boxplot(num(:, 5), t2, 'grouporder', {'DCBs', 'DCHn', ...
'DCSe', 'DDBs', 'DDHn', 'DDSe', 'DTBs', 'DTHn', 'DTSe', ...
'EDBs', 'EDHn', 'EDSo', 'EMBs', 'EMHn', 'EMSo', 'ETBs', ...
'ETHn', 'ETSo'}); figure(2);
title('Box_plots_of_pizzas_by_maker, _type, _and _topping');

```

CHAPTER 4

Looking at Relationships

We think of a dataset as a collection of d -tuples (a d -tuple is an ordered list of d elements). For example, the Chase and Dunner dataset had entries for Gender; Grade; Age; Race; Urban/Rural; School; Goals; Grades; Sports; Looks; and Money (so it consisted of 11-tuples). The previous chapter explored methods to visualize and summarize a set of values obtained by extracting a single element from each tuple. For example, I could visualize the heights or the weights of a population (as in Figure 2.7). But I could say nothing about the relationship between the height and weight. In this chapter, we will look at methods to visualize and summarize the relationships between pairs of elements of a dataset.

4.1 PLOTTING 2D DATA

We take a dataset, choose two different entries, and extract the corresponding elements from each tuple. The result is a dataset consisting of 2-tuples, and we think of this as a two dimensional dataset. The first step is to plot this dataset in a way that reveals relationships. The topic of how best to plot data fills many books, and we can only scratch the surface here. Categorical data can be particularly tricky, because there are a variety of choices we can make, and the usefulness of each tends to depend on the dataset and to some extent on one's cleverness in graphic design (section 4.1.1).

For some continuous data, we can plot the one entry as a function of the other (so, for example, our tuples might consist of the date and the number of robberies; or the year and the price of lynx pelts; and so on, section 4.1.2).

Mostly, we use a simple device, called a scatter plot. Using and thinking about scatter plots will reveal a great deal about the relationships between our data items (section 4.1.3).

4.1.1 Categorical Data, Counts, and Charts

Categorical data is a bit special. Assume we have a dataset with several categorical descriptions of each data item. One way to plot this data is to think of it as belonging to a richer set of categories. Assume the dataset has categorical descriptions, which are not ordinal. Then we can construct a new set of categories by looking at each of the cases for each of the descriptions. For example, in the Chase and Dunner data of table 2.2, our new categories would be: “boy-sports”; “girl-sports”; “boy-popular”; “girl-popular”; “boy-grades”; and “girl-grades”. A large set of categories like this can result in a poor bar chart, though, because there may be too many bars to group the bars successfully. Figure 4.1 shows such a bar chart. Notice that it is hard to group categories by eye to compare; for example, you can see that slightly more girls think grades are important than boys do, but to do so you need to compare two bars that are separated by two other bars. An alternative is a **pie**

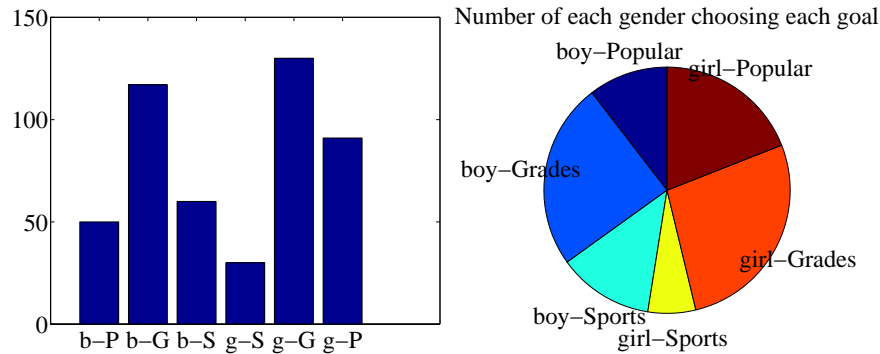


FIGURE 4.1: I sorted the children in the Chase and Dunner study into six categories (two genders by three goals), and counted the number of children that fell into each cell. I then produced the bar chart on the **left**, which shows the number of children of each gender, selecting each goal. On the **right**, a pie chart of this information. I have organized the pie chart so it is easy to compare boys and girls by eye — start at the top; going down on the left side are boy goals, and on the right side are girl goals. Comparing the size of the corresponding wedges allows you to tell what goals (resp. girls) identify with more or less often.

chart, where a circle is divided into sections whose angle is proportional to the size of the data item. You can think of the circle as a pie, and each section as a slice of pie. Figure 4.1 shows a pie chart, where each section is proportional to the number of students in its category. In this case, I’ve used my judgement to lay the categories out in a way that makes comparisons easy. I’m not aware of any tight algorithm for doing this, though.

Pie charts have problems, because it is hard to judge small differences in area accurately by eye. For example, from the pie chart in figure 4.1, it’s hard to tell that the “boy-sports” category is slightly bigger than the “boy-popular” category (try it; check using the bar chart). For either kind of chart, it is quite important to think about *what* you plot. For example, the plot of figure 4.1 shows the total number of respondents, and if you refer to figure 2.1, you will notice that there are slightly more girls in the study. Is the *percentage* of boys who think grades are important smaller (or larger) than the *percentage* of girls who think so? you can’t tell from these plots, and you’d have to plot the percentages instead.

An alternative to a pie chart that is very useful for two dimensional data is a **heat map**. This is a method of displaying a matrix as an image. Each entry of the matrix is mapped to a color, and the matrix is represented as an image. For the Chase and Dunner study, I constructed a matrix where each row corresponds to a choice of “sports”, “grades”, or “popular”, and each column corresponds to a choice of “boy” or “girl”. Each entry contains the count of data items of that type. Zero values are represented as white; the largest values as red; and as the value increases, we use an increasingly saturated pink. This plot is shown in figure 4.2

If the categorical data is ordinal, the ordering offers some hints for making a good plot. For example, imagine we are building a user interface. We build an

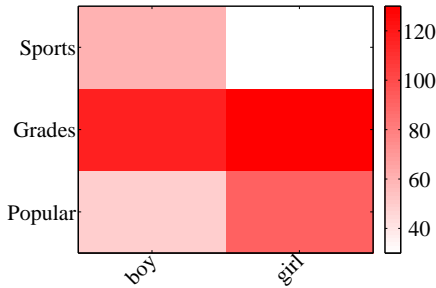


FIGURE 4.2: A heat map of the Chase and Danner data. The color of each cell corresponds to the count of the number of elements of that type. The colorbar at the side gives the correspondence between color and count. You can see at a glance that the number of boys and girls who prefer grades is about the same; that about the same number of boys prefer sports and popularity, with sports showing a mild lead; and that more girls prefer popularity to sports.

	-2	-1	0	1	2
-2	24	5	0	0	1
-1	6	12	3	0	0
0	2	4	13	6	0
1	0	0	3	13	2
2	0	0	0	1	5

TABLE 4.1: I simulated data representing user evaluations of a user interface. Each cell in the table on the **left** contains the count of users rating “ease of use” (horizontal, on a scale of -2 -very bad- to 2 -very good) vs. “enjoyability” (vertical, same scale). Users who found the interface hard to use did not like using it either. While this data is categorical, it’s also ordinal, so that the order of the cells is determined. It wouldn’t make sense, for example, to reorder the columns of the table or the rows of the table.

initial version, and collect some users, asking each to rate the interface on scales for “ease of use” (-2, -1, 0, 1, 2, running from bad to good) and “enjoyability” (again, -2, -1, 0, 1, 2, running from bad to good). It is natural to build a 5x5 table, where each cell represents a pair of “ease of use” and “enjoyability” values. We then count the number of users in each cell, and build graphical representations of this table. One natural representation is a **3D bar chart**, where each bar sits on its cell in the 2D table, and the height of the bars is given by the number of elements in the cell. Table 4.1 shows a table and figure 4.3 shows a 3D bar chart for some simulated data. The main difficulty with a 3D bar chart is that some bars are hidden behind others. This is a regular nuisance. You can improve things by using an interactive tool to rotate the chart to get a nice view, but this doesn’t always work. Heatmaps don’t suffer from this problem (Figure 4.3), another reason they are a good choice.

Counts of user responses for a user interface

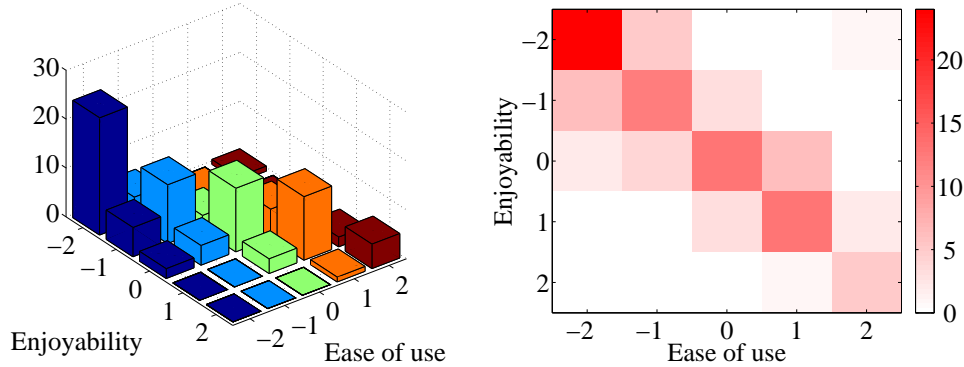


FIGURE 4.3: *On the left*, a 3D bar chart of the data. The height of each bar is given by the number of users in each cell. This figure immediately reveals that users who found the interface hard to use did not like using it either. However, some of the bars at the back are hidden, so some structure might be hard to infer. *On the right*, a heat map of this data. Again, this figure immediately reveals that users who found the interface hard to use did not like using it either. It's more apparent that everyone disliked the interface, though, and it's clear that there is no important hidden structure.

4.1.2 Series

Sometimes one component of a dataset gives a natural ordering to the data. For example, we might have a dataset giving the maximum rainfall for each day of the year. We could record this either by using a two-dimensional representation, where one dimension is the number of the day and the other is the temperature, or with a convention where the i 'th data item is the rainfall on the i 'th day. For example, at <http://lib.stat.cmu.edu/DASL/Datafiles/timeseriesdat.html>, you can find four datasets indexed in this way. It is natural to plot data like this as a function of time. From this dataset, I extracted data giving the number of burglaries each month in a Chicago suburb, Hyde Park. I have plotted part this data in Figure 4.4 (I left out the data to do with treatment effects). It is natural to plot a graph of the burglaries as a function of time (in this case, the number of the month). The plot shows each data point explicitly. I also told the plotting software to draw lines joining data points, because burglaries do not all happen on a specific day. The lines suggest, reasonably enough, the rate at which burglaries are happening between data points.

As another example, at <http://lib.stat.cmu.edu/datasets/Andrews/> you can find a dataset that records the number of lynx pelts traded to the Hudson's Bay company and the price paid for each pelt. This version of the dataset appeared first in table 3.2 of *Data: a Collection of Problems from many Fields for the Student and Research Worker* by D.F. Andrews and A.M. Herzberg, published by Springer in 1985. I have plotted it in figure 4.4. The dataset is famous, because it shows a periodic behavior in the number of pelts (which is a good proxy for the number

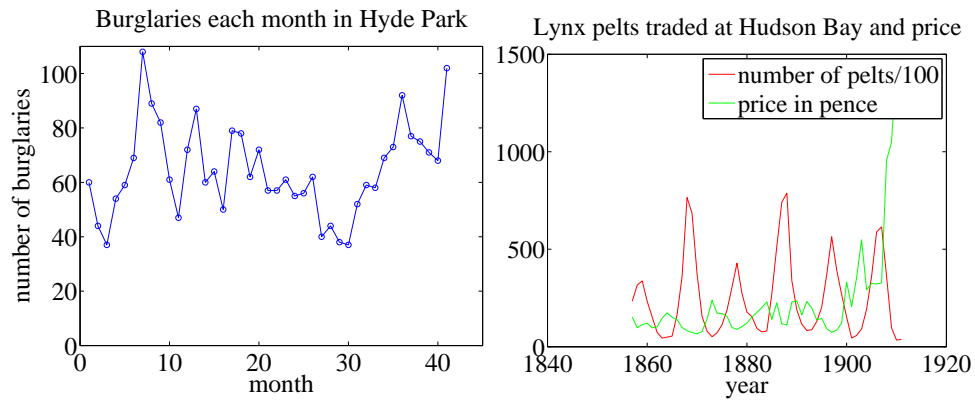


FIGURE 4.4: **Left**, the number of burglaries in Hyde Park, by month. **Right**, a plot of the number of lynx pelts traded at Hudson Bay and of the price paid per pelt, as a function of the year. Notice the scale, and the legend box (the number of pelts is scaled by 100).

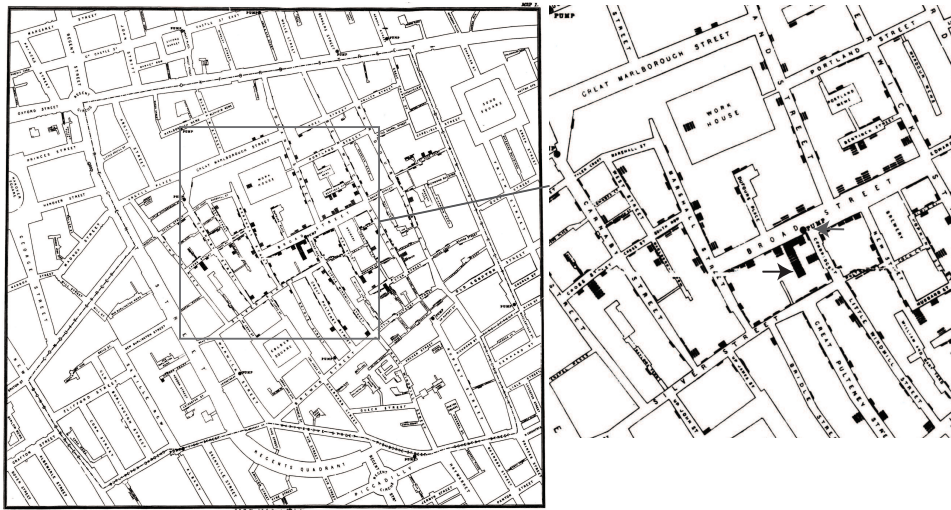


FIGURE 4.5: Snow's scatter plot of cholera deaths on the **left**. Each cholera death is plotted as a small bar on the house in which the bar occurred (for example, the black arrow points to one stack of these bars, indicating many deaths, in the detail on the **right**). Notice the fairly clear pattern of many deaths close to the Broad street pump (grey arrow in the detail), and fewer deaths further away (where it was harder to get water from the pump).

of lynx), which is interpreted as a result of predator-prey interactions. Lynx eat rabbits. When there are many rabbits, lynx kittens thrive, and soon there will be many lynx; but then they eat most of the rabbits, and starve, at which point

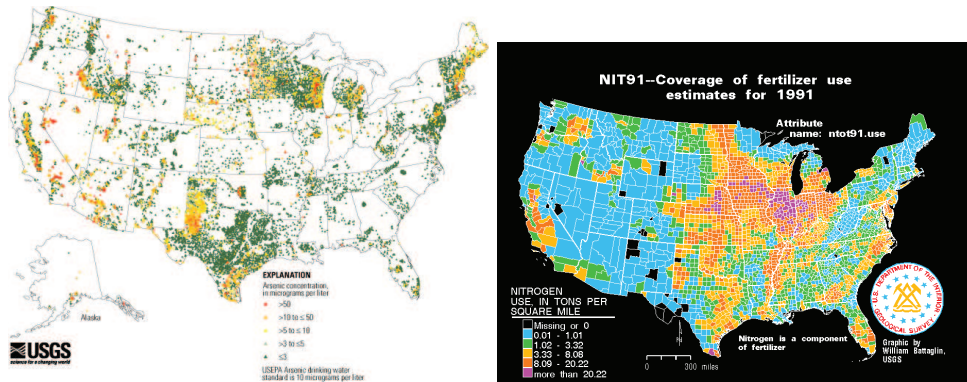


FIGURE 4.6: **Left**, a scatter plot of arsenic levels in US groundwater, prepared by the US Geological Survey (you can find the data at http://water.usgs.gov/GIS/metadata/usgswrd/XML/arsenic_map.xml). Here the shape and color of each marker shows the amount of arsenic, and the spatial distribution of the markers shows where the wells were sampled. **Right**, the usage of Nitrogen (a component of fertilizer) by US county in 1991, prepared by the US Geological Survey (you can find the data at <http://water.usgs.gov/GIS/metadata/usgswrd/XML/nit91.xml>). In this variant of a scatter plot (which usually takes specialized software to prepare) one fills each region with a color indicating the data in that region.

the rabbit population rockets. You should also notice that after about 1900, prices seem to have gone up rather quickly. I don't know why this is. There is also some suggestion, as there should be, that prices are low when there are many pelts, and high when there are few.

4.1.3 Scatter Plots for Spatial Data

It isn't always natural to plot data as a function. For example, in a dataset containing the temperature and blood pressure of a set of patients, there is no reason to believe that temperature is a function of blood pressure, or the other way round. Two people could have the same temperature, and different blood pressures, or vice-versa. As another example, we could be interested in what causes people to die of cholera. We have data indicating *where* each person died in a particular outbreak. It isn't helpful to try and plot such data as a function.

The **scatter plot** is a powerful way to deal with this situation. In the first instance, assume that our data points actually describe points on the a real map. Then, to make a scatter plot, we make a mark on the map at a place indicated by each data point. What the mark looks like, and how we place it, depends on the particular dataset, what we are looking for, how much we are willing to work with complex tools, and our sense of graphic design.

Figure 4.5 is an extremely famous scatter plot, due to John Snow. Snow — one of the founders of epidemiology — used a scatter plot to reason about a cholera outbreak centered on the Broad Street pump in London in 1854. At that time,

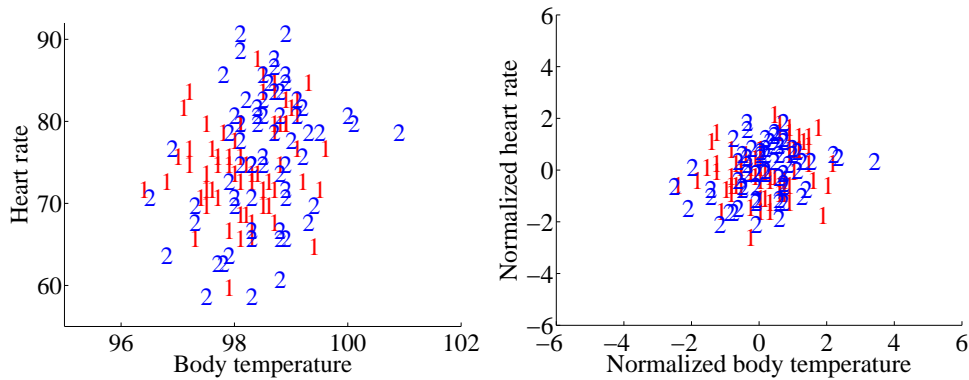


FIGURE 4.7: A scatter plot of body temperature against heart rate, from the dataset at <http://www2.stetson.edu/~jrasp/data.htm>; *normtemp.xls*. I have separated the two genders by plotting a different symbol for each (though I don't know which gender is indicated by which letter); if you view this in color, the differences in color makes for a greater separation of the scatter. This picture suggests, but doesn't conclusively establish, that there isn't much dependence between temperature and heart rate, and any dependence between temperature and heart rate isn't affected by gender.

the mechanism that causes cholera was not known. Snow plotted cholera deaths as little bars (more bars, more deaths) on the location of the house where the death occurred. More bars means more deaths, fewer bars means fewer deaths. There are more bars per block close to the pump, and few far away. This plot offers quite strong evidence of an association between the pump and death from cholera. Snow used this scatter plot as evidence that cholera was associated with water, and that the Broad Street pump was the source of the tainted water.

Figure 4.6 shows a scatter plot of arsenic levels in groundwater for the United States, prepared by the US Geological Survey. The data set was collected by Focazio and others in 2000; by Welch and others in 2000; and then updated by Ryker 2001. It can be found at http://water.usgs.gov/GIS/metadata/usgswrd/XML/arsenic_map.xml. One variant of a scatter plot that is particularly useful for geographic data occurs when one fills regions on a map with different colors, following the data in that region. Figure 4.6 shows the nitrogen usage by US county in 1991; again, this figure was prepared by the US Geological Survey.

4.1.4 Exposing Relationships with Scatter Plots

Scatter plots are natural for geographic data, but a scatter plot is a useful, simple tool for ferreting out associations in other kinds of data as well. Now we need some notation. Assume we have a dataset $\{x\}$ of N data items, x_1, \dots, x_N . Each data item is a d dimensional vector (so its components are numbers). We wish to investigate the relationship between two components of the dataset. For example, we might be interested in the 7'th and the 13'th component of the dataset. We will produce a two-dimensional plot, one dimension for each component. It does

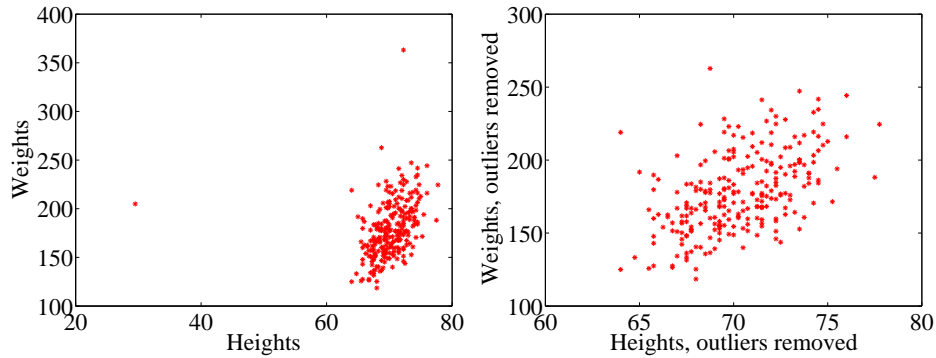


FIGURE 4.8: A scatter plots of weight against height, from the dataset at <http://www2.stetson.edu/~jrasp/data.htm>. **Left:** Notice how two outliers dominate the picture, and to show the outliers, the rest of the data has had to be bunched up. **Right** shows the data with the outliers removed. The structure is now somewhat clearer.

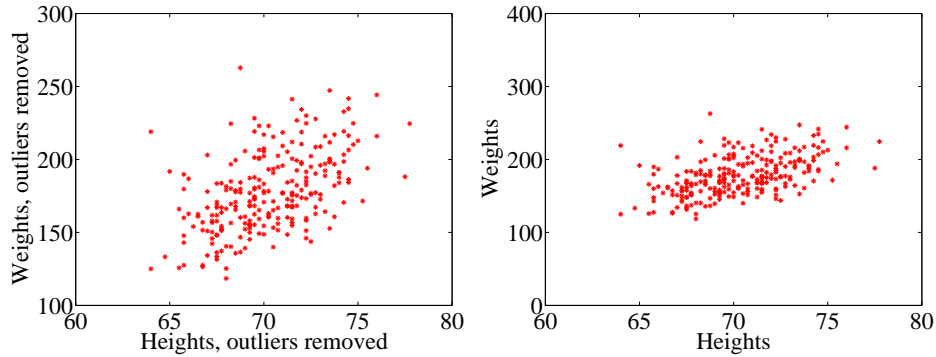


FIGURE 4.9: Scatter plots of weight against height, from the dataset at <http://www2.stetson.edu/~jrasp/data.htm>. **Left:** data with two outliers removed, as in figure 4.8. **Right:** this data, rescaled slightly. Notice how the data looks less spread out. But there is no difference between the datasets. Instead, your eye is easily confused by a change of scale.

not really matter which component is plotted on the x -coordinate and which on the y -coordinate (though it will be some pages before this is clear). But it is very difficult to write sensibly without talking about the x and y coordinates.

We will make a two-dimensional dataset out of the components that interest us. We must choose which component goes first in the resulting 2-vector. We will plot this component on the x -coordinate (and we refer to it as the x -coordinate), and to the other component as the y -coordinate. This is just to make it easier to describe what is going on; there's no important idea here. It really will not matter which is x and which is y . The two components make a dataset $\{\mathbf{x}_i\} = \{(x_i, y_i)\}$. To produce a scatter plot of this data, we plot a small shape at the location of each

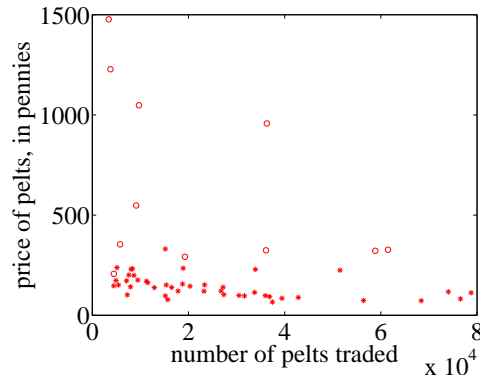


FIGURE 4.10: A scatter plot of the price of lynx pelts against the number of pelts. I have plotted data for 1901 to the end of the series as circles, and the rest of the data as *’s. It is quite hard to draw any conclusion from this data, because the scale is confusing. Furthermore, the data from 1900 on behaves quite differently from the other data.

data item.

Such scatter plots are very revealing. For example, figure 4.7 shows a scatter plot of body temperature against heart rate for humans. In this dataset, the gender of the subject was recorded (as “1” or “2” — I don’t know which is which), and so I have plotted a “1” at each data point with gender “1”, and so on. Looking at the data suggests there isn’t much difference between the blob of “1” labels and the blob of “2” labels, which suggests that females and males are about the same in this respect.

The scale used for a scatter plot matters. For example, plotting lengths in meters gives a very different scatter from plotting lengths in millimeters. Figure 4.8 shows two scatter plots of weight against height. Each plot is from the same dataset, but one is scaled so as to show two outliers. Keeping these outliers means that the rest of the data looks quite concentrated, just because the axes are in large units. In the other plot, the axis scale has changed (so you can’t see the outliers), but the data looks more scattered. This may or may not be a misrepresentation. Figure 4.9 compares the data with outliers removed, with the same plot on a somewhat different set of axes. One plot looks as though increasing height corresponds to increasing weight; the other looks as though it doesn’t. This is purely due to deceptive scaling — each plot shows the same dataset.

Dubious data can also contribute to scaling problems. Recall that, in figure 4.4, price data before and after 1900 appeared to behave differently. Figure 4.10 shows a scatter plot of the lynx data, where I have plotted number of pelts against price. I plotted the post-1900 data as circles, and the rest as asterisks. Notice how the circles seem to form a quite different figure, which supports the suggestion that something interesting happened around 1900. The scatter plot does not seem to support the idea that prices go up when supply goes down, which is puzzling, because this is a pretty reliable idea. This turns out to be a scale effect. Scale is an

important nuisance, and it's easy to get misled by scale effects. The way to avoid the problem is to plot in standard coordinates.

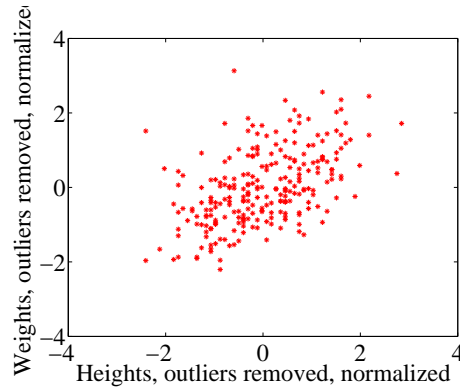


FIGURE 4.11: A normalized scatter plot of weight against height, from the dataset at <http://www2.stetson.edu/~jrasp/data.htm>. Now you can see that someone who is a standard deviation taller than the mean will tend to be somewhat heavier than the mean too.

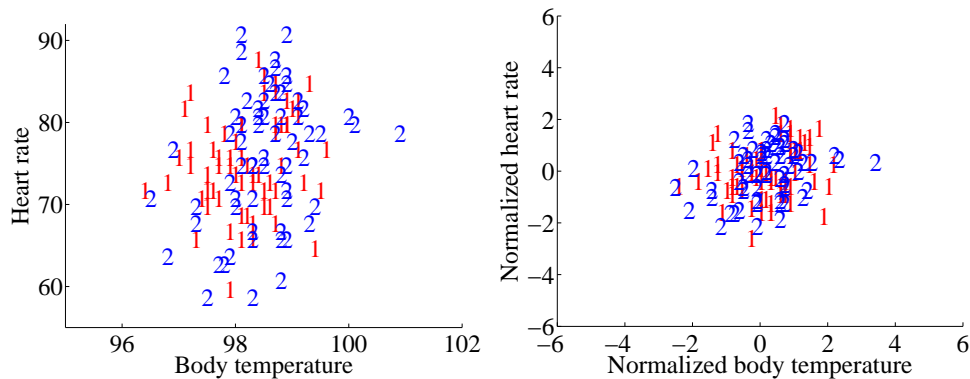


FIGURE 4.12: **Left:** A scatter plot of body temperature against heart rate, from the dataset at <http://www2.stetson.edu/~jrasp/data.htm>; `normtemp.xls`. I have separated the two genders by plotting a different symbol for each (though I don't know which gender is indicated by which letter); if you view this in color, the differences in color makes for a greater separation of the scatter. This picture suggests, but doesn't conclusively establish, that there isn't much dependence between temperature and heart rate, and any dependence between temperature and heart rate isn't affected by gender. The scatter plot of the normalized data, in standard coordinates, on the **right** supports this view.

A natural solution to problems with scale is to normalize the x and y coordinates of the two-dimensional data to standard coordinates. We can normalize

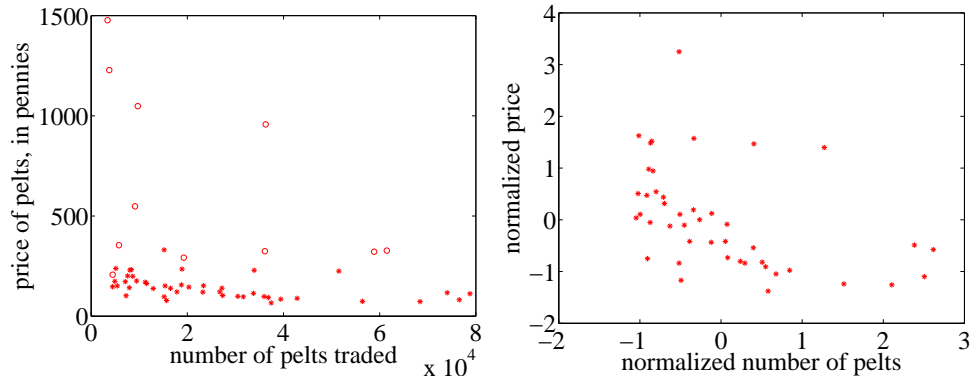


FIGURE 4.13: **Left:** A scatter plot of the price of lynx pelts against the number of pelts (this is a repeat of figure 4.10 for reference). I have plotted data for 1901 to the end of the series as circles, and the rest of the data as *'s. It is quite hard to draw any conclusion from this data, because the scale is confusing. **Right:** A scatter plot of the price of pelts against the number of pelts for lynx pelts. I excluded data for 1901 to the end of the series, and then normalized both price and number of pelts. Notice that there is now a distinct trend; when there are fewer pelts, they are more expensive, and when there are more, they are cheaper.

without worrying about the dimension of the data — we normalize each dimension independently by subtracting the mean of that dimension and dividing by the standard deviation of that dimension. We continue to use the convention of writing the normalized x coordinate as \hat{x} and the normalized y coordinate as \hat{y} . So, for example, we can write $\hat{x}_j = (x_j - \text{mean}(\{x\}))/\text{std}(x)$ for the \hat{x} value of the j 'th data item in normalized coordinates. Normalizing shows us the dataset on a standard scale. Once we have done this, it is quite straightforward to read off simple relationships between variables from a scatter plot.

4.2 CORRELATION

The simplest, and most important, relationship to look for in a scatter plot is this: when \hat{x} increases, does \hat{y} tend to increase, decrease, or stay the same? This is straightforward to spot in a normalized scatter plot, because each case produces a very clear shape on the scatter plot. Any relationship is called **correlation** (we will see later how to measure this), and the three cases are: positive correlation, which means that larger \hat{x} values tend to appear with larger \hat{y} values; zero correlation, which means no relationship; and negative correlation, which means that larger \hat{x} values tend to appear with smaller \hat{y} values. I have shown these cases together in one figure using a real data example (Figure 4.15), so you can compare the appearance of the plots.

Positive correlation occurs when larger \hat{x} values tend to appear with larger \hat{y} values. This means that data points with with small (i.e. negative with large magnitude) \hat{x} values must have small \hat{y} values, otherwise the mean of \hat{x} (resp. \hat{y}) would be too big. In turn, this means that the scatter plot should look like a “smear”

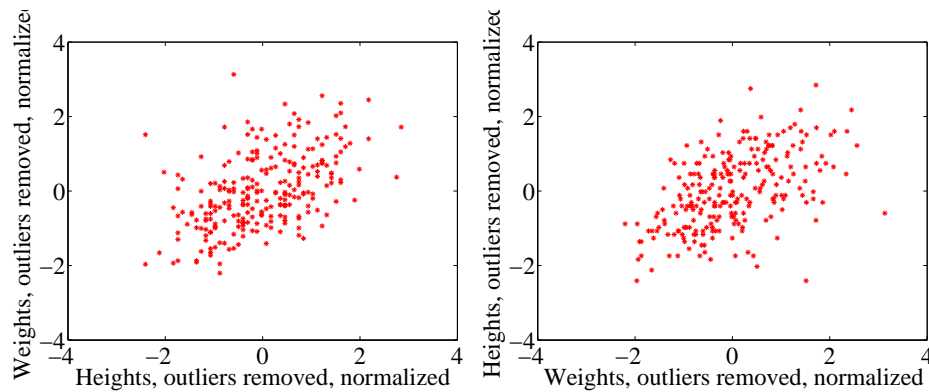


FIGURE 4.14: On the **left**, a normalized scatter plot of weight (y -coordinate) against height (x -coordinate). On the **right**, a scatter plot of height (y -coordinate) against weight (x -coordinate). I've put these plots next to one another so you don't have to mentally rotate (which is what you should usually do).

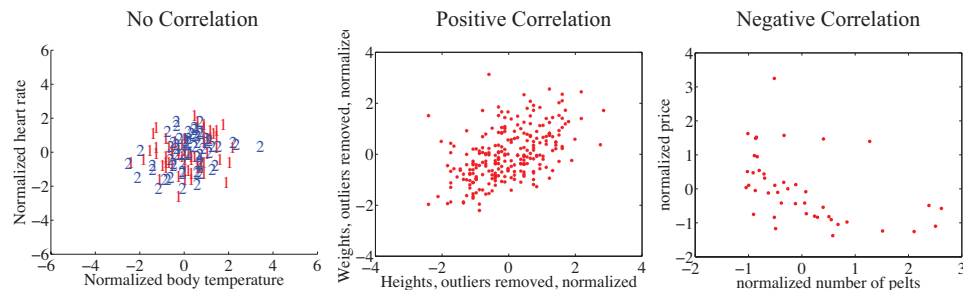


FIGURE 4.15: The three kinds of scatter plot are less clean for real data than for our idealized examples. Here I used the body temperature vs heart rate data for the zero correlation; the height-weight data for positive correlation; and the lynx data for negative correlation. The pictures aren't idealized — real data tends to be messy — but you can still see the basic structures.

of data from the bottom left of the graph to the top right. The smear might be broad or narrow, depending on some details we'll discuss below. Figure 4.11 shows normalized scatter plots of weight against height, and of body temperature against heart rate. In the weight-height plot, you can clearly see that individuals who are higher tend to weigh more. The important word here is “tend” — taller people could be lighter, but mostly they tend not to be. Notice, also, that I did NOT say that they weighed more *because* they were taller, but only that they tend to be heavier.

Zero correlation occurs when there is no relationship. This produces a characteristic shape in a scatter plot, but it takes a moment to understand why. If there really is no relationship, then knowing \hat{x} will tell you nothing about \hat{y} . All we know is that $\text{mean}(\{\hat{y}\}) = 0$, and $\text{var}(\{\hat{y}\}) = 1$. Our value of \hat{y} should have

this mean and this variance, but it doesn't depend on \hat{x} in any way. This is enough information to predict what the plot will look like. We know that $\text{mean}(\{\hat{x}\}) = 0$ and $\text{var}(\{\hat{x}\}) = 1$; so there will be many data points with \hat{x} value close to zero, and few with a much larger or much smaller \hat{x} value. The same applies to \hat{y} . Now consider the data points in a strip of \hat{x} values. If this strip is far away from the origin, there will be few data points in the strip, because there aren't many big \hat{x} values. If there is no relationship, we don't expect to see large or small \hat{y} values in this strip, because there are few data points in the strip and because large or small \hat{y} values are uncommon — we see them only if there are many data points, and then seldom. So for a strip with \hat{x} close to zero, we might see large \hat{y} values; but for one that is far away, we expect to see small \hat{y} values. We should see a blob, centered at the origin. In the temperature-heart rate plot of figure 4.12, it looks as though nothing of much significance is happening. The average heart rate seems to be about the same for people who run warm or who run cool. There is probably not much relationship here.

Negative correlation occurs when larger \hat{x} values tend to appear with smaller \hat{y} values. This means that data points with with small \hat{x} values must have large \hat{y} values, otherwise the mean of \hat{x} (resp. \hat{y}) would be too big. In turn, this means that the scatter plot should look like a “smear” of data from the top left of the graph to the bottom right. The smear might be broad or narrow, depending on some details we'll discuss below. Figure 4.13 shows a normalized scatter plot of the lynx pelt-price data, where I have excluded the data from 1901 on. I did so because there seemed to be some other effect operating to drive prices up, which was inconsistent with the rest of the series. This plot suggests that when there were more pelts, prices were lower, as one would expect.

Notice that leaving out data, as I did here, should be done with care. If you exclude every data point that might disagree with your hypothesis, you may miss the fact that you are wrong. Leaving out data is an essential component of many kinds of fraud. You should always reveal whether you have excluded data, and why, to allow the reader to judge the evidence.

The correlation is not affected by which variable is plotted on the x -axis and which is plotted on the y -axis. Figure 4.14 compares a plot of height against weight to one of weight against height. Usually, one just does this by rotating the page, or by imagining the new picture. The left plot tells you that data points with higher height value tend to have higher weight value; the right plot tells you that data points with higher weight value tend to have higher height value — i.e. the plots tell you the same thing. It doesn't really matter which one you look at. Again, the important word is “tend” — the plot doesn't tell you anything about *why*, it just tells you that when one variable is larger the other tends to be, too.

4.2.1 The Correlation Coefficient

Consider a normalized data set of N two-dimensional vectors. We can write the i 'th data point *in standard coordinates* (\hat{x}_i, \hat{y}_i) . We already know many important summaries of this data, because it is in standard coordinates. We have $\text{mean}(\{\hat{x}\}) = 0$; $\text{mean}(\{\hat{y}\}) = 0$; $\text{std}(\hat{x}) = 1$; and $\text{std}(\hat{y}) = 1$. Each of these summaries is itself the mean of some monomial. So $\text{std}(\hat{x})^2 = \text{mean}(\{\hat{x}^2\}) = 1$; $\text{std}(\hat{y})^2 =$

$\text{mean}(\{\hat{y}^2\})$ (the other two are easy). We can rewrite this information in terms of means of monomials, giving $\text{mean}(\{\hat{x}\}) = 0$; $\text{mean}(\{\hat{y}\}) = 0$; $\text{mean}(\{\hat{x}^2\}) = 1$; and $\text{mean}(\{\hat{y}^2\}) = 1$. There is one monomial missing here, which is $\hat{x}\hat{y}$.

The term $\text{mean}(\{\hat{x}\hat{y}\})$ captures correlation between x and y . The term is known as the **correlation coefficient** or **correlation**.

Definition: 4.1 *Correlation coefficient*

Assume we have N data items which are 2-vectors $(x_1, y_1), \dots, (x_N, y_N)$, where $N > 1$. These could be obtained, for example, by extracting components from larger vectors. We compute the correlation coefficient by first normalizing the x and y coordinates to obtain $\hat{x}_i = \frac{(x_i - \text{mean}(\{x\}))}{\text{std}(x)}$, $\hat{y}_i = \frac{(y_i - \text{mean}(\{y\}))}{\text{std}(y)}$. The correlation coefficient is the mean value of $\hat{x}\hat{y}$, and can be computed as:

$$\text{corr}(\{(x, y)\}) = \frac{\sum_i \hat{x}_i \hat{y}_i}{N}$$

Correlation is a measure of our ability to predict one value from another. The correlation coefficient takes values between -1 and 1 (we'll prove this below). If the correlation coefficient is close to 1 , then we are likely to predict very well. Small correlation coefficients (under about 0.5 , say, but this rather depends on what you are trying to achieve) tend not to be all that interesting, because (as we shall see) they result in rather poor predictions. Figure 4.16 gives a set of scatter plots of different real data sets with different correlation coefficients. These all come from data set of age-height-weight, which you can find at <http://www2.stetson.edu/~jrasp/data.htm> (look for bodyfat.xls). In each case, two outliers have been removed. Age and height are hardly correlated, as you can see from the figure. Younger people do tend to be slightly taller, and so the correlation coefficient is -0.25 . You should interpret this as a small correlation. However, the variable called “adiposity” (which isn't defined, but is presumably some measure of the amount of fatty tissue) is quite strongly correlated with weight, with a correlation coefficient is 0.86 . Average tissue density is quite strongly negatively correlated with adiposity, because muscle is much denser than fat, so these variables are negatively correlated — we expect high density to appear with low adiposity, and vice versa. The correlation coefficient is -0.86 . Finally, density is very strongly correlated with body weight. The correlation coefficient is -0.98 .

It's not always convenient or a good idea to produce scatter plots in standard coordinates (among other things, doing so hides the units of the data, which can be a nuisance). Fortunately, scaling or translating data does not change the value of the correlation coefficient (though it can change the sign if one scale is negative). This means that it's worth being able to spot correlation in a scatter plot that isn't in standard coordinates (even though correlation is always *defined* in standard coordinates). Figure 4.17 shows different correlated datasets plotted in their original

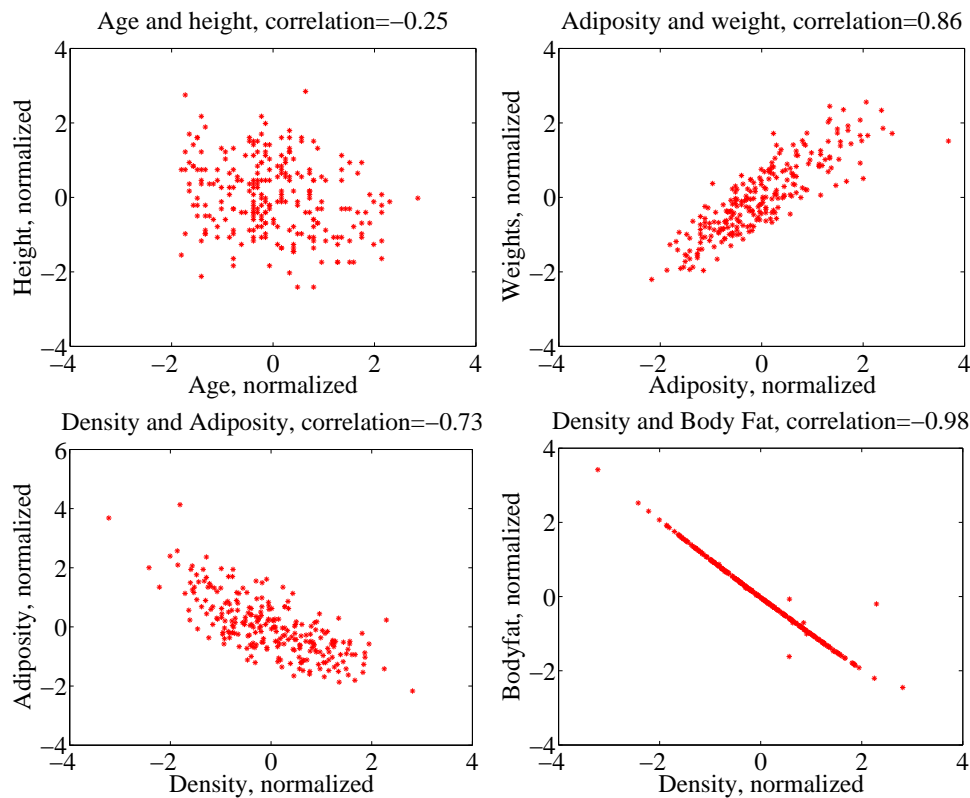


FIGURE 4.16: Scatter plots for various pairs of variables for the age-height-weight dataset from <http://www2.stetson.edu/~jrasp/data.htm>; *bodyfat.xls*. In each case, two outliers have been removed, and the plots are in standard coordinates (compare to figure 4.17, which shows these data sets plotted in their original units). The legend names the variables.

units. These data sets are the same as those used in figure 4.16

Properties of the Correlation Coefficient

You should memorize the following properties of the correlation coefficient:

- The correlation coefficient is symmetric (it doesn't depend on the order of its arguments), so

$$\text{corr}(\{(x, y)\}) = \text{corr}(\{(y, x)\})$$

- The value of the correlation coefficient is not changed by translating the data. Scaling the data can change the sign, but not the absolute value. For constants $a \neq 0$, b , $c \neq 0$, d we have

$$\text{corr}(\{(ax + b, cx + d)\}) = \text{sign}(ac)\text{corr}(\{(x, y)\})$$

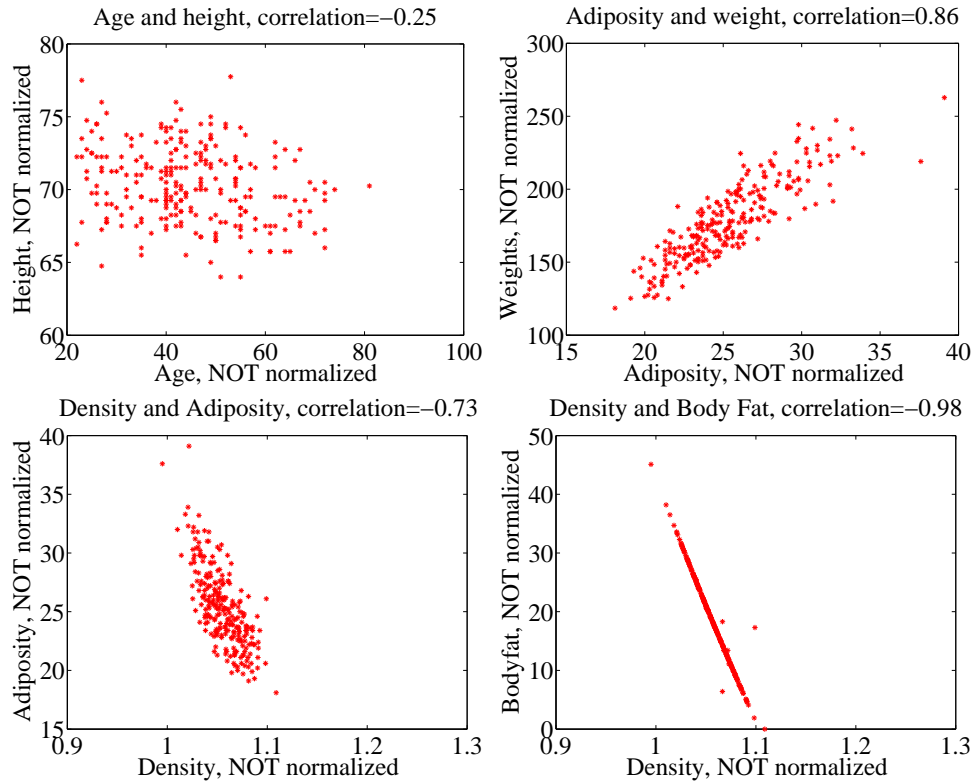


FIGURE 4.17: Scatter plots for various pairs of variables for the age-height-weight dataset from <http://www2.stetson.edu/~jrasp/data.htm>; *bodyfat.xls*. In each case, two outliers have been removed, and the plots are NOT in standard coordinates (compare to figure 4.16, which shows these data sets plotted in normalized coordinates). The legend names the variables.

- If \hat{y} tends to be large (resp. small) for large (resp. small) values of \hat{x} , then the correlation coefficient will be positive.
- If \hat{y} tends to be small (resp. large) for large (resp. small) values of \hat{x} , then the correlation coefficient will be negative.
- If \hat{y} doesn't depend on \hat{x} , then the correlation coefficient is zero (or close to zero).
- The largest possible value is 1, which happens when $\hat{x} = \hat{y}$.
- The smallest possible value is -1, which happens when $\hat{x} = -\hat{y}$.

The first property is easy, and we relegate that to the exercises. One way to see that the correlation coefficient isn't changed by translation or scale is to notice that it is defined in standard coordinates, and scaling or translating data doesn't

change those. Another way to see this is to scale and translate data, then write out the equations; notice that taking standard coordinates removes the effects of the scale and translation. In each case, notice that if the scale is negative, the sign of the correlation coefficient changes.

The property that, if \hat{y} tends to be large (resp. small) for large (resp. small) values of \hat{x} , then the correlation coefficient will be positive, doesn't really admit a formal statement. But it's relatively straightforward to see what's going on. Because $\text{mean}(\{\hat{x}\}) = 0$, small values of $\text{mean}(\{\hat{x}\})$ must be negative and large values must be positive. But $\text{corr}(\{(x, y)\}) = \frac{\sum_i \hat{x}_i \hat{y}_i}{N}$; and for this sum to be positive, it should contain mostly positive terms. It can contain few or no hugely positive (or hugely negative) terms, because $\text{std}(\hat{x}) = \text{std}(\hat{y}) = 1$ so there aren't many large (or small) numbers. For the sum to contain mostly positive terms, then the sign of \hat{x}_i should be the same as the sign \hat{y}_i for most data items. Small changes to this argument work to show that if \hat{y} tends to be small (resp. large) for large (resp. small) values of \hat{x} , then the correlation coefficient will be negative.

Showing that no relationship means zero correlation requires slightly more work. Divide the scatter plot of the dataset up into thin vertical strips. There are S strips. Each strip is narrow, so the \hat{x} value does not change much for the data points in a particular strip. For the s 'th strip, write $N(s)$ for the number of data points in the strip, $\hat{x}(s)$ for the \hat{x} value at the center of the strip, and $\bar{\hat{y}}(s)$ for the mean of the \hat{y} values within that strip. Now the strips are narrow, so we can approximate all data points within a strip as having the same value of \hat{x} . This yields

$$\text{mean}(\{\hat{x}\hat{y}\}) \approx \frac{1}{S} \sum_{s \in \text{strips}} \hat{x}(s) [N(s)\bar{\hat{y}}(s)]$$

(where you could replace \approx with $=$ if the strips were narrow enough). Now assume that $\bar{\hat{y}}(s)$ does not change from strip to strip, meaning that there is no relationship between \hat{x} and \hat{y} in this dataset (so the picture is like the left hand side in figure 4.15). Then each value of $\bar{\hat{y}}(s)$ is the same — we write $\bar{\hat{y}}$ — and we can rearrange to get

$$\text{mean}(\{\hat{x}\hat{y}\}) \approx \bar{\hat{y}} \frac{1}{S} \sum_{s \in \text{strips}} \hat{x}(s).$$

Now notice that

$$0 = \text{mean}(\{\hat{y}\}) \approx \frac{1}{S} \sum_{s \in \text{strips}} N(s)\bar{\hat{y}}(s)$$

(where again you could replace \approx with $=$ if the strips were narrow enough). This means that if every strip has the same value of $\bar{\hat{y}}(s)$, then that value must be zero. In turn, if there is no relationship between \hat{x} and \hat{y} , we must have $\text{mean}(\{\hat{x}\hat{y}\}) = 0$.

Proposition:

$$-1 \leq \text{corr}(\{(x, y)\}) \leq 1$$

Proof: Writing \hat{x} , \hat{y} for the normalized coefficients, we have

$$\text{corr}(\{(x, y)\}) = \frac{\sum_i \hat{x}_i \hat{y}_i}{N}$$

and you can think of the value as the inner product of two vectors. We write

$$\mathbf{x} = \frac{1}{\sqrt{N}} [\hat{x}_1, \hat{x}_2, \dots, \hat{x}_N] \text{ and } \mathbf{y} = \frac{1}{\sqrt{N}} [\hat{y}_1, \hat{y}_2, \dots, \hat{y}_N]$$

and we have $\text{corr}(\{(x, y)\}) = \mathbf{x}^T \mathbf{y}$. Notice $\mathbf{x}^T \mathbf{x} = \text{std}(x)^2 = 1$, and similarly for \mathbf{y} . But the inner product of two vectors is at its maximum when the two vectors are the same, and this maximum is 1. This argument is also sufficient to show that smallest possible value of the correlation is -1 , and this occurs when $\hat{x}_i = -\hat{y}_i$ for all i .

Property 4.1: The largest possible value of the correlation is 1, and this occurs when $\hat{x}_i = \hat{y}_i$ for all i . The smallest possible value of the correlation is -1 , and this occurs when $\hat{x}_i = -\hat{y}_i$ for all i .

4.2.2 Using Correlation to Predict

Assume we have N data items which are 2-vectors $(x_1, y_1), \dots, (x_N, y_N)$, where $N > 1$. These could be obtained, for example, by extracting components from larger vectors. As usual, we will write \hat{x}_i for x_i in normalized coordinates, and so on. Now assume that we know the correlation coefficient is r (this is an important, traditional notation). What does this mean?

One (very useful) interpretation is in terms of prediction. Assume we have a data point $(x_0, ?)$ where we know the x -coordinate, but not the y -coordinate. We can use the correlation coefficient to predict the y -coordinate. First, we transform to standard coordinates. Now we must obtain the best \hat{y}_0 value to predict, using the \hat{x}_0 value we have.

We want to construct a prediction function which gives a prediction for any value of \hat{x} . This predictor should behave as well as possible on our existing data. For each of the (\hat{x}_i, \hat{y}_i) pairs in our data set, the predictor should take \hat{x}_i and produce a result as close to \hat{y}_i as possible. We can choose the predictor by looking at the errors it makes at each data point.

We write \hat{y}_i^p for the value of \hat{y}_i predicted at \hat{x}_i . The simplest form of predictor is linear. If we predict using a linear function, then we have, for some unknown a, b , that $\hat{y}_i^p = a\hat{x}_i + b$. Now think about $u_i = \hat{y}_i - \hat{y}_i^p$, which is the error in our prediction. We would like to have $\text{mean}(\{u\}) = 0$ (otherwise, we could reduce the

error of the prediction just by subtracting a constant).

$$\begin{aligned}
 \text{mean}(\{u\}) &= \text{mean}(\{\hat{y} - \hat{y}^p\}) \\
 &= \text{mean}(\{\hat{y}\}) - \text{mean}(\{a\hat{x}_i + b\}) \\
 &= \text{mean}(\{\hat{y}\}) - a\text{mean}(\{\hat{x}\}) + b \\
 &= 0 - a0 + b \\
 &= 0.
 \end{aligned}$$

This means that we must have $b = 0$.

To estimate a , we need to think about $\text{var}(\{u\})$. We should like $\text{var}(\{u\})$ to be as small as possible, so that the errors are as close to zero as possible (remember, small variance means small standard deviation which means the data is close to the mean). We have

$$\begin{aligned}
 \text{var}(\{u\}) &= \text{var}(\{\hat{y} - \hat{y}^p\}) \\
 &= \text{mean}(\{(\hat{y} - a\hat{x})^2\}) \quad \text{because } \text{mean}(\{u\}) = 0 \\
 &= \text{mean}(\{(\hat{y})^2 - 2a\hat{x}\hat{y} + a^2(\hat{x})^2\}) \\
 &= \text{mean}(\{(\hat{y})^2\}) - 2a\text{mean}(\{\hat{x}\hat{y}\}) + a^2\text{mean}(\{(\hat{x})^2\}) \\
 &= 1 - 2ar + a^2,
 \end{aligned}$$

which we want to minimize by choice of a . At the minimum, we must have

$$\frac{d\text{var}(\{u_i\})}{da} = 0 = -2r + 2a$$

so that $a = r$ and the correct prediction is

$$\hat{y}_0^p = r\hat{x}_0$$

You can use a version of this argument to establish that if we have (\hat{x}_0, \hat{y}_0) , then the best prediction for \hat{x}_0 (*which is in standard coordinates*) is $r\hat{y}_0$. It is important to notice that the coefficient of \hat{y}_i is NOT $1/r$; you should work this example, which appears in the exercises. We now have a prediction procedure, outlined below.

Procedure: 4.1 *Predicting a value using correlation*

Assume we have N data items which are 2-vectors $(x_1, y_1), \dots, (x_N, y_N)$, where $N > 1$. These could be obtained, for example, by extracting components from larger vectors. Assume we have an x value x_0 for which we want to give the best prediction of a y value, based on this data. The following procedure will produce a prediction:

- Transform the data set into standard coordinates, to get

$$\begin{aligned}\hat{x}_i &= \frac{1}{\text{std}(x)}(x_i - \text{mean}(\{x\})) \\ \hat{y}_i &= \frac{1}{\text{std}(y)}(y_i - \text{mean}(\{y\})) \\ \hat{x}_0 &= \frac{1}{\text{std}(x)}(x_0 - \text{mean}(\{x\})).\end{aligned}$$

- Compute the correlation

$$r = \text{corr}(\{(x, y)\}) = \text{mean}(\{\hat{x}\hat{y}\}).$$

- Predict $\hat{y}_0 = r\hat{x}_0$.
- Transform this prediction into the original coordinate system, to get

$$y_0 = \text{std}(y)r\hat{x}_0 + \text{mean}(\{y\})$$

Now assume we have a y value y_0 , for which we want to give the best prediction of an x value, based on this data. The following procedure will produce a prediction:

- Transform the data set into standard coordinates.
- Compute the correlation.
- Predict $\hat{x}_0 = r\hat{y}_0$.
- Transform this prediction into the original coordinate system, to get

$$x_0 = \text{std}(x)r\hat{y}_0 + \text{mean}(\{x\})$$

There is another way of thinking about this prediction procedure, which is often helpful. Assume we need to predict a value for x_0 . In normalized coordinates, our prediction is $\hat{y}^P = r\hat{x}_0$; if we revert back to the original coordinate system, the

prediction becomes

$$\frac{(y^p - \text{mean}(\{y\}))}{\text{std}(y)} = r \left(\frac{(x_0 - \text{mean}(\{x\}))}{\text{std}(x)} \right).$$

This gives a really useful rule of thumb, which I have broken out in the box below.

Procedure: 4.2 *Predicting a value using correlation: Rule of thumb - 1*

If x_0 is k standard deviations from the mean of x , then the predicted value of y will be rk standard deviations away from the mean of y , and the sign of r tells whether y increases or decreases.

An even more compact version of the rule of thumb is in the following box.

Procedure: 4.3 *Predicting a value using correlation: Rule of thumb - 2*

The predicted value of y goes up by r standard deviations when the value of x goes up by one standard deviation.

We can compute the average root mean square error that this prediction procedure will make. The square of this error must be

$$\begin{aligned} \text{mean}(\{u^2\}) &= \text{mean}(\{y^2\}) - 2r\text{mean}(\{xy\}) + r^2\text{mean}(\{x^2\}) \\ &= 1 - 2r^2 + r^2 \\ &= 1 - r^2 \end{aligned}$$

so the root mean square error will be $\sqrt{1 - r^2}$. This is yet another interpretation of correlation; if x and y have correlation close to one, then predictions could have very small root mean square error, and so might be very accurate. In this case, knowing one variable is about as good as knowing the other. If they have correlation close to zero, then the root mean square error in a prediction might be as large as the root mean square error in \hat{y} — which means the prediction is nearly a pure guess.

The prediction argument means that we can spot correlations for data in other kinds of plots — one doesn't have to make a scatter plot. For example, if we were to observe a child's height from birth to their 10'th year (you can often find these observations in ballpen strokes, on kitchen walls), we could plot height as a function of year. If we also had their weight (less easily found), we could plot weight as a function of year, too. The prediction argument above says that, if you can predict the weight from the height (or vice versa) then they're correlated. One way to spot this is to look and see if one curve goes up when the other does (or goes down when the other goes up). You can see this effect in figure 4.4, where (before 19h00), prices go down when the number of pelts goes up, and vice versa. These two variables are negatively correlated.

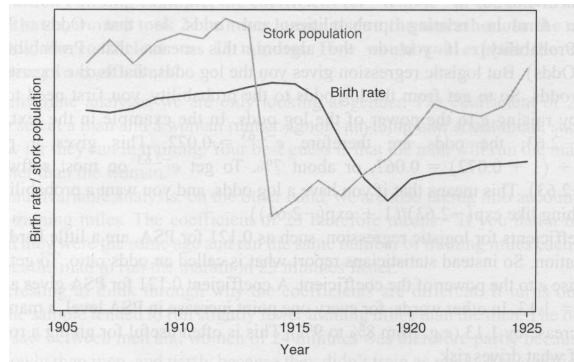


FIGURE 4.18: This figure, from Vickers (*ibid*, p184) shows a plot of the stork population as a function of time, and the human birth rate as a function of time, for some years in Germany. The correlation is fairly clear; but this does not mean that reducing the number of storks means there are fewer able to bring babies. Instead, this is the impact of the first world war — a hidden or latent variable.

4.2.3 Confusion caused by correlation

There is one very rich source of potential (often hilarious) mistakes in correlation. When two variables are correlated, they change together. If the correlation is positive, that means that, in typical data, if one is large then the other is large, and if one is small the other is small. In turn, this means that one can make a reasonable prediction of one from the other. However, correlation DOES NOT mean that changing one variable causes the other to change (sometimes known as causation).

Two variables in a dataset could be correlated for a variety of reasons. One important reason is pure accident. If you look at enough pairs of variables, you may well find a pair that appears to be correlated just because you have a small set of observations. Imagine, for example, you have a dataset consisting of only two vectors — there is a pretty good chance that there is some correlation between the coefficients. Such accidents can occur in large datasets, particularly if the dimensions are high.

Another reason variables could be correlated is that there is some causal relationship — for example, pressing the accelerator tends to make the car go faster, and so there will be some correlation between accelerator position and car acceleration. As another example, adding fertilizer does tend to make a plant grow bigger. Imagine you record the amount of fertilizer you add to each pot, and the size of the resulting potplant. There should be some correlation.

Yet another reason variables could be correlated is that there is some other background variable — often called a **latent variable** — linked causally to each of the observed variables. For example, in children (as Freedman, Pisani and Purves note in their excellent *Statistics*), shoe size is correlated with reading skills. This DOES NOT mean that making your feet grow will make you read faster, or that you can make your feet shrink by forgetting how to read. The real issue here is

the age of the child. Young children tend to have small feet, and tend to have weaker reading skills (because they've had less practice). Older children tend to have larger feet, and tend to have stronger reading skills (because they've had more practice). You can make a reasonable prediction of reading skills from foot size, because they're correlated, even though there is no direct connection.

This kind of effect can mask correlations, too. Imagine you want to study the effect of fertilizer on potplants. You collect a set of pots, put one plant in each, and add different amounts of fertilizer. After some time, you record the size of each plant. You expect to see correlation between fertilizer amount and plant size. But you might not if you had used a different species of plant in each pot. Different species of plant can react quite differently to the same fertilizer (some plants just die if over-fertilized), so the species could act as a latent variable. With an unlucky choice of the different species, you might even conclude that there was a negative correlation between fertilizer and plant size. This example illustrates why you need to take great care in setting up experiments and interpreting their results.

This sort of thing happens often, and it's an effect you should look for. Another nice example comes from Vickers (*ibid*). The graph, shown in Figure 4.18, shows a plot of (a) a dataset of the stork population in Europe over a period of years and (b) a dataset of the birth rate over those years. This isn't a scatter plot; instead, the data has been plotted on a graph. You can see by eye that these two datasets are quite strongly correlated. Even more disturbing, the stork population dropped somewhat before the birth rate dropped. Is this evidence that storks brought babies in Europe during those years? No (the usual arrangement seems to have applied). For a more sensible explanation, look at the dates. The war disturbed both stork and human breeding arrangements. Storks were disturbed immediately by bombs, etc., and the human birth rate dropped because men died at the front.

4.3 STERILE MALES IN WILD HORSE HERDS

Large herds of wild horses are (apparently) a nuisance, but keeping down numbers by simply shooting surplus animals would provoke outrage. One strategy that has been adopted is to sterilize males in the herd; if a herd contains sufficient sterile males, fewer foals should result. But catching stallions, sterilizing them, and reinserting them into a herd is a performance — does this strategy work?

We can get some insight by plotting data. At <http://lib.stat.cmu.edu/DASL/Datafiles/WildHorses.html>, you can find a dataset covering herd management in wild horses. I have plotted part of this dataset in figure 4.19. In this dataset, there are counts of all horses, sterile males, and foals made on each of a small number of days in 1986, 1987, and 1988 for each of two herds. I extracted data for one herd. I have plotted this data as a function of the count of days since the first data point, because this makes it clear that some measurements were taken at about the same time, but there are big gaps in the measurements. In this plot, the data points are shown with a marker. Joining them leads to a confusing plot because the data points vary quite strongly. However, notice that the size of the herd drifts down slowly (you could hold a ruler against the plot to see the trend), as does the number of foals, when there is a (roughly) constant number of sterile

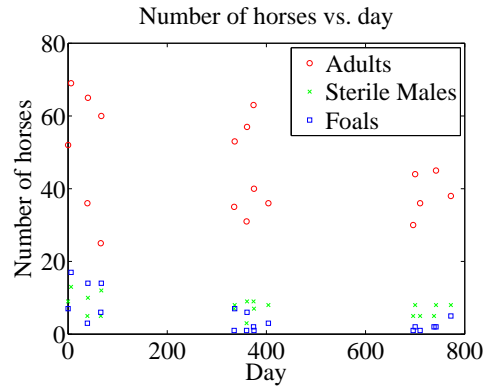


FIGURE 4.19: A plot of the number of adult horses, sterile males, and foals in horse herds over a period of three years. The plot suggests that introducing sterile males might cause the number of foals to go down. Data from <http://lib.stat.cmu.edu/DASL/Datafiles/WildHorses.html>.

males.

Does sterilizing males result in fewer foals? This is likely hard to answer for this dataset, but we could ask whether herds with more sterile males have fewer foals. A scatter plot is a natural tool to attack this question. However, the scatter plots of figure 4.20 suggest, rather surprisingly, that when there are more sterile males there are more adults (and vice versa), and when there are more sterile males there are more foals (and vice versa). This is borne out by a correlation analysis. The correlation coefficient between foals and sterile males is 0.74, and the correlation coefficient between adults and sterile males is 0.68. You should find this very surprising — how do the horses know how many sterile males there are in the herd? You might think that this is an effect of scaling the plot, but there is a scatter plot in normalized coordinates in figure 4.20 that is entirely consistent with the conclusions suggested by the unnormalized plot. What is going on here?

The answer is revealed by the scatter plots of figure 4.21. Here, rather than plotting a '*' at each data point, I have plotted the day number of the observation. This is in days from the first observation. You can see that the whole herd is shrinking — observations where there are many adults (resp. sterile adults, foals) occur with small day numbers, and observations where there are few have large day numbers. Because the whole herd is shrinking, it is true that when there are more adults and more sterile males, there are also more foals. Alternatively, you can see the plots of figure 4.19 as a scatter plot of herd size (resp. number of foals, number of sterile males) against day number. Then it becomes clear that the whole herd is shrinking, as is the size of each group. To drive this point home, we can look at the correlation coefficient between adults and days (-0.24), between sterile adults and days (-0.37), and between foals and days (-0.61). We can use the rule of thumb in box 3 to interpret this. This means that every 282 days, the herd loses about three adults; about one sterile adult; and about three foals. For the herd to have a stable size, it needs to gain by birth as many foals as it loses both to growing up

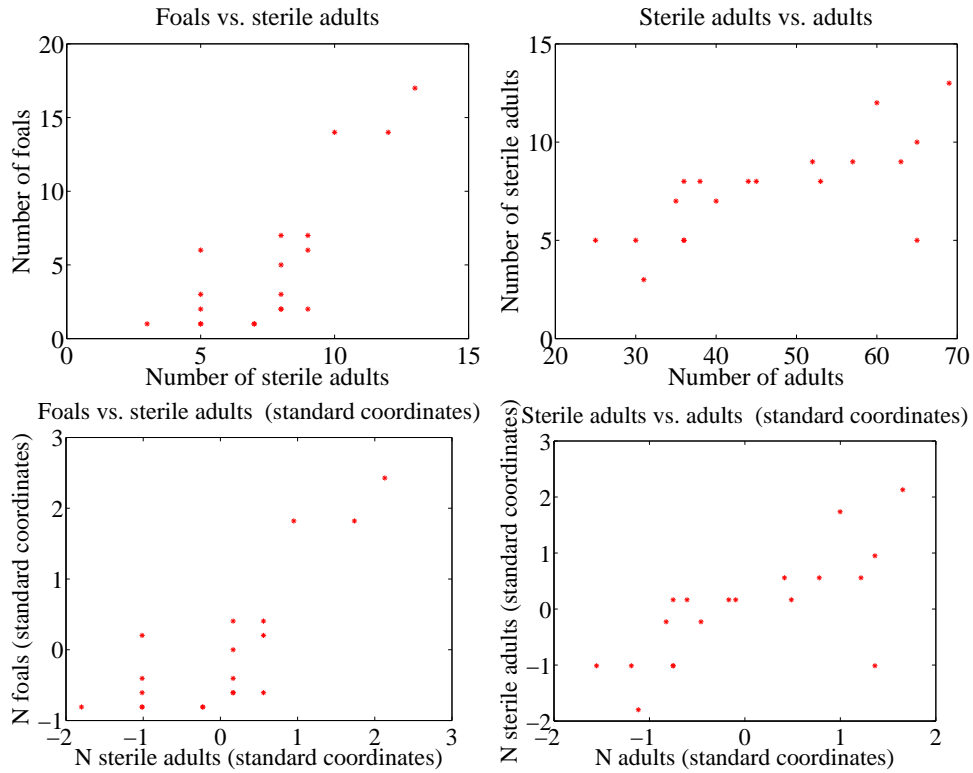


FIGURE 4.20: Scatter plots of the number of sterile males in a horse herd against the number of adults, and the number of foals against the number of sterile males, from data of <http://lib.stat.cmu.edu/DASL/Datafiles/WildHorses.html>. **Top:** unnormalized; **bottom:** standard coordinates.

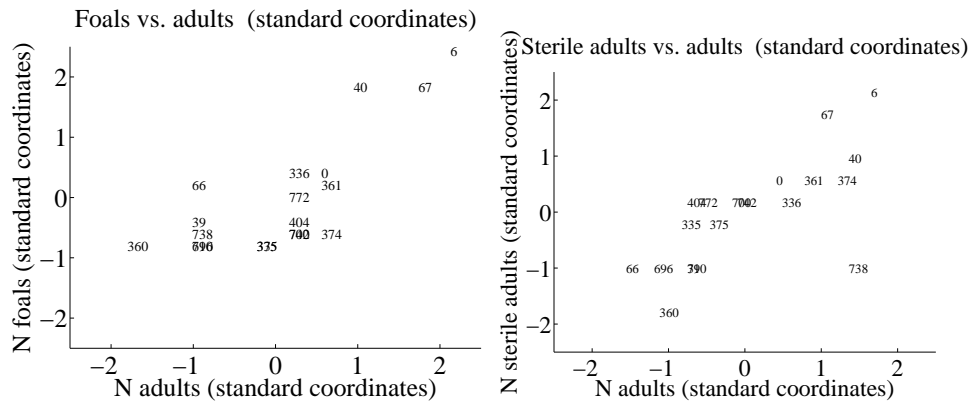


FIGURE 4.21:

and to death. If the herd is losing three foals every 282 days, then if they all grow up to replace the missing adults, the herd will be shrinking slightly (because it is losing four adults in this time); but if it loses foals to natural accidents, etc., then it is shrinking rather fast.

The message of this example is important. To understand a simple dataset, you might need to plot it several ways. You should make a plot, look at it and ask what it says, and then try to use another type of plot to confirm or refute what you think might be going on.

4.4 WHAT YOU MUST REMEMBER

You should be able to:

- Plot a bar chart, a heat map, and a pie chart for a categorical dataset.
- Plot a dataset as a graph, making sensible choices about markers, lines and the like.
- Plot a scatter plot for a dataset.
- Plot a normalized scatter plot for a dataset.
- Interpret the scatter plot to tell the sign of the correlation between two variables, and estimate the size of the correlation coefficient.
- Compute a correlation coefficient.
- Interpret a correlation coefficient.
- Use correlation to make predictions.

You should remember:

- The definition and properties of the correlation coefficient.

PROBLEMS

- 4.1. Show that $\text{corr}(\{(x, y)\}) = \text{corr}(\{(y, x)\})$ by substituting into the definition.
- 4.2. Show that if \hat{y} tends to be small (resp. large) for large (resp. small) values of \hat{x} , then the correlation coefficient will be negative.
- 4.3. We have a 2D dataset consisting of N pairs (\hat{x}_i, \hat{y}_i) in normalized coordinates. This data has correlation coefficient r . We observe a new \hat{y} value \hat{y}_0 , and wish to predict the (unknown) x value. We will do so with a linear prediction, choosing a, b , to predict an \hat{x} for any \hat{y} using the rule $\hat{x}^P = a\hat{y}^P + b$. Write $u_i = \hat{x}_i - \hat{x}_i^P$ for the error that this rule makes on each data item.
 - (a) We require $\text{mean}(\{u\}) = 0$. Show that this means that $b = 0$.
 - (b) We require that $\text{var}(\{u\})$ is minimized. Show that this means that $a = r$.
 - (c) We now have a result that seems paradoxical — if I have $(\hat{x}_0, ?)$ I predict $(\hat{x}_0, r\hat{x}_0)$ and if I have $(?, y_0)$, I predict $(r\hat{y}_0, \hat{y}_0)$. Use figure 4.22 to explain why this is right. The important difference between the two lines is that lies (approximately) in the middle of each vertical span of data, and the other lies (approximately) in the middle of each horizontal span of data.

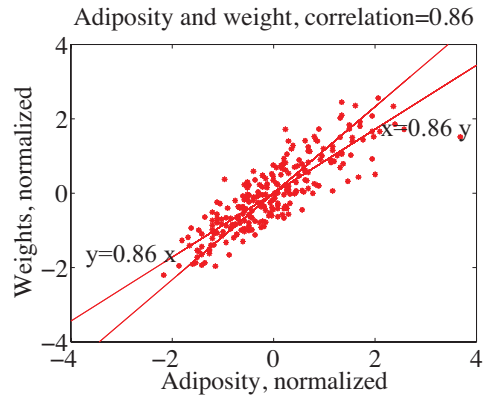


FIGURE 4.22: This figure shows two lines, $y = 0.86x$ and $x = 0.86y$, superimposed on the normalized adiposity-weight scatter plot.

- 4.4. I did the programming exercise about the earth temperature below. I looked at the years 1965-2012. Write $\{(y, T)\}$ for the dataset giving the temperature (T) of the earth in year y . I computed: $\text{mean}(\{y\}) = 1988.5$, $\text{std}(y) = 14$, $\text{mean}(\{T\}) = 0.175$, $\text{std}(T) = 0.231$ and $\text{corr}(\{y\})T = 0.892$. What is the best prediction using this information for the temperature in mid 2014? in mid 2028? in mid 2042?
- 4.5. I did the programming exercise about the earth temperature below. It is straightforward to build a dataset $\{(T, n_t)\}$ where each entry contains the temperature of the earth (T) and the number of counties where FEMA declared tornadoes n_t (for each year, you look up T and n_t , and make a data item). I computed: $\text{mean}(\{T\}) = 0.175$, $\text{std}(T) = 0.231$, $\text{mean}(\{n_t\}) = 31.6$, $\text{std}(n_t) = 30.8$, and $\text{corr}(\{T\})n_t = 0.471$. What is the best prediction using this information for the number of tornadoes if the global earth temperature is 0.5? 0.6? 0.7?

PROGRAMMING EXERCISES

- 4.6. At <http://lib.stat.cmu.edu/DASL/Datafiles/cigcancerdat.html>, you will find a dataset recording per capita cigarette sales and cancer deaths per 100K population for a variety of cancers, recorded for 43 states and the District of Columbia in 1960.
- Plot a scatter plot of lung cancer deaths against cigarette sales, using the two letter abbreviation for each state as a marker. You should see two fairly obvious outliers. The backstory at <http://lib.stat.cmu.edu/DASL/Stories/cigcancer.html> suggests that the unusual sales in Nevada are generated by tourism (tourists go home, and die there) and the unusual sales in DC are generated by commuting workers (who also die at home).
 - What is the correlation coefficient between per capita cigarette sales and lung cancer deaths per 100K population? Compute this with, and without the outliers. What effect did the outliers have? Why?
 - What is the correlation coefficient between per capita cigarette sales and bladder cancer deaths per 100K population? Compute this with, and

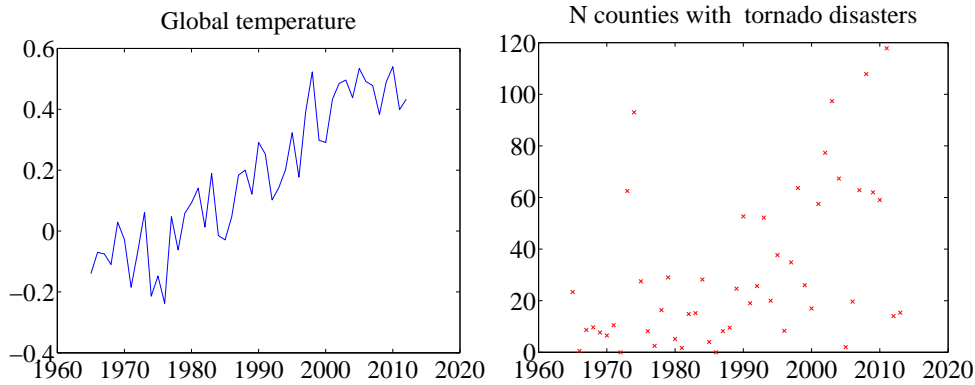


FIGURE 4.23: Plots I prepared from **left** *uea* data on temperature and **right** *FEMA* data on tornadoes by county. These should help you tell if you're on the right track.

- without the outliers. What effect did the outliers have? Why?
- (d) What is the correlation coefficient between per capita cigarette sales and kidney cancer deaths per 100K population? Compute this with, and without the outliers. What effect did the outliers have? Why?
- (e) What is the correlation coefficient between per capita cigarette sales and leukemia deaths per 100K population? Compute this with, and without the outliers. What effect did the outliers have? Why?
- (f) You should have computed a positive correlation between cigarette sales and lung cancer deaths. Does this mean that smoking causes lung cancer? Why?
- (g) You should have computed a negative correlation between cigarette sales and leukemia deaths. Does this mean that smoking cures leukemia? Why?
- 4.7. At <http://www.cru.uea.ac.uk/cru/info/warming/gtc.csv>, you can find a dataset of global temperature by year. When I accessed this, the years spanned 1880-2012. I don't know what units the temperatures are measured in. Keep in mind that measuring the temperature of the earth has non-trivial difficulties (you can't just insert an enormous thermometer!), and if you look at <http://www.cru.uea.ac.uk/cru> and <http://www.cru.uea.ac.uk/cru/data/temperature/> you can see some discussion of the choices made to get these measurements. There are two kinds of data in this dataset, smoothed and unsmoothed. I used the unsmoothed data, which should be fine for our purposes. The government publishes a great deal of data at <http://data.gov>. From there, I found a dataset, published by the Federal Emergency Management Agency (FEMA), of all federally declared disasters (which I found at <http://www.fema.gov/media-library/assets/documents/28318?id=6292>). We would like to see whether weather related disasters are correlated to global temperature.
- (a) The first step is preprocessing the data. The FEMA data has all sorts of information. From 1965 on, a disaster was declared per county (so you get one line in the data set for each county), but before 1965, it seems to have been by state. We divide the disasters into four types: TORNADO, FLOOD, STORM, HURRICANE. (FEMA seems to have a much richer

type system). We want to know how many counties declare a disaster of each type, in each year. This is a really rough estimate of the number of people affected by the disaster. If a disaster has two types (in some rows, you will see “SEVERE STORMS, HEAVY RAINS & FLOODING” we will allocate the credit evenly between the types (i.e. for this case we would count 1/2 for STORM and 1/2 for FLOOD). You should write code that will (a) read the dataset and then (b) compute a table with the count of the number of counties where a disaster of each type has occurred for each year. This takes a bit of work. Notice you only need to deal with two columns of the data (the date it was declared, and the type). Notice also that FEMA changed the way it represented dates somewhere through the first column (they added times), which can cause problems. You can tell the type of the disaster by just using a string match routine with the four keywords. Figure 4.23 shows the plot of temperature and of number of counties where FEMA declared a tornado disaster for this data.

- (b) Plot a normalized scatter plot of the number of counties where FEMA declared the disaster against temperature, for each kind.
- (c) For each kind of disaster, compute the correlation coefficient between the number of counties where FEMA declared the disaster and the year. For each kind of disaster, use this correlation coefficient to predict the number of disasters of this kind for 2013. Compare this to the true number, and explain what you see.
- (d) For each kind of disaster, compute the correlation coefficient between the number of counties where FEMA declared the disaster and the global temperature. For each kind of disaster, use this correlation coefficient to predict the number of disasters of this kind when the earth reaches 0.6 temperature units and 0.7 temperature units (on the absolute temperature scale).
- (e) Does this data show that warming of the earth causes weather disasters? Why?
- (f) Does this data suggest that more people will be affected by disasters in the US in the future? Why?
- (g) Does this data suggest that the earth will be warmer in the future? Why?

CHAPTER 5

Basic ideas in probability

We will perform experiments — which could be pretty much anything, from flipping a coin, to eating too much saturated fat, to smoking, to crossing the road without looking — and reason about the outcomes (mostly bad, for the examples I gave). But these outcomes are uncertain, and we need to weigh those uncertainties against one another. If I flip a coin, I could get heads or tails, and there’s no reason to expect to see one more often than the other. If I eat too much saturated fat or smoke, I will very likely have problems, though I might not. If I cross the road without looking, I may be squashed by a truck or I may not. Our methods need also to account for information. If I look before I cross the road, I am much less likely to be squashed.

Probability is the machinery we use to predict what will happen in an experiment. Probability measures the tendency of events to occur frequently or seldom when we repeat experiments. Building this machinery involves a formal model of potential outcomes and sets of outcomes. Once we have this, a set of quite simple axioms allows us, at least in principle, to compute probabilities in most situations.

5.1 EXPERIMENTS, OUTCOMES, EVENTS, AND PROBABILITY

If we flip a fair coin many times, we expect it to come up heads about as often as it comes up tails. If we toss a fair die many times, we expect each number to come up about the same number of times. We are performing an experiment each time we flip the coin, and each time we toss the die. We can formalize this experiment by describing the set of **outcomes** that we expect from the experiment. Every run of the experiment produces one of the set of possible outcomes. There are never two or more outcomes, and there is never no outcome.

In the case of the coin, the set of outcomes is:

$$\{H, T\}.$$

In the case of the die, the set of outcomes is:

$$\{1, 2, 3, 4, 5, 6\}.$$

We are making a modelling choice by specifying the outcomes of the experiment. For example, we are assuming that the coin can only come up heads or tails (but doesn’t stand on its edge; or fall between the floorboards; or land behind the bookcase; or whatever). We write the set of all outcomes Ω ; this is usually known as the **sample space**.

Worked example 5.1 *Find the lady*

We have three playing cards. One is a queen; one is a king, and one is a knave. All are shown face down, and one is chosen at random and turned up. What is the set of outcomes?

Solution: Write Q for queen, K for king, N for knave; the outcomes are $\{Q, K, N\}$

Worked example 5.2 *Find the lady, twice*

We play Find the Lady twice, replacing the card we have chosen. What is the sample space?

Solution: We now have $\{QQ, QK, QN, KQ, KK, KN, NQ, NK, NN\}$

Worked example 5.3 *Children*

A couple decides to have children until either (a) they have both a boy and a girl or (b) they have three children. What is the set of outcomes?

Solution: Write B for boy, G for girl, and write them in birth order; we have $\{BG, GB, BBG, BBB, GGB, GGG\}$.

Worked example 5.4 *Monty Hall (sigh!)*

There are three boxes. There is a goat, a second goat, and a car. These are placed into the boxes at random. The goats are indistinguishable for our purposes; equivalently, we do not care about the difference between goats. What is the sample space?

Solution: Write G for goat, C for car. Then we have $\{CGG, GCG, GGC\}$.

Worked example 5.5 *Monty Hall, different goats (sigh!)*

There are three boxes. There is a goat, a second goat, and a car. These are placed into the boxes at random. One goat is male, the other female, and the distinction is important. What is the sample space?

Solution: Write M for male goat, F for female goat, C for car. Then we have $\{CFM, CMF, FCM, MCF, FMC, MFC\}$. Notice how the number of outcomes has increased, because we now care about the distinction between goats.

Worked example 5.6 *A poor choice of strategy for planning a family*

A couple decides to have children. As they know no mathematics, they decide to have children until a girl then a boy are born. What is the sample space? Does this strategy bound the number of children they could be planning to have?

Solution: Write B for boy, G for girl. The sample space looks like any string of B's and G's that (a) ends in GB and (b) does not contain any other GB. In regular expression notation, you can write such strings as B^*G^+B . There is a lower bound (two), but no upper bound. As a family planning strategy, this is unrealistic, but it serves to illustrate the point that sample spaces don't have to be finite to be tractable.

5.1.1 The Probability of an Outcome

We represent our model of how often a particular outcome will occur in a repeated experiment with a **probability**, a non-negative number. This number gives the relative frequency of the outcome of interest, when an experiment is repeated a very large number of times.

Assume an outcome A has probability P . Assume we repeat the experiment a very large number of times N , and assume that the coins, dice, whatever don't communicate with one another from experiment to experiment (or, equivalently, that experiments don't "know" about one another). Then, for about $N \times P$ of those experiments the outcome will occur. Furthermore, as N gets larger, the fraction where the outcome occurs will get closer to P . We write $\#(A)$ for the number of times outcome A occurs. We interpret P as

$$\lim_{N \rightarrow \infty} \frac{\#(A)}{N}.$$

We can draw two important conclusions immediately.

- For any outcome A , $0 \leq P(A) \leq 1$.
- $\sum_{A_i \in \Omega} P(A_i) = 1$.

Remember that every run of the experiment produces exactly one outcome. The probabilities add up to one because each experiment must have one of the outcomes in the sample space.

Worked example 5.7 *A biased coin*

Assume we have a coin where the probability of getting heads is $P(H) = \frac{1}{3}$, and so the probability of getting tails is $P(T) = \frac{2}{3}$. We flip this coin three million times. How many times do we see heads?

Solution: $P(H) = \frac{1}{3}$, so we expect this coin will come up heads in $\frac{1}{3}$ of experiments. This means that we will very likely see very close to a million heads.

Some problems can be handled by building a set of outcomes and reasoning about the probability of each outcome. This is particularly useful when the outcomes *must* have the same probability, which happens rather a lot.

Worked example 5.8 *Find the Lady*

Assume that the card that is chosen is chosen fairly — that is, each card is chosen with the same probability. What is the probability of turning up a Queen?

Solution: There are three outcomes, and each is chosen with the same probability, so the probability is $1/3$.

Worked example 5.9 *Monty Hall, indistinguishable goats, again*

Each outcome has the same probability. We choose to open the first box. With what probability will we find a goat (any goat)?

Solution: There are three outcomes, each has the same probability, and two give a goat, so $2/3$

Worked example 5.10 *Monty Hall, yet again*

Each outcome has the same probability. We choose to open the first box. With what probability will we find the car?

Solution: There are three places the car could be, each has the same probability, so $1/3$

Worked example 5.11 *Monty Hall, with distinct goats, again*

Each outcome has the same probability. We choose to open the first box. With what probability will we find a female goat?

Solution: Using the reasoning of the previous example, but substituting “female goat” for “car”, $1/3$. The point of this example is that the sample space matters. If you care about the gender of the goat, then it’s important to keep track of it; if you don’t, it’s a good idea to omit it from the sample space.

5.1.2 Events

Assume we run an experiment and get an outcome. We know what the outcome is (that’s the whole point of a sample space). This means we can tell whether the outcome we get belongs to some particular known *set* of outcomes. We just look in the set and see if our outcome is there. This means that we should be able to predict the probability of a *set* of outcomes from any reasonable model of an experiment. For example, we might roll a die and ask what the probability of getting an even number is. As another example, we might flip a coin ten times, and ask what the probability of getting three heads is.

An **event** is a set of outcomes. The set of all outcomes, which we wrote Ω , must be an event. It is not a particularly interesting event, because we must have $P(\Omega) = 1$ (because we said that every run of an experiment produces one outcome, and that outcome must be in Ω). In principle, there could be no outcome, although this never happens. This means that the empty set, which we write \emptyset , is an event, and we have $P(\emptyset) = 0$. The space of events has a rich structure, which makes it possible to compute probabilities for other, more interesting, events.

Notation: Generally, we write sets like \mathcal{A} ; in principle, you could confuse this notation with the matrix notation, but it’s clear from context which is meant. We write $\mathcal{A} \cup \mathcal{B}$ for the union of two sets, $\mathcal{A} \cap \mathcal{B}$ for the intersection of two sets, and $\mathcal{A} - \mathcal{B}$ for the set theoretic difference (i.e. $\mathcal{A} - \mathcal{B} = \{x \in \mathcal{A} | x \notin \mathcal{B}\}$). We will write $\Omega - \mathcal{U}$ as \mathcal{U}^c ; read “the complement of \mathcal{U} ”.

Events have three important properties that follow from their nature as sets of outcomes:

- If \mathcal{U} and \mathcal{V} are events — sets of outcomes — then so is $\mathcal{U} \cap \mathcal{V}$. You should interpret this as the event that we have an outcome that is in \mathcal{U} and also in \mathcal{V} .
- If \mathcal{U} and \mathcal{V} are events, then $\mathcal{U} \cup \mathcal{V}$ is also an event. You should interpret this as the event that we have an outcome that is either in \mathcal{U} or in \mathcal{V} (or in both).
- If \mathcal{U} is an event, then $\mathcal{U}^c = \Omega - \mathcal{U}$ is also an event. You should think of this as the event we get an outcome that is not in \mathcal{U} .

This means that the set of all possible events Σ has a very important structure.

- \emptyset is in Σ .
- Ω is in Σ .
- If $\mathcal{U} \in \Sigma$ and $\mathcal{V} \in \Sigma$ then $\mathcal{U} \cup \mathcal{V} \in \Sigma$.
- If $\mathcal{U} \in \Sigma$ and $\mathcal{V} \in \Sigma$ then $\mathcal{U} \cap \mathcal{V} \in \Sigma$.
- If $\mathcal{U} \in \Sigma$ then $\mathcal{U}^c \in \Sigma$.

This means that the space of events can be quite big. For a single flip of a coin, the only possible space of events looks like:

$$\{\emptyset, \{H\}, \{T\}, \{H, T\}\}$$

Many experiments admit more than one space of events. For example, if we flip two coins, one natural event space is

$$\left\{ \begin{array}{l} \emptyset, \Omega, \\ \{HH\}, \{HT\}, \{TH\}, \{TT\}, \\ \{HH, HT\}, \{HH, TH\}, \{HH, TT\}, \{HT, TT\}, \{HT, TH\}, \{TT, HH\}, \\ \{HT, TH, TT\}, \{HH, TH, TT\}, \{HH, HT, TT\}, \{HH, HT, TH\} \end{array} \right\}$$

which can represent any possible event that can be built out of two coin flips. But this is not the only event space possible with these outcomes.

Worked example 5.12 *The structure of event spaces*

I flip two coins. Is the following collection of sets an event space?

$$\Sigma = \{\emptyset, \Omega, \{HH\}, \{TT, HT, TH\}\}$$

Solution: Yes, because: $\emptyset \in \Sigma$; $\Omega \in \Sigma$; if $\mathcal{A} \in \Sigma$, $\mathcal{A}^c \in \Sigma$; if $\mathcal{A} \in \Sigma$ and $\mathcal{B} \in \Sigma$, $\mathcal{A} \cup \mathcal{B} \in \Sigma$; and if $\mathcal{A} \in \Sigma$ and $\mathcal{B} \in \Sigma$, $\mathcal{A} \cap \mathcal{B} \in \Sigma$.

So, from example 12, we can have different consistent collections of events built on top of the same set of outcomes. This makes sense, because it allows us to reason about different kinds of result obtained with the same experimental equipment. You can interpret the event space in example 12 as encoding the events “two heads” and “anything other than two heads”.

For a single throw of the die, the set of every possible event is

$$\left\{ \begin{array}{cccccc} \emptyset, & \{1, 2, 3, 4, 5, 6\}, & & & & \\ \{1\}, & \{2\}, & \{3\}, & \{4\}, & \{5\}, & \{6\}, \\ & \{1, 2\}, & \{1, 3\}, & \{1, 4\}, & \{1, 5\}, & \{1, 6\}, \\ & & \{2, 3\}, & \{2, 4\}, & \{2, 5\}, & \{2, 6\}, \\ & & & \{3, 4\}, & \{3, 5\}, & \{3, 6\}, \\ & & & & \{4, 5\}, & \{4, 6\}, \\ & & & & & \{5, 6\}, \\ \\ & \{1, 2, 3\}, & \{1, 2, 4\}, & \{1, 2, 5\}, & \{1, 2, 6\}, & \\ & & \{1, 3, 4\}, & \{1, 3, 5\}, & \{1, 3, 6\}, & \\ & & & \{1, 4, 5\}, & \{1, 4, 6\}, & \\ & & & & \{1, 5, 6\}, & \\ & & \{2, 3, 4\}, & \{2, 3, 5\}, & \{2, 3, 6\}, & \\ & & \{2, 4, 5\}, & \{2, 4, 6\}, & \{2, 5, 6\}, & \\ & & & \{3, 4, 5\}, & \{3, 4, 6\}, & \\ & & & & \{3, 5, 6\}, & \\ & & & & \{4, 5, 6\}, & \\ \\ & & \{1, 2, 3, 4\}, & \{1, 2, 3, 5\}, & \{1, 2, 3, 6\}, & \\ & & & \{1, 3, 4, 5\}, & \{1, 3, 4, 6\}, & \\ & & & \{2, 3, 4, 5\}, & \{2, 3, 4, 6\}, & \\ & & & & \{3, 4, 5, 6\}, & \\ \{2, 3, 4, 5, 6\}, & \{1, 3, 4, 5, 6\}, & \{1, 2, 4, 5, 6\}, & \{1, 2, 3, 5, 6\}, & \{1, 2, 3, 4, 6\}, & \{1, 2, 3, 4, 5\} \end{array} \right\}$$

(which gives some explanation as to why we don't usually write out the whole space of events). In fact, it is seldom necessary to explain which event space one is working with. We usually assume that the event space consists of all events that can be obtained with the outcomes (as in the event space shown for a die).

5.1.3 The Probability of Events

So far, we have described the probability of each outcome with a non-negative number. This number represents the relative frequency of the outcome. Because we can tell when an event has occurred, we can compute the relative frequency of events, too. Because it is a relative frequency, the probability of an event is a non-negative number, and is no greater than one. But the probability of events must be consistent with the probability of outcomes. This implies a set of quite straightforward properties:

- **The probability of every event is non-negative**, which we write $P(\mathcal{A}) \geq 0$ for all \mathcal{A} in the collection of events.
- **Every experiment has an outcome**, which we write $P(\Omega) = 1$.
- **The probability of disjoint events is additive**, which requires more notation. Assume that we have a collection of events \mathcal{A}_i , indexed by i . We require that these have the property $\mathcal{A}_i \cap \mathcal{A}_j = \emptyset$ when $i \neq j$. This means that there is no outcome that appears in more than one \mathcal{A}_i . In turn, if we interpret probability as relative frequency, we must have that $P(\cup_i \mathcal{A}_i) = \sum_i P(\mathcal{A}_i)$.

Any function P taking events to numbers that has these properties is a probability. These very simple properties imply a series of other very important properties.

Useful Facts: 5.1 *The probability of events*

- $P(\mathcal{A}^c) = 1 - P(\mathcal{A})$
- $P(\emptyset) = 0$
- $P(\mathcal{A} - \mathcal{B}) = P(\mathcal{A}) - P(\mathcal{A} \cap \mathcal{B})$
- $P(\mathcal{A} \cup \mathcal{B}) = P(\mathcal{A}) + P(\mathcal{B}) - P(\mathcal{A} \cap \mathcal{B})$
- $P(\cup_1^n \mathcal{A}_i) = \sum_i P(\mathcal{A}_i) - \sum_{i < j} P(\mathcal{A}_i \cap \mathcal{A}_j) + \sum_{i < j < k} P(\mathcal{A}_i \cap \mathcal{A}_j \cap \mathcal{A}_k) + \dots + (-1)^{(n+1)} P(\mathcal{A}_1 \cap \mathcal{A}_2 \cap \dots \cap \mathcal{A}_n)$

I prove each of these below. Looking at the useful facts should suggest a helpful analogy. Think about the probability of an event as the “size” of that event. This “size” is relative to Ω , which has “size” 1. I find this a good way to remember equations.

For example, $P(\mathcal{A}) + P(\mathcal{A}^c) = 1$ has to be true, by this analogy, because \mathcal{A} and \mathcal{A}^c don’t overlap, and together make up all of Ω . Similarly, $P(\mathcal{A} - \mathcal{B}) = P(\mathcal{A}) - P(\mathcal{A} \cap \mathcal{B})$ is easily captured — the “size” of the part of \mathcal{A} that isn’t \mathcal{B} is obtained by taking the “size” of \mathcal{A} and subtracting the “size” of the part that is also in \mathcal{B} . Similarly, $P(\mathcal{A} \cup \mathcal{B}) = P(\mathcal{A}) + P(\mathcal{B}) - P(\mathcal{A} \cap \mathcal{B})$ says — you can get the “size” of $\mathcal{A} \cup \mathcal{B}$ by adding the two “sizes”, then subtracting the “size” of the intersection because otherwise you would count these terms twice. Some people find Venn diagrams a useful way to keep track of this argument, and Figure 5.1 is for them.

Proposition: $P(\mathcal{A}^c) = 1 - P(\mathcal{A})$

Proof: \mathcal{A}^c and \mathcal{A} are disjoint, so that $P(\mathcal{A}^c \cup \mathcal{A}) = P(\mathcal{A}^c) + P(\mathcal{A}) = P(\Omega) = 1$.

Proposition: $P(\emptyset) = 0$

Proof: $P(\emptyset) = P(\Omega^c) = P(\Omega - \Omega) = 1 - P(\Omega) = 1 - 1 = 0$.

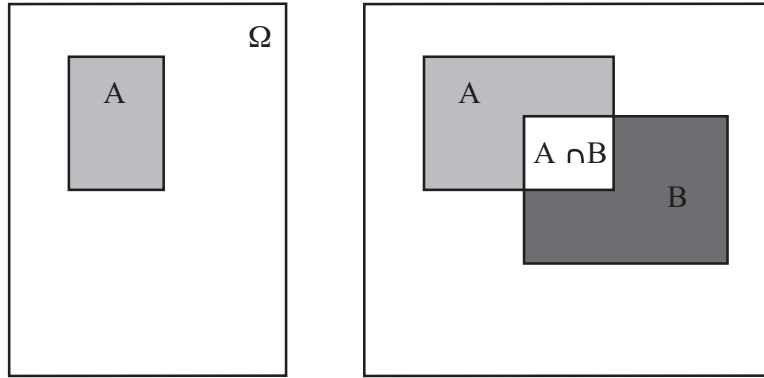


FIGURE 5.1: If you think of the probability of an event as measuring its “size”, many of the rules are quite straightforward to remember. Venn diagrams can sometimes help. On the **left**, a Venn diagram to help remember that $P(\mathcal{A}) + P(\mathcal{A}^c) = 1$. The “size” of Ω is 1, outcomes lie either in \mathcal{A} or \mathcal{A}^c , and the two don’t intersect. On the **right**, you can see that $P(\mathcal{A} - \mathcal{B}) = P(\mathcal{A}) - P(\mathcal{A} \cap \mathcal{B})$ by noticing that $P(\mathcal{A} - \mathcal{B})$ is the “size” of the part of \mathcal{A} that isn’t \mathcal{B} . This is obtained by taking the “size” of \mathcal{A} and subtracting the “size” of the part that is also in \mathcal{B} , i.e. the “size” of $\mathcal{A} \cap \mathcal{B}$. Similarly, you can see that $P(\mathcal{A} \cup \mathcal{B}) = P(\mathcal{A}) + P(\mathcal{B}) - P(\mathcal{A} \cap \mathcal{B})$ by noticing that you can get the “size” of $\mathcal{A} \cup \mathcal{B}$ by adding the “sizes” of \mathcal{A} and \mathcal{B} , then subtracting the “size” of the intersection to avoid double counting.

Proposition: $P(\mathcal{A} - \mathcal{B}) = P(\mathcal{A}) - P(\mathcal{A} \cap \mathcal{B})$

Proof: $\mathcal{A} - \mathcal{B}$ is disjoint from $\mathcal{A} \cap \mathcal{B}$, and $(\mathcal{A} - \mathcal{B}) \cup (\mathcal{A} \cap \mathcal{B}) = \mathcal{A}$. This means that $P(\mathcal{A} - \mathcal{B}) + P(\mathcal{A} \cap \mathcal{B}) = P(\mathcal{A})$.

Proposition: $P(\mathcal{A} \cup \mathcal{B}) = P(\mathcal{A}) + P(\mathcal{B}) - P(\mathcal{A} \cap \mathcal{B})$

Proof: $P(\mathcal{A} \cup \mathcal{B}) = P(\mathcal{A} \cup (\mathcal{B} \cap \mathcal{A}^c)) = P(\mathcal{A}) + P((\mathcal{B} \cap \mathcal{A}^c))$. Now $\mathcal{B} = (\mathcal{B} \cap \mathcal{A}) \cup (\mathcal{B} \cap \mathcal{A}^c)$. Furthermore, $(\mathcal{B} \cap \mathcal{A})$ is disjoint from $(\mathcal{B} \cap \mathcal{A}^c)$, so we have $P(\mathcal{B}) = P((\mathcal{B} \cap \mathcal{A})) + P((\mathcal{B} \cap \mathcal{A}^c))$. This means that $P(\mathcal{A} \cup \mathcal{B}) = P(\mathcal{A}) + P((\mathcal{B} \cap \mathcal{A}^c)) = P(\mathcal{A}) + P(\mathcal{B}) - P((\mathcal{B} \cap \mathcal{A}))$.

Proposition: $P(\cup_1^n \mathcal{A}_i) = \sum_i P(\mathcal{A}_i) - \sum_{i < j} P(\mathcal{A}_i \cap \mathcal{A}_j) + \sum_{i < j < k} P(\mathcal{A}_i \cap \mathcal{A}_j \cap \mathcal{A}_k) + \dots (-1)^{(n+1)} P(\mathcal{A}_1 \cap \mathcal{A}_2 \cap \dots \cap \mathcal{A}_n)$

Proof: This can be proven by repeated application of the previous result. As an example, we show how to work the case where there are three sets (you can get the rest by induction).

$$\begin{aligned}
 P(\mathcal{A}_1 \cup \mathcal{A}_2 \cup \mathcal{A}_3) &= P(\mathcal{A}_1 \cup (\mathcal{A}_2 \cup \mathcal{A}_3)) \\
 &= P(\mathcal{A}_1) + P(\mathcal{A}_2 \cup \mathcal{A}_3) - P(\mathcal{A}_1 \cap (\mathcal{A}_2 \cup \mathcal{A}_3)) \\
 &= P(\mathcal{A}_1) + (P(\mathcal{A}_2) + P(\mathcal{A}_3) - P(\mathcal{A}_2 \cap \mathcal{A}_3)) - \\
 &\quad P((\mathcal{A}_1 \cap \mathcal{A}_2) \cup (\mathcal{A}_1 \cap \mathcal{A}_3)) \\
 &= P(\mathcal{A}_1) + (P(\mathcal{A}_2) + P(\mathcal{A}_3) - P(\mathcal{A}_2 \cap \mathcal{A}_3)) - \\
 &\quad P(\mathcal{A}_1 \cap \mathcal{A}_2) - P(\mathcal{A}_1 \cap \mathcal{A}_3) \\
 &\quad - (-P((\mathcal{A}_1 \cap \mathcal{A}_2) \cap (\mathcal{A}_1 \cap \mathcal{A}_3))) \\
 &= P(\mathcal{A}_1) + P(\mathcal{A}_2) + P(\mathcal{A}_3) - \\
 &\quad P(\mathcal{A}_2 \cap \mathcal{A}_3) - P(\mathcal{A}_1 \cap \mathcal{A}_2) - P(\mathcal{A}_1 \cap \mathcal{A}_3) + \\
 &\quad P(\mathcal{A}_1 \cap \mathcal{A}_2 \cap \mathcal{A}_3)
 \end{aligned}$$

5.1.4 Computing Probabilities by Counting Outcomes

The rule $P(\mathcal{A} \cup \mathcal{B}) = P(\mathcal{A}) + P(\mathcal{B}) - P(\mathcal{A} \cap \mathcal{B})$ yields a very useful procedure for computing the probability of some events. Imagine \mathcal{A} and \mathcal{B} are disjoint (so $\mathcal{A} \cap \mathcal{B} = \emptyset$). Then $P(\mathcal{A} \cup \mathcal{B})$ is just $P(\mathcal{A}) + P(\mathcal{B})$. So imagine some event \mathcal{E} that consists of a set of outcomes. Each singleton set — containing just one outcome — is disjoint from each other one. I can make the set \mathcal{E} by: starting with an empty set; unioning this with one of the outcomes in \mathcal{E} ; then repeatedly unioning the result a new outcome until I get \mathcal{E} . I am always unioning disjoint sets, so the probability of this event is the sum of probabilities of the outcomes. So we have

$$P(\mathcal{E}) = \sum_{O \in \mathcal{E}} P(O)$$

where O ranges over the outcomes in \mathcal{E} .

This is particularly useful when you know each outcome in Ω has the same probability. In this case, you can show

$$P(\mathcal{E}) = \frac{\text{Number of outcomes in } \mathcal{E}}{\text{Total number of outcomes in } \Omega}$$

(exercises). Such problems become, basically, advanced counting exercises.

Worked example 5.13 *Odd numbers with fair dice*

We throw a fair (each number has the same probability) die twice, then add the two numbers. What is the probability of getting an odd number?

Solution: There are 36 outcomes. Each has the same probability ($1/36$). 18 of them give an odd number, and the other 18 give an even number, so the probability is $18/36 = 1/2$

Worked example 5.14 *Drawing a red ten*

I shuffle a standard pack of cards, and draw one card. What is the probability that it is a red ten?

Solution: There are 52 cards, and each is an outcome. Two of these outcomes are red tens; so $1/26$.

Worked example 5.15 *Numbers divisible by five with fair dice*

We throw a fair (each number has the same probability) die twice, then add the two numbers. What is the probability of getting a number divisible by five?

Solution: There are 36 outcomes. Each has the same probability ($1/36$). For this event, the spots must add to either 5 or to 10. There are 4 ways to get 5. There are 3 ways to get 10, so the probability is $7/36$.

Worked example 5.16 *Children - 1*

This example is a version of of example 1.12, p44, in Stirzaker, "Elementary Probability". A couple decides to have children. They decide simply to have three children. Assume that each gender is equally likely at each birth. Let B_i be the event that there are i boys, and C be the event there are more girls than boys. Compute $P(B_1)$ and $P(C)$.

Solution: There are eight outcomes. Each has the same probability. Three of them have a single boy, so $P(B_1) = 3/8$. $P(C) = P(C^c)$ (because C^c is the event there are more boys than than girls, AND the number of children is odd), so that $P(C) = 1/2$; you can also get this by counting outcomes.

Sometimes a bit of fiddling with the space of outcomes makes it easy to

compute what we want.

Worked example 5.17 *Children - 2*

This example is a version of of example 1.12, p44, in Stirzaker, “Elementary Probability”. A couple decides to have children. They decide to have children until the first girl is born, or until there are three, and then stop. Assume that each gender is equally likely at each birth. Let B_i be the event that there are i boys, and C be the event there are more girls than boys. Compute $P(B_1)$ and $P(C)$.

Solution: In this case, we could write the outcomes as $\{G, BG, BBG\}$, but if we think about them like this, we have no simple way to compute their probability. Instead, we could use the sample space from the previous answer, but assume that some of the later births are fictitious. This gives us natural collection of events for which it is easy to compute probabilities. Having one girl corresponds to the event $\{Gbb, Gbg, Ggb, Ggg\}$, where I have used lowercase letters to write the fictitious later births; the probability is $1/2$. Having a boy then a girl corresponds to the event $\{BGb, BGg\}$ (and so has probability $1/4$). Having two boys then a girl corresponds to the event $\{BBG\}$ (and so has probability $1/8$). Finally, having three boys corresponds to the event $\{BBB\}$ (and so has probability $1/8$). This means that $P(B_1) = 1/4$ and $P(C) = 1/2$.

Worked example 5.18 *Children - 2*

This example is a version of of example 1.12, p44, in Stirzaker, “Elementary Probability”. A couple decides to have children. They decide to have children until there is one of each gender, or until there are three, and then stop. Assume that each gender is equally likely at each birth. Let B_i be the event that there are i boys, and C be the event there are more girls than boys. Compute $P(B_1)$ and $P(C)$.

Solution: We could write the outcomes as $\{GB, BG, GGB, GGG, BBG, BBB\}$. Again, if we think about them like this, we have no simple way to compute their probability; so we use the sample space from the previous example with the device of the fictitious births again. The important events are $\{Gbb, Gbg\}$; $\{BGb, BGg\}$; $\{GGB\}$; $\{GGG\}$; $\{BBG\}$; and $\{BBB\}$. Like this, we get $P(B_1) = 5/8$ and $P(C) = 1/4$.

Worked example 5.19 *Birthdays in succession*

We stop three people at random, and ask the day of the week on which they are born. What is the probability that they are born on three days of the week in succession (for example, the first on Monday; the second on Tuesday; the third on Wednesday; or Saturday-Sunday-Monday; and so on).

Solution: We assume that births are equally common on each day of the week. The space of outcomes consists of triples of days, and each outcome has the same probability. The event is the set of triples of three days in succession (which has seven elements, one for each starting day). The space of outcomes has 7^3 elements in it, so the probability is

$$\frac{\text{Number of outcomes in the event}}{\text{Total number of outcomes}} = \frac{7}{7^3} = \frac{1}{49}.$$

Worked example 5.20 *Shared birthdays*

We stop two people at random. What is the probability that they were born on the same day of the week?

Solution: The day the first person was born doesn't matter; the probability the second person was born on that day is $1/7$. Or you could count outcomes explicitly to get

$$\frac{\text{Number of outcomes in the event}}{\text{Total number of outcomes}} = \frac{7}{7 \times 7} = \frac{1}{7}.$$

An important feature of this class of problem is that your intuition can be quite misleading. This is because, although each outcome can have very small probability, the number of outcomes in an event can be big.

5.1.5 Computing Probabilities by Reasoning about Sets

The rule $P(\mathcal{A}^c) = 1 - P(\mathcal{A})$ is occasionally useful for computing probabilities on its own; more commonly, you need other reasoning as well.

Worked example 5.21 *Shared birthdays*

What is the probability that, in a room of 30 people, there is a pair of people who have the same birthday?

Solution: We simplify, and assume that each year has 365 days, and that none of them are special (i.e. each day has the same probability of being chosen as a birthday). This model isn't perfect (there tend to be slightly more births roughly 9 months after: the start of spring; blackouts; major disasters; and so on) but it's workable. The easy way to attack this question is to notice that our probability, $P(\{\text{shared birthday}\})$, is

$$1 - P(\{\text{all birthdays different}\}).$$

This second probability is rather easy to compute. Each outcome in the sample space is a list of 30 days (one birthday per person). Each outcome has the same probability. So

$$P(\{\text{all birthdays different}\}) = \frac{\text{Number of outcomes in the event}}{\text{Total number of outcomes}}.$$

The total number of outcomes is easily seen to be 365^{30} , which is the total number of possible lists of 30 days. The number of outcomes in the event is the number of lists of 30 days, all different. To count these, we notice that there are 365 choices for the first day; 364 for the second; and so on. So we have

$$P(\{\text{shared birthday}\}) = 1 - \frac{365 \times 364 \times \dots \times 336}{365^{30}} = 1 - 0.2937 = 0.7063$$

which means there's really a pretty good chance that two people in a room of 30 share a birthday.

If we change the birthday example slightly, the problem changes drastically. If you stand up in a room of 30 people and bet that two people in the room have the same birthday, you have a probability of winning of about 0.71. If you bet that there is someone else in the room who has the same birthday that you do, your probability of winning is very different.

Worked example 5.22 *Shared birthdays*

You bet there is someone else in a room of 30 people who has the same birthday that you do. Assuming you know nothing about the other 29 people, what is the probability of winning?

Solution: The easy way to do this is

$$P(\{\text{winning}\}) = 1 - P(\{\text{losing}\}).$$

Now you will lose if everyone has a birthday different from you. You can think of the birthdays of the others in the room as a list of 29 days of the year. If your birthday is on the list, you win; if it's not, you lose. The number of losing lists is the number of lists of 29 days of the year such that your birthday is not in the list. This number is easy to get. We have 364 days of the year to choose from for each of 29 locations in the list. The total number of lists is the number of lists of 29 days of the year. Each list has the same probability. So

$$P(\{\text{losing}\}) = \frac{364^{29}}{365^{29}}$$

and

$$P(\{\text{winning}\}) \approx 0.0765.$$

There is a wide variety of problems like this; if you're so inclined, you can make a small but quite reliable profit off people's inability to estimate probabilities for this kind of problem correctly (examples 21 and 22 are reliably profitable; you could probably do quite well out of examples 19 and 20).

The rule $P(\mathcal{A} - \mathcal{B}) = P(\mathcal{A}) - P(\mathcal{A} \cap \mathcal{B})$ is also occasionally useful for computing probabilities on its own; more commonly, you need other reasoning as well.

Worked example 5.23 *Dice*

You flip two fair six-sided dice, and add the number of spots. What is the probability of getting a number divisible by 2, but not by 5?

Solution: There is an interesting way to work the problem. Write \mathcal{D}_n for the event the number is divisible by n . Now $P(\mathcal{D}_2) = 1/2$ (count the cases; or, more elegantly, notice that each die has the same number of odd and even faces, and work from there). Now $P(\mathcal{D}_2 - \mathcal{D}_5) = P(\mathcal{D}_2) - P(\mathcal{D}_2 \cap \mathcal{D}_5)$. But $\mathcal{D}_2 \cap \mathcal{D}_5$ contains only two outcomes (6, 4 and 4, 6), so $P(\mathcal{D}_2 - \mathcal{D}_5) = 18/36 - 2/36 = 4/9$

Sometimes it is easier to reason about unions than to count outcomes directly.

Worked example 5.24 *Two fair dice*

I roll two fair dice. What is the probability that the result is divisible by either 2 or 5, or both?

Solution: Write \mathcal{D}_n for the event the number is divisible by n . We want $P(\mathcal{D}_2 \cup \mathcal{D}_5) = P(\mathcal{D}_2) + P(\mathcal{D}_5) - P(\mathcal{D}_2 \cap \mathcal{D}_5)$. From example 23, we know $P(\mathcal{D}_2) = 1/2$ and $P(\mathcal{D}_2 \cap \mathcal{D}_5) = 2/36$. By counting outcomes, $P(\mathcal{D}_5) = 6/36$. So $P(\mathcal{D}_2 \cup \mathcal{D}_5) = (18 + 6 - 2)/36 = 22/36$.

5.1.6 Independence

Some experimental results do not affect others. For example, if I flip a coin twice, whether I get heads on the first flip has no effect on whether I get heads on the second flip. As another example, I flip a coin; the outcome does not affect whether I get hit on the head by a falling apple later in the day. We refer to events with this property as independent.

Useful Facts: 5.2 *Independent events*

Two events \mathcal{A} and \mathcal{B} are **independent** if and only if

$$P(\mathcal{A} \cap \mathcal{B}) = P(\mathcal{A})P(\mathcal{B})$$

The “size” analogy helps motivate this expression. We think of $P(\mathcal{A})$ as the “size” of \mathcal{A} relative to Ω , and so on. Now $P(\mathcal{A} \cap \mathcal{B})$ measures the “size” of $\mathcal{A} \cap \mathcal{B}$ — that is, the part of \mathcal{A} that lies inside \mathcal{B} . But if \mathcal{A} and \mathcal{B} are independent, then the size of $\mathcal{A} \cap \mathcal{B}$ relative to \mathcal{B} should be the same as the size of \mathcal{A} relative to Ω (Figure 5.2). Otherwise, \mathcal{B} affects \mathcal{A} , because \mathcal{A} is more (or less) likely when \mathcal{B} has occurred.

So for \mathcal{A} and \mathcal{B} to be independent, we must have

$$\text{“Size” of } \mathcal{A} \cap \mathcal{B} = \frac{\text{“Size” of piece of } \mathcal{A} \text{ in } \mathcal{B}}{\text{“Size” of } \mathcal{B}},$$

or, equivalently,

$$P(\mathcal{A}) = \frac{P(\mathcal{A} \cap \mathcal{B})}{P(\mathcal{B})}$$

which yields our expression. Independence is important, because it is straightforward to compute probabilities for sequences of independent outcomes.

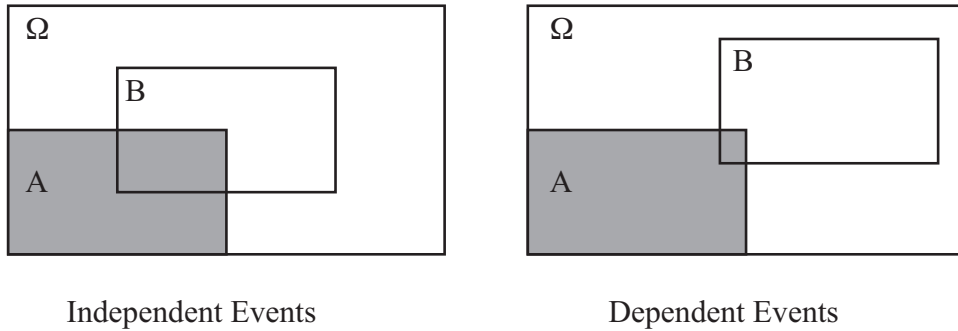


FIGURE 5.2: On the **left**, \mathcal{A} and \mathcal{B} are independent. \mathcal{A} spans $1/4$ of Ω , and $\mathcal{A} \cap \mathcal{B}$ spans $1/4$ of \mathcal{B} . This means that \mathcal{B} can't affect \mathcal{A} . $1/4$ of the outcomes of Ω lie in \mathcal{A} , and $1/4$ of the outcomes in \mathcal{B} lie in $\mathcal{A} \cap \mathcal{B}$. On the **right**, they are not. Very few of the outcomes in \mathcal{B} lie in $\mathcal{B} \cap \mathcal{A}$, so that observing \mathcal{B} means that \mathcal{A} becomes less likely, because very few of the outcomes in \mathcal{B} also lie in $\mathcal{A} \cap \mathcal{B}$.

Worked example 5.25 *A fair coin*

A **fair** coin (or die, or whatever) is one where each outcome has the same probability. A fair coin has two sides, so the probability of each outcome must be $1/2$. We flip this coin twice - what is the probability we see two heads?

Solution: The flips are independent. Write \mathcal{H}_1 for the event that the first flip comes up heads, \mathcal{H}_2 for the event the second comes up heads. We have $\mathcal{H}_1 = \{HH, HT\}$ and $\mathcal{H}_2 = \{HH, TH\}$. We seek $P(\mathcal{H}_1 \cap \mathcal{H}_2) = P(\mathcal{H}_1)P(\mathcal{H}_2)$. The coin is fair, so $P(\mathcal{H}_1) = P(\mathcal{H}_2) = 1/2$. So the probability is $\frac{1}{4}$.

The reasoning of example 25 is moderately rigorous (if you want real rigor, you should specify the event space, etc.; I assumed that it would be obvious), but it's a bit clumsy for everyday use. The rule to remember is this: the probability that both events occur is $P(\mathcal{A} \cap \mathcal{B})$ and, if \mathcal{A} and \mathcal{B} are independent, then this is $P(\mathcal{A})P(\mathcal{B})$.

Worked example 5.26 *A fair die*

The space of outcomes for a fair six-sided die is

$$\{1, 2, 3, 4, 5, 6\}.$$

The die is fair means that each outcome has the same probability. Now we toss two fair dice — with what probability do we get two threes?

Solution:

$$\begin{aligned} P(\text{first toss yields } 3 \cap \text{second toss yields } 3) &= P(\text{first toss yields } 3) \times \\ &\quad P(\text{second toss yields } 3) \\ &= (1/6)(1/6) \\ &= 1/36 \end{aligned}$$

Worked example 5.27 *Find the Lady, twice*

Assume that the card that is chosen is chosen fairly — that is, each card is chosen with the same probability. The game is played twice, and the cards are reshuffled between games. What is the probability of turning up a Queen and then a Queen again?

Solution: The events are independent, so $1/9$.

Worked example 5.28 *Children*

A couple decides to have two children. Genders are assigned to children at random, fairly, independently and at birth (our models have to abstract a little!). What is the probability of having a boy and then a girl?

Solution:

$$\begin{aligned} P(\text{first is boy} \cap \text{second is girl}) &= P(\text{first is boy})P(\text{second is girl}) \\ &= (1/2)(1/2) \\ &= 1/4 \end{aligned}$$

You can use the expression to tell whether events are independent or not. Quite small changes to a problem affect whether events are independent. For ex-

ample, simply removing a card from a deck can make some events dependent.

Worked example 5.29 *Independent cards*

We shuffle a standard deck of 52 cards and draw one card. The event \mathcal{A} is “the card is a red suit” and the event \mathcal{B} is “the card is a 10”. Are they independent?

Solution: $P(\mathcal{A}) = 1/2$, $P(\mathcal{B}) = 1/13$ and in example 14 we determined $P(\mathcal{A} \cap \mathcal{B}) = 2/52$. But $2/52 = 1/26 = P(\mathcal{A})P(\mathcal{B})$, so they are independent.

Worked example 5.30 *Independent cards*

We take a standard deck of cards, and remove the ten of hearts. We now shuffle this deck, and draw one card. The event \mathcal{A} is “the card is a red suit” and the event \mathcal{B} is “the card is a 10”. Are they independent?

Solution: These are not independent because $P(\mathcal{A}) = 25/51$, $P(\mathcal{B}) = 3/51$ and $P(\mathcal{A} \cap \mathcal{B}) = 1/51 \neq P(\mathcal{A})P(\mathcal{B}) = 75/(51^2)$

The probability of a sequence of independent events can become very small very quickly, and this often misleads people.

Worked example 5.31 *Accidental DNA Matches*

I search a DNA database with a sample. Each time I attempt to match this sample to an entry in the database, there is a probability of an accidental chance match of $1e - 4$. Chance matches are independent. There are 20, 000 people in the database. What is the probability I get at least one match, purely by chance?

Solution: This is $1 - P(\text{no chance matches})$. But $P(\text{no chance matches})$ is much smaller than you think. We have

$$\begin{aligned} P(\text{no chance matches}) &= P \left(\begin{array}{c} \text{no chance match to record 1} \cap \\ \text{no chance match to record 2} \cap \\ \dots \cap \\ \text{no chance match to record 20, 000} \end{array} \right) \\ &= P(\text{no chance match to a record})^{20,000} \\ &= (1 - 1e - 4)^{20,000} \end{aligned}$$

so the probability is about 0.86 that you get at least one match by chance. Notice that if the database gets bigger, the probability grows; so at 40, 000 the probability of one match by chance is 0.98.

5.1.7 Permutations and Combinations

Counting outcomes in an event can require pretty elaborate combinatorial arguments. One form of argument that is particularly important is to reason about permutations. You should recall that the number of permutations of k items is $k!$.

Worked example 5.32 *Counting outcomes with permutations*

I flip a coin k times. How many of the possible outcomes have exactly r heads?

Solution: Here is one natural way to think about this problem. Each outcome we are interested in is a string of r H's and $k - r$ T's, and we need to count the number of such strings. We can do so with permutations. Write down any string of r H's and $k - r$ T's. Any other string of r H's and $k - r$ T's is a permutation of this one. However, many of these permutations simply swap one H with another, or one T with another. The total number of permutations of a string of k entries is $k!$. The number of permutations that swap H's with one another is $r!$ (because there are r H's), and the number of permutations that swap T's with one another is $(k - r)!$. The total number of strings must then be

$$\frac{k!}{r!(k - r)!} = \binom{k}{r}$$

There is another way to think about example 32, which is more natural if you've seen combinations before. You start with a string of k T's, then you choose r distinct elements to turn into H's. The number of choices of r distinct elements in k items is:

$$\frac{k!}{r!(k - r)!} = \binom{k}{r},$$

so there must be this number of strings. We can use this result to investigate our model of probability as frequency.

Worked example 5.33 *A fair coin, revisited*

We conduct N experiments, where each experiment is to flip a fair coin twice. In what fraction of these experiments do we see both sides of the coin?

Solution: The sample space is $\{HH, HT, TH, TT\}$. The coin is fair, so each outcome has the same probability. This means that $(1/4)N$ experiments produce HT ; $(1/4)N$ produce TH ; and so on. We see both sides of the coin in an experiment for about $(1/2)N$ experiments.

Example 33 might seem like a contradiction to you. I claimed that we could

interpret P as the relative frequency of outcomes, and that my coin was fair; but, in only half of my experiments did I see both sides of the coin. In the other half, the coin behaved as if the probability of seeing one side is 1, and the probability of seeing the other side is 0. This occurs because the number of flips *in each experiment* is very small. If the number of flips is larger, you are much more likely to see about the right frequencies.

Worked example 5.34 *A fair coin, yet again*

We conduct N experiments, where each experiment is to flip a fair coin six times. In what fraction of these experiments do we get three heads and three tails?

Solution: There are $64 = 2^6$ outcomes in total for six coin flips. Each has the same probability. By the argument of example 32, there are

$$\frac{6!}{3!3!} = 20$$

outcomes with three heads. So the fraction is $20/64 = 1/3$.

At first glance, example 34 suggests that making the number of experiments get bigger doesn't help. But in the definition of probability, I said that there would be "about" $N \times P$ experiments with the outcome of interest.

Worked example 5.35 *A fair coin, yet again*

We conduct N experiments, where each experiment is to flip a fair coin 10 times. In what fraction of these experiments do we get between 4 and 6 heads?

Solution: There are $1024 = 2^{10}$ outcomes for an experiment. We are interested in the four head outcomes, the five head outcomes and the six head outcomes. Using the argument of example 34 gives

$$\begin{aligned} \text{total outcomes} &= 4\text{H outcomes} + 5\text{H outcomes} + 6\text{H outcomes} \\ &= \frac{10!}{4!6!} + \frac{10!}{5!5!} + \frac{10!}{6!4!} \\ &= 210 + 252 + 210 \\ &= 672 \end{aligned}$$

so in $672/1024 \approx 0.66$ of the experiments, we will see between four and six heads

Worked example 5.36 *A fair coin, and a lot of flipping*

We conduct N experiments, where each experiment is to flip a fair coin 100 times. In what fraction of these experiments do we get between 45 and 65 heads?

Solution: There are 2^{100} outcomes for an experiment. Using the argument of example 35 gives

$$\text{total outcomes} = \sum_{i=45}^{65} \frac{100!}{i!(100-i)!}$$

which isn't particularly easy to evaluate.

As these examples suggest, if an experiment consists of flipping a large number of coins, then a high fraction of those experiments will show heads with a frequency very close to the right number. We will establish this later, with rather more powerful machinery.

Worked example 5.37 *Overbooking - 1*

An airline has a regular flight with six seats. It always sells seven tickets. Passengers turn up for the flight with probability p , and do so independent of other passengers. What is the probability that the flight is overbooked?

Solution: This is like a coin-flip problem; think of each passenger as a biased coin. With probability p , it comes up T (for *turn up*) and with probability $(1-p)$, it turns up N (for *no-show*). This coin is flipped seven times, and we are interested in the probability that there are seven T 's. This is p^7 , because the flips are independent.

Worked example 5.38 *Overbooking - 2*

An airline has a regular flight with six seats. It always sells eight tickets. Passengers turn up for the flight with probability p , and do so independent of other passengers. What is the probability that six passengers arrive? (i.e. the flight is not overbooked or underbooked).

Solution: Now we flip the coin eight times, and are interested in the probability of getting exactly six T 's. We get, by the reasoning of worked example 32,

$$\frac{8!}{2!6!}p^6(1-p)^2 = 28p^6(1-p)^2$$

Worked example 5.39 *Overbooking - 3*

An airline has a regular flight with six seats. It always sells eight tickets. Passengers turn up for the flight with probability p , and do so independent of other passengers. What is the probability that the flight is overbooked?

Solution: Now we flip the coin eight times, and are interested in the probability of getting more than six T 's. This is the union of two disjoint events (seven T 's and eight T 's). For the case of seven T 's, one flip must be N ; there are eight choices. For the case of eight T 's, all eight flips must be T , and there is only one way to achieve this. So the probability the flight is overbooked is

$$\begin{aligned} P(\text{overbooked}) &= P(7 T\text{'s} \cup 8 T\text{'s}) \\ &= P(7 T\text{'s}) + P(8 T\text{'s}) \\ &= 8p^7(1-p) + p^8 \end{aligned}$$

Worked example 5.40 *Overbooking - 4*

An airline has a regular flight with s seats. It always sells t tickets. Passengers turn up for the flight with probability p , and do so independent of other passengers. What is the probability that u passengers turn up?

Solution: Now we flip the coin t times, and are interested in the probability of getting u T 's. By the argument of worked example 32, there are

$$\frac{t!}{u!(t-u)!}$$

disjoint outcomes with u T 's and $t-u$ N 's. Each such outcome is independent, and has probability $p^u(1-p)^{t-u}$. So

$$P(u \text{ passengers turn up}) = \frac{t!}{u!(t-u)!} p^u (1-p)^{t-u}$$

Worked example 5.41 *Overbooking - 5*

An airline has a regular flight with s seats. It always sells t tickets. Passengers turn up for the flight with probability p , and do so independent of other passengers. What is the probability that the flight is oversold?

Solution: We need $P(\{s+1 \text{ turn up}\} \cup \{s+2 \text{ turn up}\} \cup \dots \cup \{t \text{ turn up}\})$. But the events $\{i \text{ turn up}\}$ and $\{j \text{ turn up}\}$ are disjoint if $i \neq j$. So we can exploit example 40, and write

$$\begin{aligned} P(\text{oversold}) &= P(\{s+1 \text{ turn up}\}) + P(\{s+2 \text{ turn up}\}) + \\ &\quad \dots P(\{t \text{ turn up}\}) \\ &= \sum_{i=s+1}^t P(\{i \text{ turn up}\}) \\ &= \sum_{i=s+1}^t \frac{t!}{i!(t-i)!} p^i (1-p)^{t-i} \end{aligned}$$

5.2 CONDITIONAL PROBABILITY

If you throw a fair die twice and add the numbers, then the probability of getting a number less than six is $\frac{10}{36}$. Now imagine you know that the first die came up

three. In this case, the probability that the sum will be less than six is $\frac{1}{3}$, which is slightly larger. If the first die came up four, then the probability the sum will be less than six is $\frac{1}{6}$, which is rather less than $\frac{10}{36}$. If the first die came up one, then the probability that the sum is less than six becomes $\frac{2}{3}$, which is much larger.

Each of these probabilities is an example of a **conditional probability**. We assume we have a space of outcomes and a collection of events. The conditional probability of \mathcal{B} , conditioned on \mathcal{A} , is the probability that \mathcal{B} occurs given that \mathcal{A} has definitely occurred. We write this as

$$P(\mathcal{B}|\mathcal{A})$$

5.2.1 Evaluating Conditional Probabilities

Now to get an expression for $P(\mathcal{B}|\mathcal{A})$, notice that, because \mathcal{A} is known to have occurred, our space of outcomes or sample space is now reduced to \mathcal{A} . We know that our outcome lies in \mathcal{A} ; $P(\mathcal{B}|\mathcal{A})$ is the probability that it also lies in $\mathcal{B} \cap \mathcal{A}$.

The outcome lies in \mathcal{A} , and so it must lie in either $P(\mathcal{B} \cap \mathcal{A})$ or in $P(\mathcal{B}^c \cap \mathcal{A})$. This means that

$$P(\mathcal{B}|\mathcal{A}) + P(\mathcal{B}^c|\mathcal{A}) = 1.$$

Now recall the idea of probabilities as relative frequencies. If $P(\mathcal{C} \cap \mathcal{A}) = kP(\mathcal{B} \cap \mathcal{A})$, this means that we will see outcomes in $\mathcal{C} \cap \mathcal{A}$ k times as often as we will see outcomes in $\mathcal{B} \cap \mathcal{A}$. But this must apply even if we know in advance that the outcome is in \mathcal{A} . So we must have

$$P(\mathcal{B}|\mathcal{A}) \propto P(\mathcal{B} \cap \mathcal{A}).$$

Now we need to determine the constant of proportionality; write c for this constant, meaning

$$P(\mathcal{B}|\mathcal{A}) = cP(\mathcal{B} \cap \mathcal{A}).$$

We have that

$$P(\mathcal{B}|\mathcal{A}) + P(\mathcal{B}^c|\mathcal{A}) = cP(\mathcal{B} \cap \mathcal{A}) + cP(\mathcal{B}^c \cap \mathcal{A}) = cP(\mathcal{A}) = 1,$$

so that

$$P(\mathcal{B}|\mathcal{A}) = \frac{P(\mathcal{B} \cap \mathcal{A})}{P(\mathcal{A})}.$$

Using the “size” metaphor, this says the probability that an outcome, *which is known to be in \mathcal{A}* , is also in \mathcal{B} is the fraction of \mathcal{A} that is also in \mathcal{B} .

Another, very useful, way to write this expression is

$$P(\mathcal{B}|\mathcal{A})P(\mathcal{A}) = P(\mathcal{B} \cap \mathcal{A}).$$

Now, since $\mathcal{B} \cap \mathcal{A} = \mathcal{A} \cap \mathcal{B}$, we must have that

$$P(\mathcal{B}|\mathcal{A}) = \frac{P(\mathcal{A}|\mathcal{B})P(\mathcal{B})}{P(\mathcal{A})}$$

Worked example 5.42 *Two dice*

We throw two fair dice. What is the conditional probability that the sum of spots on both dice is greater than six, conditioned on the event that the first die comes up five?

Solution: Write the event that the first die comes up 5 as \mathcal{F} , and the event the sum is greater than six as \mathcal{S} . There are five outcomes where the first die comes up 5 and the number is greater than 6, so $P(\mathcal{F} \cap \mathcal{S}) = 5/36$. $P(\mathcal{S}|\mathcal{F}) = P(\mathcal{F} \cap \mathcal{S})/P(\mathcal{F}) = (5/36)/(1/6) = 5/6$.

Notice that $\mathcal{A} \cap \mathcal{B}$ and $\mathcal{A} \cap \mathcal{B}^c$ are disjoint sets, and that $\mathcal{A} = (\mathcal{A} \cap \mathcal{B}) \cup (\mathcal{A} \cap \mathcal{B}^c)$. So, because $P(\mathcal{A}) = P(\mathcal{A} \cap \mathcal{B}) + P(\mathcal{A} \cap \mathcal{B}^c)$, we have

$$P(\mathcal{A}) = P(\mathcal{A}|\mathcal{B})P(\mathcal{B}) + P(\mathcal{A}|\mathcal{B}^c)P(\mathcal{B}^c)$$

a tremendously important and useful fact. Another version of this fact is also very useful. Assume we have a set of disjoint sets \mathcal{B}_i . These sets must have the property that (a) $\mathcal{B}_i \cap \mathcal{B}_j = \emptyset$ for $i \neq j$ and (b) they cover \mathcal{A} , meaning that $\mathcal{A} \cap (\cup_i \mathcal{B}_i) = \mathcal{A}$. Then, because $P(\mathcal{A}) = \sum_i P(\mathcal{A} \cap \mathcal{B}_i)$, so we have

$$P(\mathcal{A}) = \sum_i P(\mathcal{A}|\mathcal{B}_i)P(\mathcal{B}_i)$$

Worked example 5.43 *Car factories*

There are two car factories, A and B . Each year, factory A produces 1000 cars, of which 10 are lemons. Factory B produces 2 cars, each of which is a lemon. All cars go to a single lot, where they are thoroughly mixed up. I buy a car.

- What is the probability it is a lemon?
- What is the probability it came from factory B ?
- The car is now revealed to be a lemon. What is the probability it came from factory B , conditioned on the fact it is a lemon?

Solution:

- Write the event the car is a lemon as \mathcal{L} . There are 1002 cars, of which 12 are lemons. The probability that I select any given car is the same, so we have $12/1002$.
- Same argument yields $2/1002$.
- Write \mathcal{B} for the event the car comes from factory B . I need $P(\mathcal{B}|\mathcal{L})$. This is $P(\mathcal{L}|\mathcal{B})P(\mathcal{B})/P(\mathcal{L}) = (1 \times 2/1002)/(12/1002) = 1/6$.

Worked example 5.44 *Royal flushes in poker - 1*

This exercise is after Stirzaker, p. 51.

You are playing a straightforward version of poker, where you are dealt five cards face down. A royal flush is a hand of AKQJ10 all in one suit. What is the probability that you are dealt a royal flush?

Solution: This is

$$\frac{\text{number of hands that are royal flushes, ignoring card order}}{\text{total number of different five card hands, ignoring card order}}$$

There are four hands that are royal flushes (one for each suit). Now the total number of five card hands is

$$\binom{52}{5} = 2598960$$

so we have

$$\frac{4}{2598960} = \frac{1}{649740}.$$

Worked example 5.45 *Royal flushes in poker - 2*

This exercise is after Stirzaker, p. 51.

You are playing a straightforward version of poker, where you are dealt five cards face down. A royal flush is a hand of AKQJ10 all in one suit. The fifth card that you are dealt lands face up. It is the nine of spades. What now is the probability that you have been dealt a royal flush? (i.e. what is the conditional probability of getting a royal flush, conditioned on the event that one card is the nine of spades)

Solution: No hand containing a nine of spades is a royal flush, so this is easily zero.

Worked example 5.46 *Royal flushes in poker - 3*

This exercise is after Stirzaker, p. 51.

You are playing a straightforward version of poker, where you are dealt five cards face down. A royal flush is a hand of AKQJ10 all in one suit. The fifth card that you are dealt lands face up. It is the Ace of spades. What now is the probability that you have been dealt a royal flush? (i.e. what is the conditional probability of getting a royal flush, conditioned on the event that one card is the Ace of spades)

Solution: There are two ways to do this. The easiest is to notice that the answer is the probability that the other four cards are KQJ10 of spades, which is

$$1 / \binom{51}{4} = \frac{1}{249900}.$$

Harder is to consider the events

\mathcal{A} = event that you receive a royal flush and last card is the ace of spades
and

\mathcal{B} = event that the last card you receive is the ace of spades,

and the expression

$$P(\mathcal{A}|\mathcal{B}) = \frac{P(\mathcal{A} \cap \mathcal{B})}{P(\mathcal{B})}.$$

Now $P(\mathcal{A}) = \frac{1}{52}$. $P(\mathcal{A} \cap \mathcal{B})$ is given by

$$\frac{\text{number of five card royal flushes where card five is Ace of spades}}{\text{total number of different five card hands}}.$$

where we DO NOT ignore card order. This is

$$\frac{4 \times 3 \times 2 \times 1}{52 \times 51 \times 50 \times 49 \times 48}$$

yielding

$$P(\mathcal{A}|\mathcal{B}) = \frac{1}{249900}.$$

Notice the interesting part: the conditional probability is rather larger than the probability. If you see this ace, the conditional probability is $\frac{13}{5}$ times the probability that you will get a flush if you don't. Seeing this card has really made a difference.

Worked example 5.47 *False positives*

After Stirzaker, p55. You have a blood test for a rare disease that occurs by chance in 1 person in 100,000. If you have the disease, the test will report that you do with probability 0.95 (and that you do not with probability 0.05). If you do not have the disease, the test will report a false positive with probability $1e-3$. If the test says you do have the disease, what is the probability it is correct?

Solution: Write S for the event you are sick and R for the event the test reports you are sick. We need $P(S|R)$. We have

$$\begin{aligned}
 P(S|R) &= \frac{P(R|S)P(S)}{P(R)} \\
 &= \frac{P(R|S)P(S)}{P(R|S)P(S) + P(R|S^c)P(S^c)} \\
 &= \frac{0.95 \times 1e-5}{0.95 \times 1e-5 + 1e-3 \times (1 - 1e-5)} \\
 &= 0.0094
 \end{aligned}$$

which should strike you as being a bit alarming. The disease is so rare that the test is almost useless.

Worked example 5.48 *False positives -2*

After Stirzaker, p55. You want to *design* a blood test for a rare disease that occurs by chance in 1 person in 100,000. If you have the disease, the test will report that you do with probability p (and that you do not with probability $(1 - p)$). If you do not have the disease, the test will report a false positive with probability q . You want to choose the value of p so that if the test says you have the disease, there is at least a 50% probability that you do.

Solution: Write S for the event you are sick and R for the event the test reports you are sick. We need $P(S|R)$. We have

$$\begin{aligned} P(S|R) &= \frac{P(R|S)P(S)}{P(R)} \\ &= \frac{P(R|S)P(S)}{P(R|S)P(S) + P(R|S^c)P(S^c)} \\ &= \frac{p \times 1e - 5}{p \times 1e - 5 + q \times (1 - 1e - 5)} \\ &\geq 0.5 \end{aligned}$$

which means that $p \geq 99999q$ which should strike you as being very alarming indeed, because $p \leq 1$ and $q \geq 0$. One plausible pair of values is $q = 1e - 5$, $p = 1 - 1e - 5$. The test has to be spectacularly accurate to be of any use.

5.2.2 The Prosecutors Fallacy

It is quite easy to make mistakes in conditional probability. Several such mistakes have names, because they're so common. One is the **prosecutor's fallacy**. This often occurs in the following form: A prosecutor has evidence \mathcal{E} against a suspect. Write \mathcal{I} for the event that the suspect is innocent. The evidence has the property that $P(\mathcal{E}|\mathcal{I})$ is extremely small. The prosecutor argues that the suspect must be guilty, because $P(\mathcal{E}|\mathcal{I})$ is so small, and this is the fallacy.

The problem here is that the conditional probability of interest is $P(\mathcal{I}|\mathcal{E})$ (rather than $P(\mathcal{E}|\mathcal{I})$). The fact that $P(\mathcal{E}|\mathcal{I})$ is small doesn't mean that $P(\mathcal{I}|\mathcal{E})$ is small, because

$$P(\mathcal{I}|\mathcal{E}) = \frac{P(\mathcal{E}|\mathcal{I})P(\mathcal{I})}{P(\mathcal{E})} = \frac{P(\mathcal{E}|\mathcal{I})P(\mathcal{I})}{(P(\mathcal{E}|\mathcal{I})P(\mathcal{I}) + P(\mathcal{E}|\mathcal{I}^c)(1 - P(\mathcal{I})))}$$

Notice how, if $P(\mathcal{I})$ is large or if $P(\mathcal{E}|\mathcal{I}^c)$ is much smaller than $P(\mathcal{E}|\mathcal{I})$, then $P(\mathcal{I}|\mathcal{E})$ could be close to one. The question to look at is not how unlikely the evidence is if the subject is innocent; instead, the question is how likely the subject is to be guilty compared to some other source of the evidence. These are two very different questions.

One useful analogy may be helpful. If you buy a lottery ticket (\mathcal{L}), the probability of winning (\mathcal{W}) is small. So $P(\mathcal{W}|\mathcal{L})$ may be very small. But $P(\mathcal{L}|\mathcal{W})$ is 1 — the winner is always someone who bought a ticket.

The prosecutor’s fallacy has contributed to a variety of miscarriages of justice. One famous incident involved a mother, Sally Clark, convicted of murdering two of her children. Expert evidence by paediatrician Roy Meadow argued that the probability of both deaths resulting from Sudden Infant Death Syndrome was extremely small. Her first appeal cited, among other grounds, statistical error in the evidence (you should spot the prosecutors fallacy; others were involved, too). The appeals court rejected this appeal, calling the statistical point “a sideshow”. This prompted a great deal of controversy, both in the public press and various professional journals, including a letter from the then president of the Royal Statistical Society to the Lord Chancellor, pointing out that “*statistical evidence . . . (should be) . . . presented only by appropriately qualified statistical experts*”. A second appeal (on other grounds) followed, and was successful. The appellate judges specifically criticized the statistical evidence, although it was not a point of appeal. Clark never recovered from this horrific set of events and died in tragic circumstances shortly after the second appeal. Roy Meadow was then struck off the rolls for serious professional misconduct as an expert witness, a ruling he appealed successfully. You can find a more detailed account of this case, with pointers to important documents including the letter to the Lord Chancellor (which is well worth reading), at http://en.wikipedia.org/wiki/Roy_Meadow; there is further material on the prosecutors fallacy at http://en.wikipedia.org/wiki/Prosecutor%27s_fallacy.

5.2.3 Independence and Conditional Probability

As we have seen, two events are **independent** if

$$P(\mathcal{A} \cap \mathcal{B}) = P(\mathcal{A})P(\mathcal{B}).$$

If two events \mathcal{A} and \mathcal{B} are independent, then

$$P(\mathcal{A}|\mathcal{B}) = P(\mathcal{A})$$

and

$$P(\mathcal{B}|\mathcal{A}) = P(\mathcal{B}).$$

Again, this means that knowing that \mathcal{A} occurred tells you nothing about \mathcal{B} — the probability that \mathcal{B} will occur is the same whether you know that \mathcal{A} occurred or not.

Events $\mathcal{A}_1 \dots \mathcal{A}_n$ are **pairwise independent** if each pair is independent (i.e. \mathcal{A}_1 and \mathcal{A}_2 are independent, etc.). They are **independent** if for any collection of distinct indices $i_1 \dots i_k$ we have

$$P(\mathcal{A}_{i_1} \cap \dots \cap \mathcal{A}_{i_k}) = P(\mathcal{A}_{i_1}) \dots P(\mathcal{A}_{i_k})$$

Notice that independence is a much stronger assumption than pairwise independence.

Worked example 5.49 *Cards and pairwise independence*

We draw three cards from a properly shuffled standard deck, with replacement and reshuffling (i.e., draw a card, make a note, return to deck, shuffle, draw the next, make a note, shuffle, draw the third). Let \mathcal{A} be the event that “card 1 and card 2 have the same suit”; let \mathcal{B} be the event that “card 2 and card 3 have the same suit”; let \mathcal{C} be the event that “card 1 and card 3 have the same suit”. Show these events are pairwise independent, but not independent.

Solution: By counting, you can check that $P(\mathcal{A}) = 1/4$; $P(\mathcal{B}) = 1/4$; and $P(\mathcal{A} \cap \mathcal{B}) = 1/16$, so that these two are independent. This argument works for other pairs, too. But $P(\mathcal{C} \cap \mathcal{A} \cap \mathcal{B}) = 1/16$ which is not $1/4^3$, so the events are not independent; this is because the third event is logically implied by the first two.

We usually do not have the information required to prove that events are independent. Instead, we use intuition (for example, two flips of the same coin are likely to be independent unless there is something very funny going on) or simply choose to apply models in which some variables are independent.

Some events are pretty obviously independent. On other occasions, one needs to think about whether they are independent or not. Sometimes, it is reasonable to choose to model events as being independent, even though they might not be exactly independent. In cases like this, it is good practice to state your assumptions, because it helps you to keep track of what your model means. For example, we have worked with the event that a person, selected fairly and randomly from a set of people in a room, has a birthday on a particular day of the year. We assumed that, for different people, the events are independent. This seems like a fair assumption, but one might want to be cautious if you know that the people in the room are drawn from a population where multiple births are common.

Independent events can lead very quickly to very small probabilities, as we saw in example 31. This can mislead intuition quite badly, and lead to serious problems. In particular, these small probabilities can interact with the prosecutor’s fallacy in a dangerous way. In example 31, we saw how the probability of getting a chance match in a large DNA database could be quite big, even though the probability of a single match is small. One version of the prosecutors fallacy is to argue that, because the probability of a single match is small, the person who matched the DNA must have committed the crime. The fallacy is to ignore the fact that the probability of a chance match to a large database is quite high.

People quite often reason poorly about independent events. The most common problem is known as the **gambler’s fallacy**. This occurs when you reason that the probability of an independent event has been changed by previous outcomes. For example, imagine I toss a coin that is known to be fair 20 times and get 20 heads. The probability that the next toss will result in a head has not changed at all — it is still 0.5 — but many people will believe that it has changed. This idea is also sometimes referred to as **antichance**.

It might in fact be sensible to behave as if you're committing some version of the gambler's fallacy in real life, because you hardly ever know for sure that your model is right. So in the coin tossing example, if the coin wasn't known to be fair, it might be reasonable to assume that it has been weighted in some way, and so to believe that the more heads you see, the more likely you will see a head in the next toss. At time of writing, Wikipedia has some fascinating stories about the gambler's fallacy; apparently, in 1913, a roulette wheel in Monte Carlo produced black 26 times in a row, and gamblers lost an immense amount of money betting on red. Here the gambler's reasoning seems to have been that the universe should ensure that probabilities produce the right frequencies in the end, and so will adjust the outcome of the next spin of the wheel to balance the sums. This is an instance of the gambler's fallacy. However, the page also contains the story of one Joseph Jagger, who hired people to keep records of the roulette wheels, and notice that one wheel favored some numbers (presumably because of some problem with balance). He won a lot of money, until the casino started more careful maintenance on the wheels. This isn't the gambler's fallacy; instead, he noticed that the numbers implied that the wheel was not a fair randomizer. He made money because the casino's odds on the bet assumed that it was fair.

5.3 EXAMPLE: THE MONTY HALL PROBLEM

Careless thinking about probability, particularly conditional probability, can cause wonderful confusion. The Monty Hall problem is a relatively simple exercise in conditional probability. Nonetheless, it has been the subject of extensive, lively, and often quite inaccurate correspondence in various national periodicals — it seems to catch the attention, which is why we describe it in some detail. The problem works like this: There are three doors. Behind one is a car. Behind each of the others is a goat. The car and goats are placed randomly and fairly, so that the probability that there is a car behind each door is the same. You will get the object that lies behind the door you choose at the end of the game. The goats are interchangeable, and, for reasons of your own, you would prefer the car to a goat.

The game goes as follows. You select a door. The host then opens a door and shows you a goat. You must now choose to either keep your door, or switch to the other door. What should you do?

You cannot tell what to do, by the following argument. Label the door you chose at the start of the game 1; the other doors 2 and 3. Write C_i for the event that the car lies behind door i . Write G_m for the event that a goat is revealed behind door m , where m is the number of the door where the goat was revealed (which could be 1, 2, or 3). You need to know $P(C_1|G_m)$. But

$$P(C_1|G_m) = \frac{P(G_m|C_1)P(C_1)}{P(G_m|C_1)P(C_1) + P(G_m|C_2)P(C_2) + P(G_m|C_3)P(C_3)}$$

and you do not know $P(G_m|C_1)$, $P(G_m|C_2)$, $P(G_m|C_3)$, because you don't know the rule by which the host chooses which door to open to reveal a goat. Different rules lead to quite different analyses.

There are several possible rules for the host to show a goat:

- **Rule 1:** choose a door uniformly at random.

- **Rule 2:** choose from the doors with goats behind them *that are not door 1* uniformly and at random.
- **Rule 3:** if the car is at 1, then choose 2; if at 2, choose 3; if at 3, choose 1.
- **Rule 4:** choose from the doors with goats behind them uniformly and at random.

We should keep track of the rules in the conditioning, so we write $P(G_m|C_1, r_1)$ for the conditional probability that a goat was revealed behind door m when the car is behind door 1, using rule 1 (and so on). This means we are interested in

$$P(C_1|G_m, r_n) = \frac{P(G_m|C_1, r_n)P(C_1)}{P(G_m|C_1, r_n)P(C_1) + P(G_m|C_2, r_n)P(C_2) + P(G_m|C_3, r_n)P(C_3)}.$$

Worked example 5.50 *Monty Hall, rule one*

Assume the host uses rule one, and shows you a goat behind door two. What is $P(C_1|G_2, r_1)$?

Solution: To work this out, we need to know $P(G_2|C_1, r_1)$, $P(G_2|C_2, r_1)$ and $P(G_2|C_3, r_1)$. Now $P(G_2|C_2, r_1)$ must be zero, because the host could not reveal a goat behind door two if there was a car behind that door. Write O_2 for the event the host *chooses* to open door two, and B_2 for the event there happens to be a goat behind door two. These two events are independent — the host chose the door uniformly at random. We can compute

$$\begin{aligned} P(G_2|C_1, r_1) &= P(O_2 \cap B_2|C_1, r_1) \\ &= P(O_2|C_1, r_1)P(B_2|C_1, r_1) \\ &= (1/3)(1) \\ &= 1/3 \end{aligned}$$

where $P(B_2|C_1, r_1) = 1$ because we conditioned on the fact there was a car behind door one, so there is a goat behind each other door. This argument establishes $P(G_2|C_3, r_1) = 1/3$, too. So $P(C_1|G_2, r_1) = 1/2$ — the host showing you the goat does not motivate you to do anything, because if $P(C_1|G_2, r_1) = 1/2$, then $P(C_3|G_2, r_1) = 1/2$, too — there's nothing to choose between the two closed doors.

Worked example 5.51 *Monty Hall, rule two*

Assume the host uses rule two, and shows you a goat behind door two. What is $P(C_1|G_2, r_2)$?

Solution: To work this out, we need to know $P(G_2|C_1, r_2)$, $P(G_2|C_2, r_2)$ and $P(G_2|C_3, r_2)$. Now $P(G_2|C_2, r_2) = 0$, because the host chooses from doors with goats behind them. $P(G_2|C_1, r_2) = 1/2$, because the host chooses uniformly and at random from doors with goats behind them that are not door one; if the car is behind door one, there are two such doors. $P(G_2|C_3, r_2) = 1$, because there is only one door that (a) has a goat behind it and (b) isn't door one. Plug these numbers into the formula, to get $P(C_1|G_2, r_2) = 1/3$. This is the source of all the fuss. It says that, if you know the host is using rule two, you should switch doors if the host shows you a goat behind door two (because $P(C_3|G_2, r_2) = 2/3$).

Notice what is happening: if the car is behind door three, then the *only* choice of goat for the host is the goat behind two. So by choosing a door under rule two, the host is signalling some information to you, which you can use. By using rule three, the host can tell you precisely where the car is (exercises).

Many people find the result of example 51 counterintuitive, and object (sometimes loudly, in newspaper columns, letters to the editor, etc.). One example that some people find helpful is an extreme case. Imagine that, instead of three doors, there are 1002. The host is using rule two, modified in the following way: open all but one of the doors that are not door one, choosing only doors that have goats behind them to open. You choose door one; the host opens 1000 doors — say, all but doors one and 1002. What would you do?

5.4 WHAT YOU SHOULD REMEMBER

You should be able to:

- Write out a set of outcomes for an experiment.
- Construct an event space.
- Compute the probabilities of outcomes and events.
- Determine when events are independent.
- Compute the probabilities of outcomes by counting events, when the count is straightforward.
- Compute a conditional probability.

You should remember:

- The definition of an event space.
- The properties of a probability function.

- The definition of independence.
- The definition of conditional probability.

PROBLEMS

Outcomes

- 5.1. You roll a four sided die. What is the space of outcomes?
- 5.2. King Lear decides to allocate three provinces (1, 2, and 3) to his daughters (Goneril, Regan and Cordelia - read the book) at random. Each gets one province. What is the space of outcomes?
- 5.3. You randomly wave a flyswatter at a fly. What is the space of outcomes?
- 5.4. You read the book, so you know that King Lear had family problems. As a result, he decides to allocate two provinces to one daughter, one province to another daughter, and no provinces to the third. Because he's a bad problem solver, he does so at random. What is the space of outcomes?

Probability of outcomes

- 5.5. You roll a fair four sided die. What is the probability of getting a 3?
- 5.6. You roll a fair four sided die, and then a fair six sided die. You add the numbers on the two dice. What is the probability the result is even?
- 5.7. You roll a fair 20 sided die. What is the probability of getting an even number?
- 5.8. You roll a fair five sided die. What is the probability of getting an even number?

Events

- 5.9. At a particular University, $1/2$ of the students drink alcohol and $1/3$ of the students smoke cigarettes.
 - (a) What is the largest possible fraction of students who do neither?
 - (b) It turns out that, in fact, $1/3$ of the students do neither. What fraction of the students does both?
- 5.10. I flip two coins. What one set needs to be added to this collection of sets to form an event space?

$$\Sigma = \{\emptyset, \Omega, \{TH\}, \{HT, TH, TT\}, \{HH\}, \{HT, TT\}, \{HH, TH\}\}$$

Probability of Events

- 5.11. Assume each outcome in Ω has the same probability. In this case, show

$$P(\mathcal{E}) = \frac{\text{Number of outcomes in } \mathcal{E}}{\text{Total number of outcomes in } \Omega}$$

- 5.12. You flip a fair coin three times. What is the probability of seeing HTH? (i.e. Heads, then Tails, then Heads)
- 5.13. You flip a fair coin three times. What is the probability of seeing two heads and one tail?
- 5.14. You remove the king of hearts from a standard deck of cards, then shuffle it and draw a card.
 - (a) What is the probability this card is a king?
 - (b) What is the probability this card is a heart?

- 5.15. You shuffle a standard deck of cards, then draw four cards.
- What is the probability all four are the same suit?
 - What is the probability all four are red?
 - What is the probability each has a different suit?
- 5.16. You roll three fair six-sided dice and add the numbers. What is the probability the result is even?
- 5.17. You roll three fair six-sided dice and add the numbers. What is the probability the result is even *and* not divisible by 20?
- 5.18. You shuffle a standard deck of cards, then draw seven cards. What is the probability that you see no aces?
- 5.19. Show that $P(\mathcal{A} - (\mathcal{B} \cup \mathcal{C})) = P(\mathcal{A}) - P(\mathcal{A} \cap \mathcal{B}) - P(\mathcal{A} \cap \mathcal{C}) + P(\mathcal{A} \cap \mathcal{B} \cap \mathcal{C})$.
- 5.20. You draw a single card from a standard 52 card deck. What is the probability that it is red?
- 5.21. You remove all heart cards from a standard 52 card deck, then draw a single card from the result. What is the probability that the card you draw is red?

Conditional Probability

- 5.22. You roll two fair six-sided dice. What is the conditional probability the sum of numbers is greater than three, conditioned on the first die coming up even.
- 5.23. You take a standard deck of cards, shuffle it, and remove one card. You then draw a card.
- What is the conditional probability that the card you draw is a red king, conditioned on the removed card being a king?
 - What is the conditional probability that the card you draw is a red king, conditioned on the removed card being a red king?
 - What is the conditional probability that the card you draw is a red king, conditioned on the removed card being a black ace?
- 5.24. A royal flush is a hand of five cards, consisting of Ace, King, Queen, Jack and 10 of a single suit. Poker players like this hand, but don't see it all that often.
- You draw five cards from a standard deck of playing cards. What is the probability of getting a royal flush?
 - You draw three cards from a standard deck of playing cards. These are Ace, King, Queen of hearts. What is the probability that the next two cards you draw will result in a getting a royal flush? (this is the conditional probability of getting a royal flush, conditioned on the first three cards being AKQ of hearts).
- 5.25. You roll a fair five-sided die, and a fair six-sided die.
- What is the probability that the sum of numbers is even?
 - What is the conditional probability that the sum of numbers is even, conditioned on the six-sided die producing an odd number?

Independence

- 5.26. You take a standard deck of cards, shuffle it, and remove both red kings. You then draw a card.
- Is the event {card is red} independent of the event {card is a queen}?
 - Is the event {card is black} independent of the event {card is a king}?

The Monty Hall Problem

- 5.27. Monty Hall, Rule 3:** If the host uses rule 3, then what is $P(C_1|G_2, r_3)$? Do this by computing conditional probabilities.
- 5.28. Monty Hall, Rule 4:** If the host uses rule 4, and shows you a goat behind door 2, what is $P(C_1|G_2, r_4)$? Do this by computing conditional probabilities.

CHAPTER 6

Random Variables and Expectations

6.1 RANDOM VARIABLES

Quite commonly, we would like to deal with numbers that are random. We can do so by linking numbers to the outcome of an experiment. We define a **random variable**:

Definition: 6.1 *Discrete random variable*

Given a sample space Ω , a set of events \mathcal{F} , and a probability function P , and a countable set of real numbers D , a discrete random variable is a function with domain Ω and range D .

This means that for any outcome ω there is a number $X(\omega)$. P will play an important role, but first we give some examples.

Example: 6.1 *Numbers from coins*

We flip a coin. Whenever the coin comes up heads, we report 1; when it comes up tails, we report 0. This is a random variable.

Example: 6.2 *Numbers from coins II*

We flip a coin 32 times. We record a 1 when it comes up heads, and when it comes up tails, we record a 0. This produces a 32 bit random number, which is a random variable.

Example: 6.3 *The number of pairs in a poker hand*

(from Stirzaker). We draw a hand of five cards. The number of pairs in this hand is a random variable, which takes the values 0, 1, 2 (depending on which hand we draw)

A function of a discrete random variable is also a discrete random variable.

Example: 6.4 *Parity of coin flips*

We flip a coin 32 times. We record a 1 when it comes up heads, and when it comes up tails, we record a 0. This produces a 32 bit random number, which is a random variable. The parity of this number is also a random variable.

Associated with any value x of the random variable X are a series of events. The most important is the set of outcomes such that $X = x$, which we can write $\{\omega : X(\omega) = x\}$; it is usual to simplify to $\{X = x\}$, and we will do so. The probability that a random variable X takes the value x is given by $P(\{X = x\})$. This is sometimes written as $P(X = x)$, and rather often written as $P(x)$.

We could also be interested in the set of outcomes such that $X \leq x$ (i.e. in $\{\omega : X(\omega) \leq x\}$), which we will write $\{X \leq x\}$; The probability that X takes the value x is given by $P(\{X \leq x\})$. This is sometimes written as $P(X \leq x)$. Similarly, we could be interested in $\{X > x\}$, and so on.

Definition: 6.2 *The probability distribution of a discrete random variable*

The probability distribution of a discrete random variable is the set of numbers $P(\{X = x\})$ for each value x that X can take. The distribution takes the value 0 at all other numbers. Notice that this is non-negative.

Definition: 6.3 *The cumulative distribution of a discrete random variable*

The cumulative distribution of a discrete random variable is the set of numbers $P(\{X \leq x\})$ for each value x that X can take. Notice that this is a non-decreasing function of x . Cumulative distributions are often written with an f , so that $f(x)$ might mean $P(\{X \leq x\})$.

Worked example 6.1 *Numbers from coins III*

We flip a biased coin 2 times. The flips are independent. The coin has $P(H) = p$, $P(T) = 1 - p$. We record a 1 when it comes up heads, and when it comes up tails, we record a 0. This produces a 2 bit random number, which is a random variable taking the values 0, 1, 2, 3. What is the probability distribution and cumulative distribution of this random variable?

Solution: Probability distribution: $P(0) = (1 - p)^2$; $P(1) = (1 - p)p$; $P(2) = p(1 - p)$; $P(3) = p^2$. Cumulative distribution: $f(0) = (1 - p)^2$; $f(1) = (1 - p)$; $f(2) = p(1 - p) + (1 - p) = (1 - p^2)$; $f(3) = 1$.

Worked example 6.2 *Betting on coins*

One way to get a random variable is to think about the reward for a bet. We agree to play the following game. I flip a coin. The coin has $P(H) = p$, $P(T) = 1 - p$. If the coin comes up heads, you pay me q ; if the coin comes up tails, I pay you r . The number of dollars that change hands is a random variable. What is its probability distribution?

Solution: We see this problem from my perspective. If the coin comes up heads, I get q ; if it comes up tails, I get $-r$. So we have $P(X = q) = p$ and $P(X = -r) = (1 - p)$, and all other probabilities are zero.

6.1.1 Joint and Conditional Probability for Random Variables

All the concepts of probability that we described for events carry over to random variables. This is as it should be, because random variables are really just a way of getting numbers out of events. However, terminology and notation change a bit.

Assume we have two random variables X and Y . The probability that X takes the value x and Y takes the value y could be written as $P(\{X = x\} \cap \{Y = y\})$. It is more usual to write it as $P(x, y)$. You can think of this as a table of values, one for each possible pair of x and y values. This table is usually referred to as the **joint probability distribution** of the random variables. Nothing (except notation) has really changed here, but the change of notation is useful.

We will simplify notation further. Usually, we are interested in random variables, rather than potentially arbitrary outcomes or sets of outcomes. We will write $P(X)$ to denote the probability distribution of a random variable, and $P(x)$ or $P(X = x)$ to denote the probability that that random variable takes a particular value. This means that, for example, the rule we could write as

$$P(\{X = x\} | \{Y = y\})P(\{Y = y\}) = P(\{X = x\} \cap \{Y = y\})$$

will be written as

$$P(x|y)P(y) = P(x, y).$$

This yields **Bayes' rule**, which is important enough to appear in its own box.

Definition: 6.4 *Bayes' rule*

$$P(x|y) = \frac{P(y|x)P(x)}{P(y)}$$

Random variables have another useful property. If $x_0 \neq x_1$, then the event $\{X = x_0\}$ must be disjoint from the event $\{X = x_1\}$. This means that

$$\sum_x P(x) = 1$$

and that, for any y ,

$$\sum_x P(x|y) = 1$$

(if you're uncertain on either of these points, check them by writing them out in the language of events).

Now assume we have the joint probability distribution of two random variables, X and Y . Recall that we write $P(\{X = x\} \cap \{Y = y\})$ as $P(x, y)$. Now consider the sets of outcomes $\{Y = y\}$ for each different value of y . These sets must be disjoint, because y cannot take two values at the same time. Furthermore, each element of the set of outcomes $\{X = x\}$ must lie in one of the sets $\{Y = y\}$. So we have

$$\sum_y P(\{X = x\} \cap \{Y = y\}) = P(\{X = x\})$$

which is usually written as

$$\sum_y P(x, y) = P(x)$$

and is often referred to as the **marginal probability** of X .

Definition: 6.5 *Independent random variables*

The random variables X and Y are **independent** if the events $\{X = x\}$ and $\{Y = y\}$ are independent. This means that

$$P(\{X = x\} \cap \{Y = y\}) = P(\{X = x\})P(\{Y = y\}),$$

which we can rewrite as

$$P(x, y) = P(x)P(y)$$

Worked example 6.3 *Sums and differences of dice*

You throw two dice. The number of spots on the first die is a random variable (call it X); so is the number of spots on the second die (Y). Now define $S = X + Y$ and $D = X - Y$. What is the probability distribution of S and of D ?

Solution: S can have values in the range $2, \dots, 12$. There is only one way to get a $S = 2$; two ways to get $S = 3$; and so on. Using the methods of chapter 1 for each case, the probabilities for $[2, 3, 4, 5, 6, 7, 8, 9, 10, 11, 12]$ are $[1, 2, 3, 4, 5, 6, 5, 4, 3, 2, 1]/36$. Similarly, D can have values in the range $-5, \dots, 5$. Again, using the methods of chapter 1, the probabilities for $[-5, -4, -3, -2, -1, 0, 1, 2, 3, 4, 5]$ are $[1, 2, 3, 4, 5, 6, 5, 4, 3, 2, 1]/36$.

Worked example 6.4 *Sums and differences of dice, II*

Using the terminology of example 3, what is the joint probability distribution of S and D ?

Solution: This is more interesting to display, because it's an 11×11 table. Each entry of the table represents a pair of S, D values. Many pairs can't occur (for example, for $S = 2$, D can only be zero; if S is even, then D must be even; and so on). You can work out the table by checking each case; it's in Table 6.1.

Worked example 6.5 *Sums and differences of dice, III*

Using the terminology of example 3, are X and Y independent? are S and D independent?

Solution: X and Y are clearly independent. But S and D are not. There are several ways to see this. One way is to notice that, if you know $S = 2$, then you know the value of D precisely; but if you know $S = 3$, D could be either 1 or -1 . This means that $P(S|D)$ depends on D , so they're not independent. Another way is to notice that the rank of the table, as a matrix, is 6, which means that it can't be the outer product of two vectors.

$$\frac{1}{36} \times \begin{pmatrix} 0 & 0 & 0 & 0 & 0 & 1 & 0 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 1 & 0 & 1 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 1 & 0 & 1 & 0 & 1 & 0 & 0 & 0 \\ 0 & 0 & 1 & 0 & 1 & 0 & 1 & 0 & 1 & 0 & 0 \\ 0 & 1 & 0 & 1 & 0 & 1 & 0 & 1 & 0 & 1 & 0 \\ 1 & 0 & 1 & 0 & 1 & 0 & 1 & 0 & 1 & 0 & 1 \\ 0 & 1 & 0 & 1 & 0 & 1 & 0 & 1 & 0 & 1 & 0 \\ 0 & 0 & 1 & 0 & 1 & 0 & 1 & 0 & 1 & 0 & 0 \\ 0 & 0 & 0 & 1 & 0 & 1 & 0 & 1 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 1 & 0 & 1 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & 1 & 0 & 0 & 0 & 0 & 0 \end{pmatrix}$$

TABLE 6.1: A table of the joint probability distribution of S (vertical axis; scale $2, \dots, 12$) and D (horizontal axis; scale $-5, \dots, 5$) from example 4

Worked example 6.6 Sums and differences of dice, IV

Using the terminology of example 3, what is $P(S|D = 0)$? what is $P(D|S = 11)$?

Solution: You could work it out either of these from the table, or by first principles. If $D = 0$, S can have values 2, 4, 6, 8, 10, 12, and each value has conditional probability $1/6$. If $S = 11$, D can have values 1, or -1 , and each value has conditional probability $1/2$.

6.1.2 Just a Little Continuous Probability

Our random variables take values from a discrete set of numbers D . This makes the underlying machinery somewhat simpler to describe, and is often, but not always, enough for model building. Some phenomena are more naturally modelled as being continuous — for example, human height; human weight; the mass of a distant star; and so on. Giving a complete formal description of probability on a continuous space is surprisingly tricky, and would involve us in issues that do not arise much in practice.

These issues are caused by two interrelated facts: real numbers have infinite precision; and you can't count real numbers. A continuous random variable is still a random variable, and comes with all the stuff that a random variable comes with. We will not speculate on what the underlying sample space is, nor on the underlying events. This can all be sorted out, but requires moderately heavy lifting that isn't particularly illuminating for us. The most interesting thing for us is specifying the probability distribution. Rather than talk about the probability that a real number takes a particular value (which we can't really do satisfactorily most of the time), we will instead talk about the probability that it lies in some interval. So we can specify a probability distribution for a continuous random variable by giving a set

of (very small) intervals, and for each interval providing the probability that the random variable lies in this interval.

The easiest way to do this is to supply a **probability density function**. Let $p(x)$ be a probability density function for a continuous random variable X . We interpret this function by thinking in terms of small intervals. Assume that dx is an infinitesimally small interval. Then

$$p(x)dx = P(\{\text{event that } X \text{ takes a value in the range } [x, x + dx]\}).$$

Important properties of probability density functions follow from this definition.

Useful Facts: 6.1 *Probability density functions*

- Probability density functions are non-negative. This follows from the definition; a negative value at some x would imply a negative probability.
- For $a < b$

$$P(\{\text{event that } X \text{ takes a value in the range } [a, b]\}) = \int_a^b p(x)dx.$$

which we obtain by summing $p(x)dx$ over all the infinitesimal intervals between a and b .

- We must have that

$$\int_{-\infty}^{\infty} p(x)dx = 1.$$

This is because

$$P(\{X \text{ takes a value in the range } [-\infty, \infty]\}) = 1 = \int_{-\infty}^{\infty} p(x)dx$$

- Probability density functions are usually called pdf's.
- It is quite usual to write all pdf's as lower-case p 's. If one specifically wishes to refer to probability, one writes an upper case P , as in the previous points.

One good way to think about pdf's is as the limit of a histogram. Imagine you collect an arbitrarily large dataset of data items, each of which is independent. You build a histogram of that dataset, using arbitrarily narrow boxes. You scale the histogram so that the sum of the box areas is one. The result is a probability density function.

The pdf doesn't represent the probability that a random variable takes a value. Instead, you should think of $p(x)$ as being the limit of a ratio (which is why

it's called a density):

$$\frac{\text{the probability that the random variable will lie in a small interval centered on } x}{\text{the length of the small interval centered on } x}$$

Notice that, while a pdf has to be non-negative, and it has to integrate to 1, it does *not* have to be smaller than one. A ratio like this could be a lot larger than one, as long as it isn't larger than one for too many x (because the integral must be one).

Probability density functions can be moderately strange functions.

Worked example 6.7 *Strange probability density functions*

There is some (small!) voltage over the terminals of a warm resistor caused by noise (electrons moving around in the heat and banging into one another). This is a good example of a continuous random variable, and we can assume there is some probability density function for it, say $p(x)$. We assume that $p(x)$ has the property that

$$\lim_{\epsilon \rightarrow 0} \int_{v-\epsilon}^{v+\epsilon} p(x) dx = 0$$

which is what you'd expect for any function you're likely to have dealt with. Now imagine I define a new random variable by the following procedure: I flip a coin; if it comes up heads, I report 0; if tails, I report the voltage over the resistor. This random variable, u , has a probability 1/2 of taking the value 0, and 1/2 of taking a value from $p(x)$. Write this random variable's probability density function $q(u)$. Compute

$$\lim_{\epsilon \rightarrow 0} \int_{-\epsilon}^{\epsilon} q(u) du$$

Solution: We can do this from the definition. We have

$$P(\{u \in [-\epsilon, \epsilon]\}) = \int_{-\epsilon}^{\epsilon} q(u) du.$$

But u will take the value 0 with probability 1/2, and otherwise takes the value over the resistor. So

$$P(\{u \in [-\epsilon, \epsilon]\}) = 1/2 + \int_{-\epsilon}^{\epsilon} p(x) dx.$$

and

$$\lim_{\epsilon \rightarrow 0} \int_{-\epsilon}^{\epsilon} q(u) du = \lim_{\epsilon \rightarrow 0} P(\{u \in [-\epsilon, \epsilon]\}) = 1/2.$$

This means $q(x)$ has the property that

$$\lim_{\epsilon \rightarrow 0} \int_{-\epsilon}^{\epsilon} q(u) du = 1/2,$$

which means that $q(u)$ is displaying quite unusual behavior at $u = 0$. We will not need to deal with probability density functions that behave like this; but you should be aware of the possibility.

Every probability density function $p(x)$ has the property that $\int_{-\infty}^{\infty} p(x) dx = 1$; this is useful, because when we are trying to determine a probability density

function, we can ignore a constant factor. So if $g(x)$ is a non-negative function that is proportional to the probability density function (often pdf) we are interested in, we can recover the pdf by computing

$$p(x) = \frac{1}{\int_{-\infty}^{\infty} g(x)dx} g(x).$$

This procedure is sometimes known as **normalizing**, and $\int_{-\infty}^{\infty} g(x)dx$ is the **normalizing constant**.

6.2 EXPECTATIONS AND EXPECTED VALUES

Imagine we play the game of example 2 multiple times. Our frequency definition of probability means that in N games, we expect to see about pN heads and $(1-p)N$ tails. In turn, this means that my total income from these N games should be about $(pN)q - ((1-p)N)r$. The N in this expression is inconvenient; instead, we could say that for any single game, my income is

$$pq - (1-p)r.$$

This isn't the actual income from a single game (which would be either q or $-r$, depending on what the coin did). Instead, it's an estimate of what would happen over a large number of games, on a per-game basis. This is an example of an **expected value**.

6.2.1 Expected Values of Discrete Random Variables

Definition: 6.6 *Expected value*

Given a discrete random variable X which takes values in the set \mathcal{D} and which has probability distribution P , we define the expected value

$$\mathbb{E}[X] = \sum_{x \in \mathcal{D}} xP(X = x).$$

This is sometimes written which is $\mathbb{E}_P[X]$, to clarify which distribution one has in mind

Notice that an expected value could take a value that the random variable doesn't take.

Example: 6.5 *Betting on coins*

We agree to play the following game. I flip a fair coin (i.e. $P(H) = P(T) = 1/2$). If the coin comes up heads, you pay me 1; if the coin comes up tails, I pay you 1. The expected value of my income is 0, even though the random variable never takes that value.

Definition: 6.7 *Expectation*

Assume we have a function f that maps a discrete random variable X into a set of numbers \mathcal{D}_f . Then $f(x)$ is a discrete random variable, too, which we write F . The expected value of this random variable is written

$$\mathbb{E}[f] = \sum_{u \in \mathcal{D}_f} uP(F = u) = \sum_{x \in \mathcal{D}} f(x)P(X = x)$$

which is sometimes referred to as “the expectation of f ”. The process of computing an expected value is sometimes referred to as “taking expectations”.

Expectations are linear, so that $\mathbb{E}[0] = 0$ and $\mathbb{E}[A + B] = \mathbb{E}[A] + \mathbb{E}[B]$. The expectation of a constant is that constant (or, in notation, $\mathbb{E}[k] = k$), because probabilities sum to 1. Because probabilities are non-negative, the expectation of a non-negative random variable must be non-negative.

6.2.2 Expected Values of Continuous Random Variables

We can compute expectations for continuous random variables, too, though summing over all values now turns into an integral. This should be expected. Imagine you choose a set of closely spaced values for x which are x_i , and then think about x as a discrete random variable. The values are separated by steps of width Δx . Then the expected value of this discrete random variable is

$$\mathbb{E}[X] = \sum_i x_i P(X \in \text{interval centered on } x_i) = \sum_i x_i p(x_i) \Delta x$$

and, as the values get closer together and Δx gets smaller, the sum limits to an integral.

Definition: 6.8 *Expected value*

Given a continuous random variable X which takes values in the set \mathcal{D} and which has probability distribution P , we define the expected value

$$\mathbb{E}[X] = \int_{x \in \mathcal{D}} xp(x)dx.$$

This is sometimes written $\mathbb{E}_p[X]$, to clarify which distribution one has in mind

The expected value of a continuous random variable could be a value that the random variable doesn't take, too. Notice one attractive feature of the $\mathbb{E}[X]$ notation; we don't need to make any commitment to whether X is a discrete random variable (where we would write a sum) or a continuous random variable (where we would write an integral). The reasoning by which we turned a sum into an integral works for functions of continuous random variables, too.

Definition: 6.9 *Expectation*

Assume we have a function f that maps a continuous random variable X into a set of numbers \mathcal{D}_f . Then $f(x)$ is a continuous random variable, too, which we write F . The expected value of this random variable is

$$\mathbb{E}[f] = \int_{x \in \mathcal{D}} f(x)p(x)dx$$

which is sometimes referred to as “the expectation of f ”. The process of computing an expected value is sometimes referred to as “taking expectations”.

Again, for continuous random variables, expectations are linear, so that $\mathbb{E}[0] = 0$ and $\mathbb{E}[A + B] = \mathbb{E}[A] + \mathbb{E}[B]$. The expectation of a constant is that constant (or, in notation, $\mathbb{E}[k] = k$), because probabilities sum to 1. Because probabilities are non-negative, the expectation of a non-negative random variable must be non-negative.

6.2.3 Mean, Variance and Covariance

There are three very important expectations with special names.

Definition: 6.10 *The mean or expected value*

The mean or expected value of a random variable X is

$$\mathbb{E}[X]$$

Definition: 6.11 *The variance*

The variance of a random variable X is

$$\text{var}[X] = \mathbb{E}[(X - \mathbb{E}[X])^2]$$

Notice that

$$\begin{aligned} \mathbb{E}[(X - \mathbb{E}[X])^2] &= \mathbb{E}[X^2 - 2X\mathbb{E}[X] + \mathbb{E}[X]^2] \\ &= \mathbb{E}[X^2] - 2\mathbb{E}[X]\mathbb{E}[X] + \mathbb{E}[X]^2 \\ &= \mathbb{E}[X^2] - (\mathbb{E}[X])^2 \end{aligned}$$

Definition: 6.12 *The covariance*

The covariance of two random variables X and Y is

$$\text{cov}(X, Y) = \mathbb{E}[(X - \mathbb{E}[X])(Y - \mathbb{E}[Y])]$$

Notice that

$$\begin{aligned} \mathbb{E}[(X - \mathbb{E}[X])(Y - \mathbb{E}[Y])] &= \mathbb{E}[(XY - Y\mathbb{E}[X] - X\mathbb{E}[Y] + \mathbb{E}[X]\mathbb{E}[Y])] \\ &= \mathbb{E}[XY] - 2\mathbb{E}[Y]\mathbb{E}[X] + \mathbb{E}[X]\mathbb{E}[Y] \\ &= \mathbb{E}[XY] - \mathbb{E}[X]\mathbb{E}[Y]. \end{aligned}$$

We also have $\text{var}[X] = \text{cov}(X, X)$.

Now assume that we have a probability distribution $P(X)$ defined on some discrete set of numbers. There is some random variable that produced this probability distribution. This means that we could talk about the mean of a probability

distribution P (rather than the mean of a random variable whose probability distribution is $P(X)$). It is quite usual to talk about the mean of a probability distribution. Furthermore, we could talk about the variance of a probability distribution P (rather than the variance of a random variable whose probability distribution is $P(X)$).

Worked example 6.8 *Variance*

Can a random variable have $\mathbb{E}[X] > \sqrt{\mathbb{E}[X^2]}$?

Solution: No, because that would mean that $\mathbb{E}[(X - \mathbb{E}[X])^2] < 0$. But this is the expected value of a non-negative quantity; it must be non-negative.

Worked example 6.9 *More variance*

We just saw that a random variable can't have $\mathbb{E}[X] > \sqrt{\mathbb{E}[X^2]}$. But I can easily have a random variable with large mean and small variance - isn't this a contradiction?

Solution: No, you're confused. Your question means you think that the variance of X is given by $\mathbb{E}[X^2]$; but actually $\text{var}[X] = \mathbb{E}[X^2] - \mathbb{E}[X]^2$

Worked example 6.10 *Mean of a coin flip*

We flip a biased coin, with $P(H) = p$. The random variable X has value 1 if the coin comes up heads, 0 otherwise. What is the mean of X ? (i.e. $\mathbb{E}[X]$).

Solution: $\mathbb{E}[X] = \sum_{x \in D} xP(X = x) = 1p + 0(1 - p) = p$

Useful Facts: 6.2 *Expectations*

1. $\mathbb{E}[0] = 0$
2. $\mathbb{E}[X + Y] = \mathbb{E}[X] + \mathbb{E}[Y]$
3. $\mathbb{E}[kX] = k\mathbb{E}[X]$
4. $\mathbb{E}[1] = 1$
5. if X and Y are independent, then $\mathbb{E}[XY] = \mathbb{E}[X]\mathbb{E}[Y]$.
6. if X and Y are independent, then $\text{cov}(X, Y) = 0$.

All but 5 and 6 are obvious from the definition. If 5 is true, then 6 is obviously true. I prove 5.

Proposition: *If X and Y are independent random variables, then $\mathbb{E}[XY] = \mathbb{E}[X]\mathbb{E}[Y]$.*

Proof: Recall that $\mathbb{E}[X] = \sum_{x \in D} xP(X = x)$, so that

$$\begin{aligned}
 \mathbb{E}[XY] &= \sum_{(x,y) \in D_x \times D_y} xyP(X = x, Y = y) \\
 &= \sum_{x \in D_x} \sum_{y \in D_y} (xyP(X = x, Y = y)) \\
 &= \sum_{x \in D_x} \sum_{y \in D_y} (xyP(X = x)P(Y = y)) \\
 &\quad \text{because } X \text{ and } Y \text{ are independent} \\
 &= \sum_{x \in D_x} \sum_{y \in D_y} (xP(X = x))(yP(Y = y)) \\
 &= \left(\sum_{x \in D_x} xP(X = x) \right) \left(\sum_{y \in D_y} yP(Y = y) \right) \\
 &= (\mathbb{E}[X])(\mathbb{E}[Y]).
 \end{aligned}$$

This is certainly not true when X and Y are not independent (try $Y = -X$).

Useful Facts: 6.3 *Variance*

It is quite usual to write $\text{var}[X]$ for the variance of the random variable X .

1. $\text{var}[0] = 0$
2. $\text{var}[1] = 0$
3. $\text{var}[X] \geq 0$
4. $\text{var}[kX] = k^2\text{var}[X]$
5. if X and Y are independent, then $\text{var}[X + Y] = \text{var}[X] + \text{var}[Y]$

1, 2, 3 are obvious. You will prove 4 and 5 in the exercises.

Worked example 6.11 *Variance of a coin flip*

We flip a biased coin, with $P(H) = p$. The random variable X has value 1 if the coin comes up heads, 0 otherwise. What is the variance of X ? (i.e. $\text{var}[X]$).

Solution: $\text{var}[X] = \mathbb{E}[(X - \mathbb{E}[X])^2] = \mathbb{E}[X^2] - \mathbb{E}[X]^2 = (1p - 0(1-p)) - p^2 = p(1-p)$

The variance of a random variable is often inconvenient, because its units are the square of the units of the random variable. Instead, we could use the **standard deviation**.

Definition: 6.13 *Standard deviation*

The **standard deviation** of a random variable X is defined as

$$\text{std}(X) = \sqrt{\text{var}[X]}$$

You do need to be careful with standard deviations. If X and Y are independent random variables, then $\text{var}[X + Y] = \text{var}[X] + \text{var}[Y]$, but $\text{std}(X + Y) = \sqrt{\text{std}(X)^2 + \text{std}(Y)^2}$. One way to avoid getting mixed up is to remember that variances add, and derive expressions for standard deviations from that.

6.2.4 Expectations and Statistics

You should have noticed we now have two notions each for mean, variance, covariance, and standard deviation. One, which we expounded in sections 1, describes

datasets. We will call these **descriptive statistics**. The other, described above, is a property of probability distributions. We will call these **expectations**. In each case, the reason we have one name for two notions is that the notions are not really all that different.

Imagine we have a dataset $\{\mathbf{x}\}$ of N items, where the i 'th item is \mathbf{x}_i . We can build a probability distribution out of this dataset, by placing a probability on each data item. We will give each data item the same probability (which must be $1/N$, so all probabilities add to 1). Write $\mathbb{E}[\mathbf{x}]$ for the mean of this distribution. We have

$$\mathbb{E}[\mathbf{x}] = \sum_i \mathbf{x}_i p(\mathbf{x}_i) = \frac{1}{N} \sum_i \mathbf{x}_i = \text{mean}(\{\mathbf{x}\}).$$

The variances, standard deviations and covariance have the same property: For this particular distribution (sometimes called the **empirical distribution**), the expectations have the same value as the descriptive statistics (exercises).

In section 1, we will see a form of converse to this fact. Imagine we have a dataset that consists of independent, identically distributed samples from a probability distribution. That is, we know that each data item was obtained independently from the distribution. For example, we might have a count of heads in each of a number of coin flip experiments. Then the descriptive statistics will turn out to be accurate estimates of the expectations.

6.2.5 Indicator Functions

It is sometimes convenient when working with random variables to use **indicator functions**. This is a function that is one when some condition is true, and zero otherwise. The reason they are useful is that their expected values have interesting properties.

Definition: 6.14 *Indicator functions*

An indicator function for an event is a function that takes the value zero for values of X where the event does not occur, and one where the event occurs. For the event \mathcal{E} , we write

$$\mathbb{I}_{[\mathcal{E}]}(X)$$

for the relevant indicator function.

For example,

$$\mathbb{I}_{\{|X| \leq a\}}(X) = \begin{cases} 1 & \text{if } -a < X < a \\ 0 & \text{otherwise} \end{cases}$$

Indicator functions have one useful property.

$$\mathbb{E}[\mathbb{I}_{[\mathcal{E}]}] = P(\mathcal{E})$$

which you can establish by checking the definition of expectations.

6.2.6 Two Inequalities

Mean and variance tell us quite a lot about a random variable, as two important inequalities show.

Definition: 6.15 *Markov's inequality*

Markov's inequality is

$$P(\{\|X\| \geq a\}) \leq \frac{\mathbb{E}[\|X\|]}{a}.$$

You should read this as indicating that a random variable is most unlikely to have an absolute value a lot larger than the mean of its absolute value. This should seem fairly intuitive from the definition of expectation. Recall that

$$\mathbb{E}[X] = \sum_{x \in D} xP(\{X = x\})$$

Assume that D contains only non-negative numbers (that absolute value). Then the only way to have a small value of $\mathbb{E}[X]$ is to be sure that, when x is large, $P(\{X = x\})$ is small. The proof is a rather more formal version of this observation, below.

Definition: 6.16 *Chebyshev's inequality*

Chebyshev's inequality is

$$P(\{|X - \mathbb{E}[X]| \geq a\}) \leq \frac{\text{var}[X]}{a^2}.$$

It is common to see this in another form, obtained by writing σ for the standard deviation of X , substituting $k\sigma$ for a , and rearranging

$$P(\{|X - \mathbb{E}[X]| \geq k\sigma\}) \leq \frac{1}{k^2}$$

This means that the probability of a random variable taking a particular value must fall off rather fast as that value moves away from the mean, in units scaled to the variance. This probably doesn't seem intuitive from the definition of expectation. But think about it this way: values of a random variable that are many standard deviations above the mean must have low probability, otherwise the standard deviation would be bigger. The proof, again, is a rather more formal version of this observation, and appears below.

Proposition: *Markov's inequality*

$$P(\{\|X\| \geq a\}) \leq \frac{\mathbb{E}[\|X\|]}{a}.$$

Proof: (from Wikipedia). Notice that, for $a > 0$,

$$a\mathbb{I}_{\{|X| \geq a\}}(X) \leq |X|$$

(because if $|X| \geq a$, the LHS is a ; otherwise it is zero). Now we have

$$\mathbb{E}[a\mathbb{I}_{\{|X| \geq a\}}] \leq \mathbb{E}[|X|]$$

but, because expectations are linear, we have

$$\mathbb{E}[a\mathbb{I}_{\{|X| \geq a\}}] = a\mathbb{E}[\mathbb{I}_{\{|X| \geq a\}}] = aP(\{|X| \geq a\})$$

and so we have

$$aP(\{|X| \geq a\}) \leq \mathbb{E}[|X|]$$

and we get the inequality by division, which we can do because $a > 0$.

Proposition: *Chebyshev's inequality*

$$P(\{|X - \mathbb{E}[X]| \geq a\}) \leq \frac{\text{var}[X]}{a^2}.$$

Proof: Write U for the random variable $(X - \mathbb{E}[X])^2$. Markov's inequality gives us

$$P(\{|U| \geq w\}) \leq \frac{\mathbb{E}[|U|]}{w}$$

Now notice that, if $a^2 = w$,

$$P(\{|U| \geq w\}) = P(\{|X - \mathbb{E}[X]| \geq a\})$$

so we have

$$P(\{|U| \geq w\}) = P(\{|X - \mathbb{E}[X]| \geq a\}) \leq \frac{\mathbb{E}[|U|]}{w} = \frac{\text{var}[X]}{a^2}$$

6.2.7 IID Samples and the Weak Law of Large Numbers

Imagine a random variable X , obtained by flipping a fair coin and reporting 1 for an H and -1 for a T . We can talk about the probability distribution $P(X)$ of this random variable; we can talk about the expected value that the random variable takes; but the random variable itself doesn't have a value. However, if we actually flip a coin, we get either a 1 or a -1 . This number is often called a **sample** of the random variable (or of its probability distribution). Similarly, if we flipped a coin many times, we'd have a set of numbers (or samples). These numbers would be independent. Their histogram would look like $P(X)$. Collections of data items like this are important enough to have their own name.

Assume we have a set of data items x_i such that (a) they are independent; (b) each is produced from the same process; and (c) the histogram of a very large set of data items looks increasingly like the probability distribution $P(X)$ as the number of data items increases. Then we refer to these data items as **independent identically distributed samples** of $P(X)$; for short, **iid samples** or even just **samples**. For all of the cases we will deal with, it will be obvious how to get IID samples. However, it's worth knowing that obtaining IID samples from arbitrary probability distributions is very difficult.

Now assume we have a set of N IID samples x_i of a probability distribution $P(X)$. Write

$$X_N = \frac{\sum_{i=1}^N x_i}{N}.$$

Now X_N is a random variable (the x_i are IID samples, and for a different set of samples you will get a different, random, X_N). Notice that $P(X = x_1, X = x_2, \dots, X = x_n) = P(X = x_1)P(X = x_2) \dots P(X = x_n)$, because the samples are independent and each is a sample of $P(X)$. Now

$$\mathbb{E}[X_N] = \mathbb{E}[X]$$

because

$$\mathbb{E}[X_N] = \left(\frac{1}{N}\right) \sum_{i=1}^N \mathbb{E}[X].$$

This means that

$$\frac{\sum_{i=1}^N x_i}{N}$$

should be give an accurate estimate of $\mathbb{E}[X]$. In fact, as N gets large, the estimate becomes more accurate.

Definition: 6.17 *The Weak Law of Large Numbers*

The **weak law of large numbers** states that, if $P(X)$ has finite variance, then for any positive number ϵ

$$\lim_{N \rightarrow \infty} P(\{\|X_N - \mathbb{E}[X]\| \geq \epsilon\}) = 0.$$

Equivalently, we have

$$\lim_{N \rightarrow \infty} P(\{\|X_N - \mathbb{E}[X]\| < \epsilon\}) = 1.$$

Each form means that, for a large enough set of IID samples, the average of the samples (i.e. X_N) will, with high probability, be very close to the expectation $\mathbb{E}[X]$.

Proposition: *Weak law of large numbers*

$$\lim_{N \rightarrow \infty} P(\{\|X_N - \mathbb{E}[X]\| \geq \epsilon\}) = 0.$$

Proof: Assume that $P(X)$ has finite variance; that is, $\text{var}(\{X\}) = \sigma^2$. Choose $\epsilon > 0$. Now we have that

$$\begin{aligned} \text{var}(\{X_N\}) &= \text{var}\left(\left\{\frac{\sum_{i=1}^N Nx_i}{N}\right\}\right) \\ &= \left(\frac{1}{N^2}\right)\text{var}\left(\left\{\sum_{i=1}^N Nx_i\right\}\right) \\ &= \left(\frac{1}{N^2}\right)(N\sigma^2) \\ &\quad \text{because the } x_i \text{ are independent} \\ &= \frac{\sigma^2}{N} \end{aligned}$$

and that

$$\mathbb{E}[X_N] = \mathbb{E}[X].$$

Now Chebyshev's inequality gives

$$P(\{\|X_N - \mathbb{E}[X]\| \geq \epsilon\}) \leq \frac{\sigma^2}{N\epsilon^2}$$

so

$$\lim_{N \rightarrow \infty} P(\{\|X_N - \mathbb{E}[X]\| \geq \epsilon\}) = \lim_{N \rightarrow \infty} \frac{\sigma^2}{N\epsilon^2} = 0.$$

Notice that

$$1 - P(\{\|X_N - \mathbb{E}[X]\| \geq \epsilon\}) = P(\{\|X_N - \mathbb{E}[X]\| < \epsilon\}) \geq 1 - \frac{\sigma^2}{N\epsilon^2}.$$

so that

$$\lim_{N \rightarrow \infty} P(\{\|X_N - \mathbb{E}[X]\| < \epsilon\}) = \lim_{N \rightarrow \infty} \left(1 - \frac{\sigma^2}{N\epsilon^2}\right) = 1.$$

6.3 USING EXPECTATIONS

The weak law of large numbers gives us a very valuable way of thinking about expectations. Assume we have a random variable X . Then the weak law says that, if you observe this random variable over a large number of trials, the mean value

you observe should be very close to $\mathbb{E}[X]$. Notice that this extends to functions of random variables (because they are random variables, too). For example, I observe values x_i of a random variable X over a large number N of trials, and compute

$$\frac{1}{N} \sum_{i=1}^N f(x_i).$$

The weak law says that the value I get should be very close to $\mathbb{E}[f]$. You can show this by defining a new random variable $F = f(X)$. This has a probability distribution $P(F)$, which might be difficult to know — but we don't need to. $\mathbb{E}[f]$, the expected value of the function f under the distribution $P(X)$. This is the same as $\mathbb{E}[F]$, and the weak law applies.

Remember: the average over repeated trials of a random variable is very close to the expectation. You can use this information to make many kinds of decision in uncertain circumstances.

6.3.1 Should you accept a bet?

We can't answer this as a moral question, but we can as a practical question, using expectations. Generally, a bet involves an agreement that amounts of money will change hands, depending on the outcome of an experiment. Mostly, you are interested in how much you get from the bet, so it is natural to give sums of money you receive a positive sign, and sums of money you pay out a negative sign. Under this convention, the practical answer is easy: accept a bet enthusiastically if its expected value is positive, otherwise decline it. It is interesting to notice how poorly this advice describes actual human behavior.

Worked example 6.12 *Red or Black?*

A roulette wheel has 36 numbers, 18 of which are red and 18 of which are black. Different wheels also have one, two, or even three zeros, which are colorless. A ball is thrown at the wheel when it is spinning, and it falls into a hole corresponding to one of the numbers (when the number is said to “come up”). The wheel is set up so that there is the same probability of each number coming up. You can bet on (among other things) whether a red number or a black number comes up. If you bet 1 on red, and a red number comes up, you keep your stake and get 1, otherwise you get -1 (i.e. the house keeps your bet).

- On a wheel with one zero, what is the expected value of a 1 bet on red?
- On a wheel with two zeros, what is the expected value of a 1 bet on red?
- On a wheel with three zeros, what is the expected value of a 1 bet on red?

Solution: Write p_r for the probability a red number comes up. The expected value is $1 \times p_r + (-1)(1 - p_r)$ which is $2p_r - 1$.

- In this case, $p_r = (\text{number of red numbers})/(\text{total number of numbers}) = 18/37$. So the expected value is $-1/37$ (you lose about 3 cents each time you bet).
- In this case, $p_r = 18/38$. So the expected value is $-2/38 = -1/19$ (you lose slightly more than five cents each time you bet).
- In this case, $p_r = 18/39$. So the expected value is $-3/39 = -1/13$ (you lose slightly less than 8 cents each time you bet).

Notice that in the roulette game, the money you lose will go to the house. So the expected value to the house is just the negative of the expected value to you. This is positive, which is a partial explanation of why there are lots of roulette wheels, and usually free food nearby. Not all bets are like this, though.

Worked example 6.13 *Coin game*

In this game, P1 flips a fair coin and P2 calls “H” or “T”. If P2 calls right, then P1 throws the coin into the river; otherwise, P1 keeps the coin. What is the expected value of this game to P2? and to P1?

Solution: To P2, which we do first, because it’s easiest: P2 gets 0 if P2 calls right, and 0 if P2 calls wrong; these are the only cases, so the expected value is 0. To P1: P1 gets -1 if P2 calls right, and 0 if P1 calls wrong. The coin is fair, so the probability P2 calls right is $1/2$. The expected value is $-1/2$. While I can’t explain why people would play such a game, I’ve actually seen this done.

We call a bet fair when its expected value is zero. Taking a bet with a negative expected value is unwise, because, on average, you will lose money. Worse, the more times you play, the more you lose. Taking a bet with a positive expected value is likely to be profitable. However, you do need to be careful you computed the expected value right.

Worked example 6.14 *Birthdays in succession*

P1 and P2 agree to the following bet. P1 gives P2 a stake of 1. If three people, stopped at random on the street, have birthdays in succession (i.e. Mon-Tue-Wed, and so on), then P2 gives P1 100. Otherwise, P1 loses the stake. What is the expected value of this bet to P1?

Solution: Write p for the probability of winning. Then the expected value is $p \times 100 - (1 - p) \times 1$. We computed p in example 1 (it was $1/49$). So the bet is worth $(52/49)$, or slightly more than a dollar, to P1. P1 should be happy to agree to this as often as possible.

The reason P2 agrees to bets like that of example 14 is most likely that P2 can’t compute the probability exactly. P2 thinks the event is quite unlikely, so the expected value is negative; but it isn’t as unlikely as P2 thought it was, and this is how P1 makes a profit. This is one of the reasons you should be careful accepting a bet from a stranger: they might be able to compute better than you.

6.3.2 Odds, Expectations and Bookmaking — a Cultural Diversion

Gamblers sometimes use a terminology that is a bit different from ours. In particular, the term **odds** is important. The term comes from the following idea: P1 pays a bookmaker b (the stake) to make a bet; if the bet is successful, P1 receives a , and if not, loses the original stake.

Assume the bet is fair, so that the expected value is zero. Write p for the probability of winning. The net income to P1 is $ap - b(1 - p)$. If this is zero, then $p = b/(a + b)$. So you can interpret odds in terms of probability, *if* you assume the

bet is fair.

A bookmaker sets odds at which to accept bets from gamblers. The bookmaker does not wish to lose money at this business, and so must set odds which are potentially profitable. Doing so is not simple (bookmakers can, and occasionally do, lose catastrophically, and go out of business). In the simplest case, assume that the bookmaker knows the probability p that a particular bet will win. Then the bookmaker could set odds of $(1 - p)/p : 1$. In this case, the expected value of the bet is zero; this is fair, but not attractive business, so the bookmaker will set odds assuming that the probability is a bit higher than it really is. There are other bookmakers out there, so there is some reason for the bookmaker to try to set odds that are close to fair.

In some cases, you can tell when you are dealing with a bookmaker who is likely to go out of business soon. For example, imagine there are two horses running in a race, both at 10 : 1 odds — whatever happens, you could win by betting 1 on each. There is a more general version of this phenomenon. Assume the bet is placed on a horse race, and that bets pay off only for the winning horse. Assume also that exactly one horse will win (i.e. the race is never scratched, there aren't any ties, etc.), and write the probability that the i 'th horse will win as p_i . Then $\sum_{i \in \text{horses}} p_i$ must be 1. Now if the bookmaker's odds yield a set of probabilities that is less than 1, their business should fail, because there is at least one horse on which they are paying out too much. Bookmakers deal with this possibility by writing odds so that $\sum_{i \in \text{horses}} p_i$ is larger than one.

But this is not the only problem a bookmaker must deal with. The bookmaker doesn't actually know the probability that a particular horse will win, and must account for errors in this estimate. One way to do so is to collect as much information as possible (talk to grooms, jockeys, etc.). Another is to look at the pattern of bets that have been placed already. If the bookmaker and the gamblers agree on the probability that each horse will win, then there should be no expected advantage to choosing one horse over another — each should pay out slightly less than zero to the gambler (otherwise the bookmaker doesn't eat). But if the bookmaker has underestimated the probability that a particular horse will win, a gambler may get a positive expected payout by betting on that horse. This means that if one particular horse attracts a lot of money from bettors, it is wise for the bookmaker to offer less generous odds on that horse. There are two reasons: first, the bettors might know something the bookmaker doesn't, and they're signalling it; second, if the bets on this horse are very large and it wins, the bookmaker may not have enough capital left to pay out or to stay in business. All this means that real bookmaking is a complex, skilled business.

6.3.3 Ending a Game Early

Imagine two people are playing a game for a stake, but must stop early — who should get what percentage of the stake? One way to do this is to give each player what they put in at the start, but this is (mildly) unfair if one has an advantage over the other. The alternative is to give each player the expected value of the game at that state for that player. Sometimes one can compute that expectation quite easily.

Worked example 6.15 *Ending a game early*

(from Durrett), two players each pay 25 to play the following game. They toss a fair coin. If it comes up heads, player H wins that toss; if tails, player T wins. The first player to reach 10 wins takes the stake of 50. But one player is called away when the state is 8-7 (H-T) — how should the stake be divided?

Solution: In this state, each player can either win — and so get 50 — or lose — and so get 0. The expectation for H is $50P(\{\text{H wins from 8-7}\}) + 0P(\{\text{T wins from 8-7}\})$, so we need to compute $P(\{\text{H wins from 8-7}\})$. Similarly, the expectation for T is $50P(\{\text{T wins from 8-7}\}) + 0P(\{\text{H wins from 8-7}\})$, so we need to compute $P(\{\text{T wins from 8-7}\})$; but $P(\{\text{T wins from 8-7}\}) = 1 - P(\{\text{H wins from 8-7}\})$. Now it is slightly easier to compute $P(\{\text{T wins from 8-7}\})$, because T can only win in two ways: 8-10 or 9-10. These are independent. For T to win 8-10, the next three flips must come up T, so that event has probability $1/8$. For T to win 9-10, the next four flips must have one H in them, but the last flip may not be H (or else H wins); so the next four flips could be H T T T, T H T T, or T T H T. The probability of this is $3/16$. This means the total probability that T wins is $5/16$. So T should get 16.625 and H should get the rest (although they might have to flip for the odd half cent).

6.3.4 Making a Decision with Decision Trees and Expectations

Imagine we have to choose an action. Once we have chosen, a sequence of random events occurs, and we get a reward with some probability. Which action should we choose? A good answer is to choose the action with the best expected outcome. In fact, choosing any other action is unwise, because if we encounter this situation repeatedly and make a choice that is even only slightly worse than the best, we could lose heavily. This is a very common recipe, and it can be applied to many situations. Usually, but not always, the reward is in money, and we will compute with money rewards for the first few examples.

For such problems, it can be useful to draw a **decision tree**. A decision tree is a drawing of possible outcomes of decisions, which makes costs, benefits and random elements explicit. Each node of the tree represents a test of an attribute (which could be either a decision, or a random variable), and each edge represents a possible outcome of a test. The final outcomes are leaves. Usually, decision nodes are drawn as squares, chance elements as circles, and leaves as triangles.

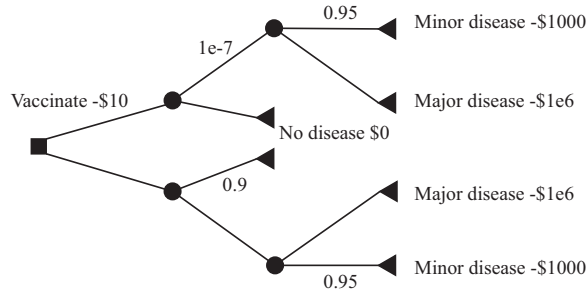


FIGURE 6.1: A decision tree for the vaccination problem. The only decision is whether to vaccinate or not (the box at the root of the tree). I have only labelled edges where this is essential, so I did not annotate the “no vaccination” edge with zero cost. Once you decide whether to vaccinate or not, there is a random node (whether you get the disease or not), and, if you get it, another (minor or major).

Worked example 6.16 *Vaccination*

It costs 10 to be vaccinated against a common disease. If you have the vaccination, the probability you will get the disease is $1e - 7$. If you do not, the probability is 0.1. The disease is unpleasant; with probability 0.95, you will experience effects that cost you 1000 (eg several days in bed), but with probability 0.05, you will experience effects that cost you $1e6$. Should you be vaccinated?

Solution: Figure 6.1 shows a decision tree for this problem. I have annotated some edges with the choices represented, and some edges with probabilities; the sum of probabilities over all rightward (downgoing) edges leaving a random node is 1. It is straightforward to compute expectations. The expected cost of the disease is $0.95 \times 1000 + 0.05 \times 1e6 = 50,950$. If you are vaccinated, your expected income will be $-(10 + 1e - 7 \times 50,950) = -10.01$ (rounding to the nearest cent). If you are not, your expected income is $-5,095$. You should be vaccinated.

Sometimes there is more than one decision. We can still do simple examples, though drawing a decision tree is now quite important, because it allows us to keep track of cases and avoid missing anything. For example, assume I wish to buy a cupboard. Two nearby towns have used furniture shops (usually called antique shops these days). One is further away than the other. If I go to town A, I will have time to look in two (of three) shops; if I go to town B, I will have time to look in one (of two) shops. I could lay out this sequence of decisions (which town to go to; which shop to visit when I get there) as Figure 6.2.

You should notice that this figure is missing a lot of information. What is the probability that I will find what I’m looking for in the shops? What is the value of

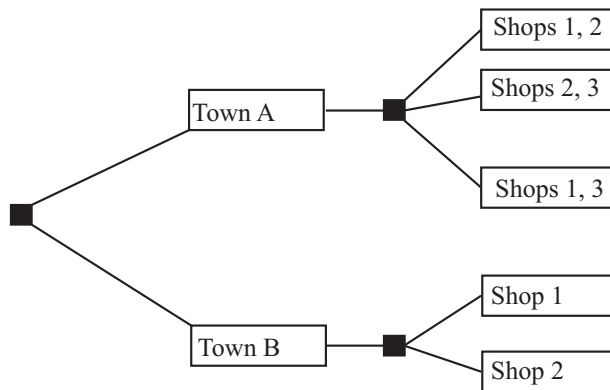


FIGURE 6.2: *The decision tree for the example of visiting furniture shops. Town A is nearer than town B, so if I go there I can choose to visit two of the three shops there; if I go to town B, I can visit only one of the two shops there. To decide what to do, I could fill in the probabilities and values of outcomes, compute the expected value of each pair of decisions, and choose the best. This could be tricky to do (where do I get the probabilities from?) but offers a rational and principled way to make the decision.*

finding it? What is the cost of going to each town? and so on. This information is not always easy to obtain. In fact, I might simply need to give my best subjective guess of these numbers. Furthermore, particularly if there are several decisions, computing the expected value of each possible sequence could get difficult. There are some kinds of model where one can compute expected values easily, but a good viable hypothesis about why people don't make optimal decisions is that optimal decisions are actually too hard to compute.

6.3.5 Utility

Sometimes it is hard to work with money. For example, in the case of a serious disease, choosing treatments often boils down to expected survival times, rather than money.

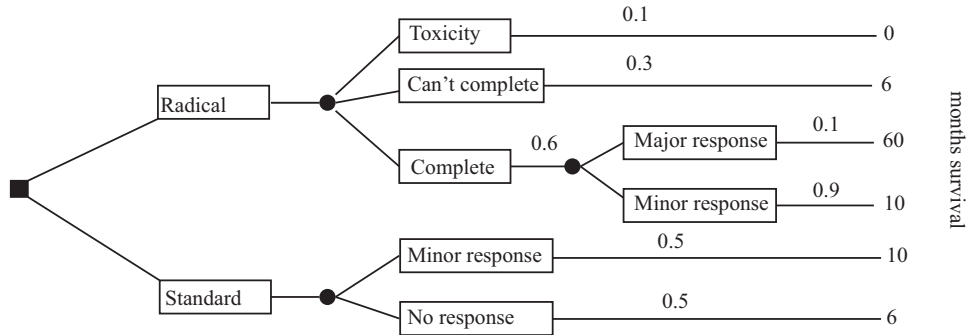


FIGURE 6.3: A decision tree for example 17. Notations vary a bit, and here I have put boxes around the labels for the edges.

Worked example 6.17 *Radical treatment*

(This example largely after Vickers, p97). Imagine you have a nasty disease. There are two kinds of treatment: standard, and radical. Radical treatment might kill you (with probability 0.1); might be so damaging that doctors stop (with probability 0.3); but otherwise you will complete the treatment. If you do complete radical treatment, there could be a major response (probability 0.1) or a minor response. If you follow standard treatment, there could be a major response (probability 0.5) or a minor response, but the outcomes are less good. All this is best summarized in a decision tree (Figure 6.3). What gives the longest expected survival time?

Solution: In this case, expected survival time with radical treatment is $(0.1 \times 0 + 0.3 \times 6 + 0.6 \times (0.1 \times 60 + 0.9 \times 10)) = 10.8$ months; expected survival time without radical treatment is $0.5 \times 10 + 0.5 \times 6 = 8$ months.

Working with money values is not always a good idea. For example, many people play state lotteries. The expected value of a 1 bet on a state lottery is well below 1 — why do people play? It’s easy to assume that all players just can’t do sums, but many players are well aware that the expected value of a bet is below the cost. It seems to be the case that people value money in a way that doesn’t depend linearly on the amount of money. So, for example, people may value a million dollars rather more than a million times the value they place on one dollar. If this is true, we need some other way to keep track of value; this is sometimes called **utility**. It turns out to be quite hard to know how people value things, and there is quite good evidence that (a) human utility is complicated and (b) it is difficult to explain human decision making in terms of expected utility.

Worked example 6.18 *Human utility is not expected payoff*

Here are four games:

- **Game 1:** The player is given 1. A biased coin is flipped, and the money is taken back with probability p ; otherwise, the player keeps it.
- **Game 2:** The player stakes 1, and a fair coin is flipped; if the coin comes up heads, the player gets r and the stake back, but otherwise loses the original stake.
- **Game 3:** The player bets nothing; a biased coin is flipped, and if it comes up heads (probability q), the player gets $1e6$.
- **Game 4:** The player stakes 1000; a fair coin is flipped, and if it comes up heads, the player gets s and the stake back, but otherwise loses the original stake.

In particular, what happens if $r = 3 - 2p$ and $q = (1 - p)/1e6$ and $s = 2 - 2p + 1000$?

Solution: Game 1 has expected value $(1 - p)1$. Game 2 has expected value $(1/2)(r - 1)$. Game 3 has expected value $q1e6$. Game 4 has expected value $(1/2)s - 500$.

In the case given, each game has the same expected value. Nonetheless, people usually have decided preferences for which game they would play. Generally, 4 is unattractive (seems expensive to play); 3 seems like free money, and so a good thing; 2 might be OK but is often seen as uninteresting; and 1 is unattractive. This should suggest to you that people's reasoning about money and utility is not what simple expectations can predict.

6.4 WHAT YOU SHOULD REMEMBER

You should be able to:

- Interpret notation for joint and conditional probability for random variables; in particular, understand notation such as: $P(\{X\})$, $P(\{X = x\})$, $p(x)$, $p(x, y)$, $p(x|y)$
- Interpret a probability density function $p(x)$ as $P(\{X \in [x, x + dx]\})$.
- Interpret the expected value of a discrete random variable.
- Interpret the expected value of a continuous random variable.
- Compute expected values of random variables for straightforward cases.

- Write down expressions for mean, variance and covariance for random variables.
- Write out a decision tree.

You should remember:

- The definition of a random variable.
- The definition of an expected value.
- The definitions of mean, variance and covariance.
- The definition of an indicator function.
- Bayes rule.
- The definition of marginalization.
- The Markov inequality.
- The Chebyshev Inequality.
- The weak law of large numbers.

PROBLEMS

Joint and Conditional Probability for Random Variables

- 6.1.** Define a random variable X by the following procedure. Draw a card from a standard deck of playing cards. If the card is knave, queen, or king, then $X = 11$. If the card is an ace, then $X = 1$; otherwise, X is the number of the card (i.e. two through ten). Now define a second random variable Y by the following procedure. When you evaluate X , you look at the color of the card. If the card is red, then $Y = X - 1$; otherwise, $Y = X + 1$.
- What is $P(\{X \leq 2\})$?
 - What is $P(\{X \geq 10\})$?
 - What is $P(\{X \geq Y\})$?
 - What is the probability distribution of $Y - X$?
 - What is $P(\{Y \geq 12\})$?
- 6.2.** Define a random variable by the following procedure. Flip a fair coin. If it comes up heads, the value is 1. If it comes up tails, roll a die: if the outcome is 2 or 3, the value of the random variable is 2. Otherwise, the value is 3.
- What is the probability distribution of this random variable?
 - What is the cumulative distribution of this random variable?
- 6.3.** Define three random variables, X , Y and Z by the following procedure. Roll a six-sided die and a four-sided die. Now flip a coin. If the coin comes up heads, then X takes the value of the six-sided die and Y takes the value of the four-sided die. Otherwise, X takes the value of the four-sided die and Y takes the value of the six-sided die. Z always takes the value of the sum of the dice.
- What is $P(X)$, the probability distribution of this random variable?
 - What is $P(X, Y)$, the joint probability distribution of these two random variables?
 - Are X and Y independent?

- (d) Are X and Z independent?
- 6.4. Define two random variables X and Y by the following procedure. Flip a fair coin; if it comes up heads, then $X = 1$, otherwise $X = -1$. Now roll a six-sided die, and call the value U . We define $Y = U + X$.
- (a) What is $P(Y|X = 1)$?
- (b) What is $P(X|Y = 0)$?
- (c) What is $P(X|Y = 7)$?
- (d) What is $P(X|Y = 3)$?
- (e) Are X and Y independent?

Expected Values

- 6.5. A simple coin game is as follows: we have a box, which starts empty. P1 flips a fair coin. If it comes up heads, P2 gets the contents of the box, and the game ends. If it comes up tails, P1 puts a dollar in the box and they flip again; this repeats until it comes up heads
- (a) With what probability will P2 win exactly 10 units?
- (b) What is the expected value of the game?
- (c) How much should P2 pay to play, to make the game fair?
- 6.6. A simple card game is as follows. P1 pays a stake of 1 to play. P1 and P2 then each draw a card. If both cards are the same color, P2 keeps the stake and the game ends. If they are different colors, P2 pays P1 the stake and 1 extra (a total of 2).
- (a) What is the expected value of the game to P1?
- (b) P2 modifies the game, as follows. If both cards are court cards (that is, knave, queen, king), then P2 keeps the stake and the game ends; otherwise, the game works as before. Now what is the expected value of the game to P1?
- 6.7. An airline company runs a flight that has six seats. Each passenger who buys a ticket has a probability p of turning up for the flight. These events are independent.
- (a) The airline sells six tickets. What is the expected number of passengers, if $p = 0.9$?
- (b) How many tickets should the airline sell to ensure that the expected number of passengers is greater than six, if $p = 0.7$? **Hint:** The easiest way to do this is to write a quick program that computes the expected value of passengers that turn up for each the number of tickets sold, then search the number of tickets sold.
- 6.8. An airline company runs a flight that has 10 seats. Each passenger who buys a ticket has a probability p of turning up for the flight. The gender of the passengers is not known until they turn up for a flight, and women buy tickets with the same frequency that men do. The pilot is eccentric, and will not fly unless at least two women turn up.
- (a) How many tickets should the airline sell to ensure that the expected number of passengers that turn up is greater than 10?
- (b) The airline sells 10 tickets. What is the expected number of passengers on the aircraft, given that it flies? (i.e. that at least two women turn up). Estimate this value with a simulation.

Mean, Variance and Covariance

- 6.9. Show that $\text{var}[kX] = k^2\text{var}[X]$.

- 6.10.** Show that if X and Y are independent random variables, then $\text{var}[X + Y] = \text{var}[X] + \text{var}[Y]$. You will find it helpful to remember that, for X and Y independent, $\mathbb{E}[XY] = \mathbb{E}[X]\mathbb{E}[Y]$.

Using Inequalities

- 6.11.** The random variable X takes the values $-2, -1, 0, 1, 2$, but has an unknown probability distribution. You know that $\mathbb{E}[\|X\|] = 0.2$. Use Markov's inequality to give a *lower* bound on $P(\{X = 0\})$. *Hint:* Notice that $P(\{X = 0\}) = 1 - P(\{\|X\| = 1\}) - P(\{\|X\| = 2\})$.
- 6.12.** The random variable X takes the values $1, 2, 3, 4, 5$, but has unknown probability distribution. You know that $\mathbb{E}[X] = 2$ and $\text{var}(\{X\}) = 0.01$. Use Chebychev's inequality to give a *lower* bound on $P(\{X = 2\})$.

Using Expectations

- 6.13.** Imagine we have a game with two players, who are playing for a stake. There are no draws, the winner gets the whole stake, and the loser gets nothing. The game must end early. We decide to give each player the expected value of the game for that player, from that state. Show that the expected values add up to the value of the stake (i.e. there won't be too little or too much money in the stake).

General Exercises

CHAPTER 7

Useful Probability Distributions

In most practical problems, we observe some data, then try to infer properties of the process or processes that produced the data. For example, we might ask what new data would look like, or whether two datasets come from the same or different sources. When we do this, we are usually trying to determine some properties of the probability distribution of a random variable. The data represent values that the random variable has taken. In many cases, the probability distribution involved can be modelled using one of a small set of standard probability distributions. In this chapter, I collect facts about the distributions we will use in following chapters.

7.1 DISCRETE DISTRIBUTIONS

7.1.1 The Discrete Uniform Distribution

If every value of a discrete random variable has the same probability, then the probability distribution is the discrete uniform distribution. We have seen this distribution before, numerous times. For example, I define a random variable by the number that shows face-up on the throw of a die. This has a uniform distribution. As another example, write the numbers 1-52 on the face of each card of a standard deck of playing cards. The number on the face of the first card drawn from a well-shuffled deck is a random variable with a uniform distribution.

One can construct expressions for the mean and variance of a discrete uniform distribution, but they're not usually much use (too many terms, not often used). Keep in mind that if two random variables have a uniform distribution, their sum and difference will not (recall example 3).

7.1.2 The Geometric Distribution

We have a biased coin. The probability it will land heads up, $P(\{H\})$ is given by p . We flip this coin until the first head appears. The number of flips required is a discrete random variable which takes integer values greater than or equal to one, which we shall call X . To get n flips, we must have $n - 1$ tails followed by 1 head. This event has probability $(1 - p)^{(n-1)}p$. We can now write out the probability distribution that n flips are required.

Definition: 7.1 *The Geometric Distribution*

We have an experiment with a binary outcome (i.e. heads or tails; 0 or 1; and so on), with $P(H) = p$ and $P(T) = 1 - p$. We repeat this experiment until the first head occurs. The probability distribution for n , the number of repetitions, is the geometric distribution. It has the form

$$P(\{X = n\}) = (1 - p)^{(n-1)}p.$$

for $0 \leq p \leq 1$ and $n \geq 1$; for other n it is zero. p is called the **parameter** of the distribution.

Notice that the geometric distribution is non-negative everywhere. It is straightforward to show that it sums to one, and so is a probability distribution (exercises).

Useful Facts: 7.1 *The geometric distribution*

1. The mean of the geometric distribution is $\frac{1}{p}$.
2. The variance of the geometric distribution is $\frac{1-p}{p^2}$.

The proof of these facts requires some work with series, and is relegated to the exercises.

7.1.3 Bernoulli Random Variables

A Bernoulli random variable models a biased coin with probability p of coming up heads in any one flip.

Definition: 7.2 *Bernoulli Random Variable*

A Bernoulli random variable takes the value 1 with probability p and 0 with probability $1 - p$. This is a model for a coin toss, among other things

Useful Facts: 7.2 *Bernoulli Random Variables*

1. A Bernoulli random variable has mean p .
2. A Bernoulli random variable has variance $p(1 - p)$.

Proofs are easy, and in the exercises.

7.1.4 The Binomial Probability Distribution

Assume we have a biased coin with probability p of coming up heads in any one flip. The binomial probability distribution gives the probability that it comes up heads h times in N flips.

Worked example 32 yields one way of deriving this distribution. In that example, I showed that there are

$$N!/(h!(N - h)!)$$

outcomes of N coin flips that have h heads. These outcomes are disjoint, and each has probability $p^h(1 - p)^{(N-h)}$. As a result, we must have the probability distribution below.

Definition: 7.3 *The Binomial distribution*

In N independent repetitions of an experiment with a binary outcome (ie heads or tails; 0 or 1; and so on) with $P(H) = p$ and $P(T) = 1 - p$, the probability of observing a total of h H 's and $N - h$ T 's is

$$P_b(h; N, p) = \binom{N}{h} p^h (1 - p)^{(N-h)}$$

as long as $0 \leq h \leq N$; in any other case, the probability is zero.

Useful Fact: 7.3 *The binomial distribution*

Write $P_b(i; N, p)$ for the binomial distribution that one observes i H 's in N trials.

$$\sum_{i=0}^N P_b(i; N, p) = (p + (1 - p))^N = (1)^N = 1$$

by pattern matching to the binomial theorem. As a result,

$$\sum_{i=0}^N P_b(i; N, p) = 1$$

The binomial distribution satisfies a recurrence relation. We must have that

$$P_b(h; N, p) = pP_b(h - 1; N - 1, p) + (1 - p)P_b(h; N - 1, p).$$

This is because can get h heads in N flips either by having $h - 1$ heads in $N - 1$ flips, then flipping another, or by having h heads in N flips then flipping a tail. You can verify by induction that the binomial distribution satisfies this recurrence relation.

Useful Facts: 7.4 *The binomial distribution*

1. The mean of $P_b(i; N, p)$ is Np .
2. The variance of $P_b(i; N, p)$ is $Np(1 - p)$

The proofs are informative, and so are not banished to the exercises.

Proof: 7.1 *The binomial distribution*

Notice that the number of heads in N coin tosses is can be obtained by adding the number of heads in each toss. Write Y_i for the Bernoulli random variable representing the i 'th toss. If the coin comes up heads, $Y_i = 1$, otherwise $Y_i = 0$. The Y_i are independent. Now

$$\begin{aligned}\mathbb{E}[X] &= \mathbb{E}\left[\sum_{j=1}^N Y_j\right] \\ &= \sum_{j=1}^N \mathbb{E}[Y_j] \\ &= N\mathbb{E}[Y_1] \quad \text{because the } Y_i \text{ are independent} \\ &= Np.\end{aligned}$$

The variance is easy, too. Each coin toss is independent, so the variance of the sum of coin tosses is the sum of the variances. This gives

$$\begin{aligned}\text{var}[X] &= \text{var}\left[\sum_{j=1}^N Y_j\right] \\ &= N\text{var}[Y_1] \\ &= Np(1-p)\end{aligned}$$

7.1.5 Multinomial probabilities

The binomial distribution describes what happens when a coin is flipped multiple times. But we could toss a die multiple times too. Assume this die has k sides, and we toss it N times. The distribution of outcomes is known as the **multinomial distribution**.

We can guess the form of the multinomial distribution in rather a straightforward way. The die has k sides. We toss the die N times. This gives us a sequence of N numbers. Each toss of the die is independent. Assume that side 1 appears n_1 times, side 2 appears n_2 times, ... side k appears n_k times. Any single sequence with this property will appear with probability $p_1^{n_1} p_2^{n_2} \dots p_k^{n_k}$, because the tosses are independent. However, there are

$$\frac{N!}{n_1! n_2! \dots n_k!}$$

such sequences. Using this reasoning, we arrive at the distribution below

Definition: 7.4 *Multinomial Distribution*

I perform N independent repetitions of an experiment with k possible outcomes. The i 'th such outcome has probability p_i . I see outcome 1 n_1 times, outcome 2 n_2 times, etc. Notice that $n_1 + n_2 + n_3 + \dots + n_k = N$. The probability of observing this set of outcomes is

$$P_m(n_1, \dots, n_k; N, p_1, \dots, p_k) = \frac{N!}{n_1!n_2!\dots n_k!} p_1^{n_1} p_2^{n_2} \dots p_k^{n_k}.$$

Worked example 7.1 *Dice*

I throw five fair dice. What is the probability of getting two 2's and three 3's?

Solution: $\frac{5!}{2!3!} (\frac{1}{6})^2 (\frac{1}{6})^3$

7.1.6 The Poisson Distribution

Assume we are interested in counts that occur in an interval of time (e.g. within a particular hour). Because they are counts, they are non-negative and integer valued. We know these counts have two important properties. First, they occur with some fixed average rate. Second, an observation occurs independent of the interval since the last observation. Then the Poisson distribution is an appropriate model.

There are numerous such cases. For example, the marketing phone calls you receive during the day time are likely to be well modelled by a Poisson distribution. They come at some average rate — perhaps 5 a day as I write, during the last phases of an election year — and the probability of getting one clearly doesn't depend on the time since the last one arrived. Classic examples include the number of Prussian soldiers killed by horse-kicks each year; the number of calls arriving at a call center each minute; the number of insurance claims occurring in a given time interval (outside of a special event like a hurricane, etc.).

Definition: 7.5 *The Poisson Distribution*

A non-negative, integer valued random variable X has a Poisson distribution when its probability distribution takes the form

$$P(\{X = k\}) = \frac{\lambda^k e^{-\lambda}}{k!},$$

where $\lambda > 0$ is a parameter often known as the **intensity** of the distribution.

Notice that the Poisson distribution is a probability distribution, because it is non-negative and because

$$\sum_{i=0}^{\infty} \frac{\lambda^i}{i!} = e^\lambda$$

so that

$$\sum_{k=0}^{\infty} \frac{\lambda^k e^{-\lambda}}{k!} = 1$$

Useful Facts: 7.5 *The Poisson Distribution*

1. The mean of a Poisson distribution with intensity λ is λ .
2. The variance of a Poisson distribution with intensity λ is λ (no, that's not an accidentally repeated line or typo).

The proof of these facts requires some work with series, and is relegated to the exercises.

I described the Poisson distribution as a natural model for counts of randomly distributed points along a time axis. But it doesn't really matter that this is a time axis — it could be a space axis instead. For example, you could take a length of road, divide it into even intervals, then count the number of road-killed animals in each interval. If the location of each animal is independent of the location of any other animal, then you could expect a Poisson model to apply to the count data. Assume that the Poisson model that best describes the data has parameter λ . One property of such models is that if you doubled the length of the intervals, then the resulting dataset would be described by a Poisson model with parameter 2λ ; similarly, if you halved the length of the intervals, the best model would have parameter $\lambda/2$. This corresponds to our intuition about such data; roughly, the number of road-killed animals in two miles of road should be twice the number in one mile of road. This property means that no pieces of the road are “special” — each behaves the same as the other.

We can build a really useful model of spatial randomness by observing this fact and generalizing very slightly. A **Poisson point process** with intensity λ is a set of random points with the property that the number of points in an interval of length s is a Poisson random variable with parameter λs . Notice how this captures our intuition that if points are “very randomly” distributed, there should be twice as many of them in an interval that is twice as long.

This model is easily, and very usefully, extended to points on the plane, on surfaces, and in 3D. In each case, the process is defined on a domain D (which has to meet some very minor conditions that are of no interest to us). The number of points in any subset s of D is a Poisson random variable, with intensity $\lambda m(s)$, where $m(s)$ is the area (resp. volume) of s . These models are useful, because they capture the property that (a) the points are random and (b) the probability you find a point doesn’t depend on where you are. You could reasonably believe models like this apply to, say, dead flies on windcreens; the places where you find acorns at the foot of an oak tree; the distribution of cowpats in a field; the distribution of cherries in a fruitcake; and so on.

7.2 CONTINUOUS DISTRIBUTIONS

7.2.1 The Continuous Uniform Distribution

Some continuous random variables have a natural upper bound and a natural lower bound but otherwise we know nothing about them. For example, imagine we are given a coin of unknown properties by someone who is known to be a skillful maker of unfair coins. The manufacturer makes no representations as to the behavior of the coin. The probability that this coin will come up heads is a random variable, about which we know nothing except that it has a lower bound of zero and an upper bound of one.

If we know nothing about a random variable apart from the fact that it has a lower and an upper bound, then a **uniform distribution** is a natural model. Write l for the lower bound and u for the upper bound. The probability density function for the uniform distribution is

$$p(x) = \begin{cases} 0 & x < l \\ 1/(u-l) & l \leq x \leq u \\ 0 & x > u \end{cases}$$

A continuous random variable whose probability distribution is the uniform distribution is often called a **uniform random variable**.

7.2.2 The Beta Distribution

It’s hard to explain now why the Beta (or β) distribution is useful, but it will come in useful later (section 1). The Beta distribution is a probability distribution for a continuous random variable x in the range $0 \leq x \leq 1$. There are two parameters, $\alpha > 0$ and $\beta > 0$. Recall the definition of the Γ function from section 1.1. We have that

$$P_\beta(x|\alpha, \beta) = \frac{\Gamma(\alpha + \beta)}{\Gamma(\alpha)\Gamma(\beta)} x^{(\alpha-1)}(1-x)^{(\beta-1)}.$$

From this expression, you can see that:

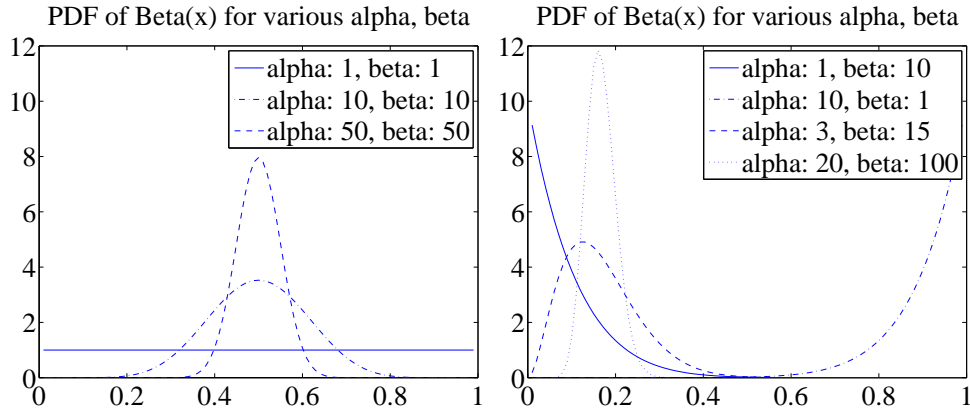


FIGURE 7.1: Probability density functions for the Beta distribution with a variety of different choices of α and β .

- $P_\beta(x|1,1)$ is a uniform distribution on the unit interval.
- $P_\beta(x|\alpha, \beta)$ has a single maximum at $x = (\alpha - 1)/(\alpha + \beta - 2)$ for $\alpha > 1, \beta > 1$ (differentiate and set to zero).
- Generally, as α and β get larger, this peak gets narrower.
- For $\alpha = 1, \beta > 1$ the largest value of $P_\beta(x|\alpha, \beta)$ is at $x = 0$.
- For $\alpha > 1, \beta = 1$ the largest value of $P_\beta(x|\alpha, \beta)$ is at $x = 1$.

Figure 7.1 shows plots of the probability density function of the Beta distribution for a variety of different values of α and β .

Useful Facts: 7.6 *The Beta Distribution*

For a Beta distribution with parameters α, β

1. The mean is $\frac{\alpha}{\alpha + \beta}$.
2. The variance is $\frac{\alpha\beta}{(\alpha + \beta)^2(\alpha + \beta + 1)}$.

7.2.3 The Gamma Distribution

The Gamma (or γ) distribution will also come in useful later on (section 1). The Gamma distribution is a probability distribution for a non-negative continuous random variable $x \geq 0$. There are two parameters, $\alpha > 0$ and $\beta > 0$. The probability density function is

$$P_\gamma(x|\alpha, \beta) = \frac{\beta^\alpha}{\Gamma(\alpha)} x^{(\alpha-1)} e^{-\beta x}.$$

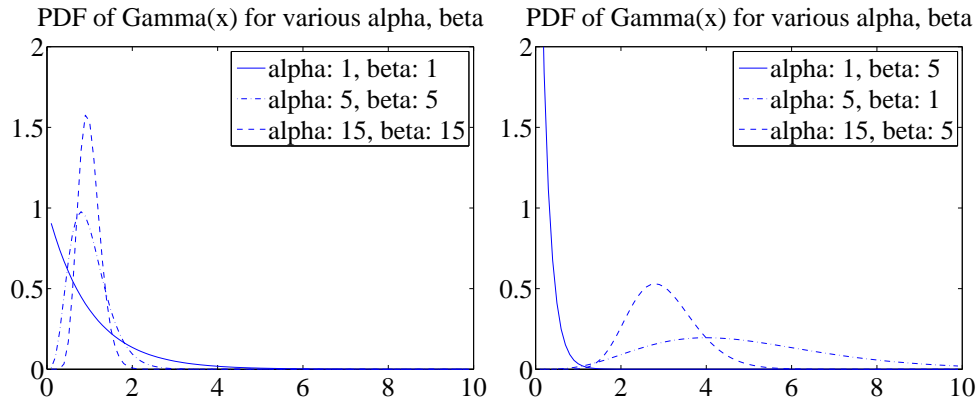


FIGURE 7.2: Probability density functions for the Gamma distribution with a variety of different choices of α and β .

Figure 7.2 shows plots of the probability density function of the Gamma distribution for a variety of different values of α and β .

Useful Facts: 7.7 *The Gamma Distribution*

For a Gamma distribution with parameters α, β

1. The mean is $\frac{\alpha}{\beta}$.
2. The variance is $\frac{\alpha}{\beta^2}$.

7.2.4 The Exponential Distribution

Assume we have an infinite interval of time or space, with points distributed on it. Assume these points form a Poisson point process, as above. For example, we might consider the times at which email arrives; or the times at which phone calls arrive at a large telephone exchange; or the locations of roadkill on a road. The distance (or span of time) between two consecutive points is a random variable X . This random variable takes an **exponential distribution**. There is a single parameter, λ . We have that

$$P_{\text{exp}}(x|\lambda) = \begin{cases} \lambda \exp^{-\lambda x} & \text{for } x \geq 0 \\ 0 & \text{otherwise} \end{cases} .$$

This distribution is often useful in modelling the failure of objects. We assume that failures form a Poisson process in time; then the time to the next failure is exponentially distributed.

Useful Facts: 7.8 *The Exponential Distribution*

For an exponential distribution with parameter λ

1. The mean is $\frac{1}{\lambda}$.
2. The variance is $\frac{1}{\lambda^2}$.

Notice the relationship between this parameter and the parameter of the Poisson distribution. If (say) the phone calls are distributed with Poisson distribution with intensity λ (per hour), then your expected number of calls per hour is λ . The time between calls will be exponentially distributed with parameter λ , and the expected time to the next call is $1/\lambda$ (in hours).

7.3 THE NORMAL DISTRIBUTION

7.3.1 The Standard Normal Distribution

Definition: 7.6 *The Standard Normal Distribution*

The probability density function

$$p(x) = \left(\frac{1}{\sqrt{2\pi}} \right) \exp \left(\frac{-x^2}{2} \right).$$

is known as the **standard normal distribution**

The first step is to plot this probability density function (Figure 7.3). You should notice it is quite familiar from work on histograms, etc. in Chapter 1. It has the shape of the histogram of standard normal data, or at least the shape that the histogram of standard normal data aspires to.

Useful Facts: 7.9 *The standard normal distribution*

1. The mean of the standard normal distribution is 0.
2. The variance of the standard normal distribution is 1.

These results are easily established by looking up (or doing!) the relevant integrals; they are relegated to the exercises.

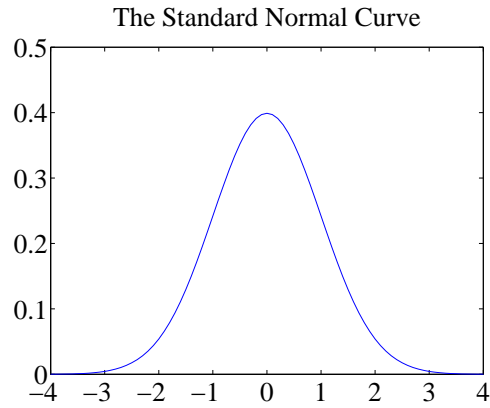


FIGURE 7.3: A plot of the probability density function of the standard normal distribution. Notice how probability is concentrated around zero, and how there is relatively little probability density for numbers with large absolute values.

A continuous random variable is a **standard normal random variable** if its probability density function is a standard normal distribution.

7.3.2 The Normal Distribution

Any probability density function that is a standard normal distribution *in standard coordinates* is a **normal distribution**. Now write μ for the mean of a random variable and σ for its standard deviation; we are saying that, if

$$\frac{x - \mu}{\sigma}$$

has a standard normal distribution, then $p(x)$ is a normal distribution. We can work out the form of the probability density function of a general normal distribution in two steps: first, we notice that for any normal distribution, we must have

$$p(x) \propto \exp \left[-\frac{(x - \mu)^2}{2\sigma^2} \right].$$

But, for this to be a probability density function, we must have $\int_{-\infty}^{\infty} p(x) dx = 1$. This yields the constant of proportionality, and we get

Definition: 7.7 *The Normal Distribution*

The probability density function

$$p(x) = \left(\frac{1}{\sqrt{2\pi}\sigma} \right) \exp \left(-\frac{(x - \mu)^2}{2\sigma^2} \right).$$

is a normal distribution.

Useful Facts: 7.10 *The normal distribution*

The probability density function

$$p(x) = \left(\frac{1}{\sqrt{2\pi}\sigma} \right) \exp \left(\frac{-(x - \mu)^2}{2\sigma^2} \right).$$

has

1. mean μ
2. and variance σ .

These results are easily established by looking up (or doing!) the relevant integrals; they are relegated to the exercises.

A continuous random variable is a **normal random variable** if its probability density function is a **normal distribution**. Notice that it is quite usual to call normal distributions **gaussian distributions**.

7.3.3 Properties of the Normal Distribution

Normal distributions are important, because one often runs into data that is well described by a normal distribution. It turns out that anything that behaves like a binomial distribution with a lot of trials — for example, the number of heads in many coin tosses; as another example, the percentage of times you get the outcome of interest in a simulation in many runs — should produce a normal distribution (Section 7.4). For this reason, pretty much any experiment where you perform a simulation, then count to estimate a probability or an expectation, should give you an answer that has a normal distribution.

It is a remarkable and deep fact, known as the **central limit theorem**, that adding many independent random variables produces a normal distribution pretty much *whatever* the distributions of those random variables. I've not shown this in detail because it's a nuisance to prove. However, if you add together many random variables, each of pretty much any distribution, then the answer has a distribution close to the normal distribution. It turns out that many of the processes we observe add up subsidiary random variables. This means that you will see normal distributions very often in practice.

A normal random variable tends to take values that are quite close to the mean, measured in standard deviation units. We can demonstrate this important fact by computing the probability that a standard normal random variable lies between u and v . We form

$$\int_u^v \frac{1}{\sqrt{2\pi}} \exp \left(-\frac{u^2}{2} \right) du.$$

It turns out that this integral can be evaluated relatively easily using a special

function. The **error function** is defined by

$$\operatorname{erf}(x) = \frac{2}{\sqrt{\pi}} \int_0^x \exp(-t^2) dt$$

so that

$$\frac{1}{2} \operatorname{erf}\left(\frac{x}{\sqrt{2}}\right) = \int_0^x \frac{1}{\sqrt{2\pi}} \exp\left(-\frac{u^2}{2}\right) du.$$

Notice that $\operatorname{erf}(x)$ is an odd function (i.e. $\operatorname{erf}(-x) = -\operatorname{erf}(x)$). From this (and tables for the error function, or Matlab) we get that, for a standard normal random variable

$$\frac{1}{\sqrt{2\pi}} \int_{-1}^1 \exp\left(-\frac{x^2}{2}\right) dx \approx 0.68$$

and

$$\frac{1}{\sqrt{2\pi}} \int_{-2}^2 \exp\left(-\frac{x^2}{2}\right) dx \approx 0.95$$

and

$$\frac{1}{\sqrt{2\pi}} \int_{-3}^3 \exp\left(-\frac{x^2}{2}\right) dx \approx 0.99.$$

These are very strong statements. They measure how often a standard normal random variable has values that are in the range $-1, 1$, $-2, 2$, and $-3, 3$ respectively. But these measurements apply to normal random variables if we recognize that they now measure how often the normal random variable is some number of standard deviations away from the mean. In particular, it is worth remembering that:

Useful Facts: 7.11 *Normal Random Variables*

- About 68% of the time, a normal random variable takes a value within one standard deviation of the mean.
- About 95% of the time, a normal random variable takes a value within one standard deviation of the mean.
- About 99% of the time, a normal random variable takes a value within one standard deviation of the mean.

7.4 APPROXIMATING BINOMIALS WITH LARGE N

The Binomial distribution appears to be a straightforward thing. We assume we flip a coin N times, where N is a very large number. The coin has probability p of coming up heads, and so probability $q = 1 - p$ of coming up tails. The number of heads h follows the binomial distribution, so

$$P(h) = \frac{N!}{h!(N-h)!} p^h q^{(N-h)}$$

The mean of this distribution is Np , the variance is Npq , and the standard deviation is \sqrt{Npq} .

Evaluating this probability distribution for large N is very difficult, because factorials grow fast. We will construct an approximation to the binomial distribution for large N that allows us to evaluate the probability that h lies in some range. This approximation will show that the probability that h is within one standard deviation of the mean is approximately 68%.

This is important, because it shows that our model of probability as frequency is consistent. Consider the probability that the number of heads you see lies within one standard deviation of the mean. The size of that interval is $2\sqrt{Npq}$. As N gets bigger, the size of that interval, *relative to the total number of flips*, gets smaller. If I flip a coin N times, in principle I could see a number of heads h that ranges from 0 to N . However, we will establish that about 68% of the time, h will lie in the interval within one standard deviation of the mean. The size of this interval, *relative to the total number of flips* is

$$2\frac{\sqrt{Npq}}{N} = 2\sqrt{\frac{pq}{N}}.$$

As a result, as $N \rightarrow \infty$,

$$\frac{h}{N} \rightarrow p$$

because h will tend to land in an interval around pN that gets narrower as N gets larger.

The main difficulty with Figure 7.4 (and with the argument above) is that the mean and standard deviation of the binomial distribution tends to infinity as the number of coin flips tends to infinity. This can confuse issues. For example, the plots of Figure 7.4 show narrowing probability distributions — but is this because the scale is compacted, or is there a real effect? It turns out there is a real effect, and a good way to see it is to consider the normalized number of heads.

7.4.1 Large N

Recall that to normalize a dataset, you subtract the mean and divide the result by the standard deviation. We can do the same for a random variable. We now consider

$$x = \frac{h - Np}{\sqrt{Npq}}.$$

The probability distribution of x can be obtained from the probability distribution for h , because $h = Np + x\sqrt{Npq}$, so

$$P(x) = \left(\frac{N!}{(Np + x\sqrt{Npq})!(Nq - x\sqrt{Npq})!} \right) p^{(Np + x\sqrt{Npq})} q^{(Nq - x\sqrt{Npq})}.$$

I have plotted this probability distribution for various values of N in Figure 7.5.

But it is hard to work with this distribution for very large N . The factorials become very difficult to evaluate. Second, it is a discrete distribution on N points, spaced $1/\sqrt{Npq}$ apart. As N becomes very large, the number of points that have

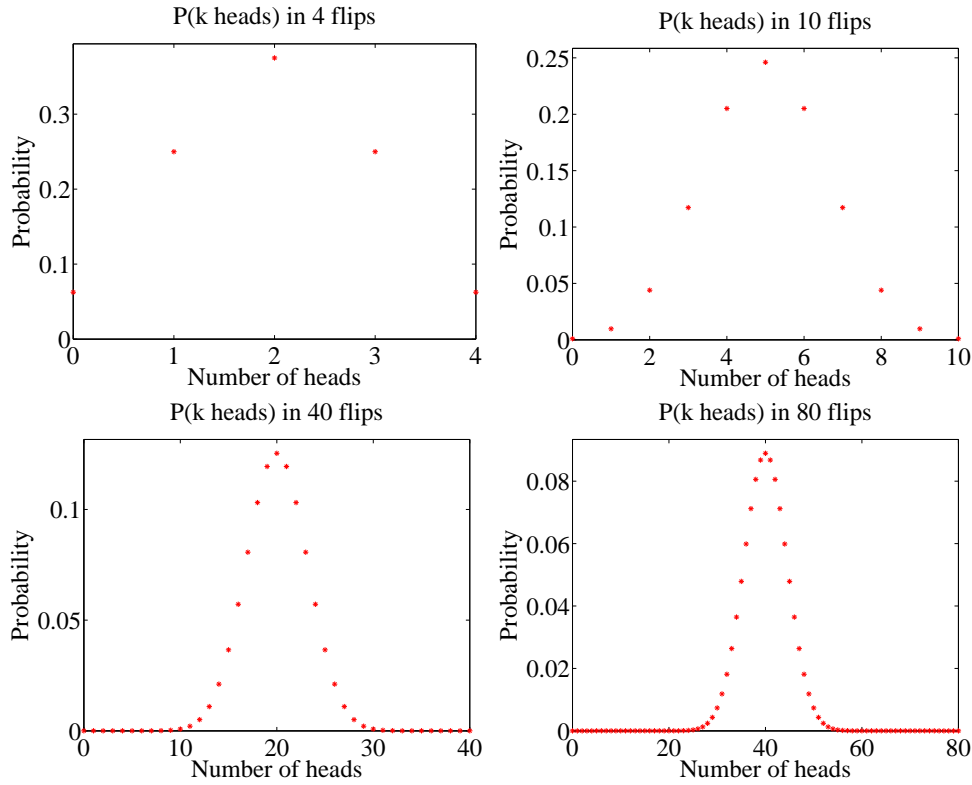


FIGURE 7.4: Plots of the binomial distribution for $p = q = 0.5$ for different values of N . You should notice that the set of values of h (the number of heads) that have substantial probability is quite narrow compared to the range of possible values. This set gets narrower as the number of flips increases. This is because the mean is pN and the standard deviation is \sqrt{Npq} — so the fraction of values that is within one standard deviation of the mean is $O(1/\sqrt{N})$.

non-zero probability becomes very large, and x can be very large, or very small. For example, there is some probability, though there may be very little indeed, on the point where $h = N$, or, equivalently, $x = N(p + \sqrt{Npq})$. For sufficiently large N , we think of this probability distribution as a probability density function. We can do so, for example, by spreading the probability for x_i (the i 'th value of x) evenly over the interval between x_i and x_{i+1} . We then have a probability density function that looks like a histogram, with bars that become narrower as N increases. But what is the limit?

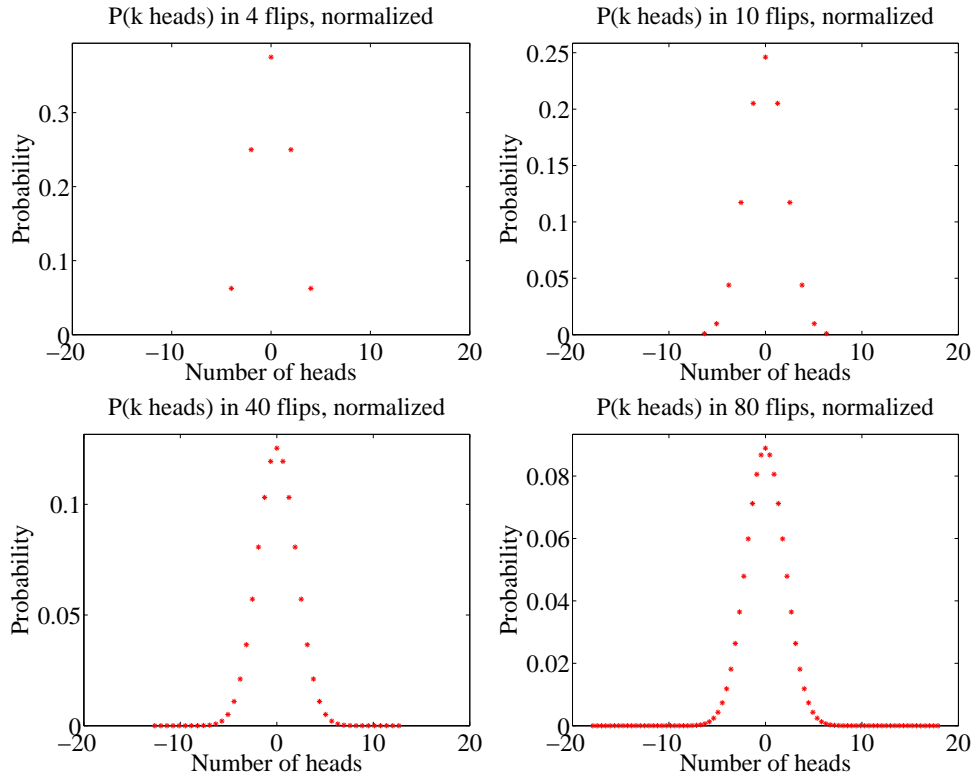


FIGURE 7.5: Plots of the distribution for the normalized variable x , with $P(x)$ given in the text, obtained from the binomial distribution with $p = q = 0.5$ for different values of N . These distributions are normalized (mean 0, variance 1). They look increasingly like a standard normal distribution EXCEPT that the value at their mode gets smaller as N gets bigger (there are more possible outcomes). In the text, we will establish that the standard normal distribution is a limit, in a useful sense.

7.4.2 Getting Normal

To proceed, we need Stirling's approximation, which says that, for large N ,

$$N! \approx \sqrt{2\pi} \sqrt{N} \left(\frac{N}{e}\right)^N.$$

This yields

$$P(h) \approx \left(\frac{Np}{h}\right)^h \left(\frac{Nq}{N-h}\right)^{(N-h)} \sqrt{\frac{N}{2\pi h(N-h)}}$$

Recall we used the normalized variable

$$x = \frac{h - Np}{\sqrt{Npq}}.$$

We will encounter the term \sqrt{Npq} often, and we use $\sigma = \sqrt{Npq}$ as a shorthand. We can compute h and $N - h$ from x by the equalities

$$h = Np + \sigma x \qquad N - h = Nq - \sigma x.$$

So the probability distribution written in this new variable x is

$$P(x) \approx \left(\frac{Np}{Np + \sigma x} \right)^{(Np + \sigma x)} \left(\frac{Nq}{Nq - \sigma x} \right)^{(Nq - \sigma x)} \sqrt{\frac{N}{2\pi(Np + \sigma x)(Nq - \sigma x)}}$$

There are three terms to deal with here. It is easiest to work with $\log P$. Now

$$\log(1 + x) = x - \frac{1}{2}x^2 + O(x^3)$$

so we have

$$\begin{aligned} \log \left(\frac{Np}{Np + \sigma x} \right) &= -\log \left(1 + \frac{\sigma x}{Np} \right) \\ &\approx -\frac{\sigma x}{Np} + \left(\frac{1}{2} \right) \left(\frac{\sigma x}{Np} \right)^2 \end{aligned}$$

and

$$\log \left(\frac{Nq}{Nq - \sigma x} \right) \approx \frac{\sigma x}{Nq} + \left(\frac{1}{2} \right) \left(\frac{\sigma x}{Nq} \right)^2.$$

From this, we have that

$$\begin{aligned} \log \left[\left(\frac{Np}{Np + \sigma x} \right)^{(Np + \sigma x)} \left(\frac{Nq}{Nq - \sigma x} \right)^{(Nq - \sigma x)} \right] &\approx [Np + \sigma x] \left[-\frac{\sigma x}{Np} + \left(\frac{1}{2} \right) \left(\frac{\sigma x}{Np} \right)^2 \right] + \\ &\quad [Nq - \sigma x] \left[\frac{\sigma x}{Nq} + \left(\frac{1}{2} \right) \left(\frac{\sigma x}{Nq} \right)^2 \right] \\ &= -\left(\frac{1}{2} \right) x^2 + O((\sigma x)^3) \end{aligned}$$

(recall $\sigma = \sqrt{Npq}$ if you're having trouble with the last step). Now we look at the square-root term. We have

$$\begin{aligned} \log \sqrt{\frac{N}{2\pi(Np + \sigma x)(Nq - \sigma x)}} &= -\frac{1}{2} (\log [Np + \sigma x] + \log [Nq - \sigma x] - \log N + \log 2\pi) \\ &= -\frac{1}{2} \left(\begin{array}{l} \log Np + O\left(\left(\frac{\sigma x}{Np}\right)\right) \\ + \log Nq - O\left(\left(\frac{\sigma x}{Nq}\right)\right) \\ - \log N + \log 2\pi \end{array} \right) \end{aligned}$$

but, since N is very large compared to σx , we can ignore the $O\left(\left(\frac{\sigma x}{Np}\right)\right)$ terms. Then this term is not a function of x . So we have

$$\log P(x) \approx \frac{-x^2}{2} + \text{constant}.$$

Now because N is very large, our probability distribution P limits to a probability density function p , with

$$p(x) \propto \exp\left(\frac{-x^2}{2}\right).$$

We can get the constant of proportionality from integrating, to

$$p(x) = \left(\frac{1}{\sqrt{2\pi}}\right) \exp\left(\frac{-x^2}{2}\right).$$

This constant of proportionality deals with the effect in figure 7.5, where the mode of the distribution gets smaller as N gets bigger. It does so because there are more points with non-zero probability to be accounted for. But we are interested in the limit where N tends to infinity. This must be a probability density function, so it must integrate to one.

Review this blizzard of terms. We started with a binomial distribution, but standardized the variables so that the mean was zero and the standard deviation was one. We then assumed there was a very large number of coin tosses, so large that that the distribution started to look like a continuous function. The function we get is the standard normal distribution.

7.4.3 So What?

I have proven an extremely useful fact, which I shall now put in a box.

Useful Fact: 7.12 *The Binomial Distribution for Large N*

Assume h follows the binomial distribution with parameters p and q . Write $x = \frac{h - Np}{\sqrt{Npq}}$. Then, for sufficiently large N , the probability distribution $P(x)$ can be approximated by the probability density function

$$\left(\frac{1}{\sqrt{2\pi}}\right) \exp\left(\frac{-x^2}{2}\right)$$

in the sense that

$$P(\{x \in [a, b]\}) \approx \int_a^b \left(\frac{1}{\sqrt{2\pi}}\right) \exp\left(\frac{-u^2}{2}\right) du$$

This justifies our model of probability as frequency. I interpreted an event having probability p to mean that, if I had a large number N of independent repetitions of the experiment, the number that produced the event would be close to Np , and would get closer as N got larger. We know that, for example, 68% of the time a standard normal random variable takes a value between 1 and -1 . In this case, the standard normal random variable is

$$\frac{h - (Np)}{\sqrt{Npq}}$$

so that 68% of the time, h must take a value in the range $[Np - \sqrt{Npq}, Np + \sqrt{Npq}]$. Equivalently, the relative frequency h/N must take a value in the range

$$\left[p - \frac{pq}{\sqrt{N}}, p + \frac{pq}{\sqrt{N}} \right]$$

but as $N \rightarrow \infty$ this range gets smaller and smaller, and h/N limits to p . So our view of probability as a frequency is consistent.

To obtain h , we added N independent Bernoulli random variables. So you can interpret the box as saying that the sum of many independent Bernoulli random variables has a probability distribution that limits to the normal distribution as the number added together gets larger. Remember that I have stated, though not precisely, but not proved the deep and useful fact that the sum of pretty much any independent random variables has a distribution that gets closer to a normal distribution as the number added together gets larger.

7.5 WHAT YOU SHOULD REMEMBER

You should remember:

- The form of the uniform distribution.
- The form of the geometric distribution, and its mean and variance.
- The form of the binomial distribution, and its mean and variance.
- The form of the Poisson distribution, and its mean and variance.
- The form of the Normal distribution, and its mean and variance.
- The fact that a sum of a large number of IID binomial random variables is normally distributed, and the mean and variance of that distribution.
- The fact that a sum of a large number of IID random variables is normally distributed for most cases you will encounter.

PROBLEMS

Sums and Differences of Discrete Random Variables

7.1. Assume X and Y are discrete random variables which take integer values in the range $1 \dots 100$ (inclusive). Write $S = X + Y$.

(a) Show that

$$P(S = k) = \sum_{u=1}^{u=100} P(\{X = k - u\} \cap \{Y = u\}).$$

(b) Show that

$$P(D = k) = \sum_{u=1}^{u=100} P(\{X = k + u\})P(\{Y = u\}).$$

(c) Now assume that both X and Y are uniform random variables. Show that S is not uniform by considering $P(S = 2)$, $P(S = 3)$, and $P(S = 100)$.

The Geometric Distribution

7.2. Recall that for $0 < r < 1$,

$$\sum_{i=0}^{\infty} r^i = \frac{1}{1-r}.$$

Now show that

$$\begin{aligned} \sum_{n=1}^{\infty} P(\{X = n\}) &= p \sum_{n=1}^{\infty} (1-p)^{(n-1)} \\ &= 1 \end{aligned}$$

7.3. Write $S_{\infty} = \sum_{i=0}^{\infty} r^i$. Show that $(1-r)S_{\infty} = 1$, so that

$$S_{\infty} = \frac{1}{1-r}$$

7.4. Show that

$$\sum_{i=0}^{\infty} ir^i = \left(\sum_{i=1}^{\infty} r^i\right) + r\left(\sum_{i=1}^{\infty} r^i\right) + r^2\left(\sum_{i=1}^{\infty} r^i\right) + \dots$$

(look carefully at the limits of the sums!) and so show that

$$\sum_{i=0}^{\infty} ir^i = \frac{r}{(1-r)^2}.$$

7.5. Write $S_{\infty} = \sum_{i=0}^{\infty} r^i$. Show that

$$\sum_{i=0}^{\infty} i^2 r^i = (\gamma - 1) + 3r(\gamma - 1) + 5r^2(\gamma - 1) + \dots$$

and so that

$$\sum_{i=0}^{\infty} i^2 r^i = \frac{r(1+r)}{(1-r)^3}$$

7.6. Show that, for a geometric distribution with parameter p , the mean is

$$\sum_{i=1}^{\infty} i(1-p)^{(i-1)}p = \sum_{i=0}^{\infty} (i+1)(1-p)^i p.$$

Now by rearranging and using the previous results, show that the mean is

$$\sum_{i=1}^{\infty} i(1-p)^{(i-1)}p = \frac{1}{p}$$

7.7. Show that, for a geometric distribution with parameter p , the variance is $\frac{1-p}{p^2}$.

To do this, note the variance is $\mathbb{E}[X^2] - \mathbb{E}[X]^2$. Now use the results of the previous exercises to show that

$$\mathbb{E}[X^2] = \sum_{i=1}^{\infty} i^2(1-p)^{(i-1)}p = \frac{p}{1-p} \frac{(1-p)(2-p)}{p^3},$$

then rearrange to get the expression for variance.

Bernoulli Random Variables

- 7.8.** Write X for a Bernoulli random variable which takes the value 1 with probability p (and 0 with probability $(1 - p)$).
- (a) Show that $\mathbb{E}[X] = p$.
- (b) Show that $\mathbb{E}[X^2] - \mathbb{E}[X]^2 = p(1 - p)$

The Binomial Distribution

- 7.9.** Show that $P_b(N - i; N, p) = P_b(i; N, p)$ for all i .
- 7.10.** Write h_r for the number of heads obtained in r flips of a coin which has probability p of coming up heads. Compare the following two ways to compute the probability of getting i heads in five coin flips:
- Flip the coin three times, count h_3 , then flip the coin twice, count h_2 , then form $w = h_3 + h_2$.
 - Flip the coin five times, and count h_5 .

Show that the probability distribution for w is the same as the probability distribution for h_5 . Do this by showing that

$$P(\{w = i\}) = \sum_{j=0}^5 P(\{h_3 = j\} \cap \{h_2 = i - j\}) = P(\{h_5 = i\}).$$

- 7.11.** Now we will do the previous exercise in a more general form. Again, write h_r for the number of heads obtained in r flips of a coin which has probability p of coming up heads. Compare the following two ways to compute the probability of getting i heads in N coin flips:
- Flip the coin t times, count h_t , then flip the coin $N - t$ times, count h_{N-t} , then form $w = h_t + h_{N-t}$.
 - Flip the coin N times, and count h_N .

Show that the probability distribution for w is the same as the probability distribution for h_N . Do this by showing that

$$P(\{w = i\}) = \sum_{j=0}^N P(\{h_t = j\} \cap \{h_{N-t} = i - j\}) = P(\{h_N = i\}).$$

You will likely find the recurrence relation

$$P_b(i; N, p) = pP_b(i - 1; N - 1, p) + (1 - p)P_b(i; N - 1, p).$$

is useful.

- 7.12.** An airline runs a regular flight with six seats on it. The airline sells six tickets. The gender of the passengers is unknown at time of sale, but women are as common as men in the population. All passengers always turn up for the flight. The pilot is eccentric, and will not fly a plane unless at least one passenger is female. What is the probability that the pilot flies?
- 7.13.** An airline runs a regular flight with s seats on it. The airline always sells t tickets for this flight. The probability a passenger turns up for departure is p , and passengers do this independently. What is the probability that the plane travels with exactly 3 empty seats?

- 7.14.** An airline runs a regular flight with s seats on it. The airline always sells t tickets for this flight. The probability a passenger turns up for departure is p , and passengers do this independently. What is the probability that the plane travels with 1 or more empty seats?
- 7.15.** An airline runs a regular flight with 10 seats on it. The probability that a passenger turns up for the flight is 0.95. What is the smallest number of seats the airline should sell to ensure that the probability the flight is full (i.e. 10 or more passengers turn up) is bigger than 0.99? (you'll probably need to use a calculator or write a program for this).

The Multinomial Distribution

- 7.16.** Show that the multinomial distribution

$$P_m(n_1, \dots, n_k; N, p_1, \dots, p_k) = \frac{N!}{n_1!n_2!\dots n_k!} p_1^{n_1} p_2^{n_2} \dots p_k^{n_k}$$

must satisfy the recurrence relation

$$\begin{aligned} P_m(n_1, \dots, n_k; N, p_1, \dots, p_k) &= p_1 P_m(n_1 - 1, \dots, n_k; N - 1, p_1, \dots, p_k) + \\ & p_2 P_m(n_1, n_2 - 1, \dots, n_k; N - 1, p_1, \dots, p_k) + \dots \\ & p_k P_m(n_1, n_2, \dots, n_k - 1; N - 1, p_1, \dots, p_k) \end{aligned}$$

The Poisson Distribution

- 7.17.** Compute the Taylor series for xe^x around $x = 0$. Use this and pattern matching to show that the mean of the Poisson distribution with intensity parameter λ is λ .
- 7.18.** Compute the Taylor series for $(x^2 + x)e^x$ around $x = 0$. Use this and pattern matching to show that the variance of the Poisson distribution with intensity parameter λ is λ .

Sums of Continuous Random Variables

- 7.19.** Write p_x for the probability density function of a continuous random variable X and p_y for the probability density function of a continuous random variable Y . Show that the probability density function of $S = X + Y$ is

$$p(s) = \int_{-\infty}^{\infty} p_x(s - u)p_y(u)du = \int_{-\infty}^{\infty} p_x(u)p_y(s - u)du$$

The Normal Distribution

- 7.20.** Write

$$f(x) = \left(\frac{1}{\sqrt{2\pi}} \right) \exp\left(\frac{-x^2}{2} \right).$$

- (a) Show that $f(x)$ is non-negative for all x .
 (b) By integration, show that

$$\int_{-\infty}^{\infty} f(x)dx = 1,$$

so that $f(x)$ is a probability density function.

(c) Show that

$$\int_{-\infty}^{\infty} xf(x)dx = 0.$$

The easiest way to do this is to notice that $f(x) = f(-x)$

(d) Show that

$$\int_{-\infty}^{\infty} xf(x - \mu)dx = \mu.$$

The easiest way to do this is to change variables, and use the previous two exercises.

(e) Show that

$$\int_{-\infty}^{\infty} x^2 f(x)dx = 1.$$

You'll need to either do, or look up, the integral to do this exercise.

7.21. Write

$$g(x) = \exp\left[-\frac{(x - \mu)^2}{2\sigma^2}\right]$$

Show that

$$\int_{-\infty}^{\infty} g(x)dx = \sqrt{2\pi}\sigma.$$

You can do this by a change of variable, and the results of the previous exercises.

7.22. Write

$$p(x) = \left(\frac{1}{\sqrt{2\pi}\sigma}\right) \exp\left(\frac{-(x - \mu)^2}{2\sigma^2}\right).$$

(a) Show that

$$\int_{-\infty}^{\infty} xp(x)dx = \mu$$

using the results of the previous exercises.

(b) Show that

$$\int_{-\infty}^{\infty} x^2 p(x)dx = \sigma^2$$

using the results of the previous exercises.

The Binomial Distribution for Large N

7.23. I flip a fair coin N times and count heads. We consider the probability that h , the fraction of heads, is in some range of numbers. *Hint:* If you know the range of numbers for h , you know the range for h/N .

(a) For $N = 1e6$, what is $P(\{h \in [49500, 50500]\})$?

(b) For $N = 1e4$, what is $P(\{h > 9000\})$?

(c) For $N = 1e2$, what is $P(\{h > 60\} \cup \{h < 40\})$?

Markov Chains and Simulation

There are many situations where one must work with a sequence of random variables. For example, consider a bus queue; people arrive at random, and so do buses. What is the probability that the queue gets to a particular length? and what is the expected length of the queue? As another example, we could have a random model of purchases from a shop — under any particular rule for restoring inventory, what is the largest (resp. smallest) amount of stock on the shelf? what is the expected amount of stock on the shelf? It turns out that there is a simple and powerful model that applies to many sequences of random variables. One can often use this model to make closed-form predictions. In cases where closed-form predictions aren't available, one can use simulation methods to estimate probabilities and expectations.

8.1 MARKOV CHAINS

A Markov chain is a sequence of random variables which has important independence properties. We will describe these properties in some detail below; first, we give some examples. Markov chains are easily represented with the language of finite state machines. For some Markov chains, it is easy to determine probabilities and expectations of interest in closed form using simple methods. For others, it is tricky (straightforward, but unreasonable amounts of straightforward work). In these cases, we can estimate the relevant probabilities and expectations by simulating the finite state machine.

8.1.1 Motivating Example: Multiple Coin Flips

We start with three examples, each of which is easy to work. Each suggests a much harder question, however, which we need new machinery to handle.

Worked example 8.1 *Multiple Coin Flips - 1*

You choose to flip a fair coin until you see two heads in a row, and then stop. What is the probability that you flip the coin twice?

Solution: Because you stopped after two flips, you must have seen two heads. So $P(2 \text{ flips}) = P(\{HH\}) = P(\{H\})^2 = 1/4$.

Worked example 8.2 *Multiple Coin Flips - 2*

You choose to flip a fair coin until you see two heads in a row, and then stop. What is the probability that you flip the coin three times?

Solution: Because you stopped after three flips, you must have seen T , then H , then H ; any other sequence either doesn't stop, or stops too early. We write this with the last flip last, so THH . So $P(3 \text{ flips}) = P(\{THH\}) = P(\{T\})P(\{H\})^2 = 1/8$.

Worked example 8.3 *Multiple Coin Flips - 3*

You choose to flip a fair coin until you see two heads in a row, and then stop. What is the probability that you flip the coin four times?

Solution: This is more interesting. The last three flips must have been THH (otherwise you'd go on too long, or end too early). But, because the second flip must be a T , the first could be either H or T . This means there are two sequences that work: $HTHH$ and $TTHH$. So $P(4 \text{ flips}) = 2/8 = 1/4$.

The harder question here is to ask what is $P(N)$, for N some number (larger than 4, because we know the answers in the other cases). It is unattractive to work this case by case (you could try this, if you don't believe me). One very helpful way to think about the coin flipping experiment is as a finite state machine (Figure 8.1). If you think of this machine as a conventional finite state machine, it accepts any string of T , H , that (a) ends with HH , and (b) has no other HH in it. Alternatively, you could think of this machine as encoding a probabilistic process. At each state, an event occurs with some probability (in this case, a coin comes up H or T , with probability $(1/2)$). The event causes the machine to follow the appropriate edge. If we take this view, this machine stops when we have flipped two heads in succession. It encodes our problem of computing the probability of flipping a coin N times then stopping — this is just the probability that the machine hits the end state after N transitions.

It turns out to be straightforward to construct a recurrence relation for $P(N)$ (i.e. an equation for $P(N)$ in terms of $P(N-1)$, $P(N-2)$, and so on). This property is quite characteristic of repeated experiments. For this example, first, notice that $P(1) = 0$. Now imagine you have flipped the coin N times and stopped. We can assume that $N > 3$, because we know what happens for the other cases. The *last* three flips must have been THH . Write Σ_N for a sequence that is: (a) N flips long; (b) ends in HH ; and (c) contains no other subsequence HH . Equivalently, this is a string accepted by the finite state machine of figure 8.1. We must have that any

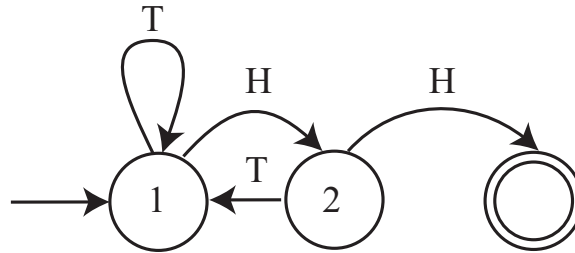


FIGURE 8.1: A finite state machine representing the coin flip example. By convention, the end state is a double circle, and the start state has an incoming arrow. I've labelled the arrows with the event that leads to the transition, but haven't bothered to put in the probabilities, because each is 0.5.

Σ_N has either the form $T\Sigma_{N-1}$ or the form $HT\Sigma_{N-2}$. But this means that

$$\begin{aligned} P(N) &= P(T)P(N-1) + P(HT)P(N-2) \\ &= (1/2)P(N-1) + (1/4)P(N-2) \end{aligned}$$

It is possible to solve this recurrence relation to get an explicit formula, but doing so would take us out of our way. You will check this recurrence relation in an exercise.

One really useful way to think of this recurrence relation is that represents an exercise in counting. We want to count all sequences that are of length N , and are accepted by the finite state machine of figure 8.1. This means they: (a) are N flips long; (b) end in HH ; and (c) contain no other subsequence HH . Now work backward along the FSM. The only way to arrive at the final state is to be in state 1, then see HH . So you can obtain an acceptable N element sequence by (a) prepending a T to an acceptable $N-1$ element sequence or (b) prepending TH (which takes you to 2, then back to 1) to an acceptable $N-2$ element sequence. This line of reasoning can be made much more elaborate. There are a few examples in the exercises.

8.1.2 Motivating Example: The Gambler's Ruin

Another useful example is known as the **gambler's ruin**. Assume you bet \$1 a tossed coin will come up heads. If you win, you get \$1 and your original stake back. If you lose, you lose your stake. But this coin has the property that $P(H) = p < 1/2$. We will study what happens when you bet repeatedly.

Assume you have \$ s when you start. You will keep betting until either (a) you have \$0 (you are ruined; you can't borrow money) or (b) the amount of money you have accumulated is \$ j , where $j > s$. The coin tosses are independent. We will compute $P(\text{ruined, starting with } s|p)$ the probability that you leave the table with nothing, when you start with \$ s . For brevity, we write $P(\text{ruined, starting with } s|p) = p_s$. You can represent the gambler's ruin problem with a finite state machine as well. Figure 8.2 shows a representation of the gambler's ruin problem.

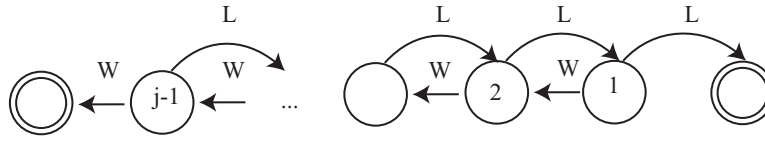


FIGURE 8.2: A finite state machine representing the gambler's ruin example. I have labelled each state with the amount of money the gambler has at that state. There are two end states, where the gambler has zero (is ruined), or has j and decides to leave the table. The problem we discuss is to compute the probability of being ruined, given the start state is s . This means that any state except the end states could be a start state. I have labelled the state transitions with "W" (for win) and "L" for lose, but have omitted the probabilities.

Worked example 8.4 *The gambler's ruin - 1*

Using the notation above, determine p_0 and p_j

Solution: We must have $p_0 = 1$, because if you have \$0, you leave the table. Similarly, if you have \$ j , you leave the table with \$ j , so you don't leave the table with nothing, so $p_j = 0$.

Worked example 8.5 *The gambler's ruin - 2*

Using the notation above, write a recurrence relation for p_s (the probability that you leave the table with nothing when you started with \$ s).

Solution: Assume that you win the first bet. Then you have \$ $s + 1$, so your probability of leaving the table with nothing now becomes p_{s+1} . If you lose the first bet, then you have \$ $s - 1$, so your probability of leaving the table with nothing now becomes p_{s-1} . The coin tosses are independent, so we can write

$$p_s = pp_{s+1} + (1-p)p_{s-1}.$$

Some fairly lively work with series, relegated to the end of the chapter as exercises, yields

$$p_s = \frac{\left(\frac{1-p}{p}\right)^j - \left(\frac{1-p}{p}\right)^s}{\left(\frac{1-p}{p}\right)^j - 1}.$$

This expression is quite informative. Notice that, if $p < 1/2$, then $(1-p)/p > 1$. This means that as $j \rightarrow \infty$, we have $p_s \rightarrow 1$. If you gamble repeatedly on an unfair

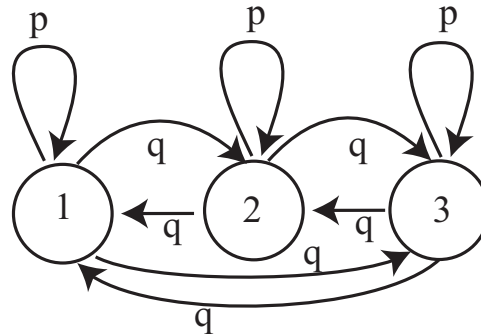


FIGURE 8.3: A virus can exist in one of 3 strains. At the end of each year, the virus mutates. With probability α , it chooses uniformly and at random from one of the 2 other strains, and turns into that; with probability $1 - \alpha$, it stays in the strain it is in. For this figure, we have transition probabilities $p = (1 - \alpha)$ and $q = (\alpha/2)$.

coin, the probability that you run out of money before you hit some threshold ($\$j$ in this case) tends to one.

8.1.3 Motivating Example: A Virus

Problems represented with a finite state machine don't need to have an end state. As one example, (which I got from ACC Coolen's lecture notes), we have a virus that can exist in one of k strains. At the end of each year, the virus mutates. With probability α , it chooses uniformly and at random from one of the $k - 1$ other strains, and turns into that; with probability $1 - \alpha$, it stays in the strain it is in (Figure 8.3 shows an example with three strains). For that figure, we have $p = (1 - \alpha)$ and $q = (\alpha/2)$. The virus just keeps on changing, and there is no end state. But there are a variety of very interesting questions we can try to answer. We would like to know, for example, the expected time to see a strain a second time, i.e. if the virus is in strain 1 at time 1, what is the expected time before it is in strain 1 again? If the virus has mutated many times, what is the probability that it is in each strain? and do these probabilities depend on the start strain?

8.1.4 Markov Chains

In Figure 8.1 and Figure 8.2, I showed the event that caused the state transitions. It is more usual to write a probability on the figure, as I did in Figure 8.3, because the probability of a state transition (rather than what caused it) is what really matters. The underlying object is now a weighted directed graph, because we have removed the events and replaced them with probabilities. These probabilities are known as **transition probabilities**; notice that the sum of transition probabilities over *outgoing* arrows must be 1.

We can now think of our process as a **biased random walk** on a weighted directed graph. A bug (or any other small object you prefer) sits on one of the graph's nodes. At each time step, the bug chooses one of the outgoing edges at

random. The probability of choosing an edge is given by the probabilities on the drawing of the graph (equivalently, the transition probabilities). The bug then follows that edge. The bug keeps doing this until it hits an end state.

This bug produces a sequence of random variables. If there are k states in the finite state machine, we can label each state with a number, $1 \dots k$. At the n 'th time step, the state of the process — the node that the bug is sitting on — is a random variable, which we write X_n . These random variables have an important property. The probability that X_n takes some particular value depends only on X_{n-1} , and not on any other previous state. If we know where the bug is at step $n - 1$ in our model, we know where it could go, and the probability of each transition. Where it was at previous times does not affect this, *as long as we know its state at step $n - 1$* .

A sequence of random variables X_n is a **Markov chain** if it has the property that, $P(X_n = j | \text{values of all previous states}) = P(X_n = j | X_{n-1})$, or, equivalently, only the last state matters in determining the probability of the current state. The probabilities $P(X_n = j | X_{n-1} = i)$ are the **transition probabilities**. Any model built by taking the transitions of a finite state machine and labelling them with probabilities must be a Markov chain. However, this is not the only way to build or represent a Markov chain.

One representation of a Markov chain uses a matrix of transition probabilities. We define the matrix \mathcal{P} with $p_{ij} = P(X_n = j | X_{n-1} = i)$. Notice that this matrix has the properties that $p_{ij} \geq 0$ and

$$\sum_j p_{ij} = 1$$

because at the end of each time step the model must be in some state. Equivalently, the sum of transition probabilities for outgoing arrows is one. Non-negative matrices with this property are **stochastic matrices**. By the way, you should look very carefully at the i 's and j 's here — Markov chains are usually written in terms of *row* vectors, and this choice makes sense in that context.

Worked example 8.6 Viruses

Write out the transition probability matrix for the virus of Figure 8.3, assuming that $\alpha = 0.2$.

Solution: We have $P(X_n = 1 | X_{n-1} = 1) = (1 - \alpha) = 0.8$, and $P(X_n = 2 | X_{n-1} = 1) = \alpha/2 = P(X_n = 3 | X_{n-1} = 1)$; so we get

$$\begin{pmatrix} 0.8 & 0.1 & 0.1 \\ 0.1 & 0.8 & 0.1 \\ 0.1 & 0.1 & 0.8 \end{pmatrix}$$

Now imagine we do not know the initial state of the chain, but instead have a probability distribution. This gives $P(X_0 = i)$ for each state i . It is usual to take these k probabilities and place them in a k -dimensional row vector, which is

usually written π . For example, we might not know what the initial strain of the virus is, but just that each strain is equally likely. So for a 3-strain virus, we would have $\pi = [1/3, 1/3, 1/3]$. From this information, we can compute the probability distribution over the states at time 1 by

$$\begin{aligned} P(X_1 = j) &= \sum_i P(X_1 = j, X_0 = i) \\ &= \sum_i P(X_1 = j | X_0 = i) P(X_0 = i) \\ &= \sum_i p_{ij} \pi_i. \end{aligned}$$

If we write $\mathbf{p}^{(n)}$ for the row vector representing the probability distribution of the state at step n , we can write this expression as

$$\mathbf{p}^{(1)} = \pi \mathcal{P}.$$

Now notice that

$$\begin{aligned} P(X_2 = j) &= \sum_i P(X_2 = j, X_1 = i) \\ &= \sum_i P(X_2 = j | X_1 = i) P(X_1 = i) \\ &= \sum_i p_{ij} \left(\sum_{ki} p_{ki} \pi_k \right). \end{aligned}$$

so that

$$\mathbf{p}^{(n)} = \pi \mathcal{P}^n.$$

This expression is useful for simulation, and also allows us to deduce a variety of interesting properties of Markov chains.

Worked example 8.7 *Viruses*

We know that the virus of Figure 8.3 started in strain 1. After two state transitions, what is the distribution of states when $\alpha = 0.2$? when $\alpha = 0.9$? What happens after 20 state transitions? If the virus starts in strain 2, what happens after 20 state transitions?

Solution: If the virus started in strain 1, then $\pi = [1, 0, 0]$. We must compute $\pi(\mathcal{P}(\alpha))^2$. This yields $[0.66, 0.17, 0.17]$ for the case $\alpha = 0.2$ and $[0.4150, 0.2925, 0.2925]$ for the case $\alpha = 0.9$. Notice that, because the virus with small α tends to stay in whatever state it is in, the distribution of states after two years is still quite peaked; when α is large, the distribution of states is quite uniform. After 20 transitions, we have $[0.3339, 0.3331, 0.3331]$ for the case $\alpha = 0.2$ and $[0.3333, 0.3333, 0.3333]$ for the case $\alpha = 0.9$; you will get similar numbers even if the virus starts in strain 2. After 20 transitions, the virus has largely “forgotten” what the initial state was.

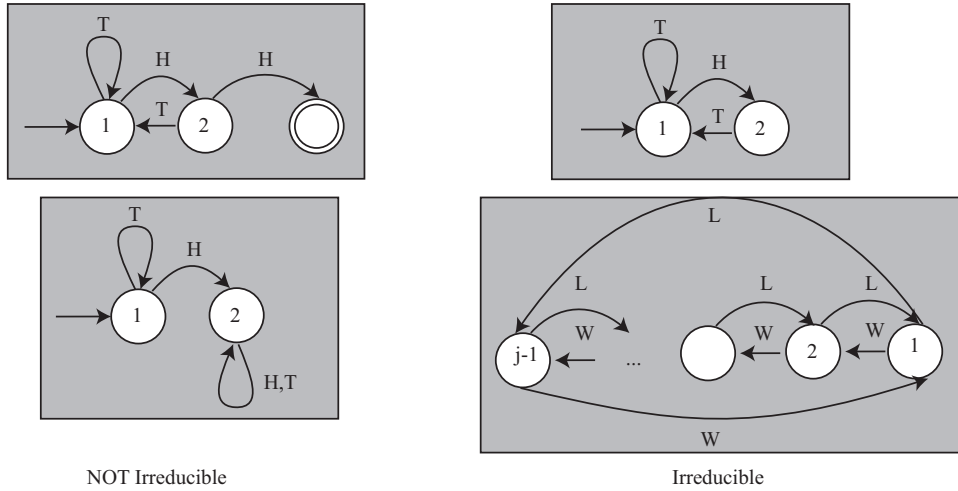


FIGURE 8.4: Examples of finite state machines that give rise to Markov chains that are NOT irreducible (left) and irreducible (right). To obtain Markov chains from these drawings, we would have to give the probabilities of the events that lead to state transitions. We'll assume that none of the probabilities are zero; after that, the values don't matter for irreducibility analysis. The **top left** FSM is not irreducible, because it has an end state; once the bug reaches this point, it can't go anywhere else. The **bottom left** FSM is not irreducible, because the bug can get stuck in state 2.

In example 7, the distribution of virus strains after a long interval appeared not to depend much on the initial strain. This property is true of many Markov chains. Assume that any state can be reached from any other state, by some sequence of transitions. Such chains are called **irreducible**; notice this means there is no end state, like the virus example. Irreducibility also means that the chain cannot get “stuck” in a state or a collection of states (Figure 8.4). Then there is a unique vector **s**, usually referred to as the **stationary distribution**, such that for *any* initial state distribution π ,

$$\lim_{n \rightarrow \infty} \pi \mathcal{P}^{(n)} = \mathbf{s}.$$

Equivalently, if the chain has run through many steps, it no longer matters what the initial distribution is. You expect that the probability distribution over states is **s**.

8.1.5 Example: Particle Motion as a Markov Chain

One can find Markov chains in quite unexpected places, often with useful consequences. In this example, I will obtain a Markov chain without reasoning about graphs or finite state machines. We will investigate a particle moving under gravity, in 3 dimensions. Write the position of the particle as $\mathbf{p}(t)$, its velocity as $\mathbf{v}(t)$, its

acceleration as $\mathbf{a}(t)$, its mass as m , and the gravitational force as \mathbf{g} . Then we know that

$$\begin{aligned}\mathbf{v}(t) &= \frac{d\mathbf{p}}{dt} \\ \mathbf{a}(t) &= \frac{d\mathbf{v}}{dt} \\ &= \mathbf{g}.\end{aligned}$$

Now stack the position and the acceleration into a single vector $X(t) = (\mathbf{p}(t), \mathbf{v}(t))^T$. We could write these equations as

$$\frac{dX}{dt} = \mathcal{A}X + \mathbf{b}$$

where

$$\mathcal{A} = \begin{pmatrix} 0 & \mathcal{I} \\ 0 & 0 \end{pmatrix}$$

and

$$\mathbf{b} = \begin{pmatrix} 0 \\ \mathbf{g} \end{pmatrix}.$$

Now imagine that we look at the position, velocity and acceleration of the particle at fixed time instants, so we are really interested in $X_i = X(t_0 + i\Delta t)$. In this case, we can approximate $\frac{dX}{dt}$ by $(X_{i+1} - X_i)/\Delta t$. We then have

$$X_{i+1} = X_i + (\Delta t)(\mathcal{A}X_i + \mathbf{b}).$$

This is beginning to look like a Markov chain, because X_{i+1} depends only on X_i but not on any previous X . But there isn't any randomness here. We can fix that by assuming that the particle is moving in, say, a light turbulent wind. Then the acceleration at any time consists of (a) the acceleration of gravity and (b) a small, random force produced by the wind. Write \mathbf{w}_i for this small, random force at time i , and $P(\mathbf{w}_i|i)$ for its probability distribution, which could reasonably depend on time. We can then rewrite our equation by writing

$$X_{i+1} = X_i + (\Delta t)(\mathcal{A}X_i + \mathbf{b}_r).$$

where

$$\mathbf{b}_r = \begin{pmatrix} 0 \\ \mathbf{g} + \mathbf{w}_i \end{pmatrix}.$$

Now the X_i are clearly random variables. We could get $P(X_{i+1}|X_i)$ by rearranging terms. If I know X_{i+1} and X_i , I know the value of \mathbf{w}_i ; I could plug this into $P(\mathbf{w}_i|i)$ to yield $P(X_{i+1}|X_i)$.

Using a finite state machine in this example would be a bit unnatural. This example turns out to be surprisingly useful in applications, because (with other algorithmic machinery we can't go into here) it offers the basis for algorithms that can track moving objects by predicting where they will go next. Applications are widespread. Military applications include tracking aircraft, UFO's, missiles, etc. Civilian applications include surveillance in public places; games console interfaces that track moving people; and methods that can follow experimental mice as they move around their cages (useful for testing medicines).

8.2 SIMULATION

Many problems in probability can be worked out in closed form if one knows enough combinatorial mathematics, or can come up with the right trick. Textbooks are full of these, and we've seen some. Explicit formulas for probabilities are often extremely useful. But it isn't always easy or possible to find a formula for the probability of an event in a model. An alternative strategy is to build a simulation, run it many times, and count the fraction of outcomes where the event occurs. This is a simulation experiment.

8.2.1 Obtaining Uniform Random Numbers

Simulation is at its most useful when we try to estimate probabilities that are hard to calculate. Usually we have processes with some random component, and we want to estimate the probability of a particular outcome. To do so, we will need a supply of random numbers to simulate the random component. Here I will describe methods to get uniformly distributed random numbers, and Section 8.2.5 describes a variety of methods to get random numbers from other distributions.

I will describe features in Matlab, because I'm used to Matlab. It's a good programming environment for numerical work, particularly for working with matrices and vectors. You can expect pretty much any programming environment to provide a random number generator that returns a number uniformly distributed in the range $[0 - 1]$. If you know anything about representing numbers, you'll already have spotted something funny. The number that comes out of the random number generator must be represented by a small set of bits, but almost all numbers in the interval $[0 - 1]$ require an infinite number of bits to represent. We'll sweep this issue under the general carpet of floating point representations; it doesn't matter for anything we need to do. The Matlab function to do this is called `rand`. It can also return matrices; for example, `rand(10, 20)` will get you a 10×20 table of independent uniformly distributed random numbers in the range $[0 - 1]$.

It is useful to get a uniformly distributed random integer in a particular range (for example, you might want to choose a random element in an array). You can do so with a random number generator and a function (like Matlab's `floor`) that returns the largest integer smaller than its argument. If I want a discrete random variable with uniform distribution, maximum value 100 and minimum value 7, I habitually choose a very tiny number (for this example, say $1e - 7$) and do `floor((100-7-1e-7)*rand()+7)`. I use the $1e - 7$ because I can never remember whether `rand` produces a number no larger than one, or one that is guaranteed to be smaller than one, and I never need to care about the very tiny differences in probability caused by the $1e-7$. You might be able to do something cleaner if you bothered checking this point.

8.2.2 Computing Expectations with Simulations

Simulation is also a very good way to estimate expectations. Imagine we have a random variable X with probability distribution $P(X)$ that takes values in some domain D . Assume that we can easily produce independent simulations, and that we wish to know $\mathbb{E}[f]$, the expected value of the function f under the distribution

$P(X)$.

The weak law of large numbers tells us how to proceed. Define a new random variable $F = f(X)$. This has a probability distribution $P(F)$, which might be difficult to know. We want to estimate $\mathbb{E}[f]$, the expected value of the function f under the distribution $P(X)$. This is the same as $\mathbb{E}[F]$. Now if we have a set of IID samples of X , which we write x_i , then we can form a set of IID samples of F by forming $f(x_i) = f_i$. Write

$$F_N = \frac{\sum_{i=1}^N f_i}{N}.$$

This is a random variable, and the weak law of large numbers gives that, for any positive number ϵ

$$\lim_{N \rightarrow \infty} P(\{\|F_N - \mathbb{E}[F]\| > \epsilon\}) = 0.$$

You can interpret this as saying that, that for a set of IID random samples x_i , the probability that

$$\frac{\sum_{i=1}^N f(x_i)}{N}$$

is very close to $\mathbb{E}[f]$ is high for large N

Worked example 8.8 *Computing an Expectation*

Assume the random variable X is uniformly distributed in the range $[0 - 1]$, and the random variable Y is uniformly distributed in the range $[0 - 10]$. X and Z are independent. Write $Z = (Y - 5X)^3 - X^2$. What is $\text{var}(\{Z\})$?

Solution: With enough work, one could probably work this out in closed form. An easy program will get a good estimate. We have that $\text{var}(\{Z\}) = \mathbb{E}[Z^2] - \mathbb{E}[Z]^2$. My program computed 1000 values of Z (by drawing X and Y from the appropriate random number generator, then evaluating the function). I then computed $\mathbb{E}[Z]$ by averaging those values, and $\mathbb{E}[Z]^2$ by averaging their squares. For a run of my program, I got $\text{var}(\{Z\}) = 2.76 \times 10^4$.

8.2.3 Computing Probabilities with Simulations

You can compute a probability using a simulation, too, because a probability can be computed by taking an expectation. Recall the property of indicator functions that

$$\mathbb{E}[\mathbb{I}_{\{\mathcal{E}\}}] = P(\mathcal{E})$$

(Section 6.2.5). This means that computing the probability of an event \mathcal{E} involves writing a function that is 1 when the event occurs, and 0 otherwise; we then estimate the expected value of that function.

The weak law of large numbers justifies this procedure. An experiment involves drawing a sample from the relevant probability distribution $P(X)$, then determining

whether event \mathcal{E} occurred or not. We define a new random variable E , which takes the value 1 when the event \mathcal{E} occurs during our experiment (i.e. with probability $P(\mathcal{E})$) and 0 when it doesn't (i.e. with probability $P(\mathcal{E}^c)$). Our experiment yields a sample of this random variable. We perform N experiments yielding a set of IID samples of this distribution e_i and compute $E_N = \frac{\sum_{i=1}^N e_i}{N}$. From the weak law of large numbers, we have that for any $\epsilon > 0$,

$$\lim_{N \rightarrow \infty} P(\{\|E_N - \mathbb{E}[E]\| \geq \epsilon\}) = 0$$

meaning that for large enough N , E_N will be very close to $\mathbb{E}[E] = P(\mathcal{E})$.

Worked example 8.9 *Computing a Probability for Multiple Coin Flips*

You flip a fair coin three times. Use a simulation to estimate the probability that you see three H 's.

Solution: You really should be able to work this out in closed form. But it's amusing to check with a simulation. I wrote a simple program that obtained a 1000x3 table of uniformly distributed random numbers in the range $[0 - 1]$. For each number, if it was greater than 0.5 I recorded an H and if it was smaller, I recorded a T . Then I counted the number of rows that had 3 H 's (i.e. the expected value of the relevant indicator function). This yielded the estimate 0.127, which compares well to the right answer.

Worked example 8.10 *Computing a Probability*

Assume the random variable X is uniformly distributed in the range $[0 - 1]$, and the random variable Y is uniformly distributed in the range $[0 - 10]$. Write $Z = (Y - 5X)^3 - X^2$. What is $P(\{Z > 3\})$?

Solution: With enough work, one could probably work this out in closed form. An easy program will get a good estimate. My program computed 1000 values of Z (by drawing X and Y from the appropriate random number generator, then evaluating the function) and counted the fraction of Z values that was greater than 3 (which is the relevant indicator function). For a run of my program, I got $P(\{Z > 3\}) \approx 0.619$

8.2.4 Simulation Results as Random Variables

The estimate of a probability or of an expectation that comes out of a simulation experiment is a random variable, because it is a function of random numbers. If you run the simulation again, you'll get a different value (unless you did something silly with the random number generator). Generally, you should expect this random variable to behave like a normal random variable. You can check this by constructing a histogram over a large number of runs. The mean of this random

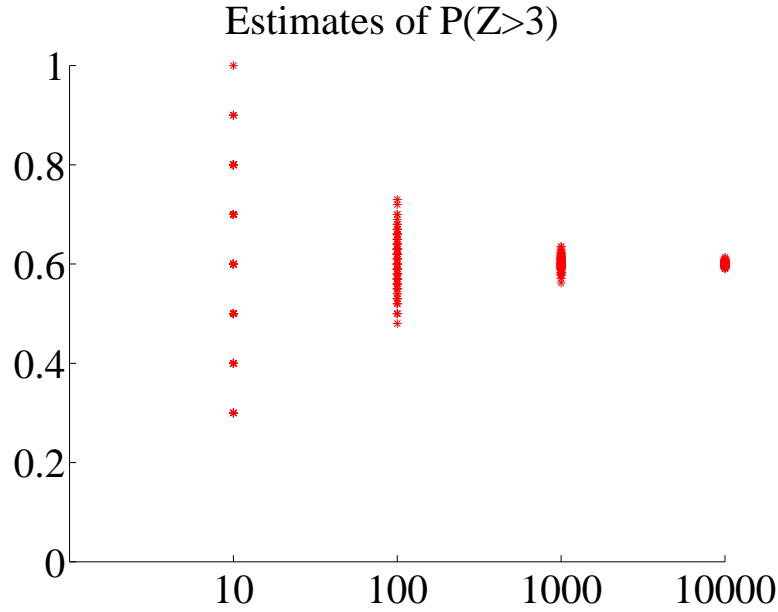


FIGURE 8.5: *Estimates of the probability from example 10, obtained from different runs of my simulator using different numbers of samples. In each case, I used 100 runs; the number of samples is shown on the horizontal axis. You should notice that the estimate varies pretty widely when there are only 10 samples, but the variance (equivalently, the size of the spread) goes down sharply as the number of samples increases to 1000. Because we expect these estimates to be roughly normally distributed, the variance gives a good idea of how accurate the original probability estimate is.*

variable is the parameter you are trying to estimate. It is useful to know that this random variable tends to be normal, because it means the standard deviation of the random variable tells you a lot about the likely values you will observe.

Another helpful rule of thumb, which is almost always right, is that the standard deviation of this random variable behaves like

$$\frac{C}{\sqrt{N}}$$

where C is a constant that depends on the problem and can be very hard to evaluate, and N is the number of runs of the simulation. What this means is that if you want to (say) double the accuracy of your estimate of the probability or the expectation, you have to run four times as many simulations. Very accurate estimates are tough to get, because they require immense numbers of simulation runs.

Figure 8.5 shows how the result of a simulation behaves when the number of runs changes. I used the simulation of example 10, and ran multiple experiments for each of a number of different samples (i.e. 100 experiments using 10 samples; 100 using 100 samples; and so on).

Small probabilities can be rather hard to estimate, as we would expect. In the case of example 10, let us estimate $P(\{Z > 950\})$. A few moments with a computer will show that this probability is of the order of $1e-3$ to $1e-4$. I obtained a million different simulated values of Z from my program, and saw 310 where $Z > 950$. This means that to know this probability to, say, three digits of numerical accuracy might involve a daunting number of samples. Notice that this does not contradict the rule of thumb that the standard deviation of the random variable defined by a simulation estimate behaves like $\frac{C}{\sqrt{N}}$; it's just that in this case, C is very large indeed.

8.2.5 Obtaining Random Samples

Building a successful simulation requires appropriate random numbers. Recall that anything we compute from a simulation can be thought of as an expectation (we used indicator functions to estimate probabilities). So we have some random variable X with distribution $P(X)$, a function f , and we wish to compute $\mathbb{E}[f]$, where the expectation is with respect to $P(X)$.

Most programming environments provide random number generators for uniform random numbers (I described the one for MATLAB briefly in Section 8.2.1) and for normally distributed random numbers. Building a really good, really fast random number generator is a major exercise. All sorts of tricks are involved, because it really matters to produce executable code that is as fast as possible on the target machine. This means that you don't have to build your own (and shouldn't, unless you can spend a lot of time and trouble on doing so).

Normal Random Variables

In pretty much any programming environment, you would also expect to find a random number generator that returns a normal random variable, with mean zero and standard deviation one. In Matlab, this function is called `randn`. Conveniently, `randn(3, 4)` will give you a 3×4 table of such numbers, which are independent of each other. As you would expect from section 1, to change the mean of this random number, you add a constant; to change the variance, you multiply by a constant. So in Matlab, if you want a normal random variable with mean 3 and standard deviation 4, you use `4*randn()+3`.

Rejection Sampling

Imagine you know a probability distribution describing a discrete random variable. In particular, you have one probability value for each of the numbers $1, 2, \dots, N$ (all the others are zero). Write $p(i)$ for the probability of i . You can generate a sample of this distribution by the following procedure: first, generate a sample x of a uniform discrete random variable in the range $1, \dots, N$; now generate a sample t of a uniformly distributed continuous random variable in the range $[0, 1]$; finally, if $t < p(x)$ report x , otherwise generate a new x and repeat. This process is known as **rejection sampling** (Algorithm 8.1).

To understand this procedure, it is best to think of it as a loop around a basic process. The basic process generates an x , then decides whether it is acceptable or not. The loop keeps invoking the basic process until it gets an acceptable x . Now

Listing 8.1: Matlab code for simple rejection sampling; this is inefficient, but simple to follow.

```

function rnum=rejectsample (pvec)
%
% pvec is a probability distribution over numbers
% 1, ..., size(pvec, 1)
%
nv=size(pvec, 1);
done=0;
while done==0
    ptr=floor(1+(nv-1e-10)*rand);
    % this gives me a uniform random
    % number in the range 1, .. nv
    pval=pvec(ptr);
    if rand<pval
        done=1;
        % i.e. accept
    end
end
rnum=ptr;

```

the probability that the basic process produces the value x_i , decides it is acceptable, and reports it is:

$$\begin{aligned}
 P(\{\text{report } x_i \text{ acceptable}\}) &= \left(\begin{array}{c} P(\{\text{accept } x_i\} | \{\text{generate } x_i\}) \\ \times \\ P(\{\text{generate } x_i\}) \end{array} \right) \\
 &= p(x_i) \times \frac{1}{N}.
 \end{aligned}$$

Now the loop keeps going until the basic process obtains an x_i that it decides is acceptable. This means the probability that the *loop* reports x_i is proportional to $p(x_i)$. But since $\sum_i p(x_i) = 1$, the probability that the *loop* reports x_i is *equal* to $p(x_i)$.

The problem here is that the basic process may not come up with an acceptable value x_i the first time; the loop might have to go on multiple times. If there are many x with small values of $p(x)$, we may find that it takes a very long time indeed to come up with a single sample. In the worst case, all values of $p(x)$ will be small. For example, for a uniform distribution on the range $1, \dots, N$, $p(x)$ is $1/N$.

We can make things more efficient by noticing that multiplying the probability distribution by a constant doesn't change the relative frequencies with which numbers are selected. So this means that we can find the largest value \hat{p} of $p(x)$, and form $q(x) = (1/\hat{p})p(x)$. Our process then becomes that shown in algorithm 8.2.

This process is not the most efficient available.

***** tree and point location algorithm

Listing 8.2: Matlab code for simple rejection sampling; this is somewhat more efficient.

```

function rnum=fasterrejectsample(pvec)
%
% pvec is a probability distribution over numbers
% 1, ..., size(pvec, 1)
%
nv=size(pvec, 1);
wv=max(pvec);
pv2=pvec/wv; % we rescale
done=0;
while done==0
    ptr=floor(1+(nv-1e-10)*rand);
    % this gives me a uniform random
    % number in the range 1, .. nv
    pval=pv2(ptr); % work with rescaled probs
    if rand<pval
        done=1;
        % i.e. accept
    end
end
rnum=ptr;

```

8.3 SIMULATION EXAMPLES

Computing probabilities and expectations from simulations should be so natural to a computer science student that it can be hard to see the magic. You estimate the probability of an event \mathcal{E} by writing a program that runs N independent simulations of an experiment, counts how many times the event occurs (which we write $\#(\mathcal{E})$), and reports

$$\frac{\#(\mathcal{E})}{N}$$

as an estimate of $P(\mathcal{E})$. This estimate will not be exact, and may be different for different runs of the program. It's often a good, simple estimate.

Choosing N depends on a lot of considerations. Generally, a larger N means a more accurate answer, and also a slower program. If N is too small, you may find that you report 1 or 0 for the probability (think of what would happen if you measured $P(H)$ for a coin with one flip). One strategy is to run several simulations, report the mean, and use the standard deviation as some guide to the accuracy.

8.3.1 Simulating Experiments

Worked example 8.11 *Getting 14's with 20-sided dice*

You throw 3 fair 20-sided dice. Estimate the probability that the sum of the faces is 14 using a simulation. Use $N = [1e1, 1e2, 1e3, 1e4, 1e5, 1e6]$. Which estimate is likely to be more accurate, and why?

Solution: You need a fairly fast computer, or this will take a long time. I ran ten versions of each experiment for $N = [1e1, 1e2, 1e3, 1e4, 1e5, 1e6]$, yielding ten probability estimates for each N . These were different for each version of the experiment, because the simulations are random. I got means of $[0, 0.0030, 0.0096, 0.0100, 0.0096, 0.0098]$, and standard deviations of $[0.0067, 0.0033, 0.0090, 0.0020, 0.0001]$. This suggests the true value is around 0.0098, and the estimate from $N = 1e6$ is best. The reason that the estimate with $N = 1e1$ is 0 is that the probability is very small, so you don't usually observe this case at all in only ten trials.

Worked example 8.12 *Comparing simulation with computation*

You throw 3 fair six-sided dice. You wish to know the probability the sum is 3. Compare the true value of this probability with estimates from six runs of a simulation using $N = 10000$. What conclusions do you draw?

Solution: I ran six simulations with $N = 10000$, and got $[0.0038, 0.0038, 0.0053, 0.0041, 0.0056, 0.0049]$. The mean is 0.00458, and the standard deviation is 0.0007, which suggests the estimate isn't that great, but the right answer should be in the range $[0.00388, 0.00528]$ with high probability. The true value is $1/216 \approx 0.00463$. The estimate is tolerable, but not super accurate.

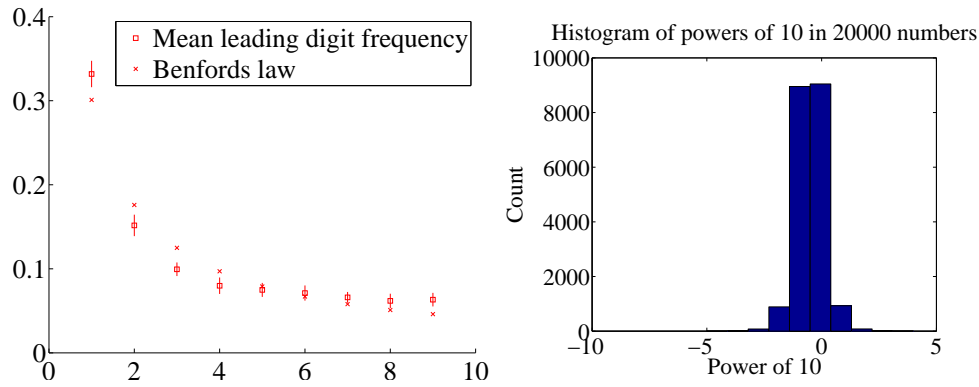


FIGURE 8.6: **Left** summarizes results of a simulation where, for 20 runs, I generated 1000 pairs of independent random numbers uniformly distributed in $[0 - 1]$ and divided one by the other. The boxes show the mean over 20 runs, and the vertical bars show one standard deviation up and down. The crosses are predictions from Benford's law. Numbers generated in this way have a wide range of orders magnitude, so a conventional histogram isn't easy to plot. On the **Right**, a histogram of the first digit of the logarithm (i.e. the power of 10, or order of magnitude) of these numbers. Notice the smallest number is eight orders of magnitude smaller than the biggest.

Worked example 8.13 *The First Digit of a Random Number*

We have seen (section ??) that adding random variables tends to produce a normal random variable. What happens if one divides one random number by another? Solving this in closed form is tricky, but simulation gives some insight.

Solution: I wrote a short program that produced 20 experiments, each consisting of 1000 numbers. I obtained the numbers by generating two uniform random numbers in the range $[0 - 1]$, then dividing one by the other. Notice that doing so can produce both very big and very small numbers, so that plotting a histogram is tricky — most boxes will be empty unless we produce an immense quantity of data. Instead, I plot a histogram of the leading digit of the number in Figure 8.6.

Figure 8.6 also shows the mean over all 20 experiments of the fraction of leading digits that is one, etc. together with one standard deviation error bars. The figure also shows a histogram of the rounded log (i.e. the power of 10) to give some idea of the range of values one obtains. It turns out that there is a widely applicable law, called **Benford's law**, that predicts the frequency of the first digit of many types of random number. The main condition for Benford's law to apply is that the numbers cover a wide range of scales. The figure shows predictions from

Benford's law as well.

8.3.2 Simulating Markov Chains

Worked example 8.14 *Coin Flips with End Conditions*

I flip a coin repeatedly until I encounter a sequence HTHT, at which point I stop. What is the probability that I flip the coin nine times?

Solution: You might well be able to construct a closed form solution to this if you follow the details of example 1 and do quite a lot of extra work. A simulation is really straightforward to write; notice you can save time by not continuing to simulate coin flips once you've flipped past nine times. I got 0.0411 as the mean probability over 10 runs of a simulation of 1000 experiments each, with a standard deviation of 0.0056.

Worked example 8.15 *A Queue*

A bus is supposed to arrive at a bus stop every hour for 10 hours each day. The number of people who arrive to queue at the bus stop each hour has a Poisson distribution, with intensity 4. If the bus stops, everyone gets on the bus and the number of people in the queue becomes zero. However, with probability 0.1 the bus driver decides not to stop, in which case people decide to wait. If the queue is ever longer than 15, the waiting passengers will riot (and then immediately get dragged off by the police, so the queue length goes down to zero). What is the expected time between riots?

Solution: I'm not sure whether one could come up with a closed form solution to this problem. A simulation is completely straightforward to write. I get a mean time of 441 hours between riots, with a standard deviation of 391. It's interesting to play around with the parameters of this problem; a less conscientious bus driver, or a higher intensity arrival distribution, lead to much more regular riots.

Worked example 8.16 *Inventory*

A store needs to control its stock of an item. It can order stocks on Friday evenings, which will be delivered on Monday mornings. The store is old-fashioned, and open only on weekdays. On each weekday, a random number of customers comes in to buy the item. This number has a Poisson distribution, with intensity 4. If the item is present, the customer buys it, and the store makes \$100; otherwise, the customer leaves. Each evening at closing, the store loses \$10 for each unsold item on its shelves. The store's supplier insists that it order a fixed number k of items (i.e. the store must order k items each week). The store opens on a Monday with 20 items on the shelf. What k should the store use to maximise profits?

Solution: I'm not sure whether one could come up with a closed form solution to this problem, either. A simulation is completely straightforward to write. To choose k , you run the simulation with different k values to see what happens. I computed accumulated profits over 100 weeks for different k values, then ran the simulation five times to see which k was predicted. Results were 21, 19, 23, 20, 21. I'd choose 21 based on this information.

For example 16, you should plot accumulated profits. If k is small, the store doesn't lose money by storing items, but it doesn't sell as much stuff as it could; if k is large, then it can fill any order but it loses money by having stock on the shelves. A little thought will convince you that k should be near 20, because that is the expected number of customers each week, so $k = 20$ means the store can expect to sell all its new stock. It may not be exactly 20, because it must depend a little on the balance between the profit in selling an item and the cost of storing it. For example, if the cost of storing items is very small compared to the profit, a very large k might be a good choice. If the cost of storage is sufficiently high, it might be better to never have anything on the shelves; this point explains the absence of small stores selling PC's.

8.3.3 Example: Ranking the Web by Simulating a Markov Chain

Perhaps the most valuable technical question of the last thirty years has been: Which web pages are interesting? Some idea of the importance of this question is that it was only really asked about 20 years ago, and at least one gigantic technology company has been spawned by a partial answer. This answer, due to Larry Page and Sergey Brin, and widely known as PageRank, starts with a Markov chain.

Figure 8.7 shows a picture of (a very small fraction of) the world wide web. I have drawn the web using a visual metaphor that should strongly suggest a finite state machine, and so a Markov chain. Each page is a state. Directed edges from page to page represent links. I count only the first link from a page to another page. Some pages are linked, others are not. I want to know how important each page is.

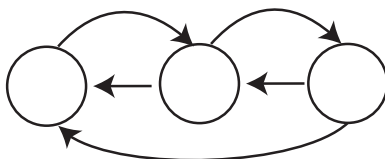


FIGURE 8.7: A very small fraction of the web, drawn to suggest a finite state machine; each state represents a page, and each directed edge represents an outgoing link. A random web surfer could either (a) follow an outgoing link, chosen at random or (b) type in the URL of any page, chosen at random. Such a surfer would see lots of pages that have many incoming links from pages that have lots of incoming links, and so on. Pages like this are likely important, so that finding important pages is analogous to simulating a random web surfer.

One way to think about importance is to think about what a random web surfer would do. The surfer can either (a) choose one of the outgoing links on a page at random, and follow it or (b) type in the URL of a new page, and go to that instead. As Figure 8.7 suggests, it is useful to think of this as a random walk on a finite state machine. We expect that this random surfer should see a lot of pages that have lots of incoming links from other pages that have lots of incoming links that (and so on). These pages are important, because lots of pages have linked to them.

For the moment, ignore the surfer's option to type in a URL. Write $r(i)$ for the importance of the i 'th page. We model importance as leaking from page to page across outgoing links (the same way the surfer jumps). Page i receives importance down each incoming link. The amount of importance is proportional to the amount of importance at the other end of the link, and inversely proportional to the number of links leaving that page. So a page with only one outgoing link transfers all its importance down that link; and the way for a page to receive a lot of importance is for it to have a lot of important pages link to it alone. We write

$$r(j) = \sum_{i \rightarrow j} \frac{r(i)}{|i|}$$

where $|i|$ means the total number of links pointing *out* of page i . We can stack the $r(j)$ values into a *row* vector \mathbf{r} , and construct a matrix \mathcal{P} , where

$$p_{ij} = \begin{cases} \frac{1}{|i|} & \text{if } i \text{ points to } j \\ 0 & \text{otherwise} \end{cases}$$

With this notation, the importance vector has the property

$$\mathbf{r} = \mathbf{r}\mathcal{P}$$

and should look a bit like the stationary distribution of a random walk to you, except that \mathcal{P} isn't stochastic — there may be some rows where the row sum of \mathcal{P} is zero, because there are *no* outgoing links from that page. We can fix this easily

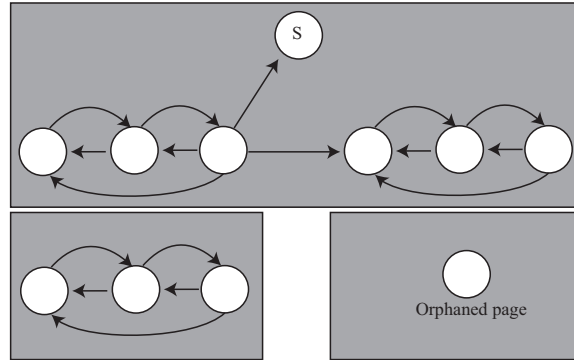


FIGURE 8.8: *The web isn't really like Figure 8.7. It's more like this figure (only bigger). In this figure, we see sections of the web that can't be reached by following links from other sections (the gray boxes); orphaned pages; and pages where a random walker would get stuck (S). Any algorithm for assigning importance must be able to deal with these effects.*

by replacing each row that sums to zero with $(1/n)\mathbf{1}$, where n is the total number of pages. Call the resulting matrix \mathcal{G} (it's quite often called the **raw Google matrix**). Notice that doing this doesn't really change anything significant; pages with no outgoing links leak a tiny fraction of their importance to every other page.

Figure 8.7 isn't a particularly good model of the web (and not because it's tiny, either). Figure 8.8 shows some of the problems that occur. There are pages with no outgoing links (which we've dealt with), pages with no incoming links, and even pages with no links at all. Worse, a random walk can get trapped (in one of the gray boxes). One way to fix all this would be to construct a process that wandered around the web inserting links that clean up the structure of the graph. This strategy is completely infeasible, because the real web is much too big. Allowing the surfer to randomly enter a URL sorts out all of these problems, because it inserts an edge of small weight from every node to every other node. Now the random walk cannot get trapped.

There are a variety of possible choices for the weight of these inserted edges. The original choice was to make each inserted edge have the same weight. Write $\mathbf{1}$ for the n dimensional column vector containing a 1 in each component, and let $0 < \alpha < 1$. We can write the matrix of transition probabilities as

$$\mathcal{G}(\alpha) = \alpha \frac{(\mathbf{1}\mathbf{1}^T)}{n} + (1 - \alpha)\mathcal{G}$$

where \mathcal{G} is the original Google matrix. An alternative choice is to choose a weight for each web page, using anything from advertising revenues to page visit statistics to thaumaturgy (Google keeps quiet about the details). Write this weight vector \mathbf{v} , and require that $\mathbf{1}^T \mathbf{v} = 1$ (i.e. the coefficients sum to one). Then we could have

$$\mathcal{G}(\alpha, \mathbf{v}) = \alpha \frac{(\mathbf{1}\mathbf{v}^T)}{n} + (1 - \alpha)\mathcal{G}.$$

Now the importance vector \mathbf{r} is the (unique, though I won't prove this) *row* vector \mathbf{r} such that

$$\mathbf{r} = \mathbf{r}\mathcal{G}(\alpha, \mathbf{v}).$$

How do we compute this vector? One natural algorithm is to start with some initial estimate, and propagate it. We write $\mathbf{r}^{(k)}$ for the estimated importance after the k 'th step. We define updates by

$$(\mathbf{r}^{(k)}) = (\mathbf{r}^{(k-1)})\mathcal{G}(\alpha, \mathbf{v}).$$

We can't compute this directly, either, because $\mathcal{G}(\alpha, \mathbf{v})$ is unreasonably big so (a) we can't form or store $\mathcal{G}(\alpha, \mathbf{v})$ and (b) we can't multiply by it either. But we could estimate \mathbf{r} with a random walk, because \mathbf{r} is the stationary distribution of a Markov chain. If we simulate this walk for many steps, the probability that the simulation is in state j should be $r(j)$, at least approximately.

This simulation is easy to build. Imagine our random walking bug sits on a web page. At each time step, it transitions to a new page by either (a) picking from all existing pages at random, using \mathbf{v} as a probability distribution on the pages (which it does with probability α); or (b) chooses one of the outgoing links uniformly and at random, and follows it (which it does with probability $1 - \alpha$). The stationary distribution of this random walk is \mathbf{r} . Another fact that I shall not prove is that, when α is sufficiently large, this random walk very quickly “forgets” its initial distribution. As a result, you can estimate the importance of web pages by starting this random walk in a random location; letting it run for a bit; then stopping it, and collecting the page you stopped on. The pages you see like this are independent, identically distributed samples from \mathbf{r} ; so the ones you see more often are more important, and the ones you see less often are less important.

8.3.4 Example: Simulating a Complicated Game

I will build several examples around a highly simplified version of a real card game. This game is Magic: The Gathering, and is protected by a variety of trademarks, etc. My version — MTGDAF — isn't very interesting as a game, but is simple enough to be studied, and interesting enough it casts some light on the real game. The game is played with decks of 60 cards. There are two types of card: Lands, and Spells. Lands can be placed on the play table and stay there permanently; Spells are played and then disappear. A Land on the table can be “tapped” or “untapped”. Players take turns. Each player draws a hand of seven cards from a shuffled deck. In each turn, a player first untaps any Lands on the table, then draws a card, then plays a land onto the table (if the player has one in hand to play), then finally can play one or more spells. Each spell has a fixed cost (of $1, \dots, 10$), and this cost is played by “tapping” a land (which is not untapped until the start of the next turn). This means that the player can cast only cheap spells in the early turns of the game, and expensive spells in the later turns.

Worked example 8.17 *MTGDFAF — The number of lands*

Assume a deck of 60 cards has 24 Lands. It is properly shuffled, and you draw seven cards. You could draw $0, \dots, 7$ Lands. Estimate the probability for each, using a simulation. Furthermore, estimate the error in your estimates.

Solution: The matlab function `randperm` produces a random permutation of given length. This means you can use it to simulate a shuffle of a deck, as in listing 8.3. I then drew 10, 000 random hands of seven cards, and counted how many times I got each number. Finally, to get an estimate of the error, I repeated this experiment 10 times and computed the standard deviation of each estimate of probability. This produced

0.0218 0.1215 0.2706 0.3082 0.1956 0.0686 0.0125 0.0012

for the probabilities (for 0 to 7, increasing number of lands to the right) and

0.0015 0.0037 0.0039 0.0058 0.0027 0.0032 0.0005 0.0004

for the standard deviations of these estimates.

Worked example 8.18 *MTGDFAF — The number of lands*

What happens to the probability of getting different numbers of lands if you put only 15 Lands in a deck of 60? It is properly shuffled, and you draw seven cards. You could draw $0, \dots, 7$ Lands. Estimate the probability for each, using a simulation. Furthermore, estimate the error in your estimates.

Solution: You can change one line in the listing to get

0.1159 0.3215 0.3308 0.1749 0.0489 0.0075 0.0006 0.0000

for the probabilities (for 0 to 7, increasing number of lands to the right) and

0.0034 0.0050 0.0054 0.0047 0.0019 0.0006 0.0003 0.0000

for the standard deviations of these estimates.

Listing 8.3: Matlab code used to simulate the number of lands

```

simcards=[ones(24, 1); zeros(36, 1)]
% 1 if land, 0 otherwise
ninsim=10000;
nsims=10;
counts=zeros(nsims, 8);
for i=1:10
    for j=1:10000
        shuffle=randperm(60);
        hand=simcards(shuffle(1:7));
        %useful matlab trick here
        nlands=sum(hand);
        %ie number of lands
        counts(i, 1+nlands)=...
            counts(i, 1+nlands)+1;
        % number of lands could be zero
    end
end
probs=counts/ninsim;
mean(probs)
std(probs)
%%

```

Worked example 8.19 *MTGDFAF — Playing spells*

Assume you have a deck of 24 Lands, 10 Spells of cost 1, 10 Spells of cost 2, 10 Spells of cost 3, 2 Spells of cost 4, 2 Spells of cost 5, and 2 Spells of cost 6. Assume you always only play the cheapest spell in your hand (i.e. you never play two spells). What is the probability you will be able to play at least one spell on each of the first four turns?

Solution: This simulation requires just a little more care. You draw the hand, then simulate the first four turns. In each turn, you can only play a spell whose cost you can pay, and only if you have it. I used the matlab of listing 8.4 and listing 8.5; I found the probability to be 0.64 with standard deviation 0.01. Of course, my code might be wrong...

Worked example 8.20 *MTGDFAF — Playing spells*

Now we use a different distribution of cards. Assume you have a deck of 20 Lands, 9 Spells of cost 1, 5 Spells of cost 2, 5 Spells of cost 3, 5 Spells of cost 4, 5 Spells of cost 5, and 11 Spells of cost 6. Assume you always only play the cheapest spell in your hand (i.e. you never play two spells). What is the probability you will be able to play at least one spell on each of the first four turns?

Solution: This simulation requires just a little more care. You draw the hand, then simulate the first four turns. In each turn, you can only play a spell whose cost you can pay, and only if you have it. I found the probability to be 0.33 with standard deviation 0.05. Of course, my code might be wrong....

One engaging feature of the real game that is revealed by these very simple simulations is the tension between a player's goals. The player would like to have few lands — so as to have lots of spells — but doing so means that there's a bigger chance of not being able to play a spell. Similarly, a player would like to have lots of powerful (=expensive) spells, but doing so means there's a bigger chance of not being able to play a spell. Players of the real game spend baffling amounts of time arguing with one another about the appropriate choices for a good set of cards.

Worked example 8.21 *MTGDFAF — How long before you can play a spell of cost 3?*

Assume you have a deck of 15 Lands, 15 Spells of cost 1, 14 Spells of cost 2, 10 Spells of cost 3, 2 Spells of cost 4, 2 Spells of cost 5, and 2 Spells of cost 6. What is the expected number of turns before you can play a spell of cost 3? Assume you always play a land if you can.

Solution: I get 6.3, with a standard deviation of 0.1. The problem is it can take quite a large number of turns to get three lands out. I used the code of listings 8.6 and 8.7

Recall that you can reasonably expect that the probability you compute from a simulation behaves like normal data. Run a simulation experiment a large number of times and construct a data set whose entries are the probability from each experiment. This data should be normal. This means that, if you subtract the mean and divide by the standard deviation, you should get a histogram that looks like the standard normal curve. Figure 8.9 shows some examples of this effect. This effect is important, because it means that the right answer should be very few standard deviations away from the mean you compute — the standard deviation gives you quite a good idea of the accuracy of your estimate.

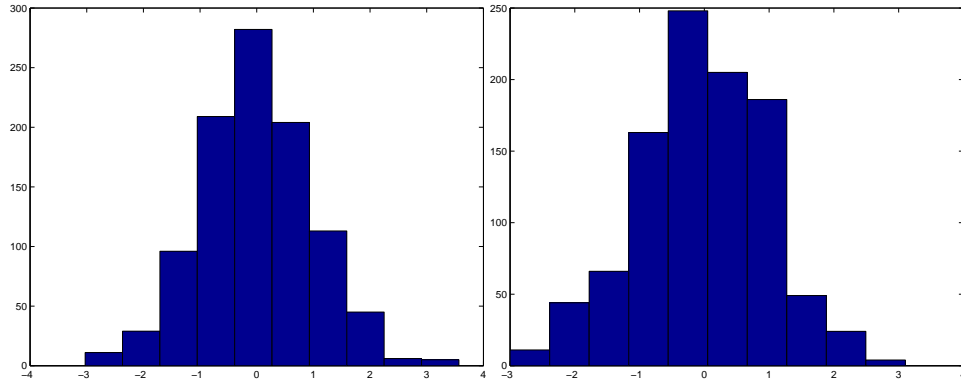


FIGURE 8.9: Estimates of probabilities produced by simulation typically behave like normal data. On the **left**, I show a histogram of probabilities of having a hand of 3 Lands in the simulation of example 17; these are plotted in standard coordinates. On the **right**, I show a histogram of probability of playing a spell in each of the first four turns (example 20), from 1000 simulation experiments; again, these are plotted in standard coordinates. Compare these to the standard normal histograms of the previous chapter.

PROBLEMS

8.1. Multiple coin flips:

- For example ??, show that $P(5) = (3/32)$ by directly writing out the sequences of flips, and summing their probabilities.
- Now check the recurrence relation $P(N) = (1/2)P(N-1) + (1/4)P(N-2)$ for the case $N = 5$.
- What is $P(7)$?

8.2. Multiple die rolls:

You roll a fair die until you see a 5, then a 6; after that, you stop. Write $P(N)$ for the probability that you roll the die N times.

- What is $P(1)$?
- Show that $P(2) = (1/36)$.
- Draw a finite state machine encoding all the sequences of die rolls that you could encounter. Don't write the events on the edges; instead, write their probabilities. There are 5 ways not to get a 5, but only one probability, so this simplifies the drawing.
- Show that $P(3) = (1/36)$.
- Now use your finite state machine to argue that $P(N) = (5/6)P(N-1) + (25/36)P(N-2)$.

8.3. More complicated multiple coin flips:

You flip a fair coin until you see either HTH or THT , and then you stop. We will compute a recurrence relation for $P(N)$.

- Figure ?? shows a finite state machine. Check that this finite state machine represents all coin sequences that you will encounter.
- Write Σ_N for some string of length N accepted by this finite state machine. Use this finite state machine to argue that Σ_N has one of four forms:

- $T\Sigma_{N-2}$

2. $HH\Sigma_{N-3}$
3. $THH\Sigma_{N-2}$
4. $HTT\Sigma_{N-3}$

(c) Now use this argument to show that $P(N) = (1/2)P(N-2) + (1/4)P(N-3)$.

8.4. The gambler's ruin:

(a) Show that you can rearrange the recurrence relation of example 1 to get

$$p_{s+1} - p_s = \frac{(1-p)}{p} (p_s - p_{s-1}).$$

Now show that this means that

$$p_{s+1} - p_s = \left(\frac{(1-p)}{p} \right)^2 (p_{s-1} - p_{s-2})$$

so that

$$\begin{aligned} p_{s+1} - p_s &= \left(\frac{(1-p)}{p} \right)^s (p_1 - p_0) \\ &= \left(\frac{(1-p)}{p} \right)^s (p_1 - 1). \end{aligned}$$

(b) Now we need a simple result about series. Assume I have a series u_k , $k \geq 0$, with the property that

$$u_k - u_{k-1} = cr^{k-1}.$$

Show that

$$u_k - u_0 = c \left(\frac{r^k - 1}{r - 1} \right).$$

(c) Use the results of the last two steps to show that

$$p_s - 1 = (p_1 - 1) \left(\frac{\left(\frac{(1-p)}{p} \right)^s - 1}{\left(\frac{(1-p)}{p} \right) - 1} \right)$$

(d) Use the fact that $p_j = 0$ and the result of the last exercise to show

$$(p_1 - 1) = \frac{-1}{\left(\frac{\left(\frac{(1-p)}{p} \right)^j - 1}{\left(\frac{(1-p)}{p} \right) - 1} \right)}$$

(e) Use the results of the previous exercises to show that

$$p_s = \frac{\left(\frac{(1-p)}{p} \right)^j - \left(\frac{(1-p)}{p} \right)^s}{\left(\frac{(1-p)}{p} \right)^j - 1}.$$

Listing 8.4: Matlab code used to simulate the four turns

```

simcards=[zeros(24, 1); ones(10, 1);...
          2*ones(10, 1);3*ones(10, 1); ...
          4*ones(2, 1); 5*ones(2, 1); 6*ones(2, 1)];
nsims=10;
ninsim=1000;
counts=zeros(nsims, 1);
for i=1:nsims
    for j=1:ninsim
        % draw a hand
        shuffle=randperm(60);
        hand=simcards(shuffle(1:7));
        %reorganize the hand
        cleanhand=zeros(7, 1);
        for k=1:7
            cleanhand(hand(k)+1)=cleanhand(hand(k)+1)+1;
            % ie count of lands, spells, by cost
        end
        landsontable=0;
        [playedspell1, landsontable, cleanhand]=...
            playground(landsontable, cleanhand, shuffle, ...
                simcards, 1);
        [playedspell2, landsontable, cleanhand]=...
            playground(landsontable, cleanhand, shuffle, ...
                simcards, 2);
        [playedspell3, landsontable, cleanhand]=...
            playground(landsontable, cleanhand, shuffle, ...
                simcards, 3);
        [playedspell4, landsontable, cleanhand]=...
            playground(landsontable, cleanhand, shuffle, ...
                simcards, 4);
        counts(i)=counts(i)+...
            playedspell1*playedspell2*...
            playedspell3*playedspell4;
    end
end
end

```

Listing 8.5: Matlab code used to simulate playing a turn

```

function [playedspell , landsontable , cleanhand]=...
    playground(landsontable , cleanhand , shuffle , simcards , ...
    turn)
    % draw
ncard=simcards( shuffle(7+turn) );
cleanhand( ncard+1)=cleanhand( ncard+1)+1;
% play land
if cleanhand(1)>0
    landsontable=landsontable+1;
    cleanhand(1)=cleanhand(1)-1;
end
playedspell=0;
if landsontable>0
    i=1; done=0;
    while done==0
        if cleanhand(i)>0
            cleanhand(i)=cleanhand(i)-1;
            playedspell=1;
            done=1;
        else
            i=i+1;
            if i>landsontable
                done=1;
            end
        end
    end
end
end

```

Listing 8.6: Matlab code used to estimate number of turns before you can play a spell of cost 3

```

simcards=[zeros(15, 1); ones(15, 1);...
          2*ones(14, 1);3*ones(10, 1); ...
          4*ones(2, 1); 5*ones(2, 1); 6*ones(2, 1)];
nsims=10;
ninsim=1000;
counts=zeros(nsims, 1);
for i=1:nsims
    for j=1:ninsim
        % draw a hand
        shuffle=randperm(60);
        hand=simcards( shuffle(1:7));
        %reorganize the hand
        cleanhand=zeros(7, 1);
        for k=1:7
            cleanhand(hand(k)+1)=cleanhand(hand(k)+1)+1;
            % ie count of lands, spells, by cost
        end
        landsontable=0;
        k=0; played3spell=0;
        while played3spell==0;
            [played3spell, landsontable, cleanhand]=...
                play3round(landsontable, cleanhand, shuffle, ...
                    simcards, k+1);
            k=k+1;
        end
        counts(i)=counts(i)+k;
    end
    counts(i)=counts(i)/ninsim;
end

```

Listing 8.7: Matlab code used to simulate a turn to estimate the number of turns before you can play a spell of cost 3

```
function [played3spell, landsontable, cleanhand]=...
    play3round(landsontable, cleanhand, shuffle, simcards, . . .
    turn)
    % draw
ncard=simcards(shuffle(7+turn));
cleanhand(ncard+1)=cleanhand(ncard+1)+1;
% play land
if cleanhand(1)>0
    landsontable=landsontable+1;
    cleanhand(1)=cleanhand(1)-1;
end
played3spell=0;
if (landsontable>=3)&&(cleanhand(4)>0)
    played3spell=1;
end
```


Inference: Making Point Estimates

Inference is the process of drawing conclusions from data. One form of inference is to estimate a number, or set of numbers, that describes a dataset. The result is known as a **point estimate**. An alternative is to estimate an interval within which a number lies, with some degree of certainty. Such estimates are known as **interval estimates**. Finally, one might wish to assess the extent to which a body of evidence supports rejecting an hypothesis — known as **hypothesis testing**. In this chapter, we deal with point estimates. In the following chapter, we deal with interval estimates and hypothesis testing, which require an understanding of point estimates. There are two, somewhat distinct, situations in which we could make point estimates.

In the first, we have a dataset $\{\mathbf{x}\}$, and a probability model we believe applies to that dataset. But we need to select appropriate values of the parameters to ensure that the model describes the data. For example, we might have a set of N coin flips which we believe to be independent and identically distributed. Of these, k flips came up H . We know that a binomial distribution with $p(H) = p$ is a good model — but what value of p should we use? Your intuition is likely to suggest using k/N , but we'd like a more robust procedure than guessing. We need an inference procedure to obtain the unknown parameter from the data. Notice that this will be an estimate, rather than the “true” value. As we shall see, there is more than one possible procedure to apply, depending to some extent on the problem. In some cases (section 9.1), we estimate parameter values based solely on data; in others (section 9.2), we are able to use prior information about the parameters to affect the estimate.

In the second situation, we want to know some property of a population. For example, we may wish to know the mean weight of a person, or the mean response of a mouse to a drug. It would be a stretch to describe this population with one of the probability models that we have seen. In principle, the number we want is not even necessarily random; in principle, we could measure everyone on the planet and average the weights. In practice, this doesn't make sense, for quite straightforward reasons. You can't really weigh every person, or dose every mouse, on the planet. Instead, to estimate this property, we obtain a sample (some people; some mice; etc.) of the population, and estimate the property from the sample. There is now an important problem. Different samples lead to different estimates of the property. We will arrange to have the sample drawn randomly from the population, so the sample we see represents the value of a set of random variables. If you followed the proof of the weak law of large numbers, you should suspect that the mean of this sample could be a good estimate of the population mean. This turns out to be the case (section 1). However, there is some random error in the estimate, and we can tell (on average) how large the error caused by random sampling could be.

9.1 ESTIMATING MODEL PARAMETERS WITH MAXIMUM LIKELIHOOD

Assume we have a dataset $\mathcal{D} = \{\mathbf{x}\}$, and a probability model we believe applies to that dataset. Generally, application logic suggests the type of model (i.e. normal probability density; Poisson probability; geometric probability; and so on). But usually, we do not know the parameters of the model — for example, the mean and standard deviation of a normal distribution; the intensity of a poisson distribution; and so on. Our model will be better or worse depending on how well we choose the parameters. We need a strategy to estimate the parameters of a model from a sample dataset. Notice how each of the following examples fits this pattern.

Example: 9.1 *Inferring p from repeated flips — binomial*

We could flip the coin N times, and count the number of heads k . We know that an appropriate probability model for a set of independent coin flips is the binomial model $P(k; N, p)$. But we do not know p , which is the parameter — we need a strategy to extract a value of p from the data.

Example: 9.2 *Inferring p from repeated flips — geometric*

We could flip the coin repeatedly until we see a head. We know that, in this case, the number of flips has the geometric distribution with parameter p . In this case, the data is a sequence of T 's with a final H from the coin flips. There are N flips (or terms) and the last flip is a head. We know that an appropriate probability model is the geometric distribution $P_g(N; p)$. But we do not know p , which is the parameter — we need a strategy to extract a value of p from the data.

Example: 9.3 *Inferring the intensity of spam — poisson*

It is reasonable to assume that the number of spam emails one gets in an hour has a Poisson distribution. But what is the intensity parameter λ ? We could count the number of spam emails that arrive in each of a set of distinct hours, giving a dataset of counts \mathcal{D} . We need a strategy to wrestle an estimate of λ from this dataset.

Example: 9.4 *Inferring the mean and standard deviation of normal data*

Imagine we know for some reason that our data is well described by a normal distribution. We could ask what is the mean and standard deviation of the normal distribution that best represents the data?

We can write that model as $P(\mathcal{D}|\theta)$, where θ are parameters of the probability mode. The model is conditioned on θ , because if we knew θ we could evaluate the model. The expression $P(\mathcal{D}|\theta)$ is known as the **likelihood** of the data, and is often written $\mathcal{L}(\theta)$ (or $\mathcal{L}(\theta; \mathcal{D})$ if you want to remember that data is involved). Notice that this is unlike our models to date. In chapter 1, we assumed that we knew θ , and could then use the model to assign a probability to a data item. Here we *know* the value of \mathcal{D} . The likelihood is a function of θ .

9.1.1 The Maximum Likelihood Principle

We need a “reasonable” procedure to choose a value of θ to report. One — and we stress this is not the only one — is the **maximum likelihood principle**. This says: Choose θ such that $\mathcal{L}(\theta) = P(\mathcal{D}|\theta)$ is maximised, as a function of θ .

For the examples we work with, the data will be **independent and identically distributed** or **IID**. This means that each data item is an independently obtained sample from the same probability distribution (see section 8.2.5). In turn, this means that the likelihood is a product of terms, one for each data item, which we can write as

$$\mathcal{L}(\theta) = P(\mathcal{D}|\theta) = \prod_{i \in \text{dataset}} P(d_i|\theta).$$

It is traditional to write θ for any set of parameters that are unknown. There are two, distinct, important concepts we must work with. One is the unknown parameter(s), which we will write θ . The other is the *estimate* of the value of that parameter, which we will write $\hat{\theta}$. This estimate is the best we can do — it may not be the “true” value of the parameter.

Worked example 9.1 *Inferring $p(H)$ for a coin from flips using a binomial model*

In N independent coin flips, you observe k heads. Use the maximum likelihood principle to infer $p(H)$.

Solution: The coin has $\theta = p(H)$, which is the unknown parameter. We know that an appropriate probability model is the binomial model $P(k; N, \theta)$. We have that

$$\mathcal{L}(\theta) = P(\mathcal{D}|\theta) = P_b(k; N, \theta) = \binom{N}{k} \theta^k (1 - \theta)^{(N-k)}$$

which is a function of θ — the unknown probability that a coin comes up heads; k and N are known. We must find the value of θ that maximizes this expression. Now the maximum occurs when

$$\frac{\partial \mathcal{L}(\theta)}{\partial \theta} = 0.$$

We have

$$\frac{\partial \mathcal{L}(\theta)}{\partial \theta} = \binom{N}{k} \left(k\theta^{k-1}(1 - \theta)^{(N-k)} - \theta^k(N - k)(1 - \theta)^{(N-k-1)} \right)$$

and this is zero when

$$k\theta^{k-1}(1 - \theta)^{(N-k)} = \theta^k(N - k)(1 - \theta)^{(N-k-1)}$$

so the maximum occurs when

$$k(1 - \theta) = \theta(N - k).$$

This means the maximum likelihood estimate is

$$\hat{\theta} = \frac{k}{N}$$

which is what we guessed would happen, but now we know why that guess “makes sense”.

Worked example 9.2 *Inferring $p(H)$ from coin flips using a geometric model*

You flip a coin N times, stopping when you see a head. Use the maximum likelihood principle to infer $p(H)$ for the coin.

Solution: The coin has $\theta = p(H)$, which is the unknown parameter. We know that an appropriate probability model is the geometric model $P_g(N; \theta)$. We have that

$$\mathcal{L}(\theta) = P(\mathcal{D}|\theta) = P_g(N; \theta) = (1 - \theta)^{(N-1)}\theta$$

which is a function of θ — the unknown probability that a coin comes up heads; N is known. We must find the value of θ that maximizes this expression. Now the maximum occurs when

$$\frac{\partial \mathcal{L}(\theta)}{\partial \theta} = 0 = ((1 - \theta)^{(N-1)} - (N - 1)(1 - \theta)^{(N-2)}\theta)$$

So the maximum likelihood estimate is

$$\hat{\theta} = \frac{1}{N}.$$

We didn't guess this.

Worked example 9.3 *Inferring die probabilities from multiple rolls and a multinomial distribution*

You throw a die N times, and see n_1 ones, ... and n_6 sixes. Write p_1, \dots, p_6 for the probabilities that the die comes up one, ..., six. Use the maximum likelihood principle to estimate p_1, \dots, p_6 .

Solution: The data are n, n_1, \dots, n_6 . The parameters are $\theta = (p_1, \dots, p_6)$. $P(\mathcal{D}|\theta)$ comes from the multinomial distribution. In particular,

$$\mathcal{L}(\theta) = P(\mathcal{D}|\theta) = \frac{n!}{n_1! \dots n_6!} p_1^{n_1} p_2^{n_2} \dots p_6^{n_6}$$

which is a function of $\theta = (p_1, \dots, p_6)$. Now we want to maximize this function by choice of θ . Notice that we could do this by simply making all p_i very large — but this omits a fact, which is that $p_1 + p_2 + p_3 + p_4 + p_5 + p_6 = 1$. So we substitute using $p_6 = 1 - p_1 - p_2 - p_3 - p_4 - p_5$ (there are other, neater, ways of dealing with this issue, but they take more background knowledge). At the maximum, we must have that for all i ,

$$\frac{\partial \mathcal{L}(\theta)}{\partial p_i} = 0$$

which means that, for p_i , we must have

$$n_i p_i^{(n_i-1)} (1 - p_1 - p_2 - p_3 - p_4 - p_5)^{n_6} - p_i^{n_i} n_6 (1 - p_1 - p_2 - p_3 - p_4 - p_5)^{(n_6-1)} = 0$$

so that, for each p_i , we have

$$n_i (1 - p_1 - p_2 - p_3 - p_4 - p_5) - n_6 p_i = 0$$

or

$$\frac{p_i}{1 - p_1 - p_2 - p_3 - p_4 - p_5} = \frac{n_i}{n_6}.$$

You can check that this equation is solved by

$$\hat{\theta} = \frac{1}{(n_1 + n_2 + n_3 + n_4 + n_5 + n_6)} (n_1, n_2, n_3, n_4, n_5, n_6)$$

The logarithm is a monotonic function (i.e. if $x > 0, y > 0, x > y$, then $\log(x) > \log(y)$). This means that the values of θ that maximise the log-likelihood are the same as the values that maximise the likelihood. This observation is very useful, because it allows us to transform a product into a sum. The derivative of a product involves numerous terms; the derivative of a sum is easy to take. We have

$$\log P(\mathcal{D}|\theta) = \log \prod_{i \in \text{dataset}} P(d_i|\theta) = \sum_{i \in \text{dataset}} \log P(d_i|\theta)$$

and in some cases, $\log P(d_i|\theta)$ takes a convenient, easy form. The log-likelihood of a dataset under a model is a function of the unknown parameters, and you will often see it written as

$$\log \mathcal{L}(\theta) = \sum_{i \in \text{dataset}} \log P(d_i|\theta).$$

Worked example 9.4 *Poisson distributions*

You observe N intervals, each of the same, fixed length (in time, or space). You know that, in these intervals, events occur with a Poisson distribution (for example, you might be observing Prussian officers being kicked by horses, or telemarketer calls...). You know also that the intensity of the Poisson distribution is the same for each observation. The number of events you observe in the i 'th interval is n_i . What is the intensity, λ ?

Solution: The likelihood is

$$\mathcal{L}(\theta) = \prod_{i \in \text{intervals}} P(\{n_i \text{ events}\}|\theta) = \prod_{i \in \text{intervals}} \frac{\theta^{n_i} e^{-\theta}}{n_i!}.$$

It will be easier to work with logs. The log-likelihood is

$$\log \mathcal{L}(\theta) = \sum_i (n_i \log \theta - \theta - \log n_i!)$$

so that we must solve

$$\frac{\partial \log \mathcal{L}(\theta)}{\partial \theta} = \sum_i \left(\frac{n_i}{\theta} - 1 \right) = 0$$

which yields a maximum likelihood estimate of

$$\hat{\theta} = \frac{\sum_i n_i}{N}$$

Worked example 9.5 *The intensity of swearing*

A famously swearsy politician gives a talk. You listen to the talk, and for each of 30 intervals 1 minute long, you record the number of swearwords. You record this as a histogram (i.e. you count the number of intervals with zero swear words, with one, etc.). For the first 10 intervals, you see

no. of swear words	no. of intervals
0	4
1	2
2	2
3	1
4	0

and for the following 20 intervals, you see

no. of swear words	no. of intervals
0	9
1	6
2	3
3	2
4	1

Assume that the politician's use of swearwords is Poisson. What is the intensity using the first 10 intervals? the second 20 intervals? all the intervals? why are they different?

Solution: Use the expression from worked example 4 to find

$$\begin{aligned}\hat{\lambda}_{10} &= \frac{\text{total number of swearwords}}{\text{number of intervals}} \\ &= \frac{7}{10}\end{aligned}$$

$$\begin{aligned}\hat{\lambda}_{20} &= \frac{\text{total number of swearwords}}{\text{number of intervals}} \\ &= \frac{22}{20}\end{aligned}$$

$$\begin{aligned}\hat{\lambda}_{30} &= \frac{\text{total number of swearwords}}{\text{number of intervals}} \\ &= \frac{29}{30}.\end{aligned}$$

These are different because the maximum likelihood estimate is an *estimate* — we can't expect to recover the exact value from a dataset. Notice, however, that the estimates are quite close.

Worked example 9.6 *Normal distributions*

Assume we have x_1, \dots, x_N which are data that can be modelled with a normal distribution. Use the maximum likelihood principle to estimate the mean of that normal distribution.

Solution: The likelihood of a set of data values under the normal distribution with unknown mean θ and standard deviation σ is

$$\begin{aligned}\mathcal{L}(\theta) &= P(x_1, \dots, x_N | \theta, \sigma) \\ &= P(x_1 | \theta, \sigma) P(x_2 | \theta, \sigma) \dots P(x_N | \theta, \sigma) \\ &= \prod_{i=1}^N \frac{1}{\sqrt{2\pi}\sigma} \exp\left(-\frac{(x_i - \theta)^2}{2\sigma^2}\right)\end{aligned}$$

and this expression is a moderate nuisance to work with. The log of the likelihood is

$$\log \mathcal{L}(\theta) = \left(\sum_{i=1}^N -\frac{(x_i - \theta)^2}{2\sigma^2} \right) + \text{term not depending on } \theta.$$

We can find the maximum by differentiating wrt θ and setting to zero, which yields

$$\begin{aligned}\frac{\partial \log \mathcal{L}(\theta)}{\partial \theta} &= \sum_{i=1}^N \frac{2(x_i - \theta)}{2\sigma^2} \\ &= 0 \\ &= \frac{1}{\sigma^2} \left(\sum_{i=1}^N x_i - N\theta \right)\end{aligned}$$

so the maximum likelihood estimate is

$$\hat{\theta} = \frac{\sum_{i=1}^N x_i}{N}$$

which probably isn't all that surprising. Notice we did not have to pay attention to σ in this derivation — we did not assume it was known, it just doesn't do anything.

Worked example 9.7 *Normal distributions -II*

Assume we have x_1, \dots, x_N which are data that can be modelled with a normal distribution. Use the maximum likelihood principle to estimate the standard deviation of that normal distribution.

Solution: Now we have to write out the log of the likelihood in more detail. Write μ for the mean of the normal distribution and θ for the unknown standard deviation of the normal distribution. We get

$$\log \mathcal{L}(\theta) = \left(\sum_{i=1}^N -\frac{(x_i - \mu)^2}{2\theta^2} \right) - N \log \theta + \text{Term not depending on } \theta$$

We can find the maximum by differentiating wrt σ and setting to zero, which yields

$$\frac{\partial \log \mathcal{L}(\theta)}{\partial \theta} = \frac{-2}{\theta^3} \sum_{i=1}^N \frac{-(x_i - \theta)}{2} - \frac{N}{\theta} = 0$$

so the maximum likelihood estimate is

$$\hat{\theta} = \sqrt{\frac{\sum_{i=1}^N (x_i - \mu)^2}{N}}$$

which probably isn't all that surprising, either.

The maximum likelihood principle has a variety of neat properties we cannot expound. One worth knowing about is **consistency**; for our purposes, this means that the maximum likelihood estimate of parameters can be made arbitrarily close to the right answer by having a sufficiently large dataset.

Another is that, in some cases, you can make online estimates. Assume, rather than seeing N elements of a dataset in one go, you get to see each one once, and you cannot store them. Assume that this dataset is modelled as normal data. Write $\hat{\mu}_k$ for the maximum likelihood estimate of the mean based on data items $1 \dots k$ (and $\hat{\sigma}_k$ for the maximum likelihood estimate of the standard deviation, etc.). Notice that

$$\hat{\mu}_{k+1} = \frac{(k\hat{\mu}_k) + x_{k+1}}{(k+1)}$$

and that

$$\hat{\sigma}_{k+1} = \sqrt{\frac{(k\hat{\sigma}_k^2) + (x_{k+1} - \hat{\mu}_{k+1})^2}{(k+1)}}$$

This means that you can incorporate new data into your estimate as it arrives without keeping all the data. This process of updating a representation of a dataset as new data arrives is known as **filtering**.

9.1.2 Cautions about Maximum Likelihood

Our examples suggest some difficulties could occur in inference. The first is that it might be hard to find the maximum of the likelihood exactly. There are strong numerical methods for maximizing functions, and these are very helpful, but even today there are likelihood functions where it is very hard to find the maximum.

The second is that small amounts of data can present nasty problems. There is a body of mathematics, well outside the scope of this book, that implies that for lots of data that is well described by our model, maximum likelihood will give an answer very close to the “right” answer. This doesn’t apply to small datasets. For example, in the binomial case, if we have only one flip we will estimate p as either 1 or 0. We should find this report unconvincing. In the geometric case, with a fair coin, there is a probability 0.5 that we will perform the estimate and then report that the coin has $p = 1$. This should also worry you. As another example, if we throw a die only a few times, we could reasonably expect that, for some i , $n_i = 0$. This doesn’t necessarily mean that $p_i = 0$, though that’s what the maximum likelihood inference procedure will tell us.

This creates a very important technical problem — how can I estimate the probability of events that haven’t occurred? This might seem like a slightly silly question to you, but it isn’t. Solving this problem has really significant practical consequences. For example, a really important part of natural language processing involves estimating the probability of groups of three words. These groups are usually known as “trigrams”. People typically know an awful lot of words (tens to hundreds of thousands, depending on what you mean by a word). This means that there are a tremendous number of trigrams, and you can expect that any real dataset lacks almost all of them, because it isn’t big enough. Some are missing because they don’t occur in real life, but others are not there simply because they are unusual (eg “Atom Heart Mother” actually occurs in real life, but you may not have seen it all that often). Modern speech recognition systems need to know how probable *every* trigram is. Worse, if a trigram is modelled as having zero probability and actually occurs, the system will make a mistake, so it is important to model all such events as having a very small, but not zero, probability.

In summary, the maximum likelihood estimate is useful, and is consistent with intuition, but small datasets present some worries because there is a real prospect that the best estimate is wrong in a way that presents problems.

9.2 INCORPORATING PRIORS WITH BAYESIAN INFERENCE

Sometimes when we wish to estimate parameters of a model we have prior information. For example, we may have good reason to believe that some parameter is close to some value. We would like to take this information into account when we estimate the model. One way to do so is to place a **prior probability distribution** $p(\theta)$ on the parameters θ . Then, rather than working with the likelihood $p(\mathcal{D}|\theta)$, we could apply Bayes’ rule, and form the **posterior** $p(\theta|\mathcal{D})$. This posterior represents the probability that θ takes various values, given the data \mathcal{D} . Extracting information from the posterior is usually called **Bayesian inference**. A natural estimate of θ is the value that maximizes the posterior. This estimate is sometimes known as a **maximum a priori estimate** or **MAP estimate**.

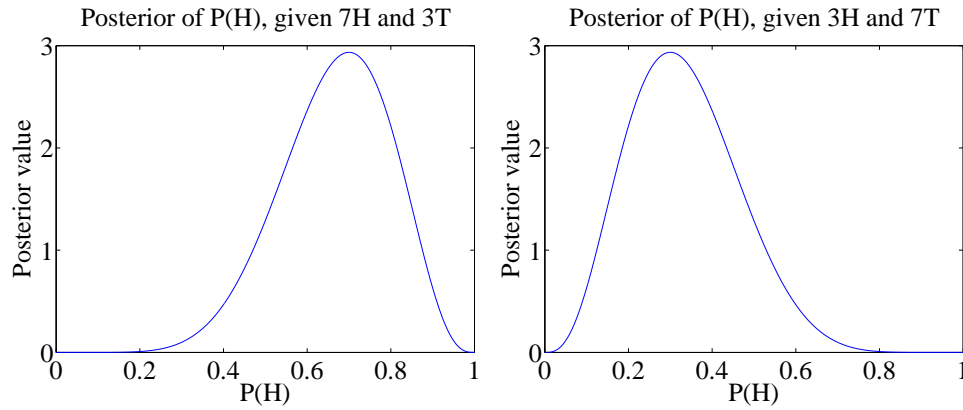


FIGURE 9.1: The curves show a function proportional to the posterior on θ , for the two cases of example ???. Notice that this information is rather richer than the single value we would get from maximum likelihood inference.

9.2.1 Constructing the Posterior

Bayes' rule tells us that

$$p(\theta|\mathcal{D}) = \frac{P(\mathcal{D}|\theta)P(\theta)}{P(\mathcal{D})}$$

but (as we shall see) it can be hard to work out $P(\mathcal{D})$. For some problems, we might not need to know it.

Worked example 9.8 *Flipping a coin*

We have a coin with probability θ of coming up heads when flipped. We start knowing nothing about θ . We then flip the coin 10 times, and see 7 heads (and 3 tails). Plot a function proportional to $p(\theta|\{7 \text{ heads and } 3 \text{ tails}\})$. What happens if there are 3 heads and 7 tails?

Solution: We know nothing about p , except that $0 \leq \theta \leq 1$, so we choose a uniform prior on p . We have that $p(\{7 \text{ heads and } 3 \text{ tails}\}|\theta)$ is binomial. The joint distribution is $p(\{7 \text{ heads and } 3 \text{ tails}\}|\theta) \times p(\theta)$ but $p(\theta)$ is uniform, so doesn't depend on θ . So the posterior is *proportional* to: $\theta^7(1-\theta)^3$ which is graphed in figure 9.1. The figure also shows $\theta^3(1-\theta)^7$ which is *proportional* to the posterior for 3 heads and 7 tails. In each case, the evidence does not rule out the possibility that $\theta = 0.5$, but tends to discourage the conclusion. Maximum likelihood would give $\theta = 0.7$ or $\theta = 0.3$, respectively.

In Example 8, it is interesting to follow how the posterior on p changes as evidence come in, which is easy to do because the posterior is proportional to a binomial distribution. Figure 9.2 shows a set of these posteriors for different sets

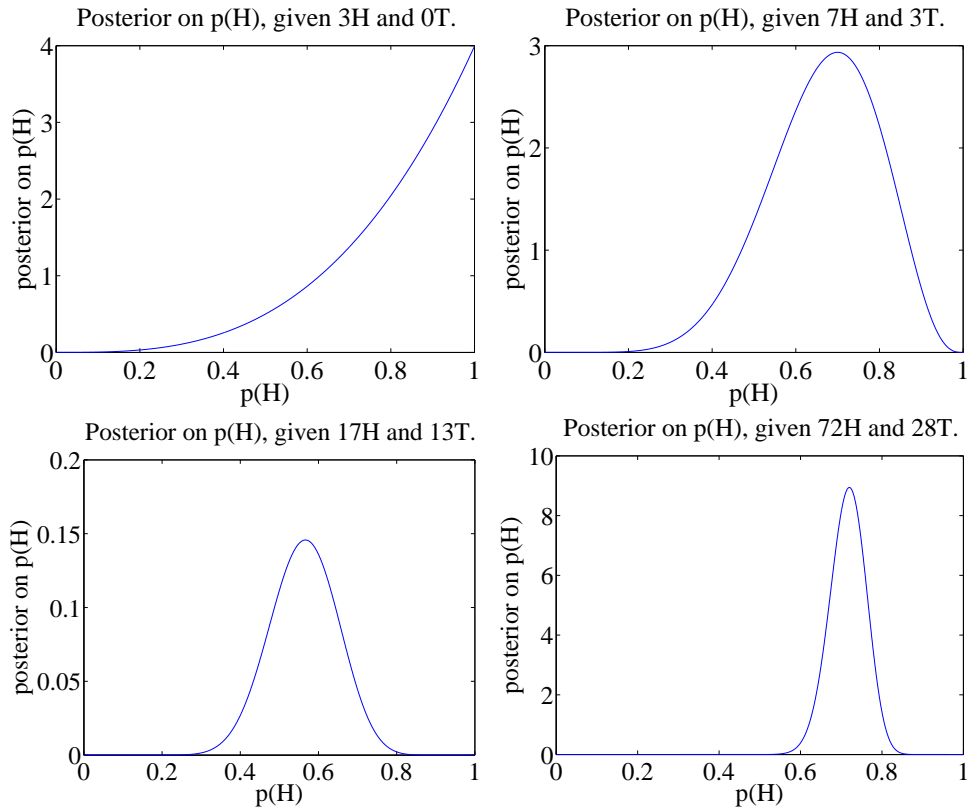


FIGURE 9.2: The probability that an unknown coin will come up heads when flipped is $p(H)$. For these figures, I simulated coin flips from a coin with $p = 0.75$. I then plotted the posterior for various data. Notice how, as we see more flips, we get more confident about p .

of evidence.

For other problems, we will need to marginalize out θ , by computing

$$P(\mathcal{D}) = \int_{\theta} P(\mathcal{D}|\theta)P(\theta)d\theta.$$

It is usually impossible to do this in closed form, so we would have to use a numerical integral. In some cases, $P(\theta)$ and $P(\mathcal{D}|\theta)$ are **conjugate**, meaning that $P(\theta|\mathcal{D})$ will take a familiar form and $P(\mathcal{D})$ follows easily.

Worked example 9.9 *Flipping a coin - II*

We have a coin with probability θ of coming up heads when flipped. We model the prior on θ with a Beta distribution, with parameters $\alpha > 0$, $\beta > 0$. We then flip the coin N times, and see h heads. What is $P(\theta|N, h, \alpha, \beta)$?

Solution: We have that $P(N, h|\theta)$ is binomial, and that $P(\theta|N, h, \alpha, \beta) \propto P(N, h|\theta)P(\theta|\alpha, \beta)$. This means that

$$P(\theta|N, h, \alpha, \beta) \propto \binom{N}{h} \theta^h (1-\theta)^{(N-h)} \frac{\Gamma(\alpha+\beta)}{\Gamma(\alpha)\Gamma(\beta)} \theta^{(\alpha-1)} (1-\theta)^{(\beta-1)}.$$

and we can write

$$P(\theta|N, h, \alpha, \beta) \propto \theta^{(\alpha+h-1)} (1-\theta)^{(\beta+N-h-1)}.$$

Notice this has the form of a Beta distribution, so it is easy to recover the constant of proportionality. We have

$$P(\theta|N, h, \alpha, \beta) = \frac{\Gamma(\alpha+\beta+N)}{\Gamma(\alpha+h)\Gamma(\beta+N-h)} \theta^{(\alpha+h-1)} (1-\theta)^{(\beta+N-h-1)}.$$

Worked example 9.10 *More swearsy politicians*

Example 5 gives some data from a swearsy politician. Assume we have only the first 10 intervals of observations, and we wish to estimate the intensity using a Poisson model. Write θ for this parameter. Use a Gamma distribution as a prior, and write out the posterior.

Solution: We have that

$$p(\theta|\alpha, \beta) = \frac{\beta^\alpha}{\Gamma(\alpha)} \theta^{(\alpha-1)} e^{-\beta\theta} \text{ and } p(\mathcal{D}|\theta) = \frac{\theta^7 e^{-\theta}}{24}.$$

This means that

$$p(\theta|\mathcal{D}) \propto \theta^{(\alpha-1+7)} e^{-(\beta+1)\theta}.$$

Notice this has the form of another Gamma distribution, so we can write

$$p(\theta|\mathcal{D}) = \frac{(\beta+1)^{(\alpha+7)}}{\Gamma(\alpha+7)} \theta^{(\alpha-1+7)} e^{-(\beta+1)\theta}$$

9.2.2 The Posterior for Normal Data

There is a very useful construction for the posterior for data where the likelihood is normal. We start with a simple example. Assume we drop a measuring device down a borehole. It is designed to stop falling and catch onto the side of the hole after it has fallen μ_0 meters. On board is a device to measure its depth. This device reports a known constant times the correct depth plus a zero mean normal random variable, which we call “noise”. The device reports depth every second.

The first question to ask is what depth do we believe the device is at *before* we receive any measurement? We designed the device to stop at μ_0 meters, so we are not completely ignorant about where it is. However, it may not have worked absolutely correctly. We choose to model the depth at which it stops as μ_0 meters plus a zero mean normal random variable. The second term could be caused by error in the braking system, etc. We could estimate the standard deviation of the second term (which we write σ_0) either by dropping devices down holes, then measuring with tape measures, or by analysis of likely errors in our braking system. The depth of the object is the unknown parameter of the model; we write this depth θ . Now the model says that θ is a normal random variable with mean μ_0 and standard deviation σ_0 .

Notice that this model probably isn’t exactly right — for example, there must be some probability in the model that the object falls beyond the bottom of the hole, which it can’t do — but it captures some important properties of our system. The device should stop at or close to μ_0 meters most of the time, and it’s unlikely to be too far away.

Now assume we receive a single measurement — what do we now know about the device’s depth? The first thing to notice is that there is something to do here. Ignoring the prior and taking the measurement might not be wise. For example, imagine that the noise in the wireless system is large, so that the measurement is often corrupted — our original guess about the device’s location might be better than the measurement. Write x_1 for the measurement. Notice that the scale of the measurement may not be the same as the scale of the depth, so the mean of the measurement is $c_1\theta$, where c_1 is a change of scale (for example, from inches to meters). We have that $p(x_1|\theta)$ is normal with mean $c_1\theta$ and standard deviation σ_{n1} . We would like to know $p(\theta|x_1)$.

We have that

$$\begin{aligned} \log p(\theta, x_1) &= \log p(x_1|\theta) + \log p(\theta) \\ &= -\frac{(x_1 - c_1\theta)^2}{2\sigma_{n1}^2} - \frac{(\theta - \mu_0)^2}{2\sigma_0^2} \\ &\quad + \text{terms not depending on } \theta \text{ or } x. \end{aligned}$$

We have two estimates of the position, θ , and we wish to come up with a representation of what we know about θ . One is x_1 , which is a *measurement* — we know its value. The expected value of x_1 is $c_1\theta$, so we could infer θ from x_1 . But we have another estimate of the position, which is μ_0 . The posterior, $p(\theta|x_1)$, is a probability distribution on the variable θ ; it depends on the known values x_1 , μ_0 , σ_0 and σ_{n1} . We need to determine its form. We can do so by some rearrangement of the expression for $\log p(\theta, x_1)$.

Notice first that this expression is of degree 2 in θ (i.e. it has terms θ^2 , θ and things that don't depend on θ). This means that $p(\theta|x_1)$ must be a normal distribution, because we can rearrange its log into the form of the log of a normal distribution. This yields a fact of crucial importance.

Useful Fact: 9.1 *Normal distributions are conjugate*

A normal prior and a normal likelihood yield a normal posterior.

Write μ_1 for the mean of this distribution, and σ_{n1} for its standard deviation. The log of the distribution must be

$$-\frac{(\theta - \mu_1)^2}{2\sigma_1^2} + \text{terms not depending on } \theta.$$

The terms not depending on θ are not interesting, because if we know σ_1 those terms must add up to

$$\log\left(\frac{1}{\sqrt{2\pi}\sigma_1}\right)$$

so that the probability density function sums to one. Our goal is to rearrange terms into the form above. Notice that

$$-\frac{(\theta - \mu_1)^2}{2\sigma_p^2} = -\theta^2\left(\frac{1}{2\sigma_1^2}\right) + 2\theta\frac{\mu_1}{2\sigma_p^2} + \text{term not depending on } \theta$$

We have

$$\begin{aligned} \log p(\theta|x_1) &= -\frac{(c_1\theta - x_1)^2}{2\sigma_{n1}^2} - \frac{(\theta - \mu_0)^2}{2\sigma_0^2} + \text{terms not depending on } \theta \\ &= -\theta^2\left(\frac{1}{2\left(\frac{\sigma_{n1}^2\sigma_0^2}{\sigma_{n1}^2 + c_1^2\sigma_0^2}\right)}\right) + 2\theta\left(c_1\frac{x_1}{2\sigma_{n1}^2} + \frac{\mu_0}{2\sigma_0^2}\right) \\ &+ \text{terms not depending on } \theta \end{aligned}$$

which means that

$$\sigma_1^2 = \frac{\sigma_{n1}^2\sigma_0^2}{\sigma_{n1}^2 + c_1^2\sigma_0^2}$$

and

$$\begin{aligned} \mu_1 &= 2\left(c_1\frac{x_1}{2\sigma_{n1}^2} + \frac{\mu_0}{2\sigma_0^2}\right)\frac{\sigma_{n1}^2\sigma_0^2}{\sigma_{n1}^2 + c_1^2\sigma_0^2} \\ &= \left(\frac{c_1x_1\sigma_0^2 + \mu_0\sigma_{n1}^2}{\sigma_{n1}^2\sigma_0^2}\right)\frac{\sigma_{n1}^2\sigma_0^2}{\sigma_{n1}^2 + c_1^2\sigma_0^2} \\ &= \frac{c_1x_1\sigma_0^2 + \mu_0\sigma_{n1}^2}{\sigma_{n1}^2 + c_1^2\sigma_0^2}. \end{aligned}$$

These equations “make sense”. Imagine that σ_0 is very small, and σ_{n1} is very big; then our new expected value of θ — which is μ_1 — is about μ_0 . Equivalently, because our prior was very accurate, and the measurement was unreliable, our expected value is about the prior value. Similarly, if the measurement is reliable (i.e. σ_{n1} is small) and the prior has high variance (i.e. σ_0 is large), then our expected value of θ is about x_1/c_1 — i.e. the measurement, rescaled. I have put these equations, in a more general form, in a box below.

Useful Fact: 9.2 *Normal posteriors*

Assume we wish to estimate a parameter θ . The prior distribution for θ is normal, with known mean μ_π and known standard deviation σ_π . We receive a single data item x . The likelihood of this data item is normal with mean $c\theta$ and standard deviation σ_m , where c and σ_m are known. Then the posterior, $p(\theta|x, c, \sigma_m, \mu_\pi, \sigma_\pi)$, is normal, with mean

$$\frac{cx\sigma_\pi^2 + \mu_\pi\sigma_m^2}{\sigma_m^2 + c^2\sigma_\pi^2}$$

and standard deviation

$$\sqrt{\frac{\sigma_m^2\sigma_\pi^2}{\sigma_m^2 + c^2\sigma_\pi^2}}.$$

Assume a second measurement, x_2 arrives. We know that $p(x_2|\theta, c_2, \sigma_{n2})$ is normal with mean $c_2\theta$ and standard deviation σ_{n2} . In the example, we have a new measurement of depth — perhaps in a new, known, scale — with new noise (which might have larger, or smaller, standard deviation than the old noise) added. Then we can use $p(\theta|x_1, c_1, \sigma_{n1})$ as a prior to get a posterior $p(\theta|x_1, x_2, c_1, c_2, \sigma_{n1}, \sigma_{n2})$. Each is normal, by useful fact 1. Not only that, but we can easily obtain the expressions for the mean μ_2 and the standard deviation σ_2 *recursively* as functions of μ_1 and σ_1 .

Applying useful fact 2, we have

$$\mu_2 = \frac{c_2x_2\sigma_1^2 + \mu_1\sigma_{n2}^2}{\sigma_{n2}^2 + c_2^2\sigma_1^2}$$

and

$$\sigma_2^2 = \frac{\sigma_{n2}^2\sigma_1^2}{\sigma_{n2}^2 + c_2^2\sigma_1^2}.$$

But what works for 2 and 1 will work for $k+1$ and k . We know the posterior after k measurements will be normal, with mean μ_k and standard deviation σ_k . The $k+1$ 'th measurement x_{k+1} arrives, and we have $p(x_{k+1}|\theta, c_{k+1}, \sigma_{n(k+1)})$ is normal. Then the posterior is normal, and we can write the mean μ_{k+1} and the standard deviation σ_{k+1} recursively as functions of μ_k and σ_k . The result is worth putting in a box.

Useful Fact: 9.3 *Online updating of normal posteriors*

Assume we wish to estimate a parameter θ . The prior distribution for θ is normal, with known mean μ_π and known standard deviation σ_π . All data is normal conditioned on θ . We have already received k data items. The posterior $p(\theta|x_1, \dots, x_k, c_1, \dots, c_k, \sigma_{n1}, \dots, \sigma_{nk}, \mu_\pi, \sigma_\pi)$ is normal, with mean μ_k and standard deviation σ_k . We receive a new data item x_{k+1} . The likelihood of this data item is normal with mean $c\theta$ and standard deviation $\sigma_{n(k+1)}$, where c_{k+1} and $\sigma_{n(k+1)}$ are known. Then the posterior, $p(\theta|x_1, \dots, x_{k+1}, c_1, \dots, c_k, c_{k+1}, \sigma_{n1}, \dots, \sigma_{n(k+1)}, \mu_\pi, \sigma_\pi)$, is normal, with mean

$$\mu_{k+1} = \frac{c_{k+1}x_{k+1}\sigma_k^2 + \mu_k\sigma_{n(k+1)}^2}{\sigma_{n(k+1)}^2 + c_{k+1}^2\sigma_k^2}$$

and

$$\sigma_{k+1}^2 = \frac{\sigma_{n(k+1)}^2\sigma_k^2}{\sigma_{n(k+1)}^2 + c_{k+1}^2\sigma_k^2}.$$

Again, notice the very useful fact that, if everything is normal, we can update our posterior representation when new data arrives using a very simple recursive form.

9.2.3 MAP Inference

Look at example 1, where we estimated the probability a coin would come up heads with maximum likelihood. We could not change our estimate just by knowing the coin was fair, but we could come up with a number for $\theta = p(H)$ (rather than, say, a posterior distribution). A natural way to produce a point estimate for θ that incorporates prior information is to choose $\hat{\theta}$ such that

$$\hat{\theta} = \operatorname{argmax}_{\theta} P(\theta|\mathcal{D}) = \operatorname{argmax}_{\theta} \frac{P(\theta, \mathcal{D})}{P(\mathcal{D})}$$

This is the MAP estimate. If we wish to perform MAP inference, $P(\mathcal{D})$ doesn't matter (it changes the value, but not the location, of the maximum). This means we can work with $P(\theta, \mathcal{D})$, often called the **joint** distribution.

Worked example 9.11 *Flipping a coin - II*

We have a coin with probability θ of coming up heads when flipped. We model the prior on θ with a Beta distribution, with parameters $\alpha > 0$, $\beta > 0$. We then flip the coin N times, and see h heads. What is the MAP estimate of θ ?

Solution: We have that

$$P(\theta|N, h, \alpha, \beta) = \frac{\Gamma(\alpha + \beta + N)}{\Gamma(\alpha + h)\Gamma(\beta + N - h)} \theta^{\alpha+h-1} (1 - \theta)^{\beta+N-h-1}.$$

You can get the MAP estimate by differentiating and setting to 0, yielding

$$\hat{\theta} = \frac{\alpha - 1 + h}{\alpha + \beta - 2 + N}.$$

This has rather a nice interpretation. You can see α and β as extra counts of heads (resp. tails) that are added to the observed counts. So, for example, if you were fairly sure that the coin should be fair, you might make α and β large and equal. When $\alpha = 1$ and $\beta = 1$, we have a uniform prior as in the previous examples.

Worked example 9.12 *More swears politicians*

We observe our swearing politician for N intervals, seeing n_i swear words in the i 'th interval. We model the swearing with a Poisson model. We wish to estimate the intensity, which we write θ . We use a Gamma distribution for the prior on θ . What is the MAP estimate of θ ?

Solution: Write $T = \sum_{i=1}^N n_i$. We have that

$$p(\theta|\mathcal{D}) = \frac{(\beta + 1)^{(\alpha+T)}}{\Gamma(\alpha + T)} \theta^{\alpha-1+T} e^{-(\beta+1)\theta}$$

and the MAP estimate is

$$\hat{\theta} = \frac{(\alpha - 1 + T)}{(\beta + 1)}$$

(which you can get by differentiating with respect to θ , then setting to zero). Notice that if β is close to zero, you can interpret α as extra counts; if β is large, then it strongly discourages large values of $\hat{\theta}$, even if the counts are large.

Worked example 9.13 *Normal data*

Assume you see N datapoints x_i which are modelled by a normal distribution with unknown mean θ and with known standard deviation σ . You model the prior on θ using a normal distribution with mean μ_0 and standard deviation σ_0 . What is the MAP estimate of the mean?

Solution: Recall that the maximum value of a normal distribution occurs at its mean. Now problem is covered by useful fact 2, but in this case we have $c_i = 1$ for each data point, and $\sigma_i = \sigma$. We can write

$$\mu_N = \frac{x_N \sigma_{N-1}^2 + \mu_{N-1} \sigma^2}{\sigma^2 + \sigma_{N-1}^2}$$

and

$$\sigma_N^2 = \frac{\sigma^2 \sigma_{N-1}^2}{\sigma^2 + \sigma_{N-1}^2}.$$

and evaluate the recursion down to μ_0, σ_0 .

9.2.4 Cautions about Bayesian Inference

Just like maximum likelihood inference, bayesian inference is not a recipe that can be applied without thought. It turns out that, when there is a lot of data, the prior has little inference on the outcome of the inference, and the MAP solution looks a lot like the maximum likelihood solution. So the difference between the two approaches is most interesting when there is little data, where the prior matters. The difficulty is that it might be hard to know what to use as a good prior. In the examples, I emphasized mathematical convenience, choosing priors that lead to clean posteriors. There is no reason to believe that nature uses conjugate priors (even though conjugacy is a neat property). How should one choose a prior for a real problem?

This isn't an easy point. If there is little data, then the choice could really affect the inference. Sometimes we're lucky, and the logic of the problem dictates a choice of prior. Mostly, we have to choose and live with the consequences of the choice. Often, doing so is succesful in applications.

The fact we can't necessarily justify a choice of prior seems to be one of life's inconveniences, but it represents a significant philosophical problem. It's been at the core of a long series of protracted, often quite intense, arguments about the philosophical basis of statistics. I haven't followed these arguments closely enough to summarize them; they seem to have largely died down without any particular consensus being reached.

9.3 SAMPLES, URNS AND POPULATIONS

Very often the data we see is a small part of the data we could have seen, if we'd been able to collect enough data. We need to know how the measurements we make

on the dataset relate to the measurements we could have made, if we had all the data. This situation occurs very often. For example, imagine we wish to know the average weight of a rat. This isn't random; you could weigh every rat on the planet, and then average the answers. But doing so would be absurd (among other things, you'd have to weigh them all at the same time, which would be tricky). Instead, we weigh a small set of rats, chosen rather carefully. If we have chosen sufficiently carefully, then the answer from the small set is quite a good representation of the answer from the whole set.

The data we could have observed, if we could have seen everything, is the **population**. The data we actually have is the **sample**. We would like to know the mean of the population, but can see only the sample; surprisingly, we can say a great deal from the sample alone, assuming that it is chosen appropriately.

Assume we have a population $\{x\}$, for $i = 1, \dots, N_p$. Notice the subscript here — this is the number of items in the population. The population could be unreasonably big: for example, it could consist of all the people in the world. We want to know the mean of this dataset, but we do not get to see the whole dataset. Instead, we see the sample.

9.3.1 Estimating the Population Mean from a Sample

How the sample is obtained is key to describing the population. We will focus on only one model (there are lots of others). In our model, the sample is obtained by choosing a fixed number of data items. Write k for the number of data items in the sample. We expect k is a lot smaller than N_p . Each item is chosen independently, and fairly. This means that each time we choose, we choose one from the entire set of N_p data items, and each has the same probability of being chosen. This is sometimes referred to as “sampling with replacement”.

One natural way to think about sampling with replacement is to imagine the data items as being written on tickets, which are placed in an urn (old-fashioned word for a jar, now used mainly by statisticians and morticians). You obtain the sample by repeating the following experiment k times: shake the urn; take a ticket from the urn and write down the data on the ticket; put it back in the urn. Notice that, in this case, each sample is drawn from the same urn. This is important, and makes the analysis easier. If we had not put the ticket back, the urn would change between samples.

We summarize the whole dataset with its mean, which we write $\text{popmean}(\{x\})$. This is known as the **population mean**. The notation is just to drive home the facts that it's the mean of the whole population, and that we don't, and can't, know it. The whole point of this exercise is to *estimate* this mean.

We would like to estimate the mean of the whole dataset from the items that we actually see. Imagine we draw k tickets from the urn as above, and average the values. The result is a random variable, because different draws of k tickets will give us different values. Write $X^{(k)}$ for this random variable, which is referred to as the **sample mean**. Because expectations are linear, we must have that

$$\mathbb{E}[X^{(k)}] = \frac{1}{k} \left(\mathbb{E}[X^{(1)}] + \dots + \mathbb{E}[X^{(1)}] \right) = \mathbb{E}[X^{(1)}]$$

(where $X^{(1)}$ is the random variable whose value is obtained by drawing one ticket

from the urn). Now

$$\begin{aligned}\mathbb{E}[X^{(1)}] &= \sum_{i \in 1, \dots, N_p} x_i p(i) \\ &= \sum_{i \in 1, \dots, N_p} x_i \frac{1}{N_p} && \text{because we draw fairly from the urn} \\ &= \frac{\sum_{i \in 1, \dots, N_p} x_i}{N_p} \\ &= \text{popmean}(\{x\})\end{aligned}$$

which is the mean value of the items in the urn. This means that

$$\mathbb{E}[X^{(k)}] = \text{popmean}(\{x_i\}).$$

Under our sampling model, the expected value of the sample mean is the population mean.

Useful Facts: 9.4 *Sample means and population means*

The sample mean is a random variable. It is random, because different samples from the population will have different values of the sample mean. The expected value of this random variable is the population mean.

We will not get the same value of $X^{(k)}$ each time we perform the experiment, because we see different data items in each sample. So $X^{(k)}$ has variance, and this variance is important. If it is large, then each estimate is quite different. If it is small, then the estimates cluster. Knowing the variance of $X^{(k)}$ would tell us how accurate our estimate of the population mean is.

9.3.2 The Variance of the Sample Mean

We write $\text{popstd}(\{x\})$ for the standard deviation of the whole population of $\{x\}$. Again, we write it like this to keep track of the facts that (a) it's for the whole population and (b) we don't — and usually can't — know it.

We can compute the variance of $X^{(k)}$ (the sample mean) easily. We have

$$\text{var}[X^{(k)}] = \mathbb{E}[(X^{(k)})^2] - \mathbb{E}[X^{(k)}]^2 = \mathbb{E}[(X^{(k)})^2] - (\text{popmean}(\{x\}))^2$$

so we need to know $\mathbb{E}[(X^{(k)})^2]$. We can compute this by writing

$$X^{(k)} = \frac{1}{k}(X_1 + X_2 + \dots + X_k)$$

where X_1 is the value of the first ticket drawn from the urn, etc. We then have

$$X^{(k)2} = \left(\frac{1}{k}\right)^2 \left(\begin{array}{l} X_1^2 + X_2^2 + \dots + X_k^2 + X_1X_2 + \dots \\ X_1X_k + X_2X_1 + \dots + X_2X_k + \dots + X_{k-1}X_k \end{array} \right).$$

Expectations are linear, so we have that

$$\mathbb{E}\left[X^{(k)}\right]^2 = \left(\frac{1}{k}\right)^2 \left(\begin{array}{l} \mathbb{E}[X^2_1] + \mathbb{E}[X^2_2] + \dots + \mathbb{E}[X^2_k] + \mathbb{E}[X_1X_2] + \\ \dots + \mathbb{E}[X_1X_k] + \mathbb{E}[X_2X_1] + \dots + \mathbb{E}[X_2X_k] + \dots + \mathbb{E}[X_{k-1}X_k] \end{array} \right).$$

The *order* in which the tickets are drawn from the urn doesn't matter, because each time we draw a ticket we draw from the same urn. This means that $\mathbb{E}[X^2_1] = \mathbb{E}[X^2_2] = \dots = \mathbb{E}[X^2_k]$. You can think of this term as the expected value of the random variable generated by: drawing a single number out of the urn; squaring that number; and reporting the square. Notice that $\mathbb{E}[X^2_1] = \mathbb{E}[(X^{(1)})^2]$ (look at the definition of $X^{(1)}$).

Because the *order* doesn't matter, we also have that $\mathbb{E}[X_1X_2] = \mathbb{E}[X_1X_3] = \dots = \mathbb{E}[X_{k-1}X_k]$. You can think of this term as the expected value of the random variable generated by: drawing a number out of the urn; writing it down; returning it to the urn; then drawing a second number from the urn; and reporting the product of these two numbers. So we can write

$$\mathbb{E}\left[X^{(k)}\right]^2 = \left(\frac{1}{k}\right)^2 \left(k\mathbb{E}[(X^{(1)})^2] + k(k-1)\mathbb{E}[X_1X_2] \right)$$

and these two terms are quite easy to evaluate.

Worked example 9.14 *Urn variances*

Show that

$$\mathbb{E}\left[(X^{(1)})^2\right] = \frac{\sum_{i=1}^{N_p} x_i^2}{N_p} = \text{popstd}(\{x\})^2 + \text{popmean}(\{x\})^2$$

Solution: First, we have $(X^{(1)})^2$ is the number obtained by taking a ticket out of the urn and squaring its data item. Now

$$\text{popstd}(\{x\})^2 = \mathbb{E}\left[(X^{(1)})^2\right] - \mathbb{E}\left[X^{(1)}\right]^2 = \mathbb{E}\left[(X^{(1)})^2\right] - \text{popmean}(\{x\})^2$$

so

$$\mathbb{E}\left[(X^{(1)})^2\right] = \text{popstd}(\{x\})^2 + \text{popmean}(\{x\})^2$$

Worked example 9.15 *Urn variances*

Show that

$$\mathbb{E}[X_1 X_2] = \text{popmean}(\{x\})^2$$

Solution: This looks hard, but isn't. Recall from the facts in chapter 6 (useful facts 2, page 125) that if X and Y are independent random variables, $\mathbb{E}[XY] = \mathbb{E}[X]\mathbb{E}[Y]$. But X_1 and X_2 are independent — they are different random draws from the same urn. So

$$\mathbb{E}[X_1 X_2] = \mathbb{E}[X_1]\mathbb{E}[X_2]$$

but $\mathbb{E}[X_1] = \mathbb{E}[X_2]$ (they are draws from the same urn) and $\mathbb{E}[X] = \text{popmean}(\{x\})$. So

$$\mathbb{E}[X_1 X_2] = \text{popmean}(\{x\})^2.$$

Now

$$\begin{aligned} \mathbb{E}\left[(X^{(k)})^2\right] &= \frac{k\mathbb{E}[(X^{(1)})^2] + k(k-1)\mathbb{E}[X_1 X_2]}{k^2} \\ &= \frac{\mathbb{E}[(X^{(1)})^2] + (k-1)\mathbb{E}[X_1 X_2]}{k} \\ &= \frac{(\text{popsd}(\{x\})^2 + \text{popmean}(\{x\})^2) + (k-1)\text{popmean}(\{x\})^2}{k} \\ &= \frac{\text{popsd}(\{x\})^2}{k} + \text{popmean}(\{x\})^2 \end{aligned}$$

so we have

$$\begin{aligned} \text{var}[X^{(k)}] &= \mathbb{E}\left[(X^{(k)})^2\right] - \mathbb{E}[X^{(k)}]^2 \\ &= \frac{\text{popsd}(\{x\})^2}{k} + \text{popmean}(\{x\})^2 - \text{popmean}(\{x\})^2 \\ &= \frac{\text{popsd}(\{x\})^2}{k}. \end{aligned}$$

This is a very useful result which is well worth remembering together with our facts on the sample mean, so we'll put them in a box together.

Useful Fact: 9.5 *The sample mean*

The sample mean is a random variable. Write $X^{(k)}$ for the mean of k samples. We have that:

$$\begin{aligned}\mathbb{E}[X^{(k)}] &= \text{popmean}(\{x\}) \\ \text{var}[X^{(k)}] &= \frac{\text{popsd}(\{x\})^2}{k} \\ \text{std}(X^{(k)}) &= \frac{\text{popsd}(\{x\})}{\sqrt{k}}\end{aligned}$$

The consequence is this: If you draw k samples, the standard deviation of your estimate of the mean is

$$\frac{\text{popsd}(\{x\})}{\sqrt{k}}$$

which means that (a) the more samples you draw, the better your estimate becomes and (b) the estimate improves rather slowly — for example, to halve the standard deviation in your estimate, you need to draw four times as many samples. The standard deviation of the estimate of the mean is often known as the **standard error** of the mean. This allows us to draw a helpful distinction: the population has a standard deviation, and our estimate of its mean (or other things — but we won't go into this) has a standard error.

Notice we cannot state the standard error of our estimate exactly, because we do not know $\text{popsd}(\{x\})$. But we could make a good estimate of $\text{popsd}(\{x\})$, by computing the standard deviation of the examples that we have. It is now helpful to have some notation for the particular sample we have. I will write $\sum_{i \in \text{sample}}$ for a sum over the sample items, and we will use

$$\text{mean}(\{x\}) = \frac{\sum_{i \in \text{sample}} x_i}{k}$$

for the mean of the sample — that is, the mean of the data we actually see; this is consistent with our old notation, but there's a little reindexing to keep track of the fact we don't see all of the population. Similarly, I will write

$$\text{std}(x) = \sqrt{\frac{\sum_{i \in \text{sample}} (x_i - \text{mean}(\{x_i\}))^2}{k}}$$

for the sample standard deviation. Again, this is the standard deviation of the data we actually see; and again, this is consistent with our old notation, again with a little reindexing to keep track of the fact we don't see all of the population. We

could estimate

$$\text{popstd}(\{x\}) \approx \text{std}(x)$$

and as long as we have enough examples, this estimate is good. If the number of samples k is small, it is better to use

$$\text{popstd}(\{x\}) \approx \sqrt{\frac{\sum_{i \in \text{sample}} (x_i - \text{mean}(\{x\}))^2}{k-1}}.$$

In fact, much more is known about the distribution of $X^{(k)}$.

9.3.3 The Probability Distribution of the Sample Mean

The sample mean is a random variable. We know an expression for its mean, and we can estimate its variance. In fact, we can determine its probability distribution, though I won't do this rigorously. In section 7.4.3, I mentioned that adding a number of independent random variables almost always got you a normal random variable, a fact sometimes known as the central limit theorem. I didn't prove it, and I'm not going to now. But when we form $X^{(k)}$, we're adding random variables. This means that $X^{(k)}$ is a normal random variable, for sufficiently big k (for some reason, $k > 30$ is usually seen as right).

This is important, because it has the following consequence. Draw a large number of different samples of k elements from the population. Each is a dataset of k items. Compute $\text{mean}(\{x\})$ for each, and regard the resulting numbers e_1, \dots, e_r as data items. Convert the e_i to standard coordinates s_i , where

$$s_i = \frac{(e_i - \text{mean}(\{e_i\}))}{\text{std}(e_i)}$$

(i.e. by subtracting the mean of the e_i , and dividing by their standard deviation). Now construct a histogram of the s . If r is sufficiently large, the histogram will be close to the standard normal curve.

9.3.4 When The Urn Model Works

In our model, there was a population of N_p data items x_i , and we saw k of them, chosen at random. In particular, each choice was fair (in the sense that each data item had the same probability of being chosen) and independent. These assumptions are very important for our analysis to apply. If our data does not have these properties, bad things can happen.

For example, assume we wish to estimate the percentage of the population that has beards. This is a mean (the data items take the value 1 for a person with a beard, and 0 without a beard). If we select people according to our model, then ask them whether they have a beard, then our estimate of the percentage of beards should behave as above.

The first thing that should strike you is that it isn't at all easy to select people according to this model. For example, we might select phone numbers at random, then call and ask the first person to answer the phone whether they have a beard; but many children won't answer the phone because they are too small. The next

important problem is that errors in selecting people can lead to massive errors in your estimate. For example, imagine you decide to survey all of the people at a kindergarten on a particular day; or all of the people in a women's clothing store; or everyone attending a beard growing competition (they do exist). In each case, you will get an answer that is a very poor estimate of the right answer, and the standard error might look very small. Of course, it is easy to tell that these cases are a bad choice.

It may not be easy to tell what a good choice is. You should notice the similarity between estimating the percentage of the population that wears a beard, and estimating the percentage that will vote for a particular candidate. There is a famous example of a survey that mispredicted the result of the Dewey-Truman presidential election in 1948; poll-takers phoned random phone numbers, and asked for an opinion. But at that time, telephones tended to be owned by a small percentage of rather comfortable households, who tended to prefer one candidate, and so the polls mispredicted the result rather badly.

Sometimes, we don't really have a choice of samples. For example, we might be presented with a small dataset of (say) human body temperatures. If we can be satisfied that the people were selected rather randomly, we might be able to use this dataset to predict expected body temperature. But if we knew that the subjects had their temperatures measured because they presented themselves at the doctor with a suspected fever, then we most likely cannot use it to predict expected body temperature.

One important and valuable case where this model works is in simulation. If you can guarantee that your simulations are independent (which isn't always easy), this model applies to estimates obtained from a simulation. Notice that it is usually straightforward to build a simulation so that the i 'th simulation reports an x_i where $\text{popmean}(\{x\})$ gives you the thing you want to measure. For example, imagine you wish to measure the probability of winning a game; then the simulation should report one when the game is won, and zero when it is lost. As another example, imagine you wish to measure the expected number of turns before a game is won; then your simulation should report the number of turns elapsed before the game was won.

Intervals and Testing

Although a point estimate is the best estimate available, it could still be wrong. This doesn't always matter. But in some applications, it can be important to know what range a parameter could take, and still be consistent with the data. We have seen this possibility before, in example 8, where I graphed the posterior — you could look at that graph and see that p could be in a fairly large range of values.

Being able to estimate an interval of values is important, particularly when there are safety or legal considerations to worry about. Imagine you have a machine that fills cereal boxes. Each box gets a quantity of cereal that is random, but has low variance. If the weight of cereal in any box is below the amount printed on the label, you might be in trouble. When you choose the amount to print, a point estimate of the mean weight of cereal might not be particularly helpful. If that estimate is a little low, you could have problems. Instead, what you'd like to know is an interval that the mean lies in with very high probability, so you can print a number that is smaller than the smallest in the interval.

Such intervals are usually known as **confidence intervals**. There are a variety of techniques for estimating them, depending on the particular problem. In section 10.1, I describe straightforward methods to construct confidence intervals. Confidence intervals for sample means are easily constructed using our understanding of standard error. Similarly, the posterior yields confidence intervals in a straightforward way when we use Bayesian inference. In other cases we might not have a good estimate of standard error; but we can use simulation to construct one (section 10.2). That section also shows how to exploit simulation methods to estimate confidence intervals for maximum likelihood parameter estimates.

Now imagine we hypothesize that (say) a population has a given mean. We can then ask how large a confidence interval we have to draw around the sample mean to include the hypothesis. If that interval is very large, then we would have to have drawn a very unusual sample to see the sample mean that we see. This means that our data does not support the hypothesis. This is an extremely fruitful line of reasoning, because it allows us to evaluate evidence in support of a model. Section 10.3 describes methods to do so. In fact, we can evaluate evidence in support of a *model*, too, and section 10.4 shows how to do this.

10.1 STRAIGHTFORWARD CONSTRUCTIONS OF CONFIDENCE INTERVALS

A **statistic** is a function of a dataset. One example of a statistic is the mean of a sample. Yet another is the MAP estimate of a parameter (which you get by computing some numbers from the dataset alone). We observe the value of a statistic, which we can compute from our dataset. But the dataset is random, because it is either a sample from a population or an IID sample from a distribution mode. This means that we should think of the statistic as the observed value of a

random variable — if we had a different sample from the same population (resp. IID sample from the same distribution) we would compute a different value. In these two cases, we know the observed value of the statistic *and* a model for the probability distribution of the statistic, as a random variable. This means we can construct confidence intervals relatively easily.

10.1.1 Confidence Intervals for Population Means

Assume we have a population, and we wish to know a confidence interval for its mean. We draw a sample $\{x\}$ of k items, and compute the sample mean. This is the value of a random variable — random, because it depends on the randomly drawn sample — whose probability distribution we know. This knowledge is very powerful, because it tells us how close to the true mean our estimate is likely to be. The reasoning looks like this. Our estimate of the unknown number $\text{popmean}(\{x\})$ is the mean of the sample we have, which we write $\text{mean}(\{x\})$. We know that the error — which is

$$\text{mean}(\{x\}) - \text{popmean}(\{x\})$$

— is a normal random variable, with mean 0 and standard deviation $\frac{\text{popstd}(\{x\})}{\sqrt{k}}$. We estimate this standard deviation as $\frac{\text{std}(x)}{\sqrt{k}}$. We can scale the error by the standard deviation to get

$$\hat{M} = \frac{\text{mean}(\{x\}) - \text{popmean}(\{x\})}{\left(\frac{\text{popstd}(\{x\})}{\sqrt{k}}\right)} \approx \frac{\text{mean}(\{x\}) - \text{popmean}(\{x\})}{\left(\frac{\text{std}(x)}{\sqrt{k}}\right)}$$

which is a standard normal random variable. But we know rather a lot about the behaviour of standard normal random variables. In particular, we have that:

- About 68% of the time, the value of a standard normal random variable is within one standard deviation of the mean.
- About 95% of the time, the value of a standard normal random variable is within two standard deviations of the mean.
- About 99% of the time, the value of a standard normal random variable is within three standard deviations of the mean.

In turn, this means that

- About 68% of the time, $\text{abs}(\text{mean}(\{x\}) - \text{popmean}(\{x\})) \leq \frac{\text{popstd}(\{x\})}{k}$.
- About 95% of the time, $\text{abs}(\text{mean}(\{x\}) - \text{popmean}(\{x\})) \leq 2 \frac{\text{popstd}(\{x\})}{k}$.
- About 99% of the time, $\text{abs}(\text{mean}(\{x\}) - \text{popmean}(\{x\})) \leq 3 \frac{\text{popstd}(\{x\})}{k}$.

This means that we can construct a **confidence interval** for our estimate. For about 68% of samples, the sample mean will lie in the interval between $\text{mean}(\{x\}) - \frac{\text{std}(x)}{k}$ and $\text{mean}(\{x\}) + \frac{\text{std}(x)}{k}$, and so on. Notice that we have just accepted that the approximation $\text{std}(x) \approx \text{popstd}(\{x\})$ is harmless, which it is. We can plot the confidence interval by drawing **error bars** — draw a bar one (or two, or three)

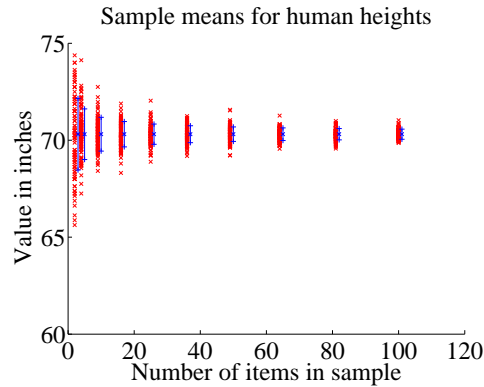


FIGURE 10.1: *I took the heights dataset. I then formed sampled elements with replacement to form random subsets of sizes (2, 4, 9, 16, ..., 100). For each of 100 subsets of each size, I computed the sample mean — these are shown as x's on the plot. I then computed the population mean, and the standard error as measured by the population standard deviation. The x to the side of each column is the population mean, and the vertical bars are one standard error above and below the population mean. Notice how (a) the sample means vary less as the sample gets bigger and (b) the sample means largely lie within the error bars.*

standard deviations up and down from the estimate. We interpret this interval as representing the effect of sampling uncertainty on our estimate. If the urn model really did apply, then the confidence intervals have the property that the true mean lies inside the interval for about 68% of possible samples (for one standard error error bars; or 95% for two; etc.).

It is quite straightforward, and informative, to verify these ideas with a simulation. I used the heights column from the bodyfat dataset (from <http://www2.stetson.edu/~jrasp/data.htm>; look for bodyfat.xls). I removed the single height outlier. I simulated the population using the whole dataset (251 items), then drew numerous samples of various sizes, with replacement. I computed the mean of each of these sets of samples. Figure 10.1 shows a scatter plot of sample means for different samples, using a set of sizes (2, 4, 9, 16, ..., 100). I have also plotted the population mean, and the true 1-standard error bars (i.e. using the population standard deviation) for each of these sample sizes. Notice how most sample means lie within the 1-standard error bars, as they should.

Figure 10.2 shows a sample mean and the computed standard error bars for a set of different sized samples. I have also shown the population mean. Notice how the population mean is usually within the computed standard error bars. Notice also how the computed standard error bars (i.e. using the sample standard deviation) are a bit small for small samples, but compare well with the true values for large samples. The problem here is that, for small samples, the sample standard deviation can be a fairly poor estimate of the population standard deviation. In the figure, I show a scatter plot of standard error estimated using sample standard deviation against standard error evaluated using population standard deviation.

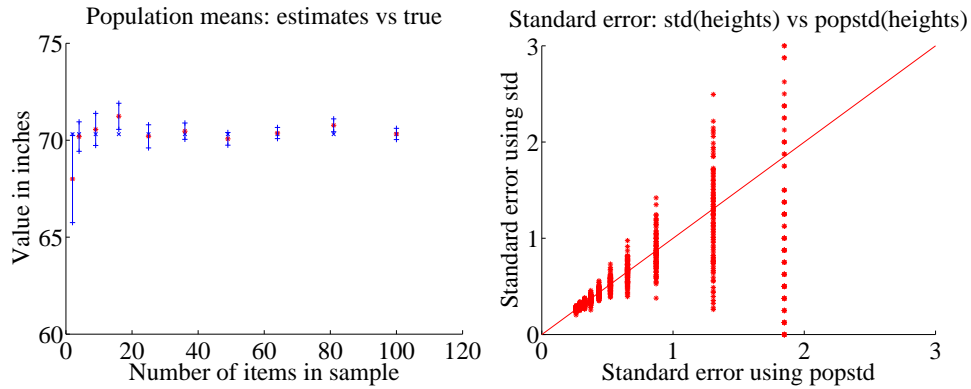


FIGURE 10.2: On the **right**, I chose one sample at random of each size; the sample mean is shown as a *. There are error bars (one standard error above and below) around the sample mean. These error bars are computed from the sample standard deviation. The population mean is the x . Notice how the population mean is within the error bars most, but not all, of the time (about 68% of the time, as they should be). The sample mean is rather a good estimate of the population mean, and the standard error is quite a reliable estimate of how well the sample mean represents the population mean when the sample is large enough. When the sample is small, the estimate is poor, as you can see on the **right**. This is a scatter plot of the true standard error (using the population standard deviation) against the estimate (using the sample standard deviation). When the sample is small, the standard error is big, and the estimate is poor.

When the standard error is small (because the sample is large), the two are quite close; when it is large (because the sample is small), they can be quite different. This suggests, correctly, that accurate inference from small samples can be tricky. The right thing to do is get more data — it is optimistic to hope for a good estimate of average human height from two people, after all.

Now we might reasonably ask another question. We should like to know a confidence interval such that the true mean lies within it for p of the samples. This is equivalent to asking what is the u such that

$$\int_{-u}^u \frac{1}{\sqrt{2\pi}} \exp(-x^2/2) dx = p$$

(i.e. what is the range of values about zero such that $p\%$ of standard normal random variables lies in this range). Such numbers can be extracted from the inverse of the error function (which is known as the **inverse error function**). An alternative is to look in statistical tables, which usually give u values for a useful set of p 's.

Worked example 10.1 *Speed of light*

Give a 95% confidence interval for the time of a light transit across a fixed distance, based on the dataset at <http://lib.stat.cmu.edu/DASL/Datafiles/SpeedofLight.html>.

Solution: The first thing to notice about this dataset is it seems to contain two negative times; but this is illusory. Look closely by tracking down on the web the paper referred to in the story (Stigler, S.M., “Do robust estimators work with real data?” *Annals of Statistics*, 5 (1977), pp. 1055–1078.). You will find that Newcomb measured the the passage of time for light to cross a fixed distance. He then subtracted a constant, and reported the result. Write t for the true measurement of time, in microseconds, and n for the number in the dataset. We have is

$$t = 24.8 + n \times 1e - 3.$$

Newcomb omitted the smallest value in the dataset, which is very different from all others. **Omitting** the smallest value gives a mean of 24.82729, a standard deviation of 0.0062, and a standard error of 0.00078. So with 95% confidence the value lies between 24.82573 and 24.82885. **Keeping** the smallest value gives a mean of 24.8262, a standard deviation of 0.0107 and a standard error of 0.0013. So with 95% confidence the value lies between 24.8236 and 24.8288.

Notice the temptation to drop an outlier in example 1- it really makes the measurement look better. These days, the speed of light is known precisely. It’s used to define the metre, so we would take the position that Newcomb was measuring the distance of the transit.

10.1.2 Bayesian Confidence Intervals

We can use the posterior to evaluate the probability that the true value of a parameter lies within an interval. Assume we want to know for $a < b$, the probability that the parameter θ lies in that interval. In many cases, we can compute

$$\int_a^b p(\theta|\mathcal{D})d\theta,$$

which supplies the relevant probability.

Worked example 10.2 *Flipping a coin - II*

We have a coin with probability θ of coming up heads when flipped. We start knowing nothing about θ except that, being a probability, $0 \leq \theta \leq 1$. We use a uniform prior on θ . We then flip the coin 10 times, and see 7 heads (and 3 tails). What is $P(\{\theta \in [0.5, 0.8]\} | \mathcal{D})$?

Solution: We could answer this question by computing

$$\int_{0.5}^{0.8} p(\theta | \mathcal{D}) d\theta.$$

Recall that a Beta distribution with $\alpha = \beta = 1$ is uniform, and look at example 11. The posterior is

$$P(\theta | 10, 7, 1, 1) = \frac{\Gamma(12)}{\Gamma(8)\Gamma(4)} \theta^7 (1 - \theta)^3.$$

I evaluated this integral numerically, and got 0.73.

In example 2, I determined that $P(\{\theta \in [0.5, 0.8]\} | \mathcal{D})$ was 0.73. Instead, I might wish to specify $0 \leq u < 0.5$, and then find an interval $[a, b]$ such that $P(\{\theta \in [a, b]\} | \mathcal{D}) = 1 - 2u$. Notice that this interval is not unique. Notice that

$$P(\{\theta \in [a, b]\} | \mathcal{D}) = 1 - P(\{\theta \leq a\} | \mathcal{D}) - P(\{\theta \geq b\} | \mathcal{D})$$

We will choose a such that

$$P(\{\theta \leq a\}) = \int_{-\infty}^a P(\theta | \mathcal{D}) d\theta = u$$

and b such that

$$P(\{\theta \geq b\} | \mathcal{D}) = \int_b^{\infty} P(\theta | \mathcal{D}) d\theta = u.$$

Actually obtaining values for a and b might be quite difficult. One strategy is to search a range of values.

Worked example 10.3 *Flipping a coin - II*

We have a coin with probability θ of coming up heads when flipped. We start knowing nothing about θ except that, being a probability, $0 \leq \theta \leq 1$. We use a uniform prior on θ . We then flip the coin 10 times, and see 7 heads (and 3 tails). Construct an interval $[a, b]$ such that $P(\{\theta \leq a\} | \mathcal{D}) \approx 0.05$ and $P(\{\theta \geq b\} | \mathcal{D}) \approx 0.05$.

Solution: I wrote a brief program to compute the relevant integrals numerically, then searched using interval halving. I found $P(\{\theta \leq 0.435\} | \mathcal{D}) \approx 0.05$ and $P(\{\theta \geq 0.865\} | \mathcal{D}) \approx 0.05$; this means the interval is $[0.435, 0.865]$.

10.2 USING SIMULATION TO CONSTRUCT INTERVALS

The analysis of some problems is difficult. But the idea of a confidence interval is naturally associated with repeated experiments. We can use simulations to estimate confidence intervals when analysis is difficult. Generally, we see the dataset as one of many possible dataset we could have seen. We must then assess what our estimate would be like for the other possible datasets. But what are those possible datasets to be?

When the data is explained by a parametric probability model, we can use that model to produce other possible datasets. If we compute a maximum likelihood estimate $\hat{\theta}$ of the parameters of the model, we can draw IID samples *from the model*. We then look at the estimate from that new dataset; the spread of the estimates yields our confidence interval.

When we have no applicable probability model of the dataset, we can resample the dataset to simulate the effects of sampling error. This strategy allows us to build confidence intervals for properties like medians.

10.2.1 Constructing Confidence Intervals for Parametric Models

Assume we have a dataset $\mathcal{D} = \{\mathbf{x}\}$, and a probability model we believe applies to that dataset. We can estimate model parameters, we can maximize the likelihood function $L(\theta)$.

A $(1 - 2\alpha)$ confidence interval for a parameter is an interval $[c_\alpha, c_{(1-\alpha)}]$. We construct the interval such that, if we were to perform a very large number of repetitions of our original experiment then estimate a parameter value for each, c_α would be the α quantile and $c_{(1-\alpha)}$ would be the $(1 - \alpha)$ quantile of those parameter values. We interpret this to mean that, with confidence $(1 - 2\alpha)$, the correct value of our parameter lies in this interval. This definition isn't really watertight. How do we perform a very large number of repetitions? If we don't, how can we tell how good the confidence interval is? Nonetheless, we can construct intervals that illustrate how sensitive our original inference is to the data that we have.

There are a variety of methods to construct confidence intervals. We will focus on simulation. The algorithm should feel so natural to you that you may already have guessed what to do. First, we compute the maximum likelihood estimate of the parameters, $\hat{\theta}$. We assume this estimate is right, but need to see how our estimates of $\hat{\theta}$ would vary with different collections of data from our model with that parameter value. So we compute a collection of simulated datasets, \mathcal{D}_i , each the same size as the original dataset. We obtain these by simulating $P(d|\hat{\theta})$. Next, we compute a maximum likelihood estimate for each of these simulated datasets, $\hat{\theta}_i$. Finally, we compute the relevant percentiles of these datasets. The result is our interval.

Procedure: 10.1 *Estimating Confidence Intervals for Maximum Likelihood Estimates using Simulation*

Assume we have a dataset $\mathcal{D} = \{x\}$ of N items. We have a parametric model of this data $p(x|\theta)$, and write the likelihood $\mathcal{L}(\theta; \mathcal{D}) = P(\mathcal{D}|\theta)$. We construct $(1 - 2\alpha)$ confidence intervals using the following steps.

1. Compute the maximum likelihood estimate of the parameter, $\hat{\theta} = \underset{\theta}{\operatorname{argmax}} \mathcal{L}(\theta; \mathcal{D})$.
2. Construct R simulated datasets \mathcal{D}_i , each consisting of N IID samples drawn from $p(x|\hat{\theta})$.
3. For each such dataset, compute $\hat{\theta}_i = \underset{\theta}{\operatorname{argmax}} \mathcal{L}(\theta; \mathcal{D}_i)$.
4. Obtain $c_\alpha(\hat{\theta}_i)$, the α 'th quantile of the collection $\hat{\theta}_i$ and $c_{(1-\alpha)}(\hat{\theta}_i)$, the $1 - \alpha$ 'th quantile of the collection $\hat{\theta}_i$.

The confidence interval is $[c_\alpha, c_{(1-\alpha)}]$.

Figure 10.3 shows an example. In this case, I worked with simulated data from a normal distribution. In each case, the normal distribution had mean 0, but there are four different standard deviations (1, 4, 7, and 10). I simulated 10 different datasets from each of these distributions, containing 10, 40, 90, . . . , 810, 1000 data items. For each, I computed the maximum likelihood estimate of the mean. This isn't zero, *even though the data was drawn from a zero-mean distribution*, because the dataset is finite. I then estimated the confidence intervals for each using 10,000 simulated datasets of the same size. I show 95% confidence intervals for the cases $\sigma = 1$, $\sigma = 4$, $\sigma = 10$ in the figure, plotted against dataset size. Notice that these aren't symmetric about zero, because the maximum likelihood estimate isn't zero. They shrink as the dataset grows, but slowly. They are bigger when the standard deviation is bigger. It should seem reasonable that you can't expect an accurate estimate of the mean of a normal distribution with large standard deviation using only a few data points. One can demonstrate using statistical theory that would take us rather out of our way that the size of these intervals should behave like

$$\frac{\sigma}{\sqrt{N}}$$

and Figure 10.3 suggests that they do. I plotted $1/\text{size}^2$; you can see this grows linearly as N grows, and falls very fast as σ grows.

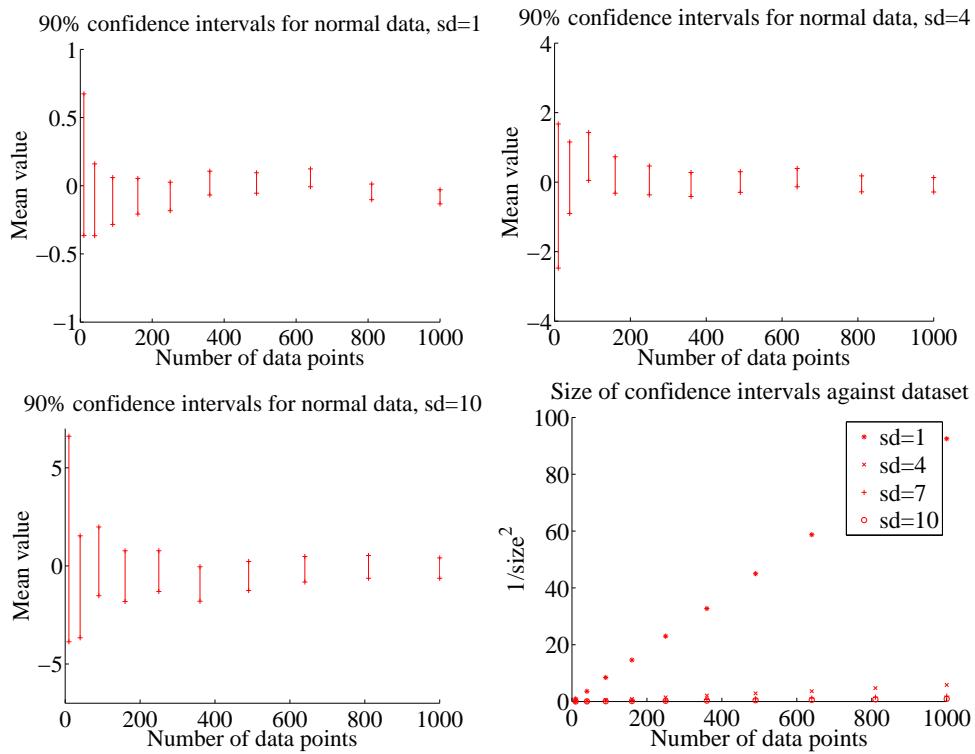


FIGURE 10.3: Confidence intervals computed for simulated normal data; details in the text.

Worked example 10.4 *Confidence Intervals by Simulation - II*

Construct a 90% confidence interval for the intensity estimate for the data of example 5 for the cases of 10 observations, 20 observations, and all 30 observations.

Solution: Recall from that example the maximum likelihood estimates of the intensity are $7/10$, $22/20$, and $29/30$ in the three cases. I used the Matlab function `poissrnd` to get 10000 replicates of a dataset of 10 (resp. 20, 30) items from a Poisson distribution with the relevant intensities. I then used `prctile` to get the 5% and 95% percentiles, yielding the intervals

$$\begin{aligned} [0.3, 1.2] & \quad \text{for 10 observations} \\ [0.75, 1.5] & \quad \text{for 20 observations} \\ [0.6667, 1.2667] & \quad \text{for 30 observations} \end{aligned}$$

Notice how having more observations makes the confidence interval smaller.

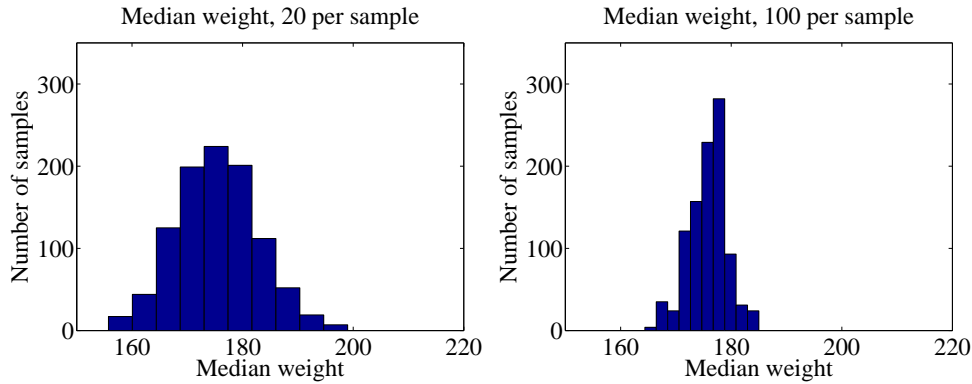


FIGURE 10.4: I took the weights dataset used all 253 measurements to represent a population. Rather than compute the median of the whole population, I chose to compute the median of a randomly chosen sample. The figures show a histogram of 1000 different values of the median, computed for 1000 different samples (of size 20 on the **left**, and of size 100 on the **right**). Notice that (a) there is a moderate amount of variation in the median of the sample; (b) these histograms look normal, and appear to have about the same median; (c) increasing the size of the sample has reduced the spread of the histogram.

10.2.2 Estimating Standard Error

We were able to produce convenient and useful estimates of standard error for sample means. But what happens if we want to, say, reason about the median of a population? It turns out that estimating standard error becomes difficult mathematically. This is an important problem, because if we had standard error estimates, we could use our methods for building confidence intervals and for testing hypotheses. It turns out that quite simple simulation methods give very good estimates.

For Figure 10.4, I assumed that all 253 weight measurements represented the entire population, then simulated what would happen for different random samples (with replacement) of different sizes. Figure 10.4 suggests that the sample median behaves quite like the sample mean as the random sample changes. Different samples have different medians, but the distribution of values looks fairly normal. When there are more elements in the sample, the standard deviation of median values is smaller. But we have no method to compute this standard deviation. The method I used here doesn't apply in practice, because we don't usually have the whole population.

There is a method, known as the **bootstrap**, which gives a very good estimate of the standard error of any statistic. Assume we wish to estimate the standard error of a statistic $S(\{\mathbf{x}\})$, which is a function of our dataset $\{\mathbf{x}\}$ of N data items. We compute r **bootstrap replicates** of this sample. Each replicate is obtained by sampling the dataset uniformly, and with replacement. One helpful way to think of this is that we are modelling our dataset as a sample of a probability distribution. This distribution, sometimes known as the **empirical distribution**,

has probability $1/N$ at each of the data items we see, and zero elsewhere. Now to obtain replicates, we simply draw new sets of IID samples from this probability distribution.

Notice that the bootstrap replicates are *not* a random permutation of the dataset; instead, we select one data item fairly and at random from the whole dataset N times. This means we expect a particular bootstrap replicate will have multiple copies of some data items, and no copies of others.

We write $\{\mathbf{x}\}_i$ for the i 'th bootstrap replicate of the dataset. We now compute

$$\bar{S} = \frac{\sum_i S(\{\mathbf{x}\}_i)}{r}$$

and the standard error estimate for S is given by:

$$\text{stderr}(\{S\}) = \sqrt{\frac{\sum_i [S(\{\mathbf{x}\}_i) - \bar{S}]^2}{r - 1}}$$

Procedure: 10.2 *The bootstrap*

Estimate the standard error for a statistic S evaluated on a dataset of N items $\{\mathbf{x}\}$.

1. Compute r bootstrap replicates of the dataset. Write the i 'th replicate $\{\mathbf{x}\}_i$. Obtain each by:
 - (a) Building a uniform probability distribution on the numbers $1 \dots N$.
 - (b) Drawing N independent samples from this distribution. Write $s(i)$ for the i 'th such sample.
 - (c) Building a new dataset $\{x_{s(1)}, \dots, x_{s(N)}\}$.

2. For each replicate, compute $\sum_i S(\{\mathbf{x}\}_i)$.

3. Compute

$$\bar{S} = \frac{\sum_i S(\{\mathbf{x}\}_i)}{r}$$

4. The standard error estimate for S is given by:

$$\text{stderr}(\{S\}) = \sqrt{\frac{\sum_i [S(\{\mathbf{x}\}_i) - \bar{S}]^2}{r - 1}}$$

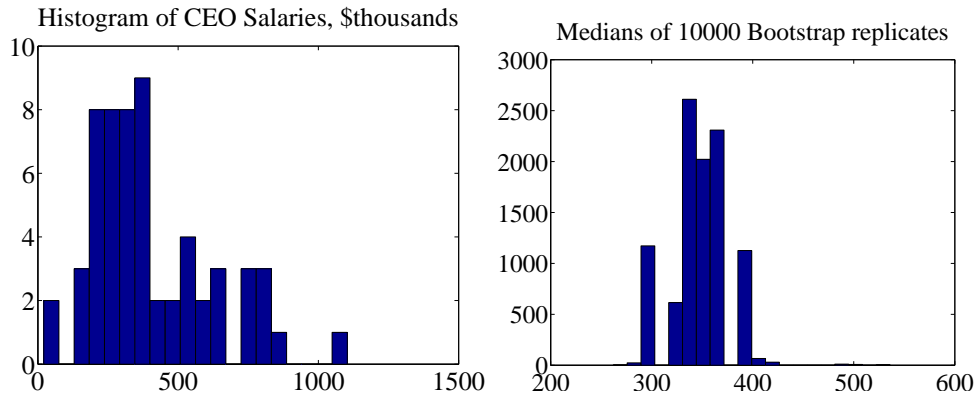


FIGURE 10.5: On the **left**, a histogram of salaries for CEO's of small companies in 1993, from the dataset of <http://lib.stat.cmu.edu/DASL/Datafiles/ceodat.html>. On the **right**, a histogram of the medians of 10,000 bootstrap replicates of this data. This simulates the effect of sampling variation on the median; see worked example 5.

Worked example 10.5 *The bootstrap standard error of the median*

You can find a dataset giving the salaries of CEO's at small firms in 1993 at <http://lib.stat.cmu.edu/DASL/Datafiles/ceodat.html>. Construct a 90% confidence interval for the median salary.

Solution: Salaries are in thousands of dollars, and one salary isn't given (we omit this value in what follows). Figure ?? shows a histogram of the salaries; notice there are some values that look like outliers. This justifies using a median. The median of the dataset is 350 (i.e. \$ 350,000 — this is 1993 data!). I constructed 10,000 bootstrap replicates. Figure 10.5 shows a histogram of the medians of the replicates. I used the matlab `prctile` function to extract the 5% and 95% percentiles of these medians, yielding the interval between 298 and 390. This means that we can expect that, for 90% of samples of CEO salaries for small companies, the median salary will be in the given range.

Worked example 10.6 *Are Brinks collectors different from city collectors? - I*

Using the dataset of <http://lib.stat.cmu.edu/DASL/Datafiles/brinkdat.html>, test whether Brinks coin collectors are different from city coin collectors (read the story, at that URL).

Solution: There are two types of parking meter in this story (which you should have read!). One is close to the city hall, and is always collected by city employees. The other is far, and for some months was collected by Brinks employees and for some months collected by city employees. I ignored the figures for month 10, because there was a missing digit in the number (don't know how or why). For the meters that are far, coins were collected for 22 months by city employees and for 24 months by Brinks employees.

For the far meters, I constructed 10,000 bootstrap replicates of each case. Figure 10.6 compares the two sets of replicates, showing the histogram of means for city collections vs Brinks collections. Notice how the means of the two histograms are quite separated, compared to their variance. This suggests, quite strongly, that the Brinks collections are different (the difference between the means is about 1.6 standard deviations). I used the nearby meter figures to check whether there was something special about the months on which Brinks employees collected coins. Again, I constructed 10,000 bootstrap replicates of each case (nearby coins in a Brinks month vs. nearby coins in a city month). This is plotted in figure 10.6. There's nothing special about the months, from the figure.

10.3 USING THE STANDARD ERROR TO EVALUATE HYPOTHESES

Assume we have a **hypothesis** to work with. For example, we believe that women and men have the same body temperature. We should like to assess the extent to which the evidence we have contradicts the hypothesis. At first glance, this may seem strange to you — surely one wants to assess the extent to which the evidence *supports* the hypothesis — but in fact it's natural. You can't prove that a scientific hypothesis is true; you can only fail to show that it's false. Just one piece of evidence can destroy a scientific hypothesis, but no amount of evidence can remove all doubt.

In our example, we measure the temperatures of a sample of women and of a sample of men. We cannot expect that the sample means are the same, because these are affected by the choice of sample. However, we have a good model of these sample means as random variables; in particular, we know the standard error. We can use this to assess the extent to which the difference in temperatures can be explained solely by the choice of sample. If it is easy to explain the difference in temperatures by choice of sample, then the evidence does not contradict the hypothesis; alternatively, if the difference in temperatures is hard to explain by

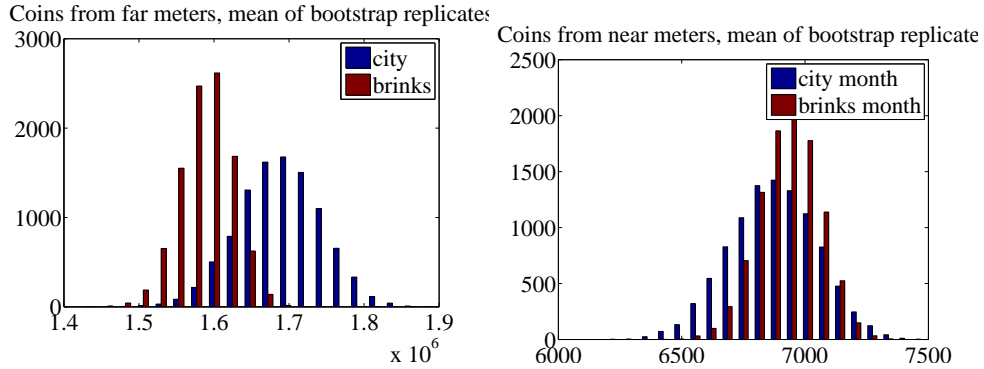


FIGURE 10.6: *On the left*, two histograms comparing the means of bootstrap replicates of the coins collected by Brinks employees and those collected by city employees. Notice how the two histograms are quite separated, compared to their width. This is fairly strong evidence that the two cases are different. The difference is not explained by the months in which the coins are collected. *On the right*, two histograms comparing the means of bootstrap replicates of coins collected from nearby meters. These coins were collected by city employees, but I have separated Brinks months from city months. There is no evidence the months are different.

choice of sample, the evidence contradicts the hypothesis.

There is an important, quite general, line of reasoning here. It is a bad idea to try and explain data using a hypothesis that makes the data you observed a rare event.

Example: 10.1 *Patriot missiles*

I got this example from “Dueling idiots”, a nice book by P.J. Nahin, Princeton University Press. Apparently in 1992, the Boston Globe of Jan 24 reported on this controversy. The pentagon claimed that the patriot missile successfully engaged SCUD missiles in 80% of encounters. An MIT physicist, Theodore Postol, pointed out there was a problem. He viewed tapes of 14 patriot/SCUD encounters, with one hit and 13 misses. We can reasonably assume each encounter is independent. We can extract the probability of getting one hit and 13 misses if $P(\text{hit}) = 0.8$ from the binomial model, to get a number around $1e-8$. Now you could look at this information and make several arguments: (a) the pentagon is right and the probability really is 0.8, but Postol looked at a really unlucky set of videotapes; (b) the probability is not 0.8, because you would need to fire 14 patriots at 14 SCUD missiles about $1e8$ times to see this set of videotapes once; (c) for some reason, the videotapes are not independent — perhaps only unsuccessful encounters get filmed. If Postol viewed tapes at random (i.e. he didn’t select only unsuccessful tapes, etc.), then argument (a) is easily dismissed, because the pentagon would have had to be unreasonably unlucky — it’s a bad idea to try to explain data with a hypothesis that makes the data very unlikely.

Example: 10.2 *MTG and Shuffling*

You build a deck of 24 lands and 36 spells. You shuffle this deck, and draw a hand of seven cards. You get no lands. You repeat the experiment, and still get no lands. On a third try, you still get no lands. By example 17, this event (three shuffles and draws with no lands in each hand) has probability about $8e-6$. You could draw several conclusions: (a) you got unlucky; (b) you miscounted lands somehow; (c) you’re shuffling the cards badly (i.e. the cards stick together, or you are bad at shuffling). Conclusion (a) is unattractive, because you would need to have been very unlucky (you’d see this outcome once in roughly $1e5$ experiments, where each consists of shuffling and drawing three hands). It’s a bad idea to try to explain data with a hypothesis that makes the data very unlikely, so conclusions (b) and (c) have become more plausible.

10.3.1 Does this Population have this Mean?

Imagine we hypothesize that the average human body temperature is 95° . Write \bar{T} for the random variable evaluated by: collecting a random sample of people, measuring their temperatures, and computing the average of these temperatures. The mean of this random variable is the population mean, and the standard deviation of this random variable is the standard error, which we write s . Now consider the random variable

$$G = \frac{(\bar{T} - 95^\circ)}{s}.$$

This is a standard normal random variable, if our hypothesis is true.

In practice, we have already chosen a single sample. We can compute the mean temperature of this sample, which we write \bar{t} . This represents an observed value of \bar{T} . If you're puzzled by the relationship between the two, use an analogy with rolling a die. The random variable is what you could get if you rolled a die; the value is the face that came up when you did roll it. Similarly, \bar{T} is what you could get by choosing a sample, and \bar{t} is the value you did get with the sample you have. In turn, we obtain a value for G , which is

$$g = \frac{(\bar{t} - 95^\circ)}{s}.$$

Now G is a standard normal random variable, if our hypothesis is true. We can interpret this to mean that for 68% of possible samples, g will take a value between -1 and 1 (etc.), if our hypothesis is true. Equivalently G will take a value between g and $-g$, for a fraction f of all possible samples of the population, where

$$f = \int_{-g}^g \exp\left(\frac{-u^2}{2}\right) du.$$

Now assume we see a very large (or very small) value of g . The value of f will be close to one, which implies that most samples from the population will have a g value closer to zero if the hypothesis were true. Equivalently, this says that, if the hypothesis were true, our sample is highly unusual, which implies the data fails to support the hypothesis.

It is easy to think about $p = 1 - f$ than about f . You should think of p as representing the fraction of samples that would give a larger absolute value of g , if the null hypothesis was true. If this fraction is very small, then there is significant evidence against the null hypothesis. We have that p — the fraction of experiments that would give us a larger absolute value of g than the one we saw — is given by

$$p = (1 - f) = \left(1 - \int_{-g}^g \exp\left(\frac{-u^2}{2}\right) du\right) = P(\{G > |\hat{g}|\}) \cup P(\{G < -|\hat{g}|\})$$

and we can compute this number easily, because we know the distribution of G is normal (or nearly so). The fraction is sometimes referred to as a **p-value**.

Once we have the p-value, testing is straightforward. A small value of the fraction means that the outcome we see would be rare if the null hypothesis is true. The fraction that we compute should be interpreted an assessment of the

significance of the evidence against the null hypothesis. The p-value is smaller when the evidence against the null hypothesis is stronger; we get to decide how small a p-value means we should reject the null hypothesis.

It is conventional to reject the null hypothesis when the p-value is less than 0.05. This is sometimes called “a significance level of 5%”. You should interpret a p-value of 0.05 as meaning that you would see evidence this unusual in about one experiment in twenty if the null hypothesis was true. Sometimes, the p-value is even smaller, and this can be interpreted as very strong evidence the null hypothesis is wrong. A p-value of less than 0.01 allows one to reject the null hypothesis at “a significance level of 1%”. Similarly, you should interpret a p-value of 0.01 as meaning that you would see evidence this unusual in about one experiment in a hundred if the null hypothesis was true. We now have a general recipe for testing whether a population has a particular mean, which belongs in a box:

Procedure: 10.3 *Testing whether a population has a given mean, based on a sample*

The initial hypothesis is that the population has a known mean, which we write μ . Write $\{x\}$ for the sample, and k for the sample size.

- Compute the sample mean, which we write $\text{mean}(\{x\})$.
- Estimate the standard error s_e using

$$s_e = \frac{\text{std}(x)}{\sqrt{k}}.$$

- Compute the **test statistic** using

$$s = \frac{(\mu - \text{mean}(\{x\}))}{s_e}.$$

- Compute the p-value, using

$$p = (1 - f) = (1 - \int_{-s}^s \exp\left(\frac{-u^2}{2}\right) du)$$

Equivalently, compute the probability that a standard normal random variable X takes a value greater than s or less than $-s$, using the expression

$$P(\{X > |s|\}) \cup P(\{X < -|s|\}).$$

- The p-value summarizes the extent to which the data contradicts the hypothesis. A small p-value implies that, *if* the hypothesis is true, the sample is very unusual. The smaller the p-value, the more strongly the evidence contradicts the hypothesis.

It is common to think that a hypothesis can be rejected only if the p-value is less than 5% (or some number). You should not think this way; the p-value summarizes the extent to which the data contradicts the hypothesis, and your particular application logic affects how you interpret it.

Worked example 10.7 *Average Human Body Weight*

Assess the null hypothesis that the average human body weight is 175 lb, using the height and weight data set of <http://www2.stetson.edu/~jrasp/data.htm> (called bodyfat.xls).

Solution: The dataset contains 252 samples; the average weight is 178.9lb. We know this average is an estimate of the mean of the original large set of all people. This estimate is a normal random variable whose mean is the mean of the population, and whose standard deviation is given by the standard error.

Our sample is large (which usually means over 30 elements) and so we can estimate the standard error as

$$\frac{\text{standard deviation of sample}}{\sqrt{\text{number in sample}}} = \frac{29.4}{15.9} = 1.9,$$

where the units are lb. The test statistic is

$$\frac{\text{sample mean} - \text{hypothesized population mean}}{\text{standard error}} = \frac{178.9 - 175}{1.9} = 2.05.$$

This is the value of a standard normal random variable. We must compute the fraction of outcomes where the test statistic would be larger than 2.05, or smaller than -2.05. This is

$$2 * \frac{1}{\sqrt{2\pi}} \int_{2.05}^{\infty} \exp\left(\frac{-x^2}{2}\right) dx = 0.02$$

so the p-value is 0.02. We can interpret this as quite strong evidence that the average human body weight is not, in fact, 175lb. This p-value says that, if (a) the average human body weight is 175lb and (b) we repeat the experiment (weigh 252 people and average their weights) 50 times, we would see a number as big as the one we see about once.

Worked example 10.8 *Average BMI*

The BMI (body mass index) is a number intended to capture how much a person's weight deviates from the usual for their height. Assess the null hypothesis that the average human BMI is 27, using the height and weight data set of <http://www2.stetson.edu/~jrasp/data.htm>; bodyfat.xls.

Solution: BMI is computed using

$$BMI = \frac{\text{weight in lb}}{(\text{height in in})^2} \times 703.$$

I did excluded two possible outliers in this dataset (entry 39 has weight 363 and height 72.25; entry 42 has weight 205 and height 29.5). I found a mean BMI of 25.3, and a standard deviation of 3.35. There are 250 items in the dataset, so the standard error is 0.21. The test statistic is

$$\frac{\text{sample mean} - \text{hypothesized population mean}}{\text{standard error}} = \frac{25.3 - 27}{2.1} = -8.1$$

so the p-value is the probability that a standard normal random variable would take a value larger than -8.1 or smaller than 8.1. There is no particular reason to care about the difference between this (extremely small, about 1e-16) number and zero. The evidence is very strongly against the hypothesis.

If one keeps the outliers, one gets a mean BMI of 25.9, and a standard deviation of 9.56. There are 252 items in the dataset, so the standard error is 0.60. Now the p-value is 0.08, suggesting that the evidence is against the hypothesis, but not overwhelmingly. Notice the significant effect of just two datapoints. We might need more data to be sure we could reject the hypothesis.

You should notice an important relationship between our test and the material of section 10.1. When we constructed a confidence interval, we knew the distribution that the sample mean would take as we randomly chose different samples from the population. In turn, this meant we could plot (for example) the range of possible values the population mean would take for (say) 95% of possible samples. When we test a hypothesis about the mean, we are asking what kind of confidence interval we would have to draw around the hypothesized mean to encompass the observed sample mean.

10.3.2 Do Two Populations have the same Mean?

We have two samples, and we need to know whether they come from the same, or different, populations. For example, we might observe people using two different interfaces, measure how quickly they perform a task, then ask are their perfor-

mances different? As another example, we might run an application with no other applications running, and test how long it takes to run some standard tasks. Because we don't know what the operating system, cache, etc. are up to, this number behaves a bit like a random variable, so it is worthwhile to do several experiments, yielding one set of samples. We now do this with other applications running as well, yielding another set of samples — is this set different from the first set? For realistic datasets, the answer is always yes, because they're random samples. A better question is could the differences be the result of chance, or are these datasets really samples of two different populations?

We will ask if the population means are different. It turns out we can answer this rather easily, using some tricks about normal random variables. Assume that the samples are large enough (30 seems to be the magic number here). Write $\{x\}$ for the first dataset, which has size k_x , and $\{y\}$ for the second, which has size k_y . These datasets need not be of the same size. Each dataset is a random sample, drawn with replacement, from a population. We should like to assess the evidence that these two populations have the same mean. To do so, we need some simple facts about normal random variables.

Useful Facts: 10.1 *Sums and differences of normal random variables*

Let X_1 be a normal random variable with mean μ_1 and standard deviation σ_1 . Let X_2 be a normal random variable with mean μ_2 and standard deviation σ_2 . Let X_1 and X_2 be independent. Then we have that:

- for any constant $c_1 \neq 0$, $c_1 X_1$ is a normal random variable with mean $c_1 \mu_1$ and standard deviation $c_1 \sigma_1$;
- for any constant c_2 , $X_1 + c_2$ is a normal random variable with mean $\mu_1 + c_2$ and standard deviation σ_1 ;
- $X_1 + X_2$ is a normal random variable with mean $\mu_1 + \mu_2$ and standard deviation $\sqrt{\sigma_1^2 + \sigma_2^2}$.

I will not prove these facts; we already know the expressions for means and standard deviations from our results on expectations. The only open question is to show that the distributions are normal. This is easy for the first two results. The third requires a bit of integration that isn't worth our trouble; you could do reconstruct the proof from section 1's notes on sums of random variables and some work with tables of integrals.

Write $X^{(k_x)}$ for the random variable obtained by: drawing a random sample with replacement of k_x elements from the first population, then averaging this sample. Write $Y^{(k_y)}$ for the random variable obtained by: drawing a random sample with replacement of k_y elements from the first population, then averaging this sample. From section 9.3, we know that each random variable is normal, and

we know the means and standard deviations of these random variables. In turn, this means that $X^{(k_x)} - Y^{(k_y)}$ is a normal random variable.

Now write D for $X^{(k_x)} - Y^{(k_y)}$. If the two populations have the same mean, then $\mathbb{E}[D] = 0$. Furthermore,

$$\text{std}(D) = \sqrt{\text{std}(X^{(k_x)})^2 + \text{std}(Y^{(k_y)})^2}.$$

We know how to estimate $\text{var}[X^{(k_x)}]$ — it is the standard error of the sample, and we can estimate it as $\text{std}(x)/\sqrt{k_x}$. So we have can estimate $\text{std}(D)$ as

$$\text{std}(D) \approx \sqrt{\left(\frac{\text{std}(x)}{\sqrt{k_x}}\right)^2 + \left(\frac{\text{std}(y)}{\sqrt{k_y}}\right)^2}.$$

We can now use the same reasoning we used to test the hypothesis that a population had a particular, known mean. We have identified a number we can compute from the two samples. We know how this number would vary under random choices of sample. If the value we observe is too many standard deviations away from the mean, the evidence is against our hypothesis. If we wanted to believe the hypothesis, we would be forced to believe that the sample is extremely strange. I have summarized this reasoning in box 4.

Procedure: 10.4 *Testing whether two populations have a given mean, based on a sample of each*

The initial hypothesis is that the populations have the same, unknown, mean. Write $\{x\}$ for the sample of the first population, $\{y\}$ for the sample of the second population, and k_x, k_y for the sample sizes.

- Compute the sample means for each population, $\text{mean}(\{x\})$ and $\text{mean}(\{y\})$.
- Estimate the standard error s_e for each population, using

$$s_{ex} = \frac{\text{std}(x)}{\sqrt{k}}, \quad s_{ey} = \frac{\text{std}(y)}{\sqrt{k}}.$$

- Compute the standard error for the difference between the means,

$$s_{ed} = \sqrt{s_{ex}^2 + s_{ey}^2}.$$

- Compute the **test statistic** using

$$s = \frac{(\text{mean}(\{x\}) - \text{mean}(\{y\}))}{s_{ed}}.$$

- Compute the p-value, using

$$p = (1 - f) = \left(1 - \int_{-s}^s \exp\left(\frac{-u^2}{2}\right) du\right)$$

Equivalently, compute the probability that a standard normal random variable X takes a value greater than s or less than $-s$, using the expression

$$P(\{X > |s|\}) \cup P(\{X < -|s|\}).$$

- The p-value summarizes the extent to which the data contradicts the hypothesis. A small p-value implies that, *if* the hypothesis is true, the sample is very unusual. The smaller the p-value, the more strongly the evidence contradicts the hypothesis.

It is common to think that a hypothesis can be rejected only if the p-value is less than 5% (or some number). You should not think this way; the p-value summarizes the extent to which the data contradicts the hypothesis, and your particular application logic affects how you interpret it.

Worked example 10.9 *Are US and Japanese cars different*

At <http://www.itl.nist.gov/div898/handbook/eda/section3/eda3531.htm> you can find a dataset, published by NIST, giving measurements of miles per gallon for Japanese and US cars. Assess the evidence these two populations have the same mean MPG.

Solution: There are 249 measurements for Japanese cars, and 79 for US cars. The mean for Japanese cars is 20.1446, and for US cars is 30.4810. The standard error for Japanese cars is 0.4065, and for US cars is 0.6872. The value of the test statistic is

$$\frac{(\text{mean}(\{x\}) - \text{mean}(\{y\}))}{s_{ed}} = 10.33/0.7984 = 12.94$$

and the p-value is the probability of encountering a standard normal random variable of this value or greater. This is so close to zero I had trouble getting sensible numbers out of MATLAB; the evidence very strongly rejects this hypothesis. A version of this example, using the more complex two-sample t-test, is worked in the NIST/SEMATECH e-Handbook of Statistical Methods, at <http://www.itl.nist.gov/div898/handbook/>, as of 2013.

Worked example 10.10 *Do women and men have the same body temperature?*

Using the dataset of human temperatures at , assess the evidence against the hypothesis that gender 1 has the same mean temperature as gender 2

Solution: We compute the mean temperatures and standard errors shown

Gender:	1	2	
Mean:	98.10	98.39	The null hypothesis is that
Std Error:	0.0867	0.0922	

in the table. these two are the same, so the test statistic is

$$\frac{\text{difference}}{\text{std error}} = \frac{98.39 - 98.10}{\sqrt{0.0867^2 + 0.0922^2}} = 2.285$$

and we must ask what is the probability of getting a number with absolute value this big, or bigger, from a normal distribution (two-sided test). This is 0.0223. The evidence in this data set does not support the null hypothesis.

10.3.3 Variants on the Basic Test

The test of procedure 3 is usually referred to as a **two-sided** test. This is because it computes

$$P(\{X > |s|\}) \cup P(\{X < -|s|\}).$$

This is the fraction of samples that would produce a value that is larger than $|s|$ or smaller than $-|s|$. In some cases, we can expect that only the larger (or smaller) value is of interest. This yields a **one-sided** test; the relevant expression is either

$$P(\{X > s\})$$

or

$$P(\{X < s\}).$$

Generally, it is more conservative to use a two-sided test, and one should do so unless there is a good reason not to. Very often, authors use one-sided tests because they result in smaller p-values, and small p-values are often a requirement for publication.

Z-Tests and T-Tests

Our procedure works because we were able to identify a random variable where: (a) the random variable represents the effect of drawing a sample of a population at random; (b) we knew the distribution of the random variable conditioned on the hypothesis; (c) we can evaluate the random variable for our particular sample. Now use the notation of procedure 3. In that procedure, we computed

$$P(\{X > |s|\}) \cup P(\{X < -|s|\}).$$

When we have a large sample, the sample mean is a normal random variable whose mean is the population mean and whose standard deviation is the standard error. This means that we know its distribution. In this case, the test is known as a **z-test**.

An important difficulty with this procedure is we assumed we know the standard error. As figure 10.2 suggested, estimating standard error with $\text{std}(x)/\sqrt{k}$ works fine for large k , but is less successful for small k . Usually, practical advice suggests that one should use a Z-test only if the sample has at least 30 elements.

When the sample is small, the sample standard deviation is a poor estimate of the population standard deviation. The value of the sample mean that we compute minimizes the sample standard deviation, which means that the estimate tends to be a little too small. In turn, the standard error is a little too small, and there is slightly more probability that the sample mean is far from the population mean than the normal model allows for. This can be corrected for. Instead of using the standard deviation of the sample to form the standard error, we use

$$\sqrt{\frac{\sum_i (x_i - \text{mean}(\{x_i\}))^2}{k - 1}}.$$

When we test, instead of using the normal distribution to compute probabilities, we use **Student's t-distribution**. This is a family of probability distributions. There are two parameters; the observed value s , and the number of degrees of freedom.

The number of degrees of freedom is $k - 1$ for our purposes. When the number of degrees of freedom is small, the t-distribution has rather heavier tails than the normal distribution, so the test takes into account that the standard error may be larger than we think (because the population standard deviation is larger than we expected). When the number of degrees of freedom is large, the t-distribution is very similar to the normal distribution. One can get p-values from tables, or by the Matlab function `ttest`.

10.4 χ^2 TESTS: IS DATA CONSISTENT WITH A MODEL?

Sometimes we have a model, and we would like to know whether the data is consistent with that model. For example, imagine we have a six-sided die. We throw it many times, and record which number comes up each time. We would like to know if the die is fair (i.e. is the data consistent with the model that the die is fair). It is highly unlikely that each face comes up the same number of times, even if the die is fair. Instead, there will be some variation in the frequencies observed; with what probability is that variation, or bigger, the result of chance effects?

As another example, we decide that the number of calls by a telemarketer in each hour is distributed with a Poisson distribution. We don't know the intensity. We could collect call data, and use maximum likelihood to determine the intensity. Once we have the best estimate of the intensity, we still want to know whether the model is consistent with the data.

In each case, the model predicts the frequencies of events. For the six-sided die case, the model tells us how often we expect to see each side. For the call case, the model predicts how often we would see no calls, one call, two calls, three calls, etc. in each hour. To tell whether the model fits the data, we need to compare the frequencies we observed with theoretical frequencies.

The appropriate test is a χ^2 (say "khi-squared") test. Assume we have a set of disjoint events $\mathcal{E}_1, \dots, \mathcal{E}_k$ which cover the space of outcomes (i.e. any outcome lies in one of these events). Assume we perform k experiments, and record the number of times each event occurs. We have a null hypothesis regarding the probability of events. We can take the probability of each event and multiply by k to get a frequency under the null hypothesis. Now write $f_o(\mathcal{E}_i)$ for the observed frequency of event i ; $f_t(\mathcal{E}_i)$ for the theoretical frequency of the event under the null hypothesis. We form the statistic

$$\sum_i \frac{(f_o(\mathcal{E}_i) - f_t(\mathcal{E}_i))^2}{f_t(\mathcal{E}_i)}$$

which compares the observed and actual frequency of events. It turns out that this statistic has a distribution very close to a known form, called the χ^2 distribution, as long as each count is 5 or more. The distribution has two parameters; the statistic, and the number of degrees of freedom. The number of degrees of freedom to use for a straightforward test is $k - 1$; if one has to estimate a total of p parameters for the null hypothesis, this number is $k - p - 1$.

After this, things follow the usual recipe. We compute the statistic; we then look at tables, or use the matlab function `chi2cdf`, to find the probability that the statistic takes this value or greater under the null hypothesis. If this is small, then we reject the null hypothesis.

Worked example 10.11 χ^2 test for dice

I throw a die 100 times. I record the outcomes, in the table below. Is this a fair die?

face	count
1	46
2	13
3	12
4	11
5	9
6	9

Solution: The expected frequency is $100/6$ for each face. The χ^2 statistic has the value 62.7, and there are 5 degrees of freedom. We get the significance as $1-\text{chi2cdf}(62.7, 5)$, which is (basically) $3\text{e-}12$. You would have to run this experiment $3\text{e}11$ times to see a table as skewed as this once, by chance. The die is not fair. Notice the most important bit of this example, which is how to get the number out of matlab.

Worked example 10.12 *Is swearing Poisson?*

A famously swearsy politician gives a talk. You listen to the talk, and for each of 30 intervals 1 minute long, you record the number of swearwords. You record this as a histogram (i.e. you count the number of intervals with zero swear words, with one, etc.).

no. of swear words	no. of intervals
0	13
1	9
2	8
3	5
4	5

The null hypothesis is that the politician's swearing is Poisson distributed, with intensity (λ) one. Can you reject this null hypothesis?

Solution: If the null hypothesis is true, then the probability of getting n swear words in a fixed length interval would be $\frac{\lambda^n e^{-\lambda}}{n!}$. There are 10 intervals, so the theoretical frequencies are 10 times the following probabilities

number of swear words	probability
0	0.368
1	0.368
2	0.184
3	0.061
4	0.015

so the χ^2 statistic takes the value 243.1 and there are 4 degrees of freedom. The significance `1-chi2cdf(243.1, 4)` is indistinguishable from zero by matlab, so you can firmly reject the null hypothesis. Of course, the intensity may be wrong; but we don't know how to deal with that, and will have to discuss that question in the next chapter.

Worked example 10.13 *Is gender 2 temperature normal?*

Recall we used body temperature data for two genders in earlier examples. The gender 2 data looked as though it might not be normal. Use a χ^2 test to tell whether it is normal with mean 98.4° and standard deviation 0.743 or not.

Solution: The null hypothesis is that the data is normal with mean 98.4° and standard deviation 0.743 . We break up the range into five buckets (less than $97.65 = 98.4 - 0.74$; between 97.65 and $98 = 98.4 - 0.743/2$; between 98 and $98.76 = 98.4 + 0.743/2$; and greater than 99.14). With a little work with error functions, we get that the theoretical frequency in each bucket under the null hypothesis is $(10.3126; 9.7423; 24.8901; 9.7423; 10.3126)$. The actual frequencies are $(7; 13; 26; 12; 7)$. The χ^2 statistic is about $4e3$, and the significance level is essentially zero. This data isn't normal with the parameters given (though it might be normal with a different mean and a different standard deviation). Looking at the frequencies suggests the problem; there are far too few temperatures far away from the mean for this to be normal, and far too many in the center bucket.

Extracting Important Relationships in High Dimensions

Chapter 4 described methods to explore the relationship between two elements in a dataset. We could extract a pair of elements and construct various plots. For vector data, we could also compute the correlation between different pairs of elements. But if each data item is d -dimensional, there could be a lot of pairs to deal with.

We will think of our dataset as a collection of d dimensional vectors. It turns out that there are easy generalizations of our summaries. However, is hard to plot d -dimensional vectors. We need to find some way to make them fit on a 2-dimensional plot. Some simple methods can offer insights, but to really get what is going on we need methods that can at all pairs of relationships in a dataset in one go.

These methods visualize the dataset as a “blob” in a d -dimensional space. Many such blobs are flattened in some directions, because components of the data are strongly correlated. Finding the directions in which the blobs are flat yields methods to compute lower dimensional representations of the dataset.

11.1 SUMMARIES AND SIMPLE PLOTS

In this chapter, we assume that our data items are vectors. This means that we can add and subtract values and multiply values by a scalar without any distress. This is an important assumption, but it doesn’t necessarily mean that data is continuous (for example, you can meaningfully add the number of children in one family to the number of children in another family). It does rule out a lot of discrete data. For example, you can’t add “sports” to “grades” and expect a sensible answer.

Notation: Our data items are vectors, and we write a vector as \mathbf{x} . The data items are d -dimensional, and there are N of them. The entire data set is $\{\mathbf{x}\}$. When we need to refer to the i ’th data item, we write \mathbf{x}_i . We write $\{\mathbf{x}_i\}$ for a new dataset made up of N items, where the i ’th item is \mathbf{x}_i . If we need to refer to the j ’th component of a vector \mathbf{x}_i , we will write $x_i^{(j)}$ (notice this isn’t in bold, because it is a component not a vector, and the j is in parentheses because it isn’t a power). Vectors are always column vectors.

11.1.1 The Mean

For one-dimensional data, we wrote

$$\text{mean}(\{x\}) = \frac{\sum_i x_i}{N}.$$

This expression is meaningful for vectors, too, because we can add vectors and divide by scalars. We write

$$\text{mean}(\{\mathbf{x}\}) = \frac{\sum_i \mathbf{x}_i}{N}$$

and call this the mean of the data. Notice that each component of $\text{mean}(\{\mathbf{x}\})$ is the mean of that component of the data. There is not an easy analogue of the median, however (how do you order high dimensional data?) and this is a nuisance. Notice that, just as for the one-dimensional mean, we have

$$\text{mean}(\{\mathbf{x} - \text{mean}(\{\mathbf{x}\})\}) = 0$$

(i.e. if you subtract the mean from a data set, the resulting data set has zero mean).

11.1.2 Parallel Plots

Parallel plots can sometimes reveal information, particularly when the dimension of the dataset is low. To construct a parallel plot, you compute a normalized representation of each component of each data item. The component is normalized by translating and scaling so that the minimum value over the dataset is zero, and the maximum value over the dataset is one. Now write the i 'th normalised data item as (n_1, n_2, \dots, n_d) . For this data item, you plot a broken line joining $(1, n_1)$ to $(2, n_2)$ to $(3, n_3)$, etc. These plots are superimposed on one another. In the case of the bodyfat dataset, this yields the plot of figure 11.1.

Some structures in the parallel plot are revealing. Outliers often stick out (in figure 11.1, it's pretty clear that there's a data point with a very low height value, and also one with a very large weight value). Outliers affect the scaling, and so make other structures difficult to spot. I have removed them for figure 11.2. In this figure, you can see that two negatively correlated components next to one another produce a butterfly like shape (bodyfat and density). In this plot, you can also see that there are still several data points that are very different from others (two data items have ankle values that are very different from the others, for example).

11.1.3 Using Covariance to encode Variance and Correlation

Variance, standard deviation and correlation can each be seen as an instance of a more general operation on data. Assume that we have two one dimensional data sets $\{x\}$ and $\{y\}$. Then we can define the **covariance** of $\{x\}$ and $\{y\}$.

Definition: 11.1 *Covariance*

Assume we have two sets of N data items, $\{x\}$ and $\{y\}$. We compute the covariance by

$$\text{cov}(\{x\}, \{y\}) = \frac{\sum_i (x_i - \text{mean}(\{x\}))(y_i - \text{mean}(\{y\}))}{N}$$

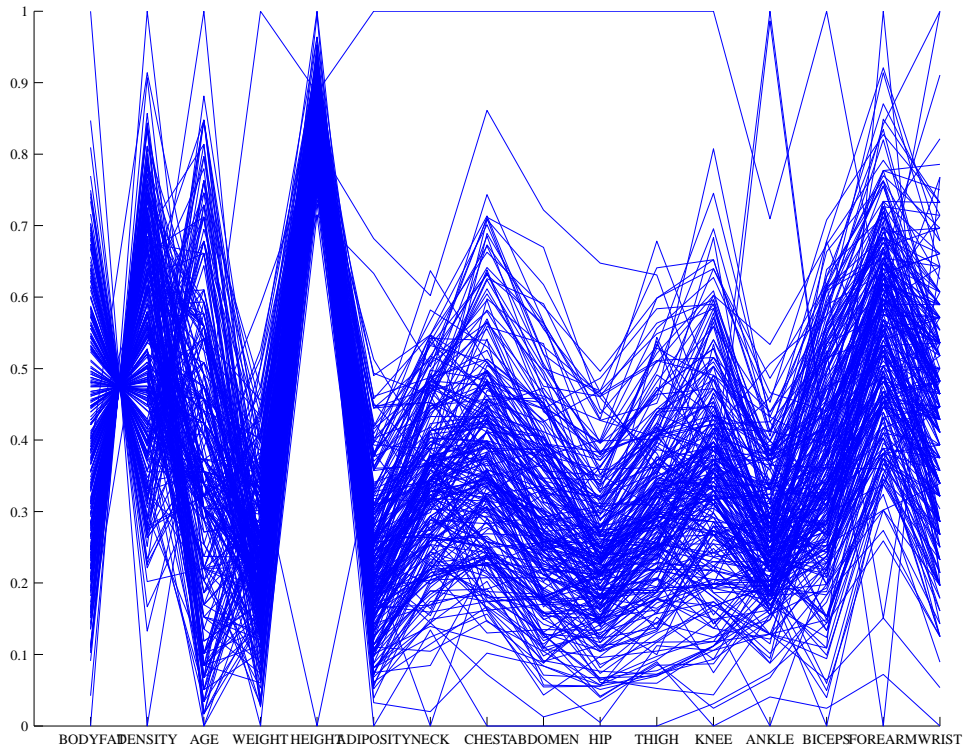


FIGURE 11.1: A parallel plot of the bodyfat dataset, including all data points. I have named the components on the horizontal axis. It is easy to see that large values of bodyfat correspond to small values of density, and vice versa. Notice that one datapoint has height very different from all others; similarly, one datapoint has weight very different from all others.

Covariance measures the tendency of corresponding elements of $\{x\}$ and of $\{y\}$ to be larger than (resp. smaller than) the mean. Just like mean, standard deviation and variance, covariance can refer either to a property of a dataset (as in the definition here) or a particular expectation (as in chapter 6). The correspondence is defined by the order of elements in the data set, so that x_1 corresponds to y_1 , x_2 corresponds to y_2 , and so on. If $\{x\}$ tends to be larger (resp. smaller) than its mean for data points where $\{y\}$ is also larger (resp. smaller) than its mean, then the covariance should be positive. If $\{x\}$ tends to be larger (resp. smaller) than its mean for data points where $\{y\}$ is smaller (resp. larger) than its mean, then the covariance should be negative.

From this description, it should be clear we have seen examples of covariance already. Notice that

$$\text{std}(x)^2 = \text{var}(\{x\}) = \text{cov}(\{x\}, \{x\})$$

which you can prove by substituting the expressions. Recall that variance measures

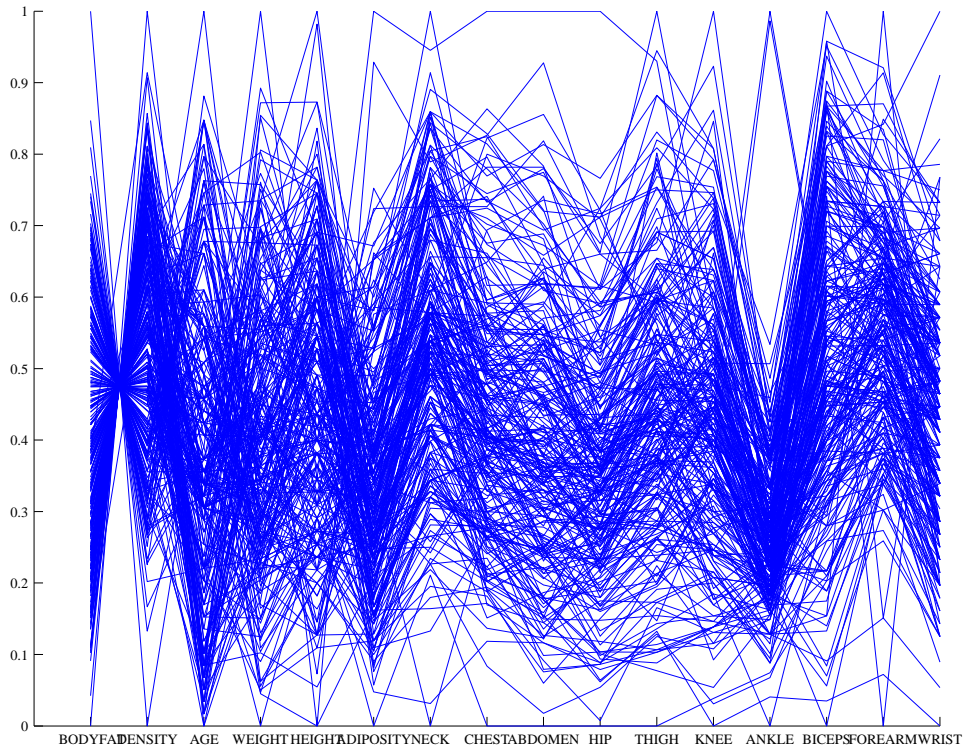


FIGURE 11.2: A plot with those data items removed, so that those components are renormalized. Two datapoints have rather distinct ankle measurements. Generally, you can see that large knees go with large ankles and large biceps (the v structure).

the tendency of a dataset to be different from the mean, so the covariance of a dataset with itself is a measure of its tendency not to be constant.

More important, notice that

$$\text{corr}(\{x, y\}) = \frac{\text{cov}(\{x\}, \{y\})}{\sqrt{\text{cov}(\{x\}, \{x\})} \sqrt{\text{cov}(\{y\}, \{y\})}}.$$

This is occasionally a useful way to think about correlation. It says that the correlation measures the tendency of $\{x\}$ and $\{y\}$ to be larger (resp. smaller) than their means for the same data points, *compared to* how much they change on their own.

Working with covariance (rather than correlation) allows us to unify some ideas. In particular, for data items which are d dimensional vectors, it is straightforward to compute a single matrix that captures all covariances between all pairs of components — this is the **covariance matrix**.

Definition: 11.2 *Covariance Matrix*

The covariance matrix is:

$$\text{Covmat}(\{\mathbf{x}\}) = \frac{\sum_i (\mathbf{x}_i - \text{mean}(\{\mathbf{x}\}))(\mathbf{x}_i - \text{mean}(\{\mathbf{x}\}))^T}{N}$$

Notice that it is quite usual to write a covariance matrix as Σ , and we will follow this convention.

Properties of the Covariance Matrix Covariance matrices are often written as Σ , whatever the dataset (you get to figure out precisely which dataset is intended, from context). Generally, when we want to refer to the j , k 'th entry of a matrix \mathcal{A} , we will write \mathcal{A}_{jk} , so Σ_{jk} is the covariance between the j 'th and k 'th components of the data.

- The j , k 'th entry of the covariance matrix is the covariance of the j 'th and the k 'th components of \mathbf{x} , which we write $\text{cov}(\{x^{(j)}\}, \{x^{(k)}\})$.
- The j , j 'th entry of the covariance matrix is the variance of the j 'th component of \mathbf{x} .
- The covariance matrix is symmetric.
- The covariance matrix is always positive semi-definite; it is positive definite, *unless* there is some vector \mathbf{a} such that $\mathbf{a}^T(\mathbf{x}_i - \text{mean}(\{\mathbf{x}_i\})) = 0$ for all i .

Proposition:

$$\text{Covmat}(\{\mathbf{x}\})_{jk} = \text{cov}(\{x^{(j)}\}, \{x^{(k)}\})$$

Proof: Recall

$$\text{Covmat}(\{\mathbf{x}\}) = \frac{\sum_i (\mathbf{x}_i - \text{mean}(\{\mathbf{x}\}))(\mathbf{x}_i - \text{mean}(\{\mathbf{x}\}))^T}{N}$$

and the j , k 'th entry in this matrix will be

$$\frac{\sum_i (x_i^{(j)} - \text{mean}(\{x^{(j)}\}))(x_i^{(k)} - \text{mean}(\{x^{(k)}\}))^T}{N}$$

which is $\text{cov}(\{x^{(j)}\}, \{x^{(k)}\})$.

Proposition:

$$\text{Covmat}(\{\mathbf{x}_i\})_{jj} = \Sigma_{jj} = \text{var}\left(\{x^{(j)}\}\right)$$

Proof:

$$\begin{aligned} \text{Covmat}(\{\mathbf{x}\})_{jj} &= \text{cov}\left(\{x^{(j)}\}, \{x^{(j)}\}\right) \\ &= \text{var}\left(\{x^{(j)}\}\right) \end{aligned}$$

Proposition:

$$\text{Covmat}(\{\mathbf{x}\}) = \text{Covmat}(\{\mathbf{x}\})^T$$

Proof: We have

$$\begin{aligned} \text{Covmat}(\{\mathbf{x}\})_{jk} &= \text{cov}\left(\{x^{(j)}\}, \{x^{(k)}\}\right) \\ &= \text{cov}\left(\{x^{(k)}\}, \{x^{(j)}\}\right) \\ &= \text{Covmat}(\{\mathbf{x}\})_{kj} \end{aligned}$$

Proposition: Write $\Sigma = \text{Covmat}(\{\mathbf{x}\})$. If there is no vector \mathbf{a} such that $\mathbf{a}^T(\mathbf{x}_i - \text{mean}(\{\mathbf{x}\})) = 0$ for all i , then for any vector \mathbf{u} , such that $\|\mathbf{u}\| > 0$,

$$\mathbf{u}^T \Sigma \mathbf{u} > 0.$$

If there is such a vector \mathbf{a} , then

$$\mathbf{u}^T \Sigma \mathbf{u} \geq 0.$$

Proof: We have

$$\begin{aligned} \mathbf{u}^T \Sigma \mathbf{u} &= \frac{1}{N} \sum_i [\mathbf{u}^T (\mathbf{x}_i - \text{mean}(\{\mathbf{x}\}))] [(\mathbf{x}_i - \text{mean}(\{\mathbf{x}\}))^T \mathbf{u}] \\ &= \frac{1}{N} \sum_i [\mathbf{u}^T (\mathbf{x}_i - \text{mean}(\{\mathbf{x}\}))]^2. \end{aligned}$$

Now this is a sum of squares. If there is some \mathbf{a} such that $\mathbf{a}^T(\mathbf{x}_i - \text{mean}(\{\mathbf{x}\})) = 0$ for every i , then the covariance matrix must be positive semidefinite (because the sum of squares could be zero in this case). Otherwise, it is positive definite, because the sum of squares will always be positive.

11.2 BLOB ANALYSIS OF HIGH-DIMENSIONAL DATA

When we plotted histograms, we saw that mean and variance were a very helpful description of data that had a unimodal histogram. If the histogram had more than one mode, one needed to be somewhat careful to interpret the mean and variance; in the pizza example, we plotted diameters for different manufacturers to try and see the data as a collection of unimodal histograms.

Generally, mean and covariance are a good description of data that lies in a “blob” (Figure 11.3). You might not believe that this is a technical term, but it’s quite widely used. This is because mean and covariance supply a natural coordinate system in which to interpret the blob. Mean and covariance are less useful as descriptions of data that forms multiple blobs (Figure 11.3). In chapter 1, we discuss methods to model data that forms multiple blobs, or other shapes that we will interpret as a set of blobs. But many datasets really are single blobs, and we concentrate on such data here. The way to understand a blob is to think about the coordinate transformations that place a blob into a particularly convenient form.

11.2.1 Understanding Blobs with Scatterplot Matrices

Plotting high dimensional data is tricky. One strategy that is very useful when there aren’t too many dimensions is to use a scatterplot matrix. To build one,

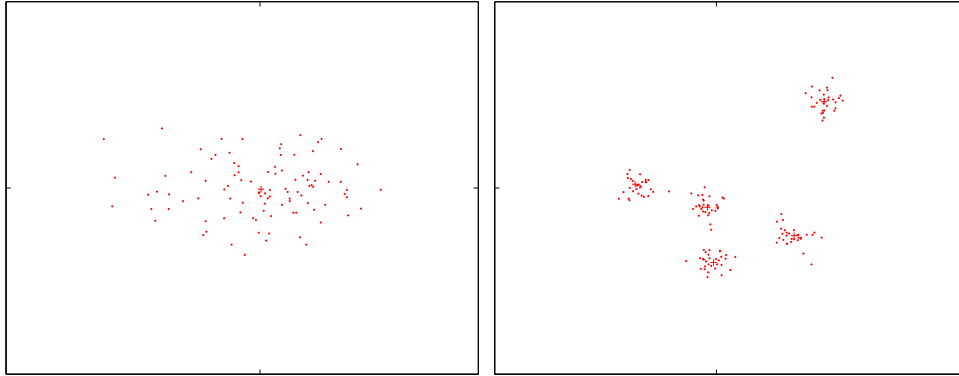


FIGURE 11.3: *On the left*, a “blob” in two dimensions. This is a set of data points that lie somewhat clustered around a single center, given by the mean. I have plotted the mean of these data points with a ‘+’. *On the right*, a data set that is best thought of as a collection of five blobs. I have plotted the mean of each with a ‘+’. We could compute the mean and covariance of this data, but it would be less revealing than the mean and covariance of a single blob. In chapter 1, I will describe automatic methods to describe this dataset as a series of blobs.

you lay out scatterplots for each pair of variables in a matrix. On the diagonal, you name the variable that is the vertical axis for each plot in the row, and the horizontal axis in the column. This sounds more complicated than it is; look at the example of figure 11.4, which shows a scatterplot matrix for four of the variables in the height weight dataset of <http://www2.stetson.edu/~jrasp/data.htm>; look for `bodyfat.xls` at that URL). This is originally a 16-dimensional dataset, but a 16 by 16 scatterplot matrix is squashed and hard to interpret.

What is nice about this kind of plot is that it’s quite easy to spot correlations between pairs of variables, though you do need to take into account the coordinates have not been normalized. For figure 11.4, you can see that weight and adiposity appear to show quite strong correlations, but weight and age are pretty weakly correlated. Height and age seem to have a low correlation. It is also easy to visualize unusual data points. Usually one has an interactive process to do so — you can move a “brush” over the plot to change the color of data points under the brush. To show what might happen, figure 11.5 shows a scatter plot matrix with some points shown as circles. Notice how they lie inside the “blob” of data in some views, and outside in others. This is an effect of projection.

11.2.2 Transforming High Dimensional Data

Assume we apply an affine transformation to our data set $\{\mathbf{x}\}$, to obtain a new dataset $\{\mathbf{u}\}$, where $\mathbf{u}_i = \mathcal{A}\mathbf{x}_i + \mathbf{b}$. Here \mathcal{A} is any matrix (it doesn’t have to be square, or symmetric, or anything else; it just has to have second dimension d). It

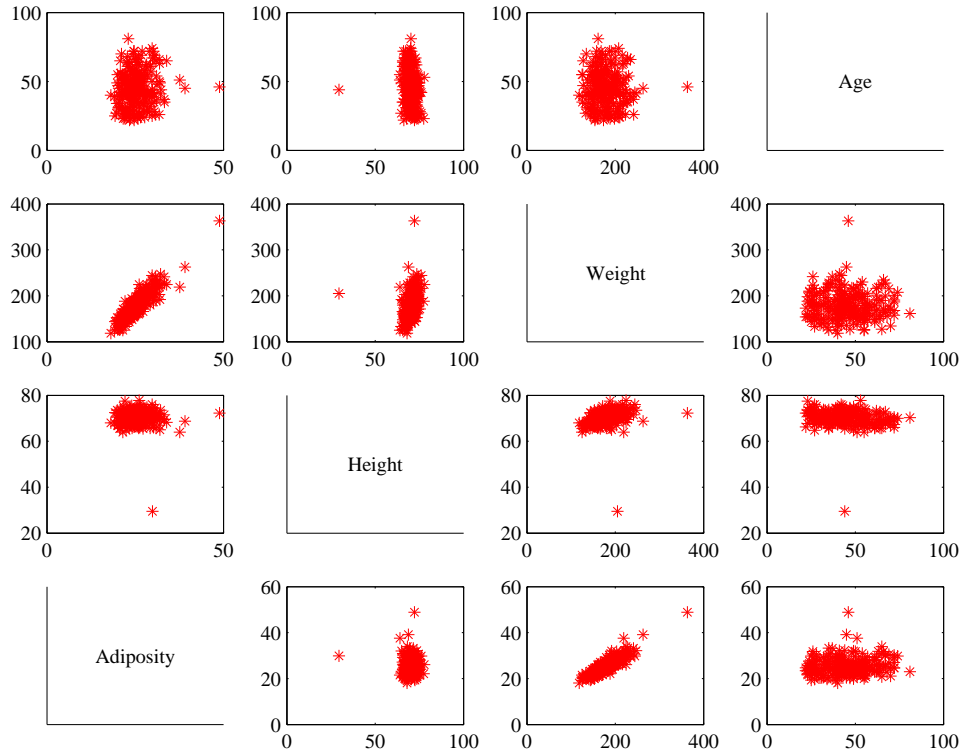


FIGURE 11.4: This is a scatterplot matrix for four of the variables in the height weight dataset of <http://www2.stetson.edu/~jrasp/data.htm>. Each plot is a scatterplot of a pair of variables. The name of the variable for the horizontal axis is obtained by running your eye down the column; for the vertical axis, along the row. Although this plot is redundant (half of the plots are just flipped versions of the other half), that redundancy makes it easier to follow points by eye. You can look at a column, move down to a row, move across to a column, etc. Notice how you can spot correlations between variables and outliers (the arrows).

is easy to compute the mean and covariance of $\{\mathbf{u}\}$. We have

$$\begin{aligned} \text{mean}(\{\mathbf{u}\}) &= \text{mean}(\{\mathcal{A}\mathbf{x} + \mathbf{b}\}) \\ &= \mathcal{A}\text{mean}(\{\mathbf{x}\}) + \mathbf{b}, \end{aligned}$$

so you get the new mean by multiplying the original mean by \mathcal{A} and adding \mathbf{b} .

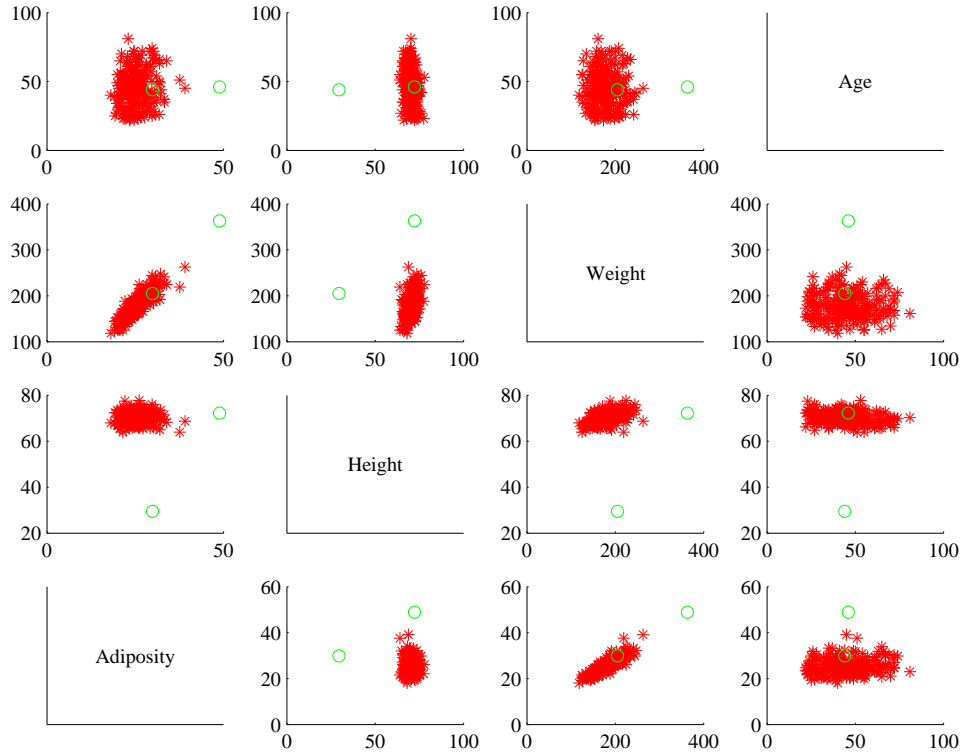


FIGURE 11.5: You should compare this figure with figure 11.4. I have marked two data points with circles in this figure; notice that in some panels these are far from the rest of the data, in others close by. A “brush” in an interactive application can be used to mark data like this to allow a user to explore a dataset.

The new covariance matrix is easy to compute as well. We have:

$$\begin{aligned}
 \text{Covmat}(\{\mathbf{u}\}) &= \text{Covmat}(\{\mathcal{A}\mathbf{x} + \mathbf{b}\}) \\
 &= \frac{\sum_i (\mathbf{u}_i - \text{mean}(\{\mathbf{u}\}))(\mathbf{u}_i - \text{mean}(\{\mathbf{u}\}))^T}{N} \\
 &= \frac{\sum_i (\mathcal{A}\mathbf{x}_i + \mathbf{b} - \mathcal{A}\text{mean}(\{\mathbf{x}\}) - \mathbf{b})(\mathcal{A}\mathbf{x}_i + \mathbf{b} - \mathcal{A}\text{mean}(\{\mathbf{x}\}) - \mathbf{b})^T}{N} \\
 &= \frac{\mathcal{A} \sum_i (\mathbf{x}_i - \text{mean}(\{\mathbf{x}\}))(\mathbf{x}_i - \text{mean}(\{\mathbf{x}\}))^T \mathcal{A}^T}{N} \\
 &= \mathcal{A} \text{Covmat}(\{\mathbf{x}\}) \mathcal{A}^T.
 \end{aligned}$$

11.2.3 Transforming Blobs

The trick to interpreting high dimensional data is to use the mean and covariance to understand the blob. Figure 11.6 shows a two-dimensional data set. Notice that there is obviously some correlation between the x and y coordinates (it’s a diagonal blob), and that neither x nor y has zero mean. We can easily compute the mean

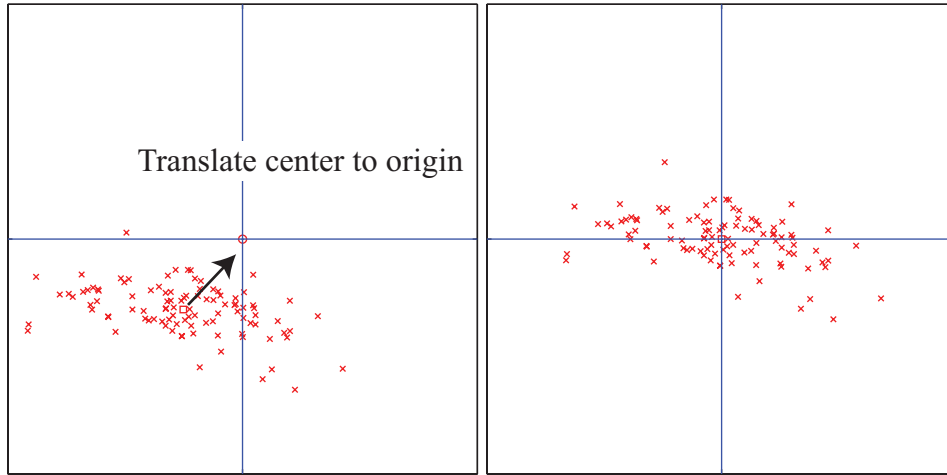


FIGURE 11.6: On the **left**, a “blob” in two dimensions. This is a set of data points that lie somewhat clustered around a single center, given by the mean. I have plotted the mean of these data points with a hollow square (it’s easier to see when there is a lot of data). To translate the blob to the origin, we just subtract the mean from each datapoint, yielding the blob on the **right**.

and subtract it from the data points, and this translates the blob so that the origin is at the center (Figure 11.6). In coordinates, this means we compute the new dataset $\{\mathbf{u}\}$ from the old dataset $\{\mathbf{x}\}$ by the rule $\mathbf{u}_i = \mathbf{x}_i - \text{mean}(\{\mathbf{x}\})$. This new dataset has been translated so that the mean is zero.

Once this blob is translated (Figure 11.7, left), we can rotate it as well. It is natural to try to rotate the blob so that there is no correlation between distinct pairs of dimensions. We can do so by diagonalizing the covariance matrix. In particular, let \mathcal{U} be the matrix formed by stacking the eigenvectors of $\text{Covmat}(\{\mathbf{x}\})$ into a matrix (i.e. $\mathcal{U} = [\mathbf{v}_1, \dots, \mathbf{v}_d]$, where \mathbf{v}_j are eigenvectors of the covariance matrix). We now form the dataset $\{\mathbf{n}\}$, using the rule

$$\mathbf{n}_i = \mathcal{U}^T \mathbf{u}_i = \mathcal{U}^T (\mathbf{x}_i - \text{mean}(\{\mathbf{x}\})).$$

The mean of this new dataset is clearly $\mathbf{0}$. The covariance of this dataset is

$$\begin{aligned} \text{Covmat}(\{\mathbf{n}\}) &= \text{Covmat}(\{\mathcal{U}^T \mathbf{x}\}) \\ &= \mathcal{U}^T \text{Covmat}(\{\mathbf{x}\}) \mathcal{U} \\ &= \Lambda, \end{aligned}$$

where Λ is a diagonal matrix of eigenvalues of $\text{Covmat}(\{\mathbf{x}\})$. Remember that, in describing diagonalization, we adopted the convention that the eigenvectors of the matrix being diagonalized were ordered so that the eigenvalues are sorted in descending order along the diagonal of Λ . We now have two very useful facts about $\{\mathbf{n}\}$: (a) every pair of distinct components has covariance zero, and so has correlation zero; (b) the first component has the highest variance, the second component

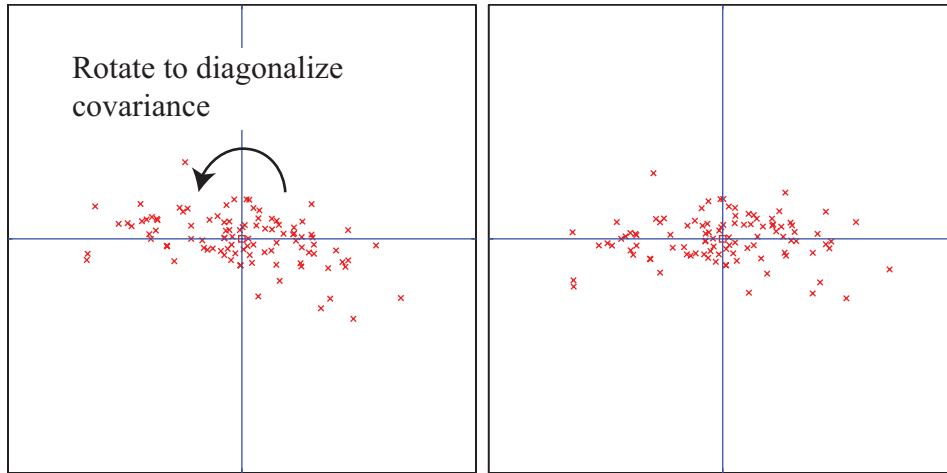


FIGURE 11.7: *On the left*, the translated blob of figure 11.6. This blob lies somewhat diagonally, because the vertical and horizontal components are correlated. *On the right*, that blob of data rotated so that there is no correlation between these components. We can now describe the blob by the vertical and horizontal variances alone, as long as we do so in the new coordinate system. In this coordinate system, the vertical variance is significantly larger than the horizontal variance — the blob is short and wide.

has the second highest variance, and so on. We can rotate and translate any blob into a coordinate system that has these properties. *In this coordinate system*, we can describe the blob simply by giving the variances of each component — the covariances are zero.

Translating a blob of data doesn't change the scatterplot matrix in any interesting way (the axes change, but the picture doesn't). Rotating a blob produces really interesting results, however. Figure 11.9 shows the dataset of figure 11.4, translated to the origin and rotated to diagonalize it. Now we do not have names for each component of the data (they're linear combinations of the original components), but each pair is now not correlated. This blob has some interesting shape features. Figure 11.9 shows the gross shape of the blob best. Each panel of this figure has the same scale in each direction. You can see the blob extends about 80 units in direction 1, but only about 15 units in direction 2, and much less in the other two directions. You should think of this blob as being rather cigar-shaped; it's long in one direction, but there isn't much in the others. The cigar metaphor isn't perfect because there aren't any 4 dimensional cigars, but it's helpful. You can think of each panel of this figure as showing views down each of the four axes of the cigar.

Now look at figure ???. This shows the same rotation of the same blob of data, but now the scales on the axis have changed to get the best look at the detailed shape of the blob. First, you can see that blob is a little curved (look at the projection onto direction 2 and direction 4). There might be some effect here

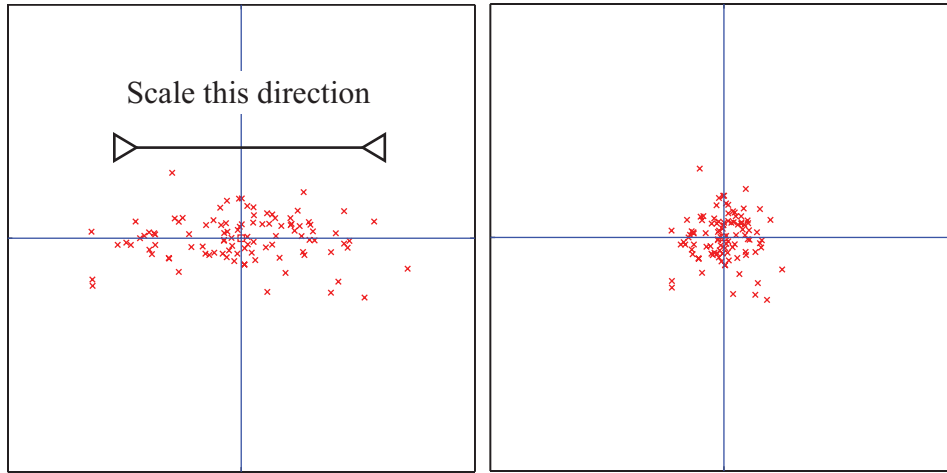


FIGURE 11.8: *On the left, the translated and rotated blob of figure 11.7. This blob is stretched — one direction has more variance than another. Because all covariances are zero, it is easy to scale the blob so that all variances are one (the blob on the right). You can think of this as a standard blob. All blobs can be reduced to a standard blob, by relatively straightforward linear algebra.*

worth studying. Second, you can see that some points seem to lie away from the main blob. I have plotted each data point with a dot, and the interesting points with a number. These points are clearly special in some way.

We could now scale the data in this new coordinate system so that all the variances are either one (if there is any variation in that direction) or zero (directions where the data doesn't vary — these occur only if some directions are functions of others). Figure 11.8 shows the final scaling. The result is a standard blob. Our approach applies to any dimension — I gave 2D figures because they're much easier to understand. There is a crucial point here: we can reduce any blob of data, in any dimension, to a standard blob of that dimension. All blobs are the same, except for some stretching, some rotation, and some translation. This is why blobs are so well-liked.

11.2.4 Whitening Data

It is sometimes useful to actually reduce a dataset to a standard blob. Doing so is known as **whitening the data** (for reasons I find obscure). This can be a sensible thing to do when we don't have a clear sense of the relative scales of the components of each data vector. For example, if we have a dataset where one component ranges from $1e5$ to $2e5$, and the other component ranges from $-1e-5$ to $1e-5$, we are likely to face numerical problems in many computations (adding small numbers to big numbers is often unwise). Often, this kind of thing follows from a poor choice of units, rather than any kind of useful property of the data. In such a case, it could be quite helpful to whiten the data. Another reason to whiten the data might be that we know relatively little about the meaning of each component. In this

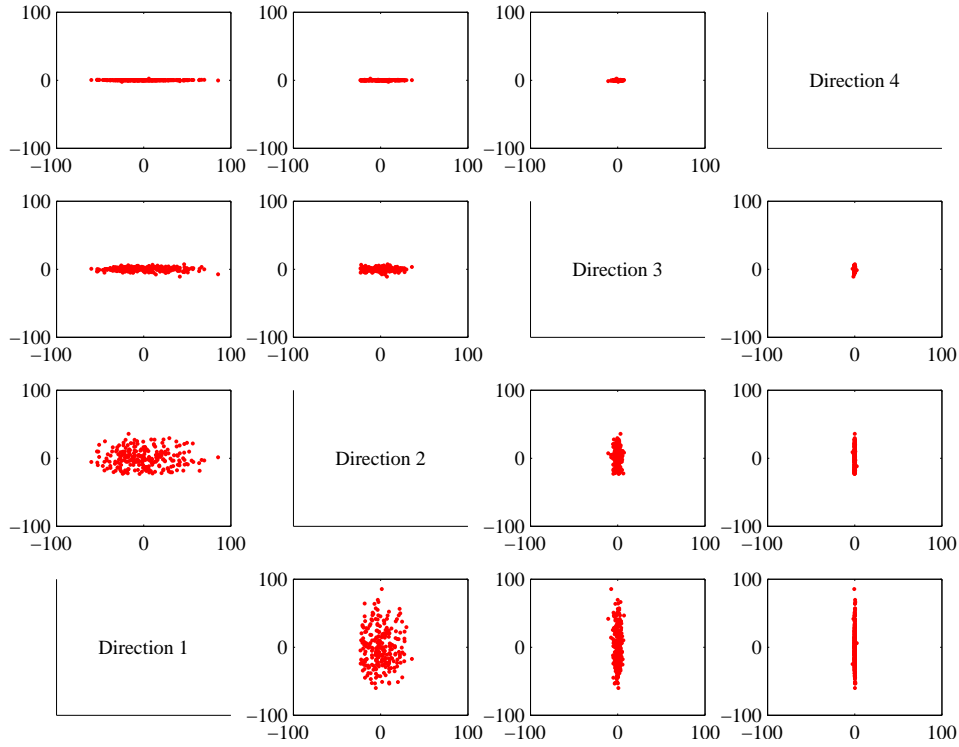


FIGURE 11.9: A panel plot of the bodyfat dataset of figure 11.4, now rotated so that the covariance between all pairs of distinct dimensions is zero. Now we do not know names for the directions — they’re linear combinations of the original variables. Each scatterplot is on the same set of axes, so you can see that the dataset extends more in some directions than in others.

case, the original choice of coordinate system was somewhat arbitrary anyhow, and transforming data to a uniform blob could be helpful.

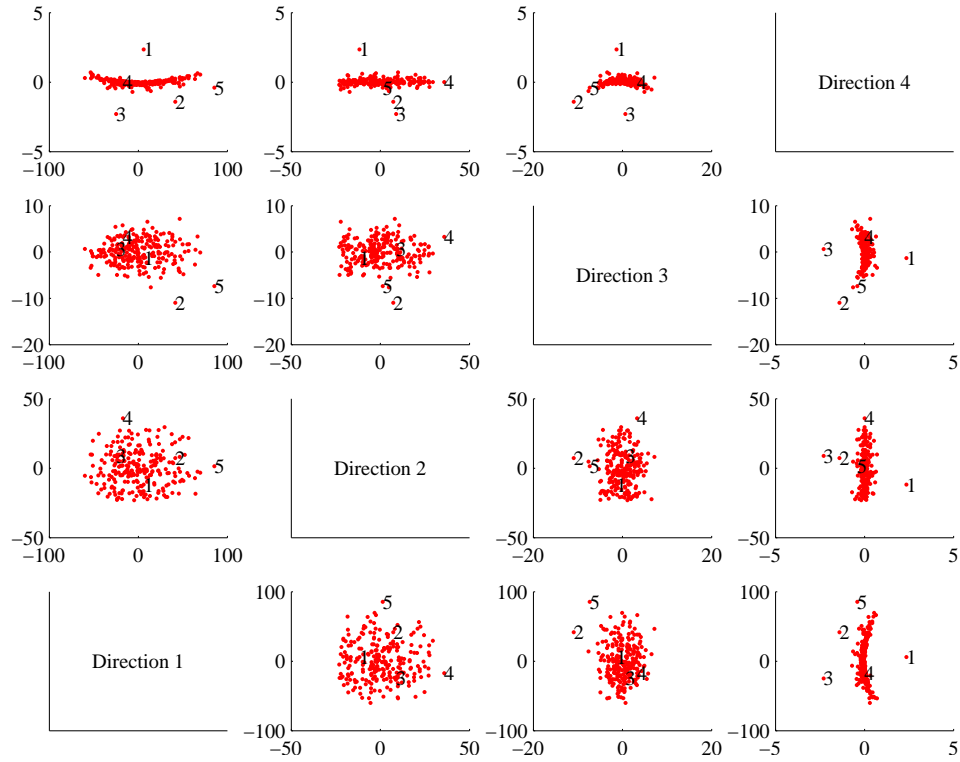


FIGURE 11.10: A panel plot of the bodyfat dataset of figure 11.4, now rotated so that the covariance between all pairs of distinct dimensions is zero. Now we do not know names for the directions — they’re linear combinations of the original variables. I have scaled the axes so you can see details; notice that the blob is a little curved, and there are several data points that seem to lie some way away from the blob, which I have numbered.

Useful Facts: 11.1 *Whitening a dataset*

For a dataset $\{\mathbf{x}\}$, compute:

- \mathcal{U} , the matrix of eigenvectors of $\text{Covmat}(\{\mathbf{x}\})$;
- and $\text{mean}(\{\mathbf{x}\})$.

Now compute $\{\mathbf{n}\}$ using the rule

$$\mathbf{n}_i = \mathcal{U}^T(\mathbf{x}_i - \text{mean}(\{\mathbf{x}\})).$$

Then $\text{mean}(\{\mathbf{n}\}) = \mathbf{0}$ and $\text{Covmat}(\{\mathbf{n}\})$ is diagonal.

Now write Λ for the diagonal matrix of eigenvalues of $\text{Covmat}(\{\mathbf{x}\})$ (so that $\text{Covmat}(\{\mathbf{x}\})\mathcal{U} = \mathcal{U}\Lambda$). Assume that each of the diagonal entries of Λ is greater than zero (otherwise there is a redundant dimension in the data). Write λ_i for the i 'th diagonal entry of Λ , and write $\Lambda^{-(1/2)}$ for the diagonal matrix whose i 'th diagonal entry is $1/\sqrt{\lambda_i}$. Compute $\{\mathbf{z}\}$ using the rule

$$\mathbf{z}_i = \Lambda^{-(1/2)}\mathcal{U}(\mathbf{x}_i - \text{mean}(\{\mathbf{x}\})).$$

We have that $\text{mean}(\{\mathbf{z}\}) = \mathbf{0}$ and $\text{Covmat}(\{\mathbf{z}\}) = \mathcal{I}$. The dataset $\{\mathbf{z}\}$ is often known as **whitened data**.

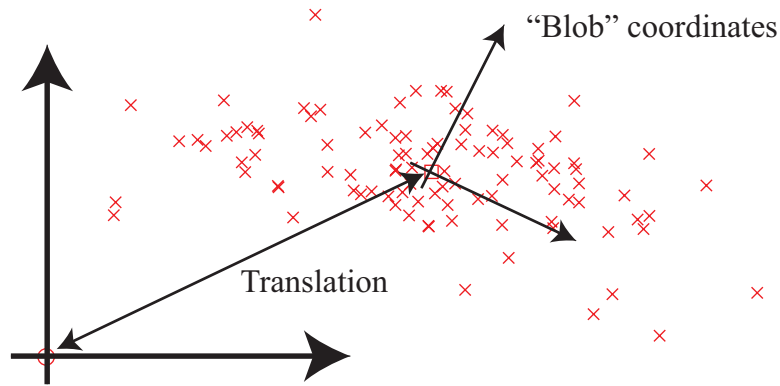


FIGURE 11.11: A 2D blob, with its natural blob coordinate system. The origin of this coordinate system is at the mean of the data. The coordinate axes are (a) at right angles to one another and (b) are directions that have no covariance.

It isn't always a good idea to whiten data. In some circumstances, each separate component is meaningful, and in a meaningful set of units. For example, one of the components might be a length using a natural scale and the other might be a time on a natural scale. When this happens, we might be reluctant to transform the data, either because we don't want to add lengths to times or because we want to preserve the scales.

11.3 PRINCIPAL COMPONENTS ANALYSIS

Mostly, when one deals with high dimensional data, it isn't clear which individual components are important. As we have seen with the height weight dataset (for example, in the case of density and weight) some components can be quite strongly correlated. Equivalently, as in Figure 1, the blob is not aligned with the coordinate axes.

11.3.1 The Blob Coordinate System and Smoothing

We can use the fact that we *could* rotate, translate and scale the blob to define a coordinate system within the blob. The origin of that coordinate system is the mean of the data, and the coordinate axes are given by the eigenvectors of the covariance matrix. These are orthonormal, so they form a set of unit basis vectors at right angles to one another (i.e. a coordinate system). You should think of these as blob coordinates; Figure 11.11 illustrates a set of blob coordinates.

The blob coordinate system is important because, *once we know the blob coordinate system*, we can identify important scales of variation in the data. For example, if you look at Figure 11.11, you can see that this blob is extended much further along one direction than along the other. We can use this information to identify the most significant forms of variation in very high dimensional data. In some directions in the blob coordinate system, the blob will be spread out — ie

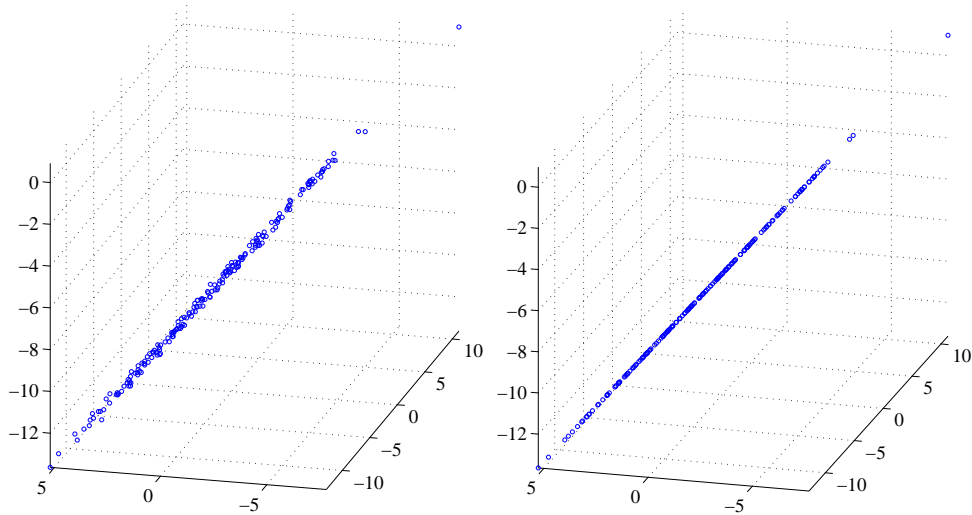


FIGURE 11.12: On the **left**, a blob of 3D data that has very low variance in two directions in the blob coordinates. As a result, all the data points are very close to a 1D blob. Experience shows that this is a common phenomenon. Although there might be many components in the data items, all data points are very close to a much lower dimensional object in the high dimensional space. When this is the case, we could obtain a lower dimensional representation of the data by working in blob coordinates, or we could smooth the data (as on the **right**), by projecting each data point onto the lower dimensional space.

have large variance — but in others, it might not be.

Equivalently, imagine we choose to represent each data item *in blob coordinates*. Then the mean over the dataset will be zero. Each pair of distinct coordinates will be uncorrelated. Some coordinates — corresponding to directions where the blob is spread out — will have a large range of values. Other coordinates — directions in which the blob is small — will have a small range of values. We could choose to replace these coordinates with zeros, with little significant loss in accuracy. The advantage of doing so is that we would have lower dimensional data to deal with.

However, it isn't particularly natural to work in blob coordinates. Each component of a data item may have a distinct meaning and scale (i.e. feet, pounds, and so on), but this is not preserved in any easy way in blob coordinates. Instead, we should like to (a) compute a lower dimensional representation in blob coordinates then (b) transform that representation into the original coordinate system of the data item. Doing so is a form of **smoothing** — suppressing small, irrelevant variations by exploiting multiple data items.

For example, look at Figure 11.12. Imagine we transform the blob on the left to blob coordinates. The covariance matrix in these coordinates is a 3×3 diagonal matrix. One of the values on the diagonal is large, because the blob is extended on one direction; but the other two are small. This means that, in blob coordinates, the

data varies significantly in one direction, but very little in the other two directions.

Now imagine we project the data points onto the high-variation direction; equivalently, we set the other two directions to zero for each data point. Each of the new data points is very close to the corresponding old data point, because by setting the small directions to zero we haven't moved the point very much. In blob coordinates, the covariance matrix of this new dataset has changed very little. It is again a 3×3 diagonal matrix, but now two of the diagonal values are zero, because there isn't any variance in those directions. The third value is large, because the blob is extended in that direction. We take the new dataset, and rotate and translate it into the original coordinate system. Each point must lie close to the corresponding point in the original dataset. However, the new dataset lies along a straight line (because it lay on a straight line in the blob coordinates). This process gets us the blob on the right in Figure 11.12. This blob is a smoothed version of the original blob.

Smoothing works because when two data items are strongly correlated, the value of one is a good guide to the value of the other. This principle works for more than two data items. Section 1 describes an example where the data items have dimension 101, but all values are extremely tightly correlated. In a case like this, there may be very few dimensions in blob coordinates that have any significant variation (3-6 for this case, depending on some details of what one believes is a small number, and so on). The components are so strongly correlated in this case that the 101-dimensional blob really looks like a slightly thickened 3 (or slightly more) dimensional blob that has been inserted into a 101-dimensional space (Figure 11.12). If we project the 101-dimensional data onto that structure in the original, 101-dimensional space, we may get much better estimates of the components of each data item than the original measuring device can supply. This occurs because each component is now estimated using correlations between all the measurements.

11.3.2 The Low-Dimensional Representation of a Blob

We wish to construct an r dimensional representation of a blob, where we have chosen r in advance. First, we compute $\{\mathbf{v}\}$ by translating the blob so its mean is at the origin, so that $\mathbf{v}_i = \mathbf{x}_i - \text{mean}(\{\mathbf{x}\})$. Now write $\mathcal{V} = [\mathbf{v}_1, \mathbf{v}_2, \dots, \mathbf{v}_N]$. The covariance matrix of $\{\mathbf{v}\}$ is then

$$\text{Covmat}(\{\mathbf{v}\}) = \frac{1}{N} \mathcal{V} \mathcal{V}^T = \text{Covmat}(\{\mathbf{x}\}).$$

Now write Λ for the diagonal matrix of eigenvalues of $\text{Covmat}(\{\mathbf{x}\})$ and \mathcal{U} for the matrix of eigenvectors, so that $\text{Covmat}(\{\mathbf{x}\})\mathcal{U} = \mathcal{U}\Lambda$. We assume that the elements of Λ are sorted in decreasing order along the diagonal. The covariance matrix for the dataset transformed into blob coordinates will be Λ . Notice that

$$\begin{aligned} \Lambda &= \mathcal{U}^T \text{Covmat}(\{\mathbf{x}\}) \mathcal{U} \\ &= \mathcal{U}^T \mathcal{V} \mathcal{V}^T \mathcal{U} \\ &= (\mathcal{U}^T \mathcal{V})(\mathcal{U}^T \mathcal{V})^T. \end{aligned}$$

This means we can interpret $(\mathcal{U}^T \mathcal{V})$ as a new dataset $\{\mathbf{b}\}$. This is our data, rotated into blob coordinates.

Now write Π_r for the $d \times d$ matrix

$$\begin{bmatrix} \mathcal{I}_r & 0 \\ 0 & 0 \end{bmatrix}$$

which projects a d dimensional vector onto its first r components, and replaces the others with zeros. Then we have that

$$\Lambda_r = \Pi_r \Lambda \Pi_r^T$$

is the covariance matrix for the reduced dimensional data in blob coordinates. Notice that Λ_r keeps the r largest eigenvalues on the diagonal of Λ , and replaces all others with zero.

We have

$$\begin{aligned} \Lambda_r &= \Pi_r \Lambda \Pi_r^T \\ &= \Pi_r \mathcal{U}^T \text{Covmat}(\{\mathbf{x}\}) \mathcal{U} \Pi_r^T \\ &= (\Pi_r \mathcal{U}^T \mathcal{V})(\mathcal{V}^T \mathcal{U} \Pi_r^T) \\ &= \mathcal{P} \mathcal{P}^T \end{aligned}$$

where $\mathcal{P} = (\Pi_r \mathcal{U}^T \mathcal{V})$. This represents our data, rotated into blob coordinates, and then projected down to r dimensions, with remaining terms replaced by zeros. Write $\{\mathbf{b}_r\}$ for this new dataset.

Occasionally, we need to refer to this representation, and we give it a special name. Write

$$\text{pcaproj}(\mathbf{x}_i, r, \{\mathbf{x}\}) = \Pi_r \mathcal{U}^T (\mathbf{x}_i - \text{mean}(\{\mathbf{x}\}))$$

where the notation seeks to explicitly keep track of the fact that the low dimensional representation of a particular data item *depends on the whole dataset* (because you have to be able to compute the mean, and the eigenvectors of the covariance). Notice that $\text{pcaproj}(\mathbf{x}_i, r, \{\mathbf{x}\})$ is a representation of the dataset with important properties:

- The representation is r -dimensional (i.e. the last $d - r$ components are zero).
- Each pair of distinct components of $\{\text{pcaproj}(\mathbf{x}_i, r, \{\mathbf{x}\})\}$ has zero covariance.
- The first component of $\{\text{pcaproj}(\mathbf{x}_i, r, \{\mathbf{x}\})\}$ has largest variance; the second component has second largest variance; and so on.

11.3.3 Smoothing Data with a Low-Dimensional Representation

We would now like to construct a low dimensional representation of the blob, in the original coordinates. We do so by rotating the low-dimensional representation back to the original coordinate system, then adding back the mean to translate the origin back to where it started. We can write this as

$$\begin{aligned} \text{pcsmooth}(\mathbf{x}_i, r, \{\mathbf{x}\}) &= \mathcal{U} \Pi_r^T (\Pi_r \mathcal{U}^T (\mathbf{x}_i - \text{mean}(\{\mathbf{x}\}))) + \text{mean}(\{\mathbf{x}\}) \\ &= \mathcal{U} \Pi_r^T \text{pcaproj}(\mathbf{x}_i, r, \{\mathbf{x}\}) + \text{mean}(\{\mathbf{x}\}) \end{aligned}$$

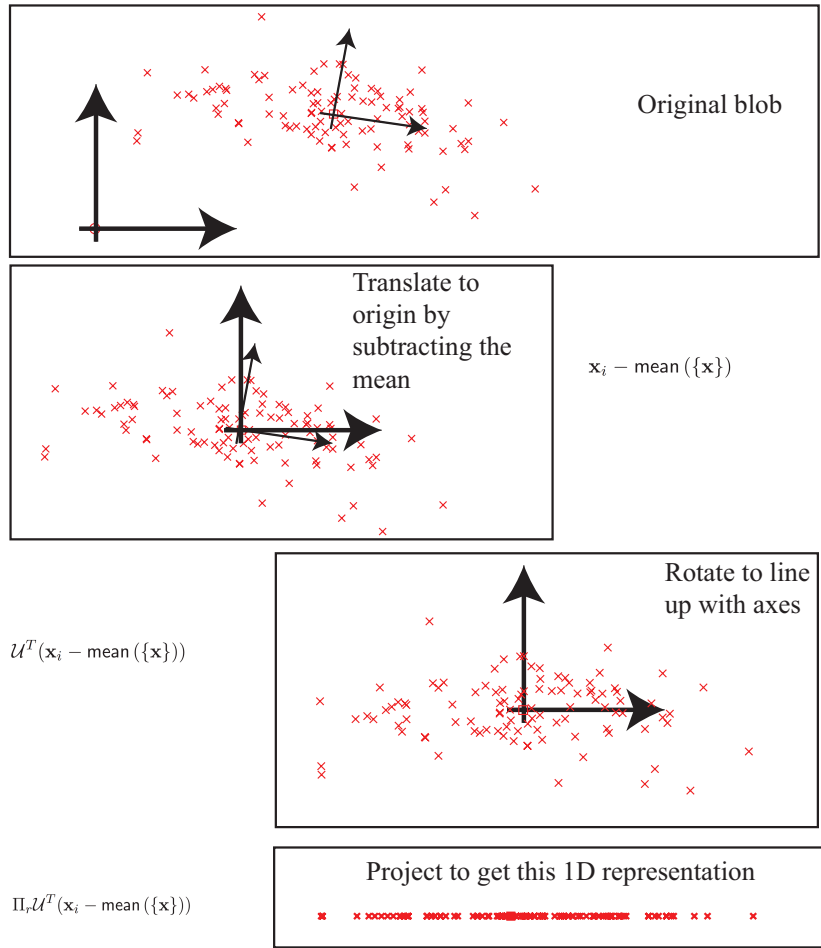


FIGURE 11.13: Computing a low dimensional representation for principal components analysis.

we have a new representation of the i 'th data item *in the original space* (Figure 1). Now consider the dataset obtained by smoothing each of our data items. We write this dataset as $\{\text{pcsmooth}(\mathbf{x}_i, r, \{\mathbf{x}\})\}$.

You should think of $\{\text{pcsmooth}(\mathbf{x}_i, r, \{\mathbf{x}\})\}$ as a smoothed version of the original data. One way to think of this process is that we have chosen a low-dimensional basis that represents the main variance in the data rather well. It is quite usual to think of a data item as being given by a the mean plus a weighted sum of these basis elements. In this view, the first weight has larger variance than the second, and so on. By construction, this dataset lies in an r dimensional affine subspace of the original space. We constructed this r -dimensional space to preserve the largest variance directions of the data. Each column of this matrix is known as a **principal component**. In particular, we have constructed this dataset so that

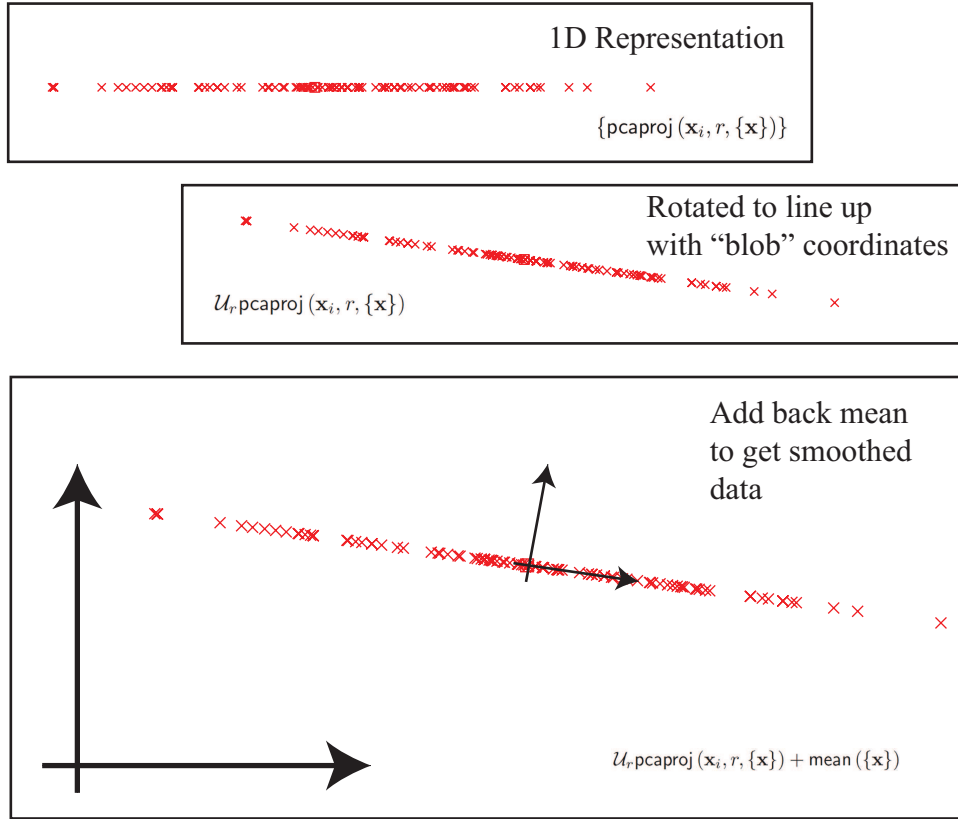


FIGURE 11.14: *Smoothing data with principal components analysis.*

- $\text{mean}(\{\text{pcsmooth}(\mathbf{x}_i, r, \{\mathbf{x}\})\}) = \text{mean}(\{\mathbf{x}\})$;
- $\text{Covmat}(\{\text{pcsmooth}(\mathbf{x}_i, r, \{\mathbf{x}\})\})$ has rank r ;
- $\text{Covmat}(\{\text{pcsmooth}(\mathbf{x}_i, r, \{\mathbf{x}\})\})$ is the best approximation of $\text{Covmat}(\{\mathbf{x}\})$ with rank r .

Figure 11.14 gives a visualization of the smoothing process. By comparing figures 11.9 and 11.15, you can see that a real dataset can lose several dimensions without much significant going wrong. As we shall see in the examples, some datasets can lose many dimensions without anything bad happening.

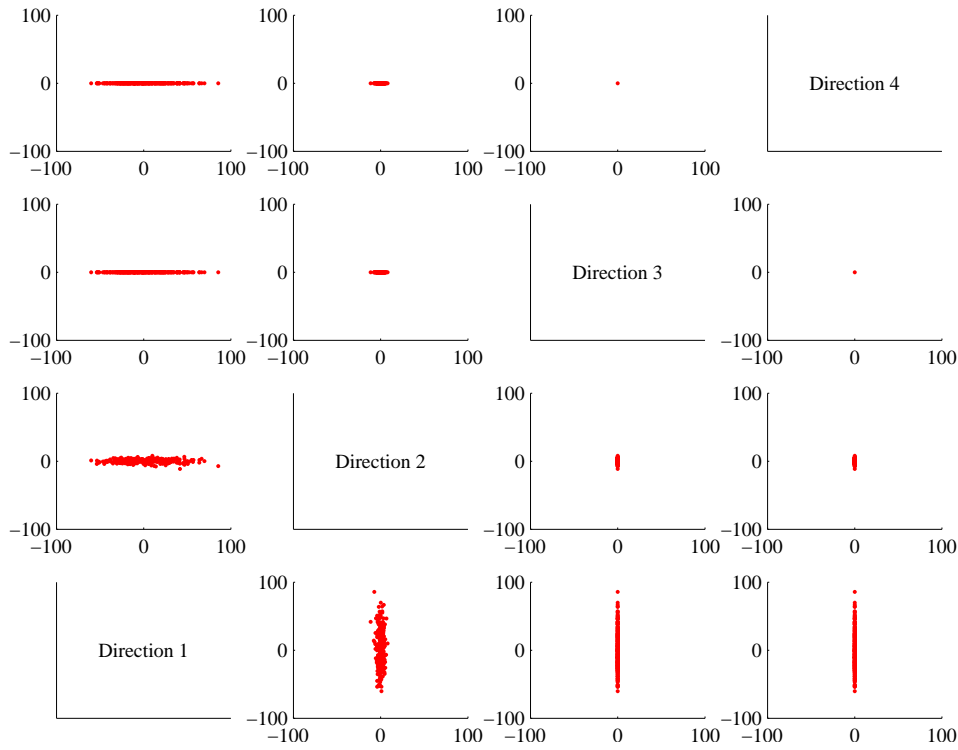


FIGURE 11.15: A panel plot of the bodyfat dataset of figure 11.4, with the dimension reduced to two using principal components analysis. Compare this figure to figure 11.9, which is on the same set of axes. You can see that the blob has been squashed in direction 3 and direction 4. But not much has really happened, because there wasn't very much variation in those directions in the first place.

Procedure: 11.1 *Principal Components Analysis*

Assume we have a general data set \mathbf{x}_i , consisting of N d -dimensional vectors. Now write $\Sigma = \text{Covmat}(\{\mathbf{x}\})$ for the covariance matrix. Form \mathcal{U}, Λ , such that $\Sigma\mathcal{U} = \mathcal{U}\Lambda$ (these are the eigenvectors and eigenvalues of Σ). Ensure that the entries of Λ are sorted in decreasing order. Choose r , the number of dimensions you wish to represent. Typically, we do this by plotting the eigenvalues and looking for a “knee” (Figure ??). It is quite usual to do this by hand.

Constructing a low-dimensional representation: Form \mathcal{U}_r , a matrix consisting of the first r columns of \mathcal{U} . Now compute $\{\text{pcproj}(\mathbf{x}_i, r, \{\mathbf{x}\})\} = \{(\Pi_r \mathcal{U}^T (\mathbf{x}_i - \text{mean}(\{\mathbf{x}\})))\}$. This is a set of data vectors which are r dimensional, and where each component is independent of each other component (i.e. the covariances of distinct components are zero).

Smoothing the data: Form $\{\text{pcsmooth}(\mathbf{x}_i, r, \{\mathbf{x}\})\} = \{\mathcal{U}_r \text{pcproj}(\mathbf{x}_i, r, \{\mathbf{x}\}) + \text{mean}(\{\mathbf{x}\})\}$. These are d dimensional vectors that lie in a r -dimensional subspace of d -dimensional space. The “missing dimensions” have the lowest variance, and are independent.

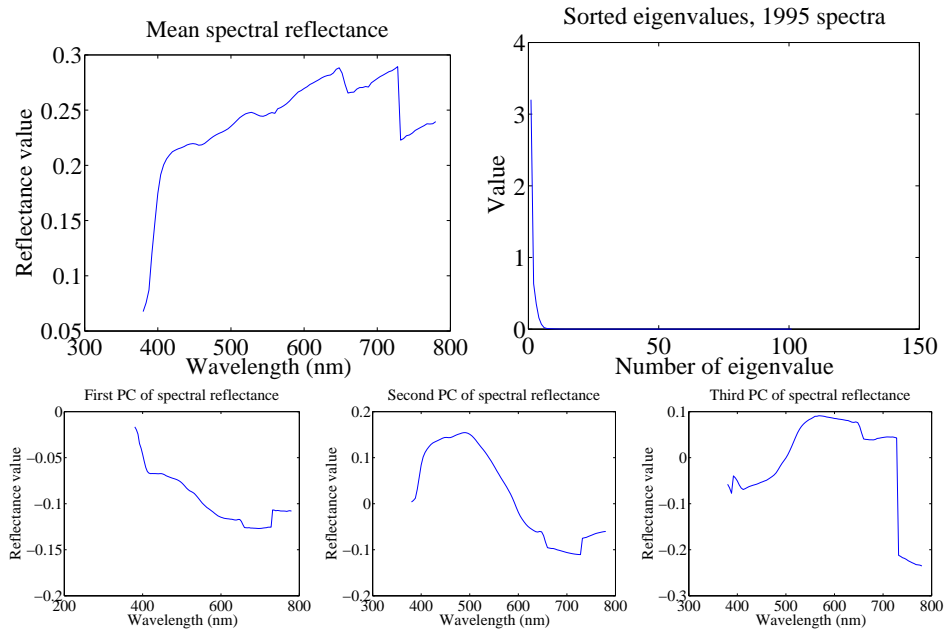


FIGURE 11.16: On the **top left**, the mean spectral reflectance of a dataset of 1995 spectral reflectances, collected by Kobus Barnard (at <http://www.cs.sfu.ca/~colour/data/>). On the **top right**, eigenvalues of the covariance matrix of spectral reflectance data, from a dataset of 1995 spectral reflectances, collected by Kobus Barnard (at <http://www.cs.sfu.ca/~colour/data/>). Notice how the first few eigenvalues are large, but most are very small; this suggests that a good representation using few principal components is available. The **bottom row** shows the first three principal components. A linear combination of these, with appropriate weights, added to the mean of figure ??, gives a good representation of the dataset.

11.3.4 The Error of the Low-Dimensional Representation

We took a dataset, $\{\mathbf{x}\}$, and constructed a d -dimensional dataset $\{\mathbf{b}\}$ in blob coordinates. We did so by translating, then rotating, the data, so no information was lost; we could reconstruct our original dataset by rotating, then translating $\{\mathbf{b}\}$. But in blob coordinates we projected each data item down to the first r components to get an r -dimensional dataset $\{\mathbf{b}_r\}$. We then reconstructed a smoothed dataset by rotating, then translating, $\{\mathbf{b}_r\}$. Information has been lost here, but how much?

The answer is easy to get if you recall that rotations and translations do not change lengths. This means that

$$\|\mathbf{x}_i - \text{pcsmooth}(\mathbf{x}_i, r, \{\mathbf{x}\})\|^2 = \|\mathbf{b}_i - \mathbf{b}_{r,i}\|^2.$$

This expression is easy to evaluate, because \mathbf{b}_i and $\mathbf{b}_{r,i}$ agree in their first r com-

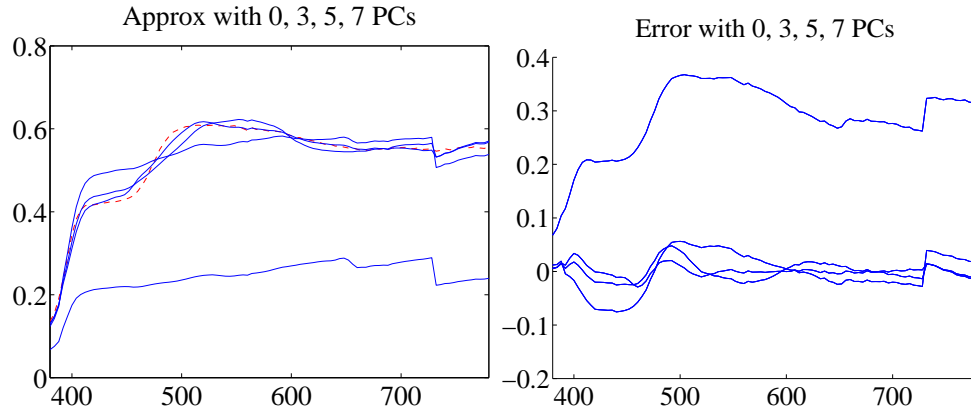


FIGURE 11.17: *On the left*, a spectral reflectance curve (dashed) and approximations using the mean, the mean and 3 principal components, the mean and 5 principal components, and the mean and 7 principal components. Notice the mean is a relatively poor approximation, but as the number of principal components goes up, the error falls rather quickly. *On the right* is the error for these approximations. Figure plotted from a dataset of 1995 spectral reflectances, collected by Kobus Barnard (at <http://www.cs.sfu.ca/~colour/data/>).

ponents. The remaining $d - r$ components of $\mathbf{b}_{r,i}$ are zero. So we can write

$$\|\mathbf{x}_i - \text{pcsmooth}(\mathbf{x}_i, r, \{\mathbf{x}\})\|^2 = \sum_{u=r+1}^d (\mathbf{b}_i^{(u)})^2.$$

Now a natural measure of error is the average over the dataset of this term. We have that

$$\frac{1}{N} \sum_{u=r+1}^d (\mathbf{b}_i^{(u)})^2 = \sum_{u=r+1}^d \text{var}(\{\mathbf{b}^{(u)}\})$$

which is easy to evaluate, because we know these variances — they are the values of the $d - r$ eigenvalues that we decided to ignore. So the mean error can be written as

$$\mathbf{1}^T (\Lambda - \Lambda_r) \mathbf{1}.$$

Now we could choose r by identifying how much error we can tolerate. More usual is to plot the eigenvalues of the covariance matrix, and look for a “knee”, like that in Figure 1. You can see that the sum of remaining eigenvalues is small.

11.3.5 Example: Representing Spectral Reflectances

Diffuse surfaces reflect light uniformly in all directions. Examples of diffuse surfaces include matte paint, many styles of cloth, many rough materials (bark, cement, stone, etc.). One way to tell a diffuse surface is that it does not look brighter (or darker) when you look at it along different directions. Diffuse surfaces can

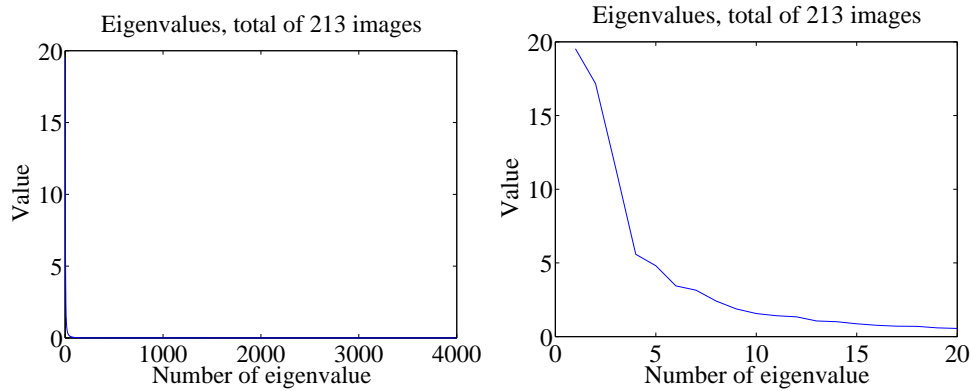


FIGURE 11.18: On the **left**, the eigenvalues of the covariance of the Japanese facial expression dataset; there are 4096, so it's hard to see the curve (which is packed to the left). On the **right**, a zoomed version of the curve, showing how quickly the values of the eigenvalues get small.

be colored, because the surface reflects different fractions of the light falling on it at different wavelengths. This effect can be represented by measuring the spectral reflectance of a surface, which is the fraction of light the surface reflects as a function of wavelength. This is usually measured in the visual range of wavelengths (about 380nm to about 770 nm). Typical measurements are every few nm, depending on the measurement device. I obtained data for 1995 different surfaces from <http://www.cs.sfu.ca/~colour/data/> (there are a variety of great datasets here, from Kobus Barnard).

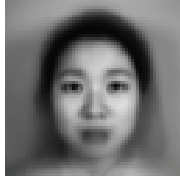
Each spectrum has 101 measurements, which are spaced 4nm apart. This represents surface properties to far greater precision than is really useful. Physical properties of surfaces suggest that the reflectance can't change too fast from wavelength to wavelength. It turns out that very few principal components are sufficient to describe almost any spectral reflectance function. Figure 11.16 shows the mean spectral reflectance of this dataset, and Figure 11.16 shows the eigenvalues of the covariance matrix.

This is tremendously useful in practice. One should think of a spectral reflectance as a function, usually written $\rho(\lambda)$. What the principal components analysis tells us is that we can represent this function rather accurately on a (really small) finite dimensional basis. This basis is shown in figure 11.16. This means that there is a mean function $r(\lambda)$ and k functions $\phi_m(\lambda)$ such that, for any $\rho(\lambda)$,

$$\rho(\lambda) = r(\lambda) + \sum_{i=1}^k c_i \phi_i(\lambda) + e(\lambda)$$

where $e(\lambda)$ is the error of the representation, which we know is small (because it consists of all the other principal components, which have tiny variance). In the case of spectral reflectances, using a value of k around 3-5 works fine for most applications (Figure 11.17). This is useful, because when we want to predict what

Mean image from Japanese Facial Expression dataset



First sixteen principal components of the Japanese Facial Expression dat



FIGURE 11.19: *The mean and first 16 principal components of the Japanese facial expression dataset.*

a particular object will look like under a particular light, we don't need to use a detailed spectral reflectance model; instead, it's enough to know the c_i for that object. This comes in useful in a variety of rendering applications in computer graphics. It is also the key step in an important computer vision problem, called **color constancy**. In this problem, we see a picture of a world of colored objects under unknown colored lights, and must determine what color the objects are. Modern color constancy systems are quite accurate, even though the problem sounds unconstrained. This is because they are able to exploit the fact that relatively few c_i are enough to accurately describe a surface reflectance.

11.3.6 Example: Representing Faces with Principal Components

An image is usually represented as an array of values. We will consider intensity images, so there is a single intensity value in each cell. You can turn the image into a vector by rearranging it, for example stacking the columns onto one another

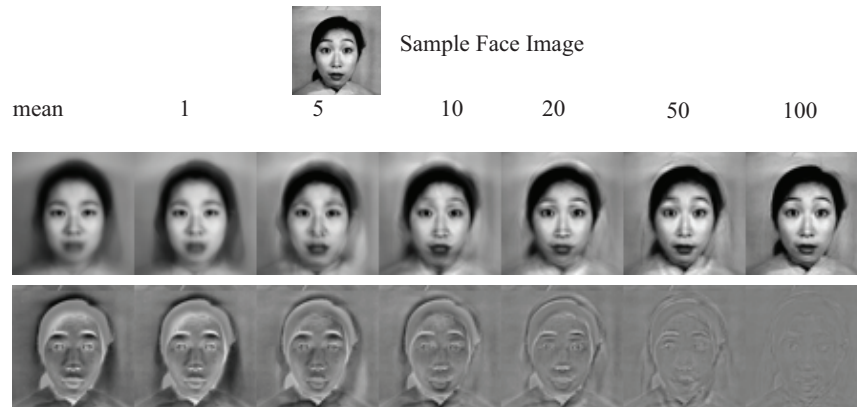


FIGURE 11.20: *Approximating a face image by the mean and some principal components; notice how good the approximation becomes with relatively few components.*

(use `reshape` in Matlab). This means you can take the principal components of a set of images. Doing so was something of a fashionable pastime in computer vision for a while, though there are some reasons that this is not a great representation of pictures. However, the representation yields pictures that can give great intuition into a dataset.

Figure ?? shows the mean of a set of face images encoding facial expressions of Japanese women (available at <http://www.kasrl.org/jaffe.html>; there are tons of face datasets at <http://www.face-rec.org/databases/>). I reduced the images to 64×64 , which gives a 4096 dimensional vector. The eigenvalues of the covariance of this dataset are shown in figure 11.18; there are 4096 of them, so it's hard to see a trend, but the zoomed figure suggests that the first couple of hundred contain most of the variance. Once we have constructed the principal components, they can be rearranged into images; these images are shown in figure 11.19. Principal components give quite good approximations to real images (figure 11.20).

The principal components sketch out the main kinds of variation in facial expression. Notice how the mean face in Figure 11.19 looks like a relaxed face, but with fuzzy boundaries. This is because the faces can't be precisely aligned, because each face has a slightly different shape. The way to interpret the components is to remember one adjusts the mean towards a data point by adding (or subtracting) some scale times the component. So the first few principal components have to do with the shape of the haircut; by the fourth, we are dealing with taller/shorter faces; then several components have to do with the height of the eyebrows, the shape of the chin, and the position of the mouth; and so on. These are all images of women who are not wearing spectacles. In face pictures taken from a wider set of models, moustaches, beards and spectacles all typically appear in the first couple of dozen principal components.

11.4 MULTI-DIMENSIONAL SCALING

One way to get insight into a dataset is to plot it. But choosing what to plot for a high dimensional dataset could be difficult. Assume we must plot the dataset in two dimensions (by far the most common choice). We wish to build a scatter plot in two dimensions — but where should we plot each data point? One natural requirement is that the points be laid out in two dimensions in a way that reflects how they sit in many dimensions. In particular, we would like points that are far apart in the high dimensional space to be far apart in the plot, and points that are close in the high dimensional space to be close in the plot.

11.4.1 Principal Coordinate Analysis

We will plot the high dimensional point \mathbf{x}_i at \mathbf{v}_i , which is a two-dimensional vector. Now the squared distance between points i and j in the high dimensional space is

$$D_{ij}^{(2)}(\mathbf{x}) = (\mathbf{x}_i - \mathbf{x}_j)^T (\mathbf{x}_i - \mathbf{x}_j)$$

(where the superscript is to remind you that this is a squared distance). We could build an $N \times N$ matrix of squared distances, which we write $\mathcal{D}^{(2)}(\mathbf{x})$. The i, j 'th entry in this matrix is $D_{ij}^{(2)}(\mathbf{x})$, and the \mathbf{x} argument means that the distances are between points in the high-dimensional space. Now we could choose the \mathbf{v}_i to make

$$\sum_{ij} \left(D_{ij}^{(2)}(\mathbf{x}) - D_{ij}^{(2)}(\mathbf{v}) \right)^2$$

as small as possible. Doing so should mean that points that are far apart in the high dimensional space are far apart in the plot, and that points that are close in the high dimensional space are close in the plot.

In its current form, the expression is difficult to deal with, but we can refine it. Because translation does not change the distances between points, it cannot change either of the $\mathcal{D}^{(2)}$ matrices. So it is enough to solve the case when the mean of the points \mathbf{x}_i is zero. We can assume that $\frac{1}{N} \sum_i \mathbf{x}_i = \mathbf{0}$. Now write $\mathbf{1}$ for the n -dimensional vector containing all ones, and \mathcal{I} for the identity matrix. Notice that

$$D_{ij}^{(2)} = (\mathbf{x}_i - \mathbf{x}_j)^T (\mathbf{x}_i - \mathbf{x}_j) = \mathbf{x}_i \cdot \mathbf{x}_i - 2\mathbf{x}_i \cdot \mathbf{x}_j + \mathbf{x}_j \cdot \mathbf{x}_j.$$

Now write

$$\mathcal{A} = \left[\mathcal{I} - \frac{1}{N} \mathbf{1}\mathbf{1}^T \right].$$

Using this expression, you can show that the matrix \mathcal{M} , defined below,

$$\mathcal{M}(\mathbf{x}) = -\frac{1}{2} \mathcal{A} \mathcal{D}^{(2)}(\mathbf{x}) \mathcal{A}^T$$

has i, j 'th entry $\mathbf{x}_i \cdot \mathbf{x}_j$ (exercises). I now argue that, to make $\mathcal{D}^{(2)}(\mathbf{v})$ close to $\mathcal{D}^{(2)}(\mathbf{x})$, it is enough to make $\mathcal{M}(\mathbf{v})$ close to $\mathcal{M}(\mathbf{x})$. Proving this will take us out of our way unnecessarily, so I omit a proof.

We can choose a set of \mathbf{v}_i that makes $\mathcal{D}^{(2)}(\mathbf{v})$ close to $\mathcal{D}^{(2)}(\mathbf{x})$ quite easily, using the method of the previous section. Take the dataset of N d -dimensional column vectors \mathbf{x}_i , and form a matrix \mathcal{X} by stacking the vectors, so

$$\mathcal{X} = [\mathbf{x}_1, \mathbf{x}_2, \dots, \mathbf{x}_N].$$

In this notation, we have

$$\mathcal{M}(\mathbf{x}) = \mathcal{X}^T \mathcal{X}.$$

This matrix is symmetric, and it is positive semidefinite. It can't be positive definite, because the data is zero mean, so $\mathcal{M}(\mathbf{x})\mathbf{1} = 0$. The $\mathcal{M}(\mathbf{v})$ we seek must (a) be as close as possible to $\mathcal{M}(\mathbf{x})$ and (b) have rank 2. It must have rank 2 because there must be some \mathcal{V} which is $2 \times N$ so that $\mathcal{M}(\mathbf{v}) = \mathcal{V}^T \mathcal{V}$. The columns of this \mathcal{V} are our \mathbf{v}_i .

We can use the method of section 15.1.1 to construct $\mathcal{M}(\mathbf{v})$ and \mathcal{V} . As usual, we write \mathcal{U} for the matrix of eigenvectors of $\mathcal{M}(\mathbf{x})$, Λ for the diagonal matrix of eigenvalues sorted in descending order, Λ_2 for the 2×2 upper left hand block of Λ , and $\Lambda_2^{(1/2)}$ for the matrix of positive square roots of the eigenvalues. Then our methods yield

$$\mathcal{M}(\mathbf{v}) = \mathcal{U}_2 \Lambda_2^{(1/2)} \Lambda_2^{(1/2)} \mathcal{U}_2^T$$

and

$$\mathcal{V} = \Lambda_2^{(1/2)} \mathcal{U}_2^T$$

and we can plot these \mathbf{v}_i (example in section 1). This method for constructing a plot is known as **principal coordinate analysis**.

This plot might not be perfect, because reducing the dimension of the data points should cause some distortions. In many cases, the distortions are tolerable. In other cases, we might need to use a more sophisticated scoring system that penalizes some kinds of distortion more strongly than others. There are many ways to do this; the general problem is known as **multidimensional scaling**.

Procedure: 11.2 *Principal Coordinate Analysis*

Assume we have a matrix $D^{(2)}$ consisting of the squared differences between each pair of N points. We do not need to know the points. We wish to compute a set of points in r dimensions, such that the distances between these points are as similar as possible to the distances in $D^{(2)}$. Form $\mathcal{A} = [\mathcal{I} - \frac{1}{N}\mathbf{1}\mathbf{1}^T]$. Form $\mathcal{W} = \frac{1}{2}\mathcal{A}D^{(2)}\mathcal{A}^T$. Form \mathcal{U} , Λ , such that $\mathcal{W}\mathcal{U} = \mathcal{U}\Lambda$ (these are the eigenvectors and eigenvalues of \mathcal{W}). Ensure that the entries of Λ are sorted in decreasing order. Choose r , the number of dimensions you wish to represent. Form Λ_r , the top left $r \times r$ block of Λ . Form $\Lambda_r^{(1/2)}$, whose entries are the positive square roots of Λ_r . Form \mathcal{U}_r , the matrix consisting of the first r columns of \mathcal{U} . Then $\mathcal{V} = \Lambda_r^{(1/2)}\mathcal{U}_r^T = [\mathbf{v}_1, \dots, \mathbf{v}_N]$ is the set of points to plot.

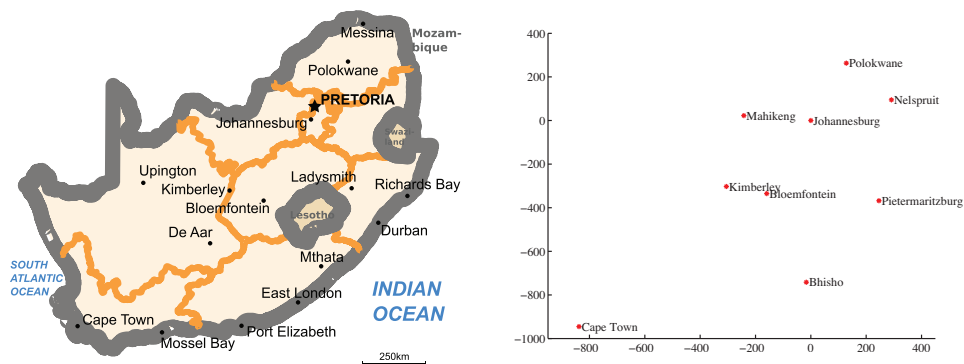


FIGURE 11.21: On the left, a public domain map of South Africa, obtained from http://commons.wikimedia.org/wiki/File:Map_of_South_Africa.svg, and edited to remove surrounding countries. On the right, the locations of the cities inferred by multidimensional scaling, rotated, translated and scaled to allow a comparison to the map by eye. The map doesn't have all the provincial capitals on it, but it's easy to see that MDS has placed the ones that are there in the right places (use a piece of ruled tracing paper to check).

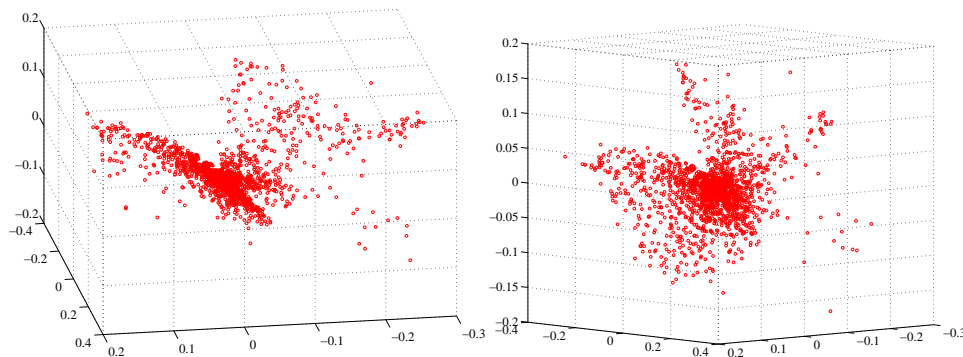


FIGURE 11.22: Two views of the spectral data of section 11.3.5, plotted as a scatter plot by applying principal coordinate analysis to obtain a 3D set of points. Notice that the data spreads out in 3D, but seems to lie on some structure; it certainly isn't a single blob. This suggests that further investigation would be fruitful.

11.4.2 Example: Mapping with Multidimensional Scaling

Multidimensional scaling gets positions (the \mathcal{V} of section 11.4.1) from distances (the $\mathcal{D}^{(2)}(\mathbf{x})$ of section 11.4.1). This means we can use the method to build maps from distances alone. I collected distance information from the web (I used <http://www.distancefromto.net>, but a google search on “city distances” yields a wide range of possible sources), then apply multidimensional scaling. Table ?? shows distances between the South African provincial capitals, in kilometers, rounded to the nearest kilometer. I then used principal coordinate analysis to find positions

for each capital, and rotated, translated and scaled the resulting plot to check it against a real map (Figure 11.21).

One natural use of principal coordinate analysis is to see if one can spot any structure in a dataset. Does the dataset form a blob, or is it clumpy? This isn't a perfect test, but it's a good way to look and see if anything interesting is happening. In figure 11.22, I show a 3D plot of the spectral data, reduced to three dimensions using principal coordinate analysis. The plot is quite interesting. You should notice that the data points are spread out in 3D, but actually seem to lie on a complicated curved surface — they very clearly don't form a uniform blob. To me, the structure looks somewhat like a butterfly. I don't know why this occurs, but it certainly suggests that something worth investigating is going on. Perhaps the choice of samples that were measured is funny; perhaps the measuring instrument doesn't make certain kinds of measurement; or perhaps there are physical processes that prevent the data from spreading out over the space.

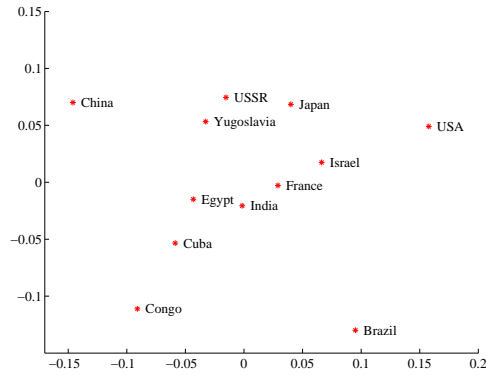


FIGURE 11.23: A map of country similarity, prepared from the data of figure ???. The map is often interpreted as showing a variation in development or wealth (poorest at bottom left to richest at top right); and freedom (most repressed at top left and freest at bottom right). I haven't plotted these axes, because the interpretation wouldn't be consistent with current intuition (the similarity data is forty years old, and quite a lot has happened in that time).

Our algorithm has one really interesting property. In some cases, we do not actually know the datapoints as vectors. Instead, we *just* know distances between the datapoints. This happens often in the social sciences, but there are important cases in computer science as well. As a rather contrived example, one could survey people about breakfast foods (say, eggs, bacon, cereal, oatmeal, pancakes, toast, muffins, kippers and sausages for a total of 9 items). We ask each person to rate the similarity of each pair of distinct items on some scale. We advise people that similar items are ones where, if they were offered both, they would have no particular preference; but, for dissimilar items, they would have a strong preference for one over the other. The scale might be “very similar”, “quite similar”, “similar”, “quite dissimilar”, and “very dissimilar” (scales like this are often called **Likert scales**). We collect these similarities from many people for each pair of distinct items, and

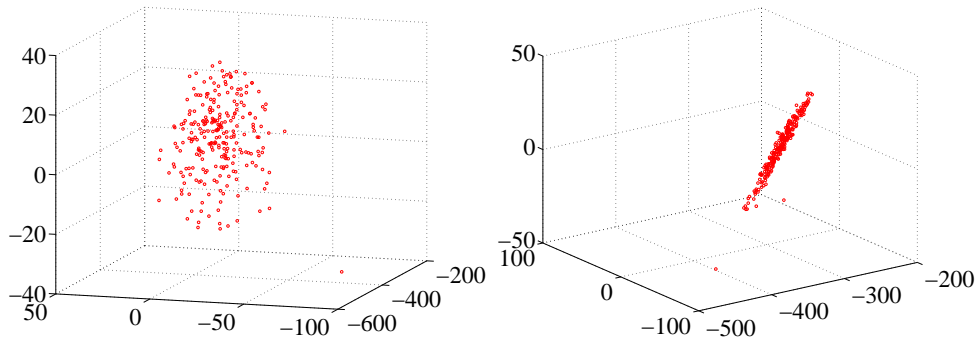


FIGURE 11.24: Two views of a multidimensional scaling to three dimensions of the height-weight dataset. Notice how the data seems to lie in a flat structure in 3D, with one outlying data point. This means that the distances between data points can be (largely) explained by a 2D representation.

then average the similarity over all respondents. We compute distances from the similarities in a way that makes very similar items close and very dissimilar items distant. Now we have a table of distances between items, and can compute a \mathbf{V} and produce a scatter plot. This plot is quite revealing, because items that most people think are easily substituted appear close together, and items that are hard to substitute are far apart. The neat trick here is that we did not start with a \mathcal{X} , but with just a set of distances; but we were able to associate a vector with “eggs”, and produce a meaningful plot.

Table ?? shows data from one such example. Students were interviewed (in 1971! things may have changed since then) about their perceptions of the similarity of countries. The averaged perceived similarity is shown in table ?. Large numbers reflect high similarity, so we can’t use these numbers directly. It is reasonable to turn these numbers into distances by (a) using 0 as the distance between a country and itself and (b) using $e^{-s_{ij}}$ as the distance between countries i and j (where s_{ij} is the similarity between them). Once we have distances, we can apply the procedure of section 11.4.1 to get points, then plot a scatter plot (Figure 11.23).

11.5 EXAMPLE: UNDERSTANDING HEIGHT AND WEIGHT

Recall the height-weight data set of section ?? (from <http://www2.stetson.edu/~jrasp/data.htm>; look for bodyfat.xls at that URL). This is, in fact, a 16-dimensional dataset. The entries are (in this order): *bodyfat*; *density*; *age*; *weight*; *height*; *adiposity*; *neck*; *chest*; *abdomen*; *hip*; *thigh*; *knee*; *ankle*; *biceps*; *forearm*; *wrist*. We know already that many of these entries are correlated, but it’s hard to grasp a 16 dimensional dataset in one go. The first step is to investigate with a multidimensional scaling.

Figure ?? shows a multidimensional scaling of this dataset down to three dimensions. The dataset seems to lie on a (fairly) flat structure in 3D, meaning that inter-point distances are relatively well explained by a 2D representation. Two points seem to be special, and lie far away from the flat structure. The structure

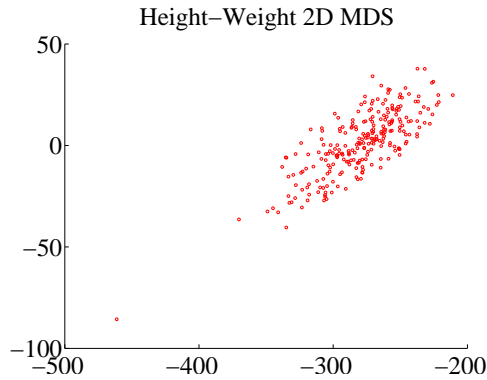


FIGURE 11.25: A multidimensional scaling to two dimensions of the height-weight dataset. One data point is clearly special, and another looks pretty special. The data seems to form a blob, with one axis quite a lot more important than another.

isn't perfectly flat, so there will be small errors in a 2D representation; but it's clear that a lot of dimensions are redundant. Figure 11.25 shows a 2D representation of these points. They form a blob that is stretched along one axis, and there is no sign of multiple blobs. There's still at least one special point, which we shall ignore but might be worth investigating further. The distortions involved in squashing this dataset down to 2D seem to have made the second special point less obvious than it was in figure ??.

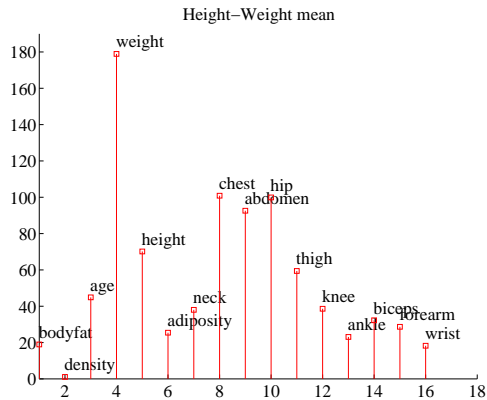


FIGURE 11.26: The mean of the bodyfat.xls dataset. Each component is likely in a different unit (though I don't know the units), making it difficult to plot the data without being misleading. I've adopted one solution here, by plotting each component as a vertical bar, and labelling the bar. You shouldn't try to compare the values to one another. Instead, think of this plot as a compact version of a table.

The next step is to try a principal component analysis. Figure 11.26 shows the mean of the dataset. The components of the dataset have different units, and

shouldn't really be compared. But it is difficult to interpret a table of 16 numbers, so I have plotted the mean by showing a vertical bar for each component. Figure 11.27 shows the eigenvalues of the covariance for this dataset. Notice how one dimension is very important, and after the third principal component, the contributions become very small. Of course, I could have said “fourth”, or “fifth”, or whatever — the precise choice depends on how small a number you think is “small”.

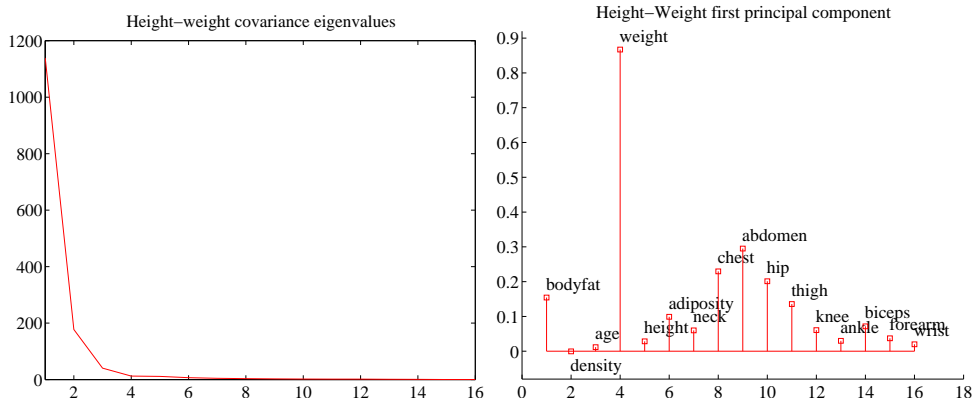


FIGURE 11.27: On the **left**, the eigenvalues of the covariance matrix for the bodyfat data set. Notice how fast the eigenvalues fall off; this means that most principal components have very small variance, so that data can be represented well with a small number of principal components. On the **right**, the first principal component for this dataset, plotted using the same convention as for figure 11.26.

Figure 11.27 also shows the first principal component. The eigenvalues justify thinking of each data item as (roughly) the mean plus some weight times this principal component. From this plot you can see that data items with a larger value of *weight* will also have larger values of most other measurements, except *age* and *density*. You can also see how much larger; if the weight goes up by 8.5 units, then the abdomen will go up by 3 units, and so on. This explains the main variation in the dataset.

In the rotated coordinate system, the components are not correlated, and they have different variances (which are the eigenvalues of the covariance matrix). You can get some sense of the data by adding these variances; in this case, we get 1404. This means that, in the translated and rotated coordinate system, the average data point is about $37 = \sqrt{1404}$ units away from the center (the origin). Translations and rotations do not change distances, so the average data point is about 37 units from the center in the original dataset, too. If we represent a datapoint by using the mean and the first three principal components, there will be some error. We can estimate the average error from the component variances. In this case, the sum of the first three eigenvalues is 1357, so the mean square error in representing a datapoint by the first three principal components is $\sqrt{(1404 - 1357)}$, or 6.8. The relative error is $6.8/37 = 0.18$. Another way to represent this information, which is more widely used, is to say that the first three principal components explain all but

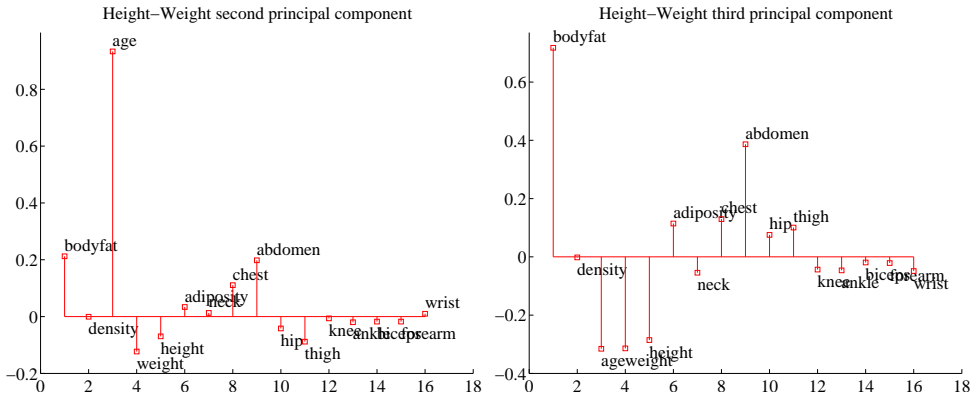


FIGURE 11.28: On the **left**, the second principal component, and on the **right** the third principal component of the height-weight dataset.

$(1404 - 1357)/1404 = 0.034$, or 3.4% of the variance; notice that this is the square of the relative error, which will be a much smaller number.

All this means that explaining a data point as the mean and the first three principal components produces relatively small errors. Figure 11.28 shows the second and third principal component of the data. These two principal components suggest some further conclusions. As *age* gets larger, *height* and *weight* get slightly smaller, but the weight is redistributed; *abdomen* gets larger, whereas *thigh* gets smaller. A smaller effect (the third principal component) links *bodyfat* and *abdomen*. As *bodyfat* goes up, so does *abdomen*.

Models of High Dimensional Data

High-dimensional data comes with problems. Data points tend not to be where you think; they can scattered quite far apart, and can be quite far from the mean. Except in special cases, the only really reliable probability model is the Gaussian (or Gaussian blob, or blob).

There is an important rule of thumb for coping with high dimensional data: **Use simple models.** A blob is another good, simple model. Modelling data as a blob involves computing its mean and its covariance. Sometimes, as we shall see, the covariance can be hard to compute. Even so, a blob model is really useful. It is natural to try and extend this model to cover datasets that don't obviously consist of a single blob.

One very good, very simple, model for high dimensional data is to assume that it consists of multiple blobs. To build models like this, we must determine (a) what the blob parameters are and (b) which datapoints belong to which blob. Generally, we will collect together data points that are close and form blobs out of them. This process is known as **clustering**.

Clustering is a somewhat puzzling activity. It is extremely useful to cluster data, and it seems to be quite important to do it reasonably well. But it surprisingly hard to give crisp criteria for a good (resp. bad) clustering of a dataset. Usually, clustering is part of building a model, and the main way to know that the clustering algorithm is bad is that the model is bad.

12.1 THE CURSE OF DIMENSION

High dimensional models display unintuitive behavior (or, rather, it can take years to make your intuition see the true behavior of high-dimensional models as natural). In these models, most data lies in places you don't expect. We will do several simple calculations with an easy high-dimensional distribution to build some intuition.

Assume our data lies within a cube, with edge length two, centered on the origin. This means that each component of \mathbf{x}_i lies in the range $[-1, 1]$. One simple model for such data is to assume that each dimension has uniform probability density in this range. In turn, this means that $P(x) = \frac{1}{2^d}$. The mean of this model is at the origin, which we write as $\mathbf{0}$.

The first surprising fact about high dimensional data is that most of the data can lie quite far away from the mean. For example, we can divide our dataset into two pieces. $\mathcal{A}(\epsilon)$ consists of all data items where *every* component of the data has a value in the range $[-(1 - \epsilon), (1 - \epsilon)]$. $\mathcal{B}(\epsilon)$ consists of all the rest of the data. If you think of the data set as forming a cubical orange, then $\mathcal{B}(\epsilon)$ is the rind (which has thickness ϵ) and $\mathcal{A}(\epsilon)$ is the fruit.

Your intuition will tell you that there is more fruit than rind. This is true, for three dimensional oranges, but not true in high dimensions. The fact that the orange is cubical just simplifies the calculations, but has nothing to do with the

real problem.

We can compute $P(\{\mathbf{x} \in \mathcal{A}(\epsilon)\})$ and $P(\{\mathbf{x} \in \mathcal{B}(\epsilon)\})$. These probabilities tell us the probability a data item lies in the fruit (resp. rind). $P(\{\mathbf{x} \in \mathcal{A}(\epsilon)\})$ is easy to compute as

$$P(\{\mathbf{x} \in \mathcal{A}(\epsilon)\}) = (2(1 - \epsilon))^d \left(\frac{1}{2^d}\right) = (1 - \epsilon)^d$$

and

$$P(\{\mathbf{x} \in \mathcal{B}(\epsilon)\}) = 1 - P(\{\mathbf{x} \in \mathcal{A}(\epsilon)\}) = 1 - (1 - \epsilon)^d.$$

But notice that, as $d \rightarrow \infty$,

$$P(\{\mathbf{x} \in \mathcal{A}(\epsilon)\}) \rightarrow 0.$$

This means that, for large d , we expect most of the data to be in $\mathcal{B}(\epsilon)$. Equivalently, for large d , we expect that at least one component of each data item is close to either 1 or -1 .

This suggests (correctly) that much data is quite far from the origin. It is easy to compute the average of the squared distance of data from the origin. We want

$$\mathbb{E}[\mathbf{x}^T \mathbf{x}] = \int_{\text{box}} \left(\sum_i x_i^2 \right) P(\mathbf{x}) d\mathbf{x}$$

but we can rearrange, so that

$$\mathbb{E}[\mathbf{x}^T \mathbf{x}] = \sum_i \mathbb{E}[x_i^2] = \sum_i \int_{\text{box}} x_i^2 P(\mathbf{x}) d\mathbf{x}.$$

Now each component of \mathbf{x} is independent, so that $P(\mathbf{x}) = P(x_1)P(x_2)\dots P(x_d)$. Now we substitute, to get

$$\mathbb{E}[\mathbf{x}^T \mathbf{x}] = \sum_i \mathbb{E}[x_i^2] = \sum_i \int_{-1}^1 x_i^2 P(x_i) dx_i = \sum_i \frac{1}{2} \int_{-1}^1 x_i^2 dx_i = \frac{d}{3},$$

so as d gets bigger, most data points will be further and further from the origin. Worse, as d gets bigger, data points tend to get further and further from one another. We can see this by computing the average of the squared distance of data points from one another. Write \mathbf{u} for one data point and \mathbf{v} ; we can compute

$$\mathbb{E}[d(\mathbf{u}, \mathbf{v})^2] = \int_{\text{box}} \int_{\text{box}} \sum_i (u_i - v_i)^2 d\mathbf{u} d\mathbf{v} = \mathbb{E}[\mathbf{u}^T \mathbf{u}] + \mathbb{E}[\mathbf{v}^T \mathbf{v}] - \mathbb{E}[\mathbf{u}^T \mathbf{v}]$$

but since \mathbf{u} and \mathbf{v} are independent, we have $\mathbb{E}[\mathbf{u}^T \mathbf{v}] = \mathbb{E}[\mathbf{u}]^T \mathbb{E}[\mathbf{v}] = 0$. This yields

$$\mathbb{E}[d(\mathbf{u}, \mathbf{v})^2] = 2\frac{d}{3}$$

meaning that, for large d , we expect our data points to be quite far apart.

It is difficult to build histogram representations for high dimensional data. The strategy of dividing the domain into boxes, then counting data into them, fails miserably because there are too many boxes. In the case of our cube, imagine we wish to divide each dimension in half (i.e. between $[-1, 0]$ and between $[0, 1]$). Then we must have 2^d boxes. This presents two problems. First, we will have difficulty representing this number of boxes. Second, unless we are exceptionally lucky, most boxes must be empty because we will not have 2^d data items.

12.2 THE MULTIVARIATE NORMAL DISTRIBUTION

All the nasty facts about high dimensional data, above, suggest that we need to use quite simple probability models. By far the most important model is the multivariate normal distribution, which is quite often known as the multivariate gaussian distribution. There are two sets of parameters in this model, the mean μ and the covariance Σ . For a d -dimensional model, the mean is a d -dimensional column vector and the covariance is a $d \times d$ dimensional matrix. The covariance is a symmetric matrix. For our definitions to be meaningful, the covariance matrix must be positive definite.

The form of the distribution $p(\mathbf{x}|\mu, \Sigma)$ is

$$p(\mathbf{x}|\mu, \Sigma) = \frac{1}{\sqrt{(2\pi)^d \det(\Sigma)}} \exp\left(-\frac{1}{2}(\mathbf{x} - \mu)^T \Sigma^{-1}(\mathbf{x} - \mu)\right).$$

The following facts explain the names of the parameters:

Useful Facts: 12.1 *Parameters of a Multivariate Normal Distribution*

Assuming a multivariate normal distribution, we have

- $\mathbb{E}[\mathbf{x}] = \mu$, meaning that the mean of the distribution is μ .
- $\mathbb{E}[(\mathbf{x} - \mu)(\mathbf{x} - \mu)^T] = \Sigma$, meaning that the entries in Σ represent covariances.

Assume I know have a dataset of items \mathbf{x}_i , where i runs from 1 to N , and we wish to model this data with a multivariate normal distribution. The maximum likelihood estimate of the mean, $\hat{\mu}$, is

$$\hat{\mu} = \frac{\sum_i \mathbf{x}_i}{N}$$

(which is quite easy to show). The maximum likelihood estimate of the covariance, $\hat{\Sigma}$, is

$$\hat{\Sigma} = \frac{\sum_i (\mathbf{x}_i - \hat{\mu})(\mathbf{x}_i - \hat{\mu})^T}{N}$$

(which is rather a nuisance to show, because you need to know how to differentiate a determinant). These facts mean that we already know most of what is interesting about multivariate normal distributions (or gaussians).

12.2.1 Affine Transformations and Gaussians

Gaussians behave very well under affine transformations. In fact, we've already worked out all the math. Assume I have a dataset \mathbf{x}_i . The mean of the maximum likelihood gaussian model is $\text{mean}(\{\mathbf{x}_i\})$, and the covariance is $\text{Covmat}(\{\mathbf{x}_i\})$. I can now transform the data with an affine transformation, to get $\mathbf{y}_i = \mathcal{A}\mathbf{x}_i + \mathbf{b}$. The mean of the maximum likelihood gaussian model for the transformed dataset is $\text{mean}(\{\mathbf{y}_i\})$, and we've dealt with this; similarly, the covariance is $\text{Covmat}(\{\mathbf{y}_i\})$, and we've dealt with this, too.

A very important point follows in an obvious way. I can apply an affine transformation to any multivariate gaussian to obtain one with (a) zero mean and (b) independent components. In turn, this means that, *in the right coordinate system*, any gaussian is a product of zero mean one-dimensional normal distributions. This fact is quite useful. For example, it means that simulating multivariate normal distributions is quite straightforward — you could simulate a standard normal distribution for each component, then apply an affine transformation.

12.3 AGGLOMERATIVE AND DIVISIVE CLUSTERING

There are two natural algorithms for clustering. In **divisive clustering**, the entire data set is regarded as a cluster, and then clusters are recursively split to yield a good clustering (Algorithm 12.2). In **agglomerative clustering**, each data item is regarded as a cluster, and clusters are recursively merged to yield a good clustering (Algorithm 12.1).

```

Make each point a separate cluster
Until the clustering is satisfactory
    Merge the two clusters with the
        smallest inter-cluster distance
end

```

Algorithm 12.1: *Agglomerative Clustering or Clustering by Merging.*

```

Construct a single cluster containing all points
Until the clustering is satisfactory
    Split the cluster that yields the two
        components with the largest inter-cluster distance
end

```

Algorithm 12.2: *Divisive Clustering, or Clustering by Splitting.*

There are two major issues in thinking about clustering:

- *What is a good inter-cluster distance?* Agglomerative clustering uses an inter-cluster distance to fuse nearby clusters; divisive clustering uses it to split insufficiently coherent clusters. Even if a natural distance between data points

is available (which might not be the case for vision problems), there is no canonical inter-cluster distance. Generally, one chooses a distance that seems appropriate for the data set. For example, one might choose the distance between the closest elements as the inter-cluster distance, which tends to yield extended clusters (statisticians call this method **single-link clustering**). Another natural choice is the maximum distance between an element of the first cluster and one of the second, which tends to yield rounded clusters (statisticians call this method **complete-link clustering**). Finally, one could use an average of distances between elements in the cluster, which also tends to yield “rounded” clusters (statisticians call this method **group average clustering**).

- *How many clusters are there?* This is an intrinsically difficult task if there is no model for the process that generated the clusters. The algorithms we have described generate a hierarchy of clusters. Usually, this hierarchy is displayed to a user in the form of a **dendrogram**—a representation of the structure of the hierarchy of clusters that displays inter-cluster distances—and an appropriate choice of clusters is made from the dendrogram (see the example in Figure 12.1).

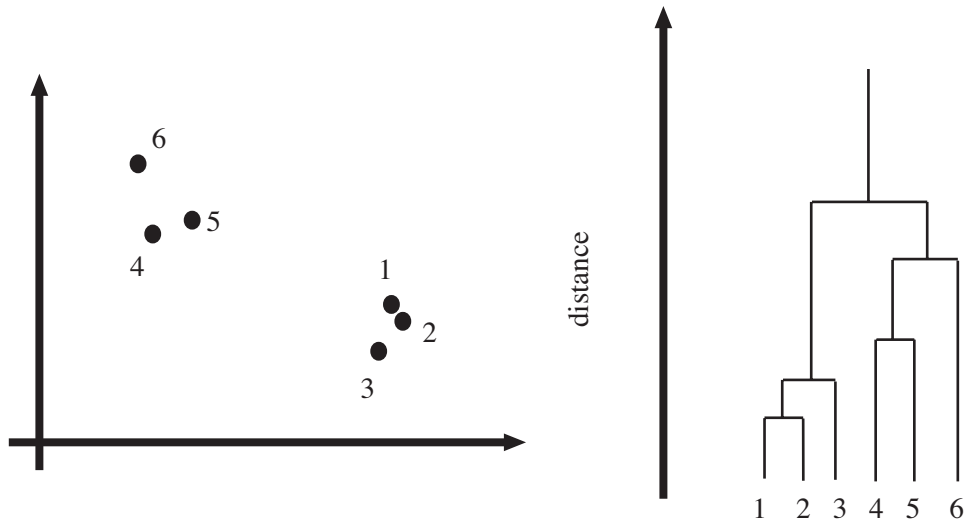


FIGURE 12.1: **Left**, a data set; **right**, a dendrogram obtained by agglomerative clustering using single-link clustering. If one selects a particular value of distance, then a horizontal line at that distance splits the dendrogram into clusters. This representation makes it possible to guess how many clusters there are and to get some insight into how good the clusters are.

The main difficulty in using a divisive model is knowing where to split. This is sometimes made easier for particular kinds of data. For example, we could segment an image by clustering pixel values. In this case, you can sometimes find good splits by constructing a histogram of intensities, or of color values.

Another important thing to notice about clustering from the example of figure 12.1 is that there is no right answer. There are a variety of different clusterings of the same data. For example, depending on what scales in that figure mean, it might be right to zoom out and regard all of the data as a single cluster, or to zoom in and regard each data point as a cluster. Each of these representations may be useful.

12.3.1 Clustering and Distance

In the algorithms above, and in what follows, we assume that the features are scaled so that distances (measured in the usual way) between data points are a good representation of their similarity. This is quite an important point. For example, imagine we are clustering data representing brick walls. The features might contain several distances: the spacing between the bricks, the length of the wall, the height of the wall, and so on. If these distances are given in the same set of units, we could have real trouble. For example, assume that the units are centimeters. Then the spacing between bricks is of the order of one or two centimeters, but the heights of the walls will be in the hundreds of centimeters. In turn, this means that the distance between two datapoints is likely to be completely dominated by the height and length data. This could be what we want, but it might also not be a good thing.

There are some ways to manage this issue. One is to know what the features measure, and know how they should be scaled. Usually, this happens because you have a deep understanding of your data. If you don't (which happens!), then it is often a good idea to try and normalize the scale of the data set. There are two good strategies. The simplest is to translate the data so that it has zero mean (this is just for neatness - translation doesn't change distances), then scale each direction so that it has unit variance. More sophisticated is to translate the data so that it has zero mean, then transform it so that each direction is independent and has unit variance. Doing so is sometimes referred to as **decorrelation** or **whitening** (because you make the data more like white noise).

12.3.2 Example: Agglomerative Clustering

Matlab provides some tools that are useful for agglomerative clustering. These functions use a scheme where one first builds the whole tree of merges, then analyzes that tree to decide which clustering to report. `linkage` will determine which pairs of clusters should be merged at which step (there are arguments that allow you to choose what type of inter-cluster distance it should use); `dendrogram` will plot you a dendrogram; and `cluster` will extract the clusters from the linkage, using a variety of options for choosing the clusters. I used these functions to prepare the dendrogram of figure 12.2 for the height-weight dataset of section ?? (from <http://www2.stetson.edu/~jrasp/data.htm>; look for bodyfat.xls). I deliberately forced Matlab to plot the whole dendrogram, which accounts for the crowded look of the figure (you can allow it to merge small leaves, etc.). I used a single-link strategy. In particular, notice that many data points are about the same distance from one another, which suggests a single big cluster with a smaller set of nearby clusters. The clustering of figure 12.3 supports this view. I plotted the data points on the first

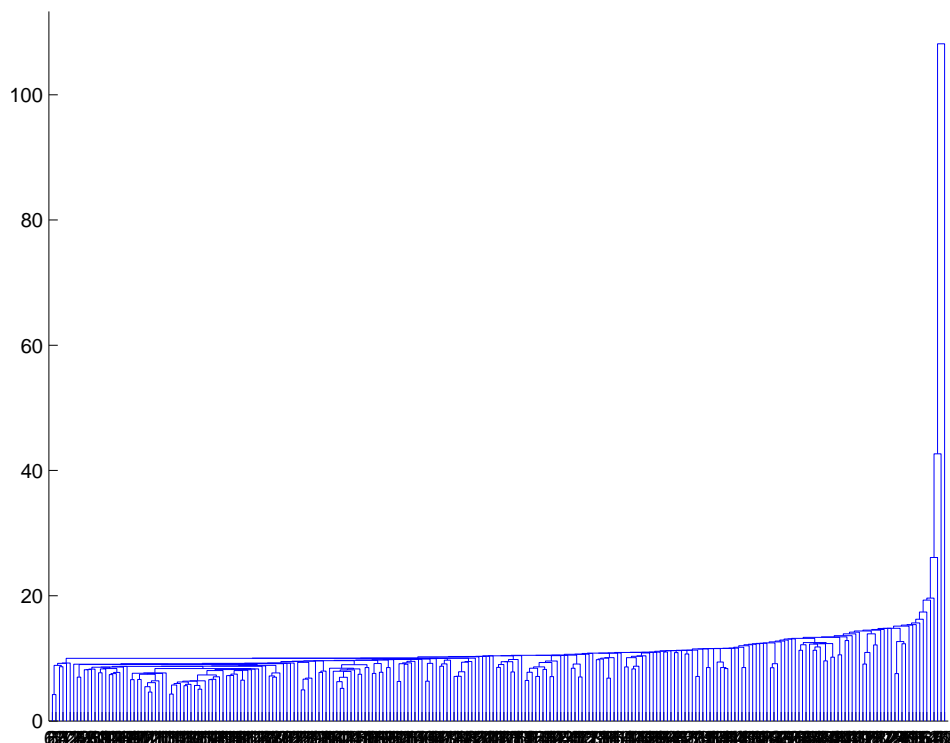


FIGURE 12.2: A dendrogram obtained from the body fat dataset, using single link clustering. Recall that the data points are on the horizontal axis, and that the vertical axis is distance; there is a horizontal line linking two clusters that get merged, established at the height at which they're merged. I have plotted the entire dendrogram, despite the fact it's a bit crowded at the bottom, because it shows that most data points are relatively close (*i.e.* there are lots of horizontal branches at about the same height).

two principal components, using different colors and shapes of marker to indicate different clusters. There are a total of 30 clusters here, though most are small.

As another example, I obtained the seed dataset from the UC Irvine Machine Learning Dataset Repository (you can find it at <http://archive.ics.uci.edu/ml/datasets/seeds>). Each item consists of seven measurements of a wheat kernel; there are three types of wheat represented in this dataset. As you can see in figures 12.4 and 12.5, this data clusters rather well.

12.4 THE K-MEANS ALGORITHM

We could use any clustering method to vector quantize (Chapter ?? describes a number of different clustering methods in the context of segmentation). However, by far the most common method used is **k-means** clustering. Assume we have a set of data items that we wish to cluster. We now assume that we know how many

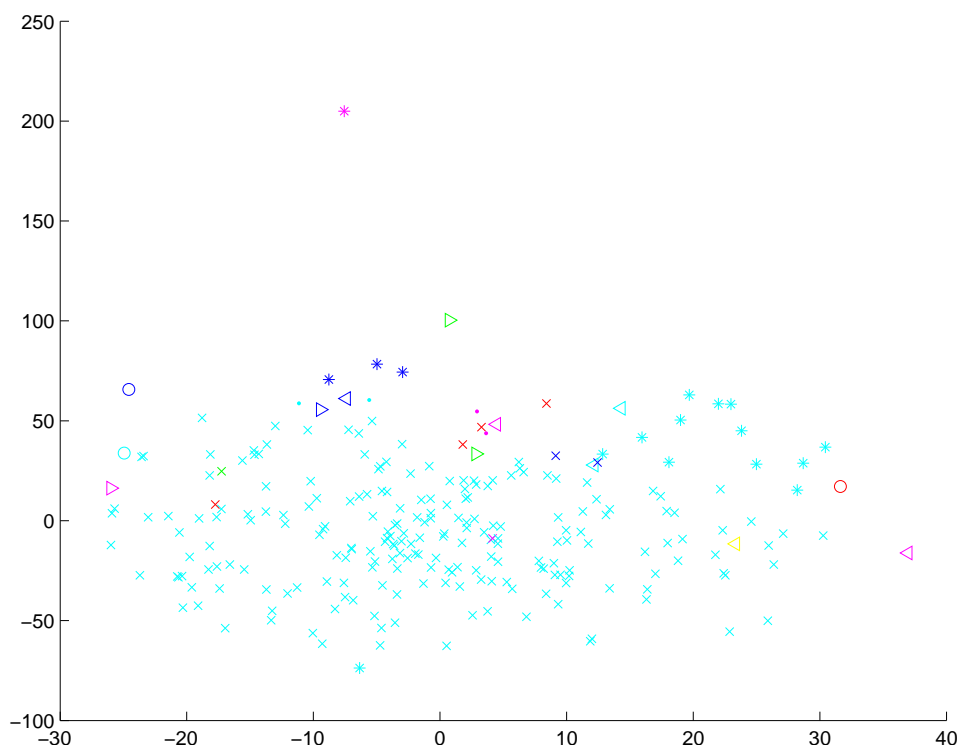


FIGURE 12.3: A clustering of the body fat dataset, using agglomerative clustering, single link distance, and requiring a maximum of 30 clusters. I have plotted each cluster with a distinct marker (though some markers differ only by color; you might need to look at the PDF version to see this figure at its best). Notice that one cluster contains much of the data, and that there are a set of small isolated clusters. The original data is 16 dimensional, which presents plotting problems; I show a scatter plot on the first two principal components (though I computed distances for clustering in the original 16 dimensional space).

clusters there are in the data, which we write k . This is equivalent to fixing the number of values we wish to quantize to. Each cluster is assumed to have a center; we write the center of the i th cluster as \mathbf{c}_i . The j th data item to be clustered is described by a feature vector \mathbf{x}_j . In our case, these items are vectors of filter responses observed at image locations.

Because pattern elements repeat, and so are common, we can assume that most data items are close to the center of their cluster. This suggests that we cluster the data by minimizing the the objective function

$$\Phi(\text{clusters, data}) = \sum_{i \in \text{clusters}} \left\{ \sum_{j \in i\text{th cluster}} (\mathbf{x}_j - \mathbf{c}_i)^T (\mathbf{x}_j - \mathbf{c}_i) \right\}.$$

Notice that if we know the center for each cluster, it is easy to determine which

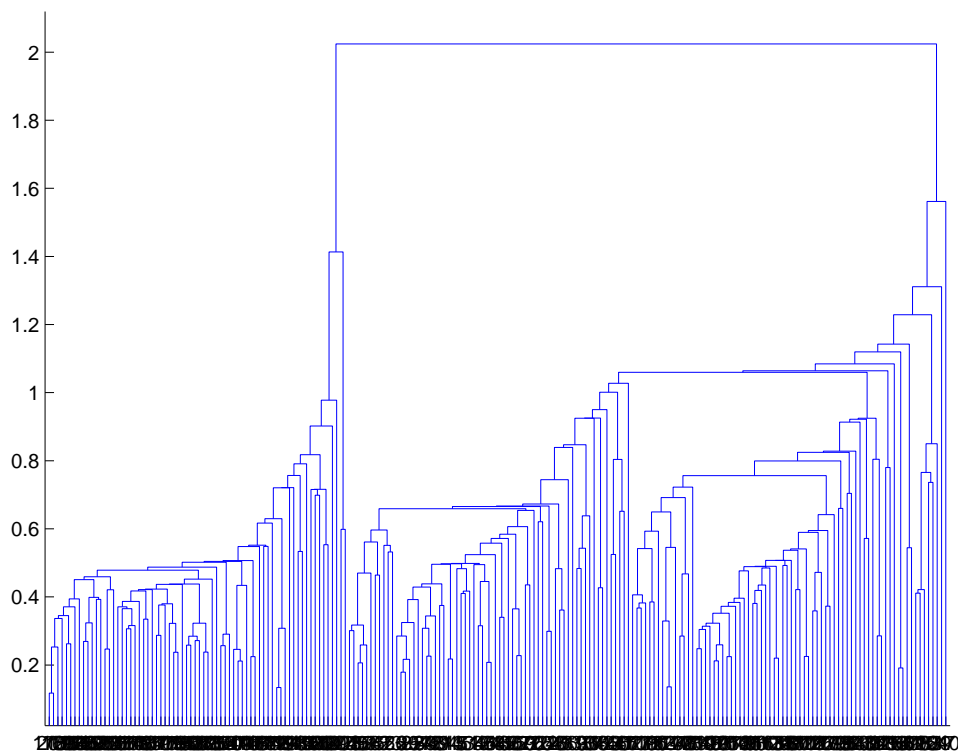


FIGURE 12.4: A dendrogram obtained from the seed dataset, using single link clustering. Recall that the data points are on the horizontal axis, and that the vertical axis is distance; there is a horizontal line linking two clusters that get merged, established at the height at which they're merged. I have plotted the entire dendrogram, despite the fact it's a bit crowded at the bottom, because you can now see how clearly the data set clusters into a small set of clusters — there are a small number of vertical “runs”.

cluster is the best choice for each point. Similarly, if the allocation of points to clusters is known, it is easy to compute the best center for each cluster. However, there are far too many possible allocations of points to clusters to search this space for a minimum. Instead, we define an algorithm that iterates through two activities:

- Assume the cluster centers are known and, allocate each point to the closest cluster center.
- Assume the allocation is known, and choose a new set of cluster centers. Each center is the mean of the points allocated to that cluster.

We then choose a start point by randomly choosing cluster centers, and then iterate these stages alternately. This process eventually converges to a local minimum of the objective function (the value either goes down or is fixed at each step, and it is bounded below). It is not guaranteed to converge to the global minimum of

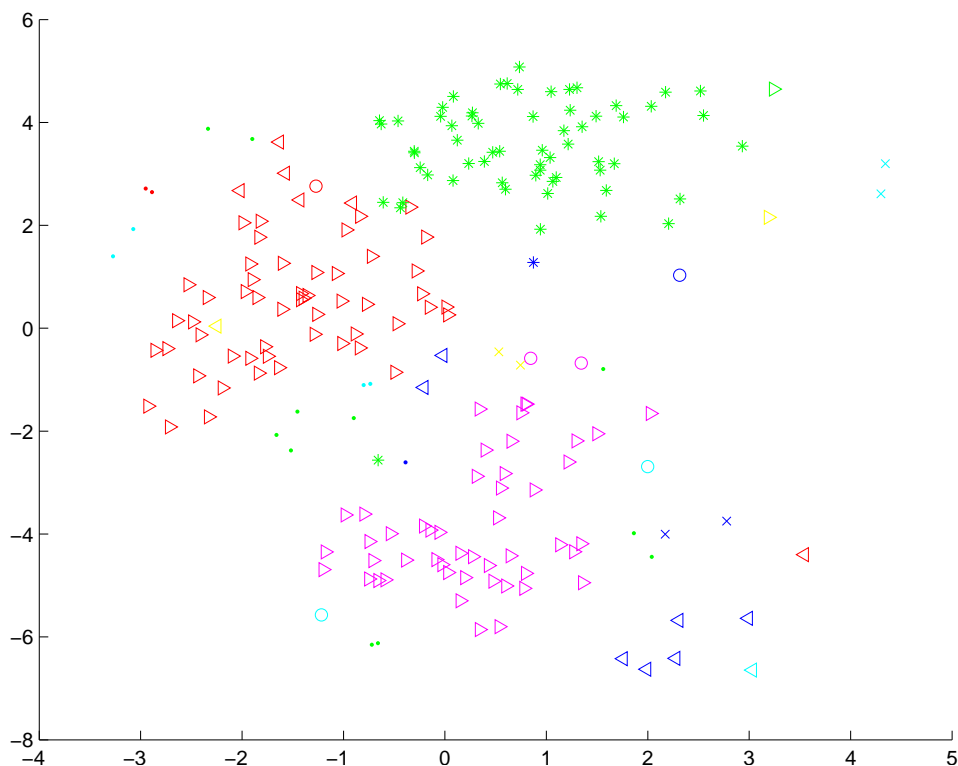


FIGURE 12.5: A clustering of the seed dataset, using agglomerative clustering, single link distance, and requiring a maximum of 30 clusters. I have plotted each cluster with a distinct marker (though some markers differ only by color; you might need to look at the PDF version to see this figure at its best). Notice that there are a set of fairly natural isolated clusters. The original data is 8 dimensional, which presents plotting problems; I show a scatter plot on the first two principal components (though I computed distances for clustering in the original 8 dimensional space).

the objective function, however. It is also not guaranteed to produce k clusters, unless we modify the allocation phase to ensure that each cluster has some nonzero number of points. This algorithm is usually referred to as **k-means** (summarized in Algorithm 12.3). It is possible to search for an appropriate number of clusters by applying k-means for different values of k and comparing the results; we defer a discussion of this issue until Section ??.

```
Choose  $k$  data points to act as cluster centers
Until the cluster centers change very little
  Allocate each data point to cluster whose center is nearest.
  Now ensure that every cluster has at least
  one data point; one way to do this is by
  supplying empty clusters with a point chosen at random from
  points far from their cluster center.
  Replace the cluster centers with the mean of the elements
  in their clusters.
end
```

Algorithm 12.3: *Clustering by K-Means.*

Regression

Classification tries to predict a class from a data item. Regression tries to predict a value. There are several reasons to do this. First, we might actually need to predict a value. For example, we know the zip code of a house, the square footage of its lot, the number of rooms and the square footage of the house, and we wish to predict its likely sale price. As another example, we know the cost and condition of a trading card for sale, and we wish to predict a likely profit in buying it and then reselling it. As yet another example, we have a picture with some missing pixels – perhaps there was text covering them, and we want to replace it – and we want to fill in the missing values. Predicting values is very useful, and so there are many examples like this.

Second, we might want to spot a trend in data. Doing so could make it clear what is really happening. Here is an example from Efron (“COMPUTER-INTENSIVE METHODS IN STATISTICAL REGRESSION”, B. Efron, SIAM Review, 1988). Table 1 shows some data from medical devices, which sit in the body and release a hormone. Each is supposed to have the same behavior. The data describes devices from three production lots (A, B, and C). The important question is: Are the lots the same?

The scatter plot of figure 13.1 suggests what might be going on here; the longer a device has been in service, the less hormone it has. Now the question of whether the lots are different boils down to whether this relationship is different for each lot. To answer this question, we need to determine the relationship between time in service and hormone in more detail. This is what regression is for.

13.1 LEAST SQUARES AND LINEAR MODELS

Assume we have a dataset consisting of a set of N pairs (\mathbf{x}_i, y_i) . We think of y_i as the value of some function evaluated at \mathbf{x}_i , with some random component added. This means there might be two data items where the \mathbf{x}_i are the same, and the y_i are different. We refer to the \mathbf{x}_i as **explanatory variables** and the y_i is a **dependent variable**. We want to build a model of the dependence between y and \mathbf{x} . This model will be used to predict values of y for new values of \mathbf{x} , or to understand the relationships between the \mathbf{x} . The model needs to have some probabilistic component; we do not expect that y is a function of \mathbf{x} , and there is likely some error in evaluating y anyhow.

A good, simple model is to assume that

$$y = \mathbf{x}^T \beta + \xi$$

where ξ is a zero mean normal random variable with unknown variance (we will be able to estimate this later). Here β is a vector of weights. This model might look as though it will have trouble predicting cases where the y intercept is not zero (for

Number	A	B	C
1.	25.8	16.3	28.8
2.	20.5	11.6	22.0
3.	14.3	11.8	29.7
4.	23.2	32.5	28.9
5.	20.6	32.0	32.8
6.	31.1	18.0	32.5
7.	20.9	24.1	25.4
8.	20.9	26.5	31.7
9.	30.4	25.8	28.5

TABLE 13.1: A table showing the amount of hormone remaining in devices from lot A, lot B and lot C. The numbering is arbitrary (i.e. there's no relationship between device 3 in lot A and device 3 in lot B). At first glance, lot C looks rather different from lots A and B.

Number	A	B	C
1.	99	376	119
2.	152	385	188
3.	293	402	115
4.	155	29	88
5.	196	76	58
6.	53	296	49
7.	184	151	150
8.	171	177	107
9.	52	209	125

TABLE 13.2: A table showing the time in service for devices from lot A, lot B and lot C. The numbering is arbitrary (i.e. there's no relationship between device 3 in lot A and device 3 in lot B), but corresponds to the numbering of table 13.2. The time may explain the differing amounts of hormone.

example, $y = ax + b$ where b is not zero). It doesn't, because we can attach a one to the end of each \mathbf{x} , so that it looks like $(x_1, x_2, \dots, x_n, 1)^T$.

We must first determine β . We must have that β minimizes

$$\sum_i (y_i - \mathbf{x}_i^T \beta)^2$$

which we can write more conveniently using vectors and matrices. Write \mathbf{y} for the vector

$$\begin{pmatrix} y_1 \\ y_2 \\ \dots \\ y_n \end{pmatrix}$$

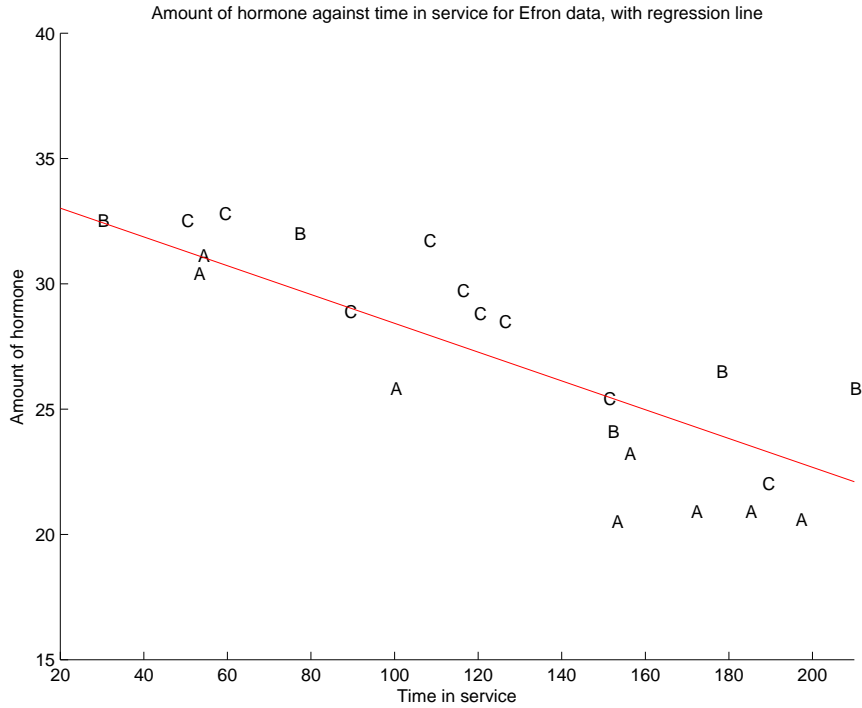


FIGURE 13.1: A scatter plot of hormone against time for devices from tables 13.1 and 13.2. Notice that there is a pretty clear relationship between time and amount of hormone (the longer the device has been in service the less hormone there is). The issue now is to understand that relationship so that we can tell whether lots A, B and C are the same or different. The best fit line to all the data is shown as well, fitted using the methods of section 13.1.

and \mathcal{X} for the matrix

$$\begin{pmatrix} \mathbf{x}_1^T \\ \mathbf{x}_2^T \\ \dots \mathbf{x}_n^T \end{pmatrix}.$$

Then we want to minimize

$$(\mathbf{y} - \mathcal{X}\beta)^T(\mathbf{y} - \mathcal{X}\beta)$$

which means that we must have

$$\mathcal{X}^T \mathcal{X}\beta - \mathcal{X}^T \mathbf{y} = 0.$$

For reasonable choices of features, we could expect that $\mathcal{X}^T \mathcal{X}$ — which should strike you as being a lot like a covariance matrix — has full rank. If it does, this equation is easy to solve. If it does not, there is more to do, which we will do below once we have done some examples.

13.1.1 Interpreting a Regression

For the data of tables 13.1 and 13.2, y is the hormone remaining in the device and $\mathbf{x} = (\text{time}, 1)$. This gives us a model of the hormone in the device as

$$y = \beta_1 \text{time} + \beta_2$$

which should look like the line in figure ???. We get $\beta = (-0.0574, 34.2)$. Now we can ask whether some lots of device behave differently than others. One way to address this is to consider the **residual**,

$$y_i - \mathbf{x}_i^T \beta$$

which is the difference between the observed value and what the model predicts. Look at figure 13.4, which shows this residual plotted against the time. Notice that, for batch A, the model always over predicts, whereas batches B and C seem about the same; this suggests that there is something different about A — any effects caused by the hormone being taken up from the device have been accounted for by our model, and the residual shows the effects that are left.

You can regress anything against anything. For example, we will take the bodyfat dataset. The dependent variable is weight, and the explanatory variables are everything else (except the number of the example). Figure ?? shows the residual for this regression, plotted against example number. Figure ?? shows a regression of boston house prices against a variety of explanatory variables (the data set is quite well known; you can find it at the UCI repository, <http://archive.ics.uci.edu/ml/datasets/Housing>).

13.1.2 How Good is a Fit?

However, it is quite important to know whether the regression is helpful. For example, it seems highly unlikely that regressing the first digit of a telephone number against some numerological score of your name will work that well (or at all). There are some simple facts we can use to make this easier. Assume we ensure that \mathcal{X} always has a column of ones (or twos, or threes, etc.) in it, so that the regression can have a non-zero x-intercept. We now fit a model

$$\mathbf{y} = \mathcal{X}\beta + \mathbf{e}$$

(where \mathbf{e} is the vector of residual values) by choosing β such that $\mathbf{e}^T \mathbf{e}$ is minimized. The facts give some useful technical results.

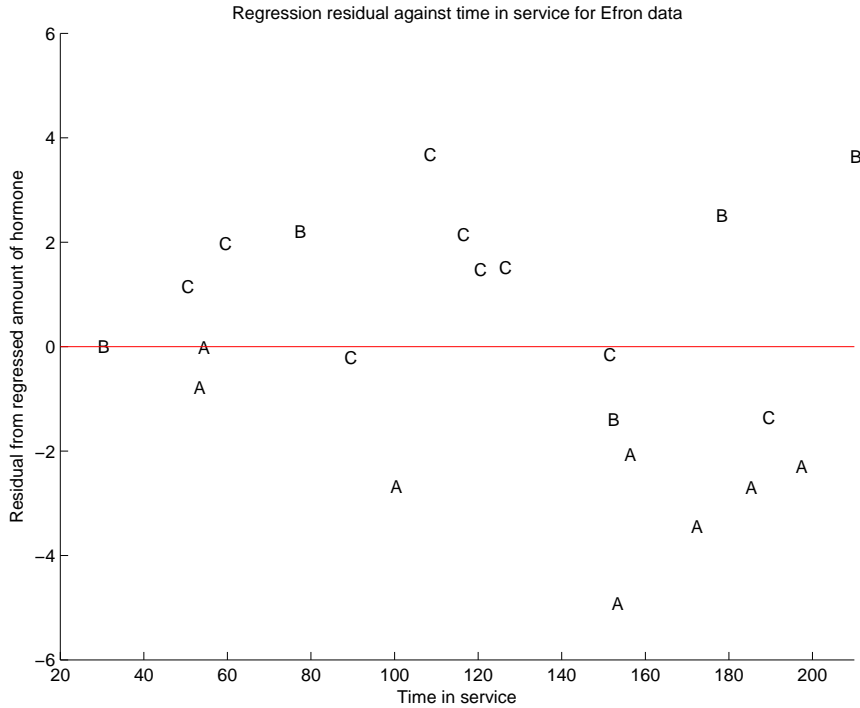


FIGURE 13.2: This is a scatter plot of residual — the distance between each data point and the best fit line — against time for the devices from tables 13.1 and 13.2. Now you should notice a clear difference; some devices from lots B and C have positive and some negative residuals, but all lot A devices have negative residuals. This means that, when we account for loss of hormone over time, lot A devices still have less hormone in them. This is pretty good evidence that there is a problem with this lot.

Useful Facts: 13.1 *Regression*

We write $\mathbf{y} = \mathcal{X}\beta + \mathbf{e}$, where \mathbf{e} is the residual. Assume \mathcal{X} has a column of ones, and β is chosen to minimize $\mathbf{e}^T \mathbf{e}$. Then we have

1. $\mathbf{e}^T \mathcal{X} = \mathbf{0}$, i.e. that \mathbf{e} is orthogonal to any column of \mathcal{X} . This is because, if \mathbf{e} is not orthogonal to some column of \mathcal{X} , we can increase or decrease the β term corresponding to that column to make the error smaller. Another way to see this is to notice that β is chosen to minimize $\mathbf{e}^T \mathbf{e}$, which is $(\mathbf{y} - \mathcal{X}\beta)^T (\mathbf{y} - \mathcal{X}\beta)$. Now because this is a minimum, the gradient with respect to β is zero, so $(\mathbf{y} - \mathcal{X}\beta)^T (-\mathcal{X}) = -\mathbf{e}^T \mathcal{X} = 0$.
2. $\mathbf{e}^T \mathbf{1} = 0$ (recall that \mathcal{X} has a column of all ones, and apply the previous result).
3. $\mathbf{1}^T (\mathbf{y} - \mathcal{X}\beta) = 0$ (same as previous result).
4. $\mathbf{e}^T \mathcal{X}\beta = 0$ (first result means that this is true).

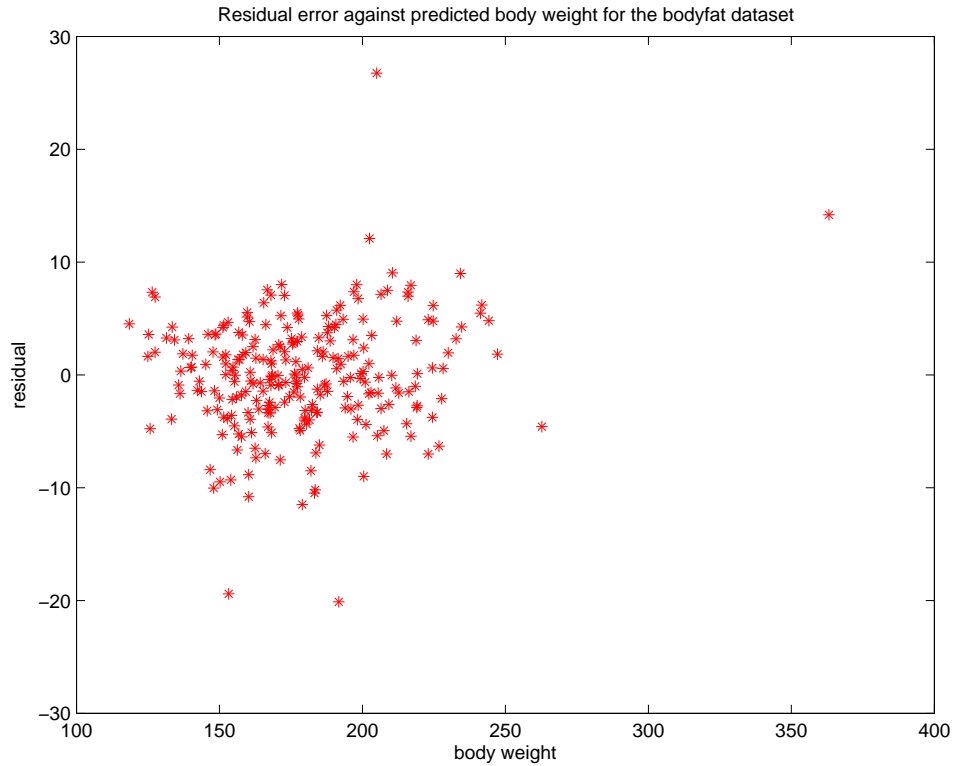


FIGURE 13.3: *This is a scatter plot of residual — the distance between each data point and the best fit line — against weight for the bodyfat dataset. You should notice that there is no particular pattern to the residuals — a good sign; if the residual had a pattern, then there is some structure in the data that the regression is not explaining. However, the residuals are moderately large. The mean weight in this dataset is 178.9, and the standard deviation of the weight is 29.4. The mean of the residual is zero — it has to be — and the standard deviation is 5. This means that the regression is explaining a lot, but not all, of the variance in the weight. This might be due to genuinely random errors, or it may be because there is some explanatory variable missing.*

Now \mathbf{y} is a one dimensional dataset arranged into a vector, so $\text{mean}(\{y_i\})$ is meaningful, as is $\text{var}[y_i]$. Similarly, $\mathcal{X}\beta$ is a one dimensional dataset arranged into a vector (its elements are $\mathbf{x}_i^T\beta$), as is \mathbf{e} , so we know the meaning of mean and variance for each. We have a particularly important result:

$$\text{var}[y_i] = \text{var}[\mathbf{x}_i^T\beta] + \text{var}[e_i].$$

This is quite easy to show, with a little more notation. Write $\bar{\mathbf{y}} = (1/N)(\mathbf{1}^T\mathbf{y})\mathbf{1}$ for

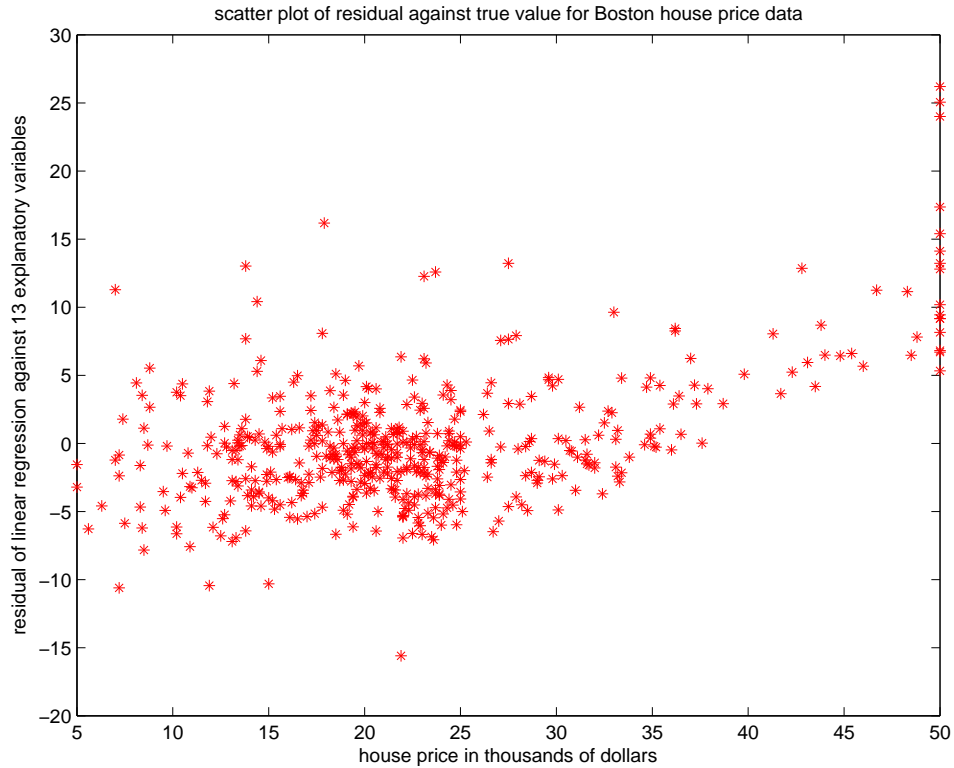


FIGURE 13.4: This is a scatter plot of residual — the distance between each data point and the best fit line — against house price for a data set of Boston house prices. There are a total of 13 explanatory variables here. You should notice that there is no particular pattern to the residuals — a good sign; if the residual had a pattern, then there is some structure in the data that the regression is not explaining. However, the residuals are moderately large. The mean house price in this dataset is 22.5 (thousand's of dollars; this is quite old data) and the standard deviation is 9.2. The mean of the residual is zero — it has to be — and the standard deviation is 4.7. This means that the regression is explaining only about half of the variance in the house prices. This might be due to genuinely random errors, or — more likely — there is some explanatory variable missing.

the vector whose entries are all $\text{mean}(\{y_i\})$; similarly for $\bar{\epsilon}$ and for $\overline{\mathcal{X}\beta}$. We have

$$\text{var}[y_i] = (1/N)(\mathbf{y} - \bar{\mathbf{y}})^T(\mathbf{y} - \bar{\mathbf{y}})$$

and so on for $\text{var}[e_i]$, etc. Notice from the facts that $\bar{\mathbf{y}} = \overline{\mathcal{X}\beta}$. Now

$$\begin{aligned} \text{var}[y_i] &= (1/N) \left([\mathcal{X}\beta - \overline{\mathcal{X}\beta}] + [\mathbf{e} - \bar{\mathbf{e}}] \right)^T \left([\mathcal{X}\beta - \overline{\mathcal{X}\beta}] + [\mathbf{e} - \bar{\mathbf{e}}] \right) \\ &= (1/N) \left([\mathcal{X}\beta - \overline{\mathcal{X}\beta}]^T [\mathcal{X}\beta - \overline{\mathcal{X}\beta}] + 2[\mathbf{e} - \bar{\mathbf{e}}]^T [\mathcal{X}\beta - \overline{\mathcal{X}\beta}] + [\mathbf{e} - \bar{\mathbf{e}}]^T [\mathbf{e} - \bar{\mathbf{e}}] \right) \\ &= (1/N) \left([\mathcal{X}\beta - \overline{\mathcal{X}\beta}]^T [\mathcal{X}\beta - \overline{\mathcal{X}\beta}] + [\mathbf{e} - \bar{\mathbf{e}}]^T [\mathbf{e} - \bar{\mathbf{e}}] \right) \\ &\quad \text{because } \bar{\mathbf{e}} = \mathbf{0} \text{ and } \mathbf{e}^T \mathcal{X}\beta = 0 \text{ and } \mathbf{e}^T \mathbf{1} = 0 \\ &= \text{var}[\mathbf{x}_i^T \beta] + \text{var}[e_i]. \end{aligned}$$

This is extremely important, because it allows us to think about a regression as explaining variance in \mathbf{y} . As we are better at explaining \mathbf{y} , $\text{var}[e_i]$ goes down. A natural measure of the goodness of a regression is what percentage of the variance of \mathbf{y} it explains. This is known as R^2 (the r-squared measure). We have

$$R^2 = \frac{\text{var}[\mathbf{x}_i^T \beta]}{\text{var}[y_i]}$$

which gives some sense of how well the regression explains the training data.

One natural use of a regression is to try and find potential explanatory variables. It is quite natural to want to look at the entries in β , and wonder what they suggest about the real world. This is something that should be done with caution. For example, you might think that a β component with large absolute value is evidence that a variable is important; but it might just be evidence that that variable is poorly scaled. Similarly, you might want to try and add an explanatory variable to see if the R^2 of the regression improves; but you will find that it always does, because β is chosen to achieve the smallest possible value of R^2 .

There is a procedure from hypothesis testing that will tell whether an improvement in a regression is significant (and so whether one model is better than another). Assume that we have two **nested models**. This means that model 2 sees the same explanatory variables as model one, and some extra explanatory variables as well. Write n_1 for the number of explanatory variables in model 1, n_2 for the number of explanatory variables in model 2, \mathbf{e}_1 for the residual under model 1, \mathbf{e}_2 for the residual under model 2, s_1 for $\mathbf{e}_1^T \mathbf{e}_1$, s_2 for $\mathbf{e}_2^T \mathbf{e}_2$, and N for the number of data points. If we compute the statistic

$$\frac{\frac{s_2 - s_1}{n_2 - n_1}}{\frac{s_2}{N - n_2}}$$

then, under the null hypothesis that model 2 is as good at explaining the data as model 1, this statistic will have an F-distribution with $(n_2 - n_1, N - n_2)$ degrees of freedom. This means you do what we are used to, after chapter 1; you compute this number, look up the probability of getting this or a larger number under the null hypothesis in tables (which is where you use the degrees of freedom), and then use the significance level to decide whether to prefer model 2 to model 1 or not.

With this trick, it is natural to use regression as a tool to try and determine what effects are important. We can do this by repeatedly (a) inserting plausible explanatory variables then (b) checking whether the regression has improved or not.

We could obtain plausible explanatory variables by looking for new measurements (for example, does the acuity of color perception affect age at death?). We could take non-linear functions of existing explanatory variables. Natural choices include powers and logarithms.

Notice also that for some problems we can expect a power law — something like $y = x_1^{a_1} x_2^{a_2} \dots x_n^{a_n}$. In this case, we have $\log y = a_1 \log x_1 + a_2 \log x_2 + \dots a_n \log x_n$, so it makes sense to take logs and then apply a linear regression.

13.1.3 Regularizing Linear Regressions

One occasionally important difficulty is that the explanatory variables might be significantly correlated. If they are, then it will generally be easy to predict one explanatory variable from another. This means that $\mathcal{X}^T \mathcal{X}$ may have some very small eigenvalues (because there is a vector \mathbf{u} so that $\mathcal{X}\mathbf{u}$ is small; this means that $\mathbf{u}^T \mathcal{X}^T \mathcal{X} \mathbf{u}$ must be small).

These small eigenvalues lead to bad predictions. If $\mathcal{X}^T \mathcal{X}$ has a small eigenvalue, then there is some vector \mathbf{v} such that $\mathcal{X}^T \mathcal{X} \mathbf{v}$ is small, or, equivalently, that the matrix can turn large vectors into small ones; but that means that $(\mathcal{X}^T \mathcal{X})^{-1}$ will turn some small vectors into big ones. In turn, this means that small errors in \mathbf{y} — which are likely inevitable — will result in big errors in β . This could cause trouble in two ways. If we are looking at β to tell which explanatory variables are important, then large errors in β will be a problem. And if we are trying to predict new y values, we expect that errors in β turn into errors in prediction.

An important and useful way to suppress these errors is to regularize, using the same trick we saw in the case of classification. Instead of choosing β to minimize

$$\sum_i (y_i - \mathbf{x}_i^T \beta)^2 = (\mathbf{y} - \mathcal{X}\beta)^T (\mathbf{y} - \mathcal{X}\beta)$$

we minimize

$$\sum_i (y_i - \mathbf{x}_i^T \beta)^2 + \lambda \beta^T \beta = (\mathbf{y} - \mathcal{X}\beta)^T (\mathbf{y} - \mathcal{X}\beta) + \lambda \beta^T \beta$$

where $\lambda > 0$ is a constant. We choose λ in the same way we used for classification; split the training set into a training piece and a validation piece, train for different values of λ , and test the resulting regressions on the validation piece. We choose the λ that yields the smallest validation error. Notice we could use multiple splits, and average over the splits.

This helps, because to solve for β we must solve the equation

$$(\mathcal{X}^T \mathcal{X} + \lambda \mathcal{I})\beta = \mathcal{X}^T \mathbf{y}$$

(obtained by differentiating with respect to β and setting to zero) and the smallest eigenvalue of the matrix $(\mathcal{X}^T \mathcal{X} + \lambda \mathcal{I})$ will be at least λ .

L_1 Regularization and the Lasso

If we are going to use linear regression to recover a set of explanatory variables, then we want some terms in β — terms corresponding to explanatory variables that

aren't relevant — to be zero. This will not happen, unless we are very lucky, with our current schemes. Usually, numerical accidents and sampling effects mean that an explanatory variable will be slightly correlated with the dependent variable. Furthermore, our regularization does not help much, because the regularizer is squared. There is no particular advantage to driving a small β to zero, because the cost of it being non-zero is tiny.

There is a very clever method to encourage β to have zeros in it. We regularize with what is known as an L_1 norm. Rather than use $\beta^T \beta$, we use $\|\beta\|_1 = \sum_j |\beta_j|$. This means we must choose *beta* to minimize

$$\sum_i (y_i - \mathbf{x}_i^T \beta)^2 + \lambda \|\beta\|_1 = (\mathbf{y} - \mathcal{X}\beta)^T (\mathbf{y} - \mathcal{X}\beta) + \lambda \|\beta\|_1 .$$

This strategy is known as the **lasso**. It encourages β to have zeros, because small non-zero entries in β are relatively expensive. There are good numerical methods for obtaining β under this scheme, which are way beyond our scope. In practice, this method is very effective when \mathbf{x} is large, and it is the method to be preferred. There are also strong mathematical results suggesting that the choice of explanatory variables one comes up with is quite good.

Learning to Classify

A **classifier** is a procedure that accepts a set of features and produces a class label for them. There could be two, or many, classes, though it is usual to produce multi-class classifiers out of two-class classifiers. Classifiers are immensely useful, and find wide application, because many problems are naturally decision problems. For example, if you wish to determine whether to place an advert on a web-page or not, you would use a classifier (i.e. look at the page, and say yes or no according to some rule). As another example, if you have a program that you found for free on the web, you would use a classifier to decide whether it was safe to run it (i.e. look at the program, and say yes or no according to some rule). As yet another example, you can think of doctors as extremely complex multi-class classifiers.

Classifiers are built by taking a set of labeled examples and using them to come up with a rule that assigns a label to any new example. In the general problem, we have a training dataset (\mathbf{x}_i, y_i) ; each of the **feature vectors** \mathbf{x}_i consists of measurements of the properties of different types of object, and the y_i are labels giving the type of the object that generated the example.

14.1 CLASSIFICATION, ERROR, AND LOSS

You should think of a classifier as a rule, though it might not be implemented that way. We pass in a feature vector, and the rule returns a class label. We know the relative costs of mislabeling each class and must come up with a rule that can take any plausible \mathbf{x} and assign a class to it, in such a way that the expected mislabeling cost is as small as possible, or at least tolerable. For most of this chapter, we will assume that there are two classes, labeled 1 and -1 . Section 14.5.2 shows methods for building multi-class classifiers from two-class classifiers.

14.1.1 Using Loss to Determine Decisions

The choice of classification rule must depend on the cost of making a mistake. A two-class classifier can make two kinds of mistake. A **false positive** occurs when a negative example is classified positive; a **false negative** occurs when a positive example is classified negative. For example, pretend there is only one disease; then doctors would be classifiers, deciding whether a patient had it or not. If this disease is dangerous, but is safely and easily treated, then false negatives are expensive errors, but false positives are cheap. Similarly, if it is not dangerous, but the treatment is difficult and unpleasant, then false positives are expensive errors and false negatives are cheap.

Generally, we write outcomes as $(i \rightarrow j)$, meaning that an item of type i is classified as an item of type j . There are four outcomes for the two-class case. Each outcome has its own cost, which is known as a **loss**. Hence, we have a loss function that we write as $L(i \rightarrow j)$, meaning the loss incurred when an object of type i is

classified as having type j . Since losses associated with correct classification should not affect the design of the classifier, $L(i \rightarrow i)$ must be zero, but the other losses could be any positive numbers.

The **risk function** of a particular classification strategy is the expected loss when using that strategy, as a function of the kind of item. The **total risk** is the total expected loss when using the classifier. The total risk depends on the strategy, but not on the examples. Write $p(-1 \rightarrow 1|\text{using } s)$ for the probability that class -1 is labeled class 1 (and so on). Then, if there were two classes, the total risk of using strategy s would be

$$R(s) = p(-1 \rightarrow 1|\text{using } s)L(-1 \rightarrow 1) + p(1 \rightarrow -1|\text{using } s)L(-1 \rightarrow 1).$$

The desirable strategy is one that minimizes this total risk.

A Two-class Classifier that Minimizes Total Risk

Assume that the classifier can choose between two classes and we have a known loss function. There is some boundary in the feature space, which we call the **decision boundary**, such that points on one side belong to class one and points on the other side to class two.

We can resort to a trick to determine where the decision boundary is. If the decision boundary is optimal, then *for points on the decision boundary*, either choice of class has the same expected loss; if this weren't so, we could obtain a better classifier by always choosing one class (and so moving the boundary). This means that, for measurements on the decision boundary, choosing label -1 yields the same expected loss as choosing label 1 .

Now write $p(-1|\mathbf{x})$ for the posterior probability of label -1 given feature vector \mathbf{x} (and so on). Although this might be very hard to know in practice, we can manipulate the abstraction and gain some insight. A choice of label $y = 1$ for a point \mathbf{x} at the decision boundary yields an expected loss

$$p(-1|\mathbf{x})L(-1 \rightarrow 1) + p(1|\mathbf{x})L(1 \rightarrow 1) = p(-1|\mathbf{x})L(-1 \rightarrow 1),$$

and if we choose the other label, the expected loss is

$$p(1|\mathbf{x})L(1 \rightarrow -1),$$

and these two terms must be equal. This means our decision boundary consists of the points \mathbf{x} , where

$$p(-1|\mathbf{x})L(-1 \rightarrow 1) = p(1|\mathbf{x})L(1 \rightarrow -1).$$

At points off the boundary, we must choose the class with the *lowest* expected loss. Recall that if we choose label 1 for a point \mathbf{x} , the expected loss is

$$p(-1|\mathbf{x})L(-1 \rightarrow 1),$$

and so on. This means that we should choose label -1 if

$$p(-1|\mathbf{x})L(-1 \rightarrow 1) > p(1|\mathbf{x})L(1 \rightarrow -1)$$

and label 1 if the inequality is reversed. Notice it does not matter which label we choose at the decision boundary.

A Multi-class Classifier that Minimizes Total Risk

Analyzing expected loss gives a strategy for choosing from any number of classes. We allow the option of refusing to decide which class an object belongs to, which is useful in some problems. Refusing to decide costs d . Conveniently, if d is larger than any misclassification loss, we will never refuse to decide. This means our analysis covers the case when we are forced to decide. The same reasoning applies as above, but there are more boundaries to consider. The simplest case, which is widely dominant in vision, is when loss is **0-1 loss**; here the correct answer has zero loss, and any error costs one.

In this case, the best strategy, known as the **Bayes classifier**, is given in Algorithm 14.1. The total risk associated with this rule is known as the **Bayes risk**; this is the smallest possible risk that we can have using a classifier for this problem. It is usually rather difficult to know what the Bayes classifier—and hence the Bayes risk—is because the probabilities involved are not known exactly. In a few cases, it is possible to write the rule out explicitly. One way to tell the effectiveness of a technique for building classifiers is to study the behavior of the risk as the number of examples increases (e.g., one might want the risk to converge to the Bayes risk in probability if the number of examples is large). The Bayes risk is seldom zero, as Figure 14.1 illustrates.

For a loss function

$$L(i \rightarrow j) = \begin{cases} 1 & i \neq j \\ 0 & i = j \\ d & \text{no decision} \end{cases}$$

the best strategy is

- if $p(k|\mathbf{x}) > p(i|\mathbf{x})$ for all i not equal to k , and if this probability is greater than $1 - d$, choose type k ;
- if there are several classes $k_1 \dots k_j$ for which $p(k_1|\mathbf{x}) = p(k_2|\mathbf{x}) = \dots = p(k_j|\mathbf{x}) = p > p(i|\mathbf{x})$ for all i not in k_1, \dots, k_j , and if $p > 1 - d$, choose uniformly and at random between k_1, \dots, k_j ;
- if for all i we have $1 - d \geq q = p(k|\mathbf{x}) \geq p(i|\mathbf{x})$, refuse to decide.

Algorithm 14.1: *The Bayes Classifier.*

14.1.2 Example: Building a Classifier out of Histograms

As we saw in chapter 1, histograms reveal a great deal about data. A histogram is a representation of a probability distribution, and so we can use histograms to build classifiers. Recall that in the pizza data of chapter 1 (you can find a version of this dataset, along with a neat backstory, at http://www.amstat.org/publications/jse/jse_data_archive.htm), there were pizzas from two manufacturers — Eagle-Boys and Dominos. For these pizza's, we had the diameter, and some information

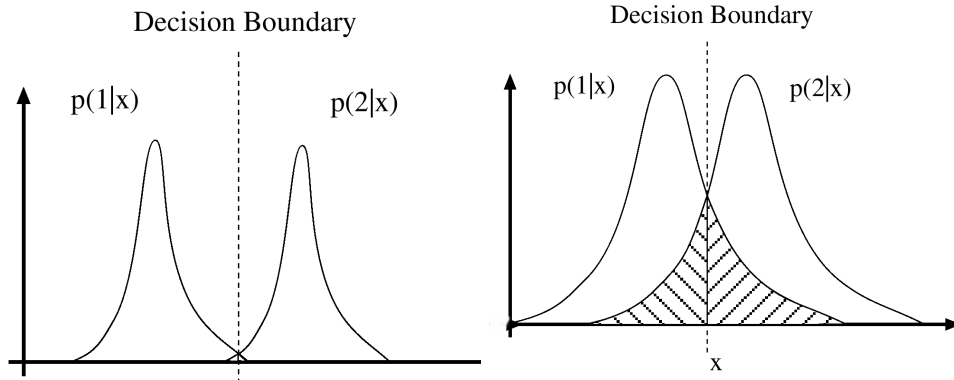


FIGURE 14.1: This figure shows typical elements of a two-class classification problem. We have plotted $p(\text{class}|x)$ as a function of the feature x . Assuming that $L(-1 \rightarrow 1) = L(1 \rightarrow -1)$, we have marked the classifier boundaries. In this case, the Bayes risk is the sum of the amount of the posterior for class one in the class two region and the amount of the posterior for class two in the class one region (the hatched area in the figures). For the case on the left, the classes are well separated, which means that the Bayes risk is small; for the case on the right, the Bayes risk is rather large.

about the topping. We could then try to predict the manufacturer from the diameter of the pizza.

We have two classes (EagleBoys and Dominos). Write x for the diameter of a pizza. As we have seen, to classify we need a model of $p(E|x)$. Since there are two classes, $p(D|x) = 1 - p(E|x)$, so the model of $p(E|x)$ is enough. A natural way to get this model is to use Bayes rule:

$$p(E|x) = \frac{p(x|E)p(E)}{p(x)}$$

We can model $p(x|E)$, $p(E)$ and $p(x)$ with histograms. We construct a set of boxes of appropriate sizes, then count data items into the boxes. The conditional probability that the diameter of a pizza lies in a box, conditioned on its coming from EagleBoys is $p(x|E)$; this can be estimated by taking the number of EagleBoys pizzas in that box, and dividing by the total number of EagleBoys pizzas. Similarly, we can estimate $p(E)$ by the fraction of all pizzas that come from EagleBoys. Finally, we could get $p(x)$ by taking the fraction of all pizzas in the relevant histogram box.

However, we don't really need $p(x)$. Write $L(E \rightarrow D)$, etc. for the loss of misclassifying pizzas. Recall that we want to test

$$p(E|x)L(E \rightarrow D) > p(D|x)L(D \rightarrow E)$$

but this is

$$\frac{p(x|E)p(E)}{p(x)}L(E \rightarrow D) > \frac{p(x|D)p(D)}{p(x)}L(D \rightarrow E).$$

The first thing to notice about this expression is that only the ratio of the losses matters; we could write an equivalent test as

$$\frac{p(x|E)}{p(x|D)} > \frac{p(D) L(D \rightarrow E)}{p(E) L(E \rightarrow D)}.$$

If we do not know the losses, we can simply test

$$\frac{p(x|E)}{p(x|D)} > \theta$$

against a variety of different thresholds θ , and then choose the one that works best (for example, in a test on a different dataset).

Listing 14.1 shows a simple histogram-based classifier for the pizza data. To train this classifier, I split the pizza data into two pools — training data, and test data (more on this in the next section). I built the histograms on the training data, then evaluated the classifier on test data for different values of the threshold in listing 14.2

Different values of the threshold give different false-positive and false-negative rates. This information can be summarized with a **receiver operating characteristic curve**, or **ROC**. This curve is a plot of the **detection rate** or **true positive rate** as a function of the **false positive rate** for a particular model as the threshold changes (Figure ??). An ideal model would detect all positive cases and produce no false positives, for any threshold value; in this case, the curve would be a single point. A model that has no information about whether an example is a positive or a negative will produce the line from $(0, 0)$ to $(1, 1)$. If the ROC lies below this line, then we can produce a better classifier by inverting the decision of the original classifier, so this line is the worst possible classifier. The detection rate never goes down as the false positive rate goes up, so the ROC is the graph of a non-decreasing function. Figure 14.2 shows the ROC for the pizza classifier.

Models of a classification problem can be compared by comparing their ROC's. Alternatively, we can build a summary of the ROC. Most commonly used in computer vision is the area under the ROC (the **AUC**), which is 1 for a perfect classifier, and 0.5 for a classifier that has no information about the problem. The area under the ROC has the following interpretation: assume we select one positive example and one negative example uniformly at random, and display them to the classifier; the AUC is the probability that the classifier tells correctly which of these two is positive.

14.1.3 Example: Classifying Using Nearest Neighbors

Example points near an unclassified point should indicate the class of that point. **Nearest neighbors** methods build classifiers using this heuristic. We could classify a point by using the class of the nearest example whose class is known, or use several example points and make them vote. It is reasonable to require that some minimum number of points vote for the class we choose.

A (k, l) nearest neighbor classifier finds the k example points closest to the point being considered, and classifies this point with the class that has the highest number of votes, as long as this class has more than l votes (otherwise, the point

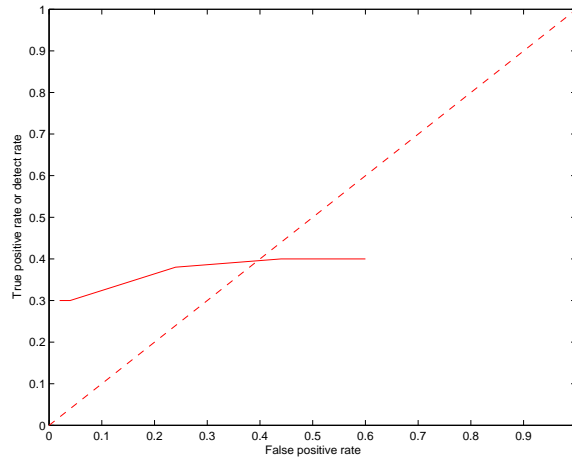


FIGURE 14.2: *The receiver operating curve for a histogram-based classifier to tell EagleBoys pizza from Domino's pizza using the diameter. This curve plots the detection rate against the false-negative rate for a variety of values of the parameter θ . A perfect classifier has an ROC that, on these axes, is a horizontal line at 100% detection. This classifier isn't very good, as you would expect if you look back at the histograms for the pizza data.*

is classified as unknown). A $(k, 0)$ -nearest neighbor classifier is usually known as a **k-nearest neighbor classifier**, and a $(1, 0)$ -nearest neighbor classifier is usually known as a **nearest neighbor classifier**.

Nearest neighbor classifiers are known to be good, in the sense that the risk of using a nearest neighbor classifier with a sufficiently large number of examples lies within quite good bounds of the Bayes risk. As k grows, the difference between the Bayes risk and the risk of using a k -nearest neighbor classifier goes down as $1/\sqrt{k}$. In practice, one seldom uses more than three nearest neighbors. Furthermore, if the Bayes risk is zero, the expected risk of using a k -nearest neighbor classifier is also zero. Finding the k nearest points for a particular query can be difficult, and Section 14.5.3 reviews this point.

A second difficulty in building such classifiers is the choice of distance. For features that are obviously of the same type, such as lengths, the usual metric may be good enough. But what if one feature is a length, one is a color, and one is an angle? One possibility is to whiten the features (section 1). This may be hard if the dimension is so large that the covariance matrix is hard to estimate. It is almost always a good idea to scale each feature independently so that the variance of each feature is the same, or at least consistent; this prevents features with very large scales dominating those with very small scales.

14.1.4 Training Error, Test Error, and Overfitting

It can be quite difficult to know a good loss function, but one can usually come up with a plausible model. If we knew the posterior probabilities, building a classifier

would be straightforward. Usually we don't, and must build a model from data. This model could be a model of the posterior probabilities, or an estimate of the decision boundaries. In either case, we have only the training data to build it with. **Training error** is the error a model makes on the training data set.

Generally, we will try to make this training error small. However, what we really want to minimize is the **test error**, the error the classifier makes on test data. We cannot minimize this error directly, because we don't know the test set (if we did, special procedures in training apply). However, classifiers that have small training error might not have small test error. One example of this problem is the (silly) classifier that takes any data point and, if it is the same as a point in the training set, emits the class of that point and otherwise chooses randomly between the classes. This classifier has been learned from data, and has a zero error rate on the training dataset; it is likely to be unhelpful on any other dataset, however.

The phenomenon that causes test error to be worse than training error is sometimes called **overfitting** (other names include **selection bias**, because the training data has been selected and so isn't exactly like the test data, and **generalizing badly**, because the classifier fails to generalize). It occurs because the classifier has been trained to perform well *on the training dataset*. The training dataset is not the same as the test dataset. First, it is quite likely smaller. Second, it might be biased through a variety of accidents. This means that small training error may have to do with quirks of the training dataset that don't occur in other sets of examples. It is quite possible that, in this case, the test error will be larger than the training error. Generally, we expect classifiers to perform somewhat better on the training set than on the test set. Overfitting can result in a substantial difference between performance on the training set and performance on the test set. One consequence of overfitting is that classifiers should always be evaluated on test data. Doing this creates other problems, which we discuss in Section 14.1.5.

A procedure called **regularization** attaches a penalty term to the training error to get a better estimate of the test error. This penalty term could take a variety of different forms, depending on the requirements of the application. Section 1 describes regularization in further detail.

14.1.5 Error Rate and Cross-Validation

There are a variety of methods to describe the performance of a classifier. Natural, straightforward choices are to report the **error rate**, the percentage of classification attempts on a test set that result in the wrong answer. This presents an important difficulty. We cannot estimate the error rate of the classifier using training data, because the classifier has been trained to do well on that data, which will mean our error rate estimate will be an underestimate. An alternative is to separate out some training data to form a **validation set**, then train the classifier on the rest of the data, and evaluate on the validation set. This has the difficulty that the classifier will not be the best estimate possible, because we have left out some training data when we trained it. This issue can become a significant nuisance when we are trying to tell which of a set of classifiers to use—did the classifier perform poorly on validation data because it is not suited to the problem representation or because

it was trained on too little data?

We can resolve this problem with **cross-validation**, which involves repeatedly: splitting data into training and validation sets uniformly and at random, training a classifier on the training set, evaluating it on the validation set, and then averaging the error over all splits. This allows an estimate of the likely future performance of a classifier, at the expense of substantial computation.

Choose some class of subsets of the training set, for example, singletons.

For each element of that class, construct a classifier by omitting that element in training, and compute the classification errors (or risk) on the omitted subset.

Average these errors over the class of subsets to estimate the risk of using the classifier trained on the entire training dataset.

Algorithm 14.2: *Cross-Validation*

The most usual form of this algorithm involves omitting single items from the dataset and is known as **leave-one-out cross-validation**. Errors are usually estimated by simply averaging over the class, but more sophisticated estimates are available. We do not justify this tool mathematically; however, it is worth noticing that leave-one-out cross-validation, in some sense, looks at the sensitivity of the classifier to a small change in the training set. If a classifier performs well under this test, then large subsets of the dataset look similar to one another, which suggests that a representation of the relevant probabilities derived from the dataset might be quite good.

For a multi-class classifier, it is often helpful to know which classes were misclassified. We can compute a **class-confusion matrix**, a table whose i, j th entry is the number of times an item of true class i was labeled j by the classifier (notice that this definition is not symmetric). If there are many classes, this matrix can be rendered as an image (Figure 14.3), where the intensity values correspond to counts; typically, larger values are lighter. Such images are quite easy to assess at a glance. One looks for a light diagonal (because the diagonal elements are the counts of correct classifications), for any row that seems dark (which means that there were few elements in that class), and for bright off-diagonal elements (which are high-frequency misclassifications).

14.2 BASIC IDEAS FOR NUMERICAL MINIMIZATION

Assume we have a function $g(\mathbf{a})$, and we wish to obtain a value of \mathbf{a} that achieves the minimum for that function. Sometimes we can solve this problem in closed form by constructing the gradient and finding a value of \mathbf{a} that makes the gradient zero. More usually we need a numerical method. Implementing these numerical

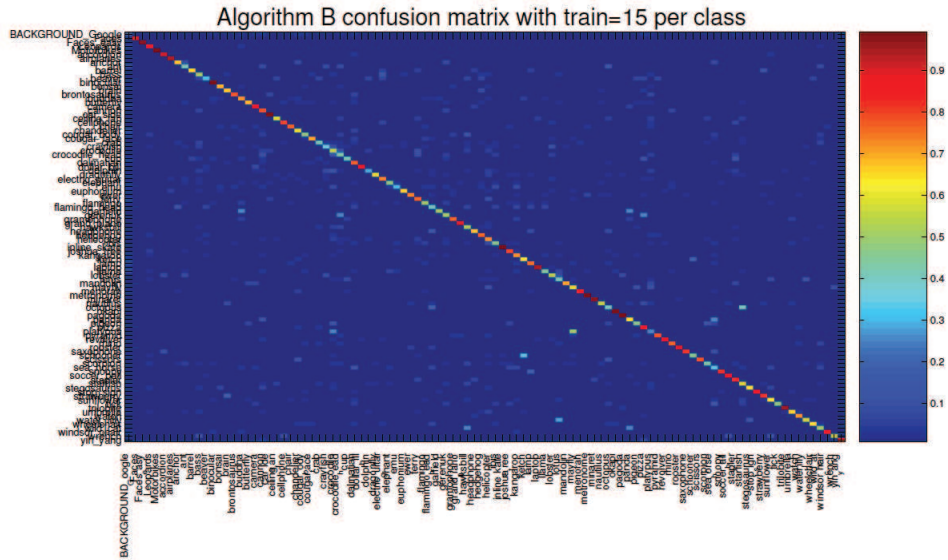


FIGURE 14.3: An example of a class confusion matrix from a recent image classification system, due to ?. The vertical bar shows the mapping of color to number (warmer colors are larger numbers). Note the redness of the diagonal; this is good, because it means the diagonal values are large. There are spots of large off-diagonal values, and these are informative, too. For example, this system confuses: schooners and ketches (understandable); waterlily and lotus (again, understandable); and platypus and mayfly (which might suggest some feature engineering would be a good idea). This figure was originally published as Figure 5 of “SVM-KNN: Discriminative Nearest Neighbor Classification for Visual Category Recognition,” by H. Zhang, A. Berg, M. Maire, and J. Malik, Proc. IEEE CVPR, 2006, © IEEE, 2006.

methods is a specialized business, and it is usual to use general optimization codes. This section is intended to sketch how such codes work, so you can read manual pages, etc. more effectively. Personally, I am a happy user of Matlab’s `fminunc`, although the many different settings take some getting used to.

14.2.1 Overview

Typical codes take a description of the objective function (typically, the name of a function), a start point for the search, and a collection of parameters. All codes take an estimate $\mathbf{a}^{(i)}$, update it to $\mathbf{a}^{(i+1)}$, then check to see whether the result is a minimum. This process is started from the start point. The update is usually obtained by computing a direction $\mathbf{p}^{(i)}$ such that for small values of h , $g(\mathbf{a}^{(i)} + h\mathbf{p}^{(i)})$ is smaller than $g(\mathbf{a}^{(i)})$. Such a direction is known as a **descent direction**.

Assume we have a descent direction. We must now choose how far to travel along that direction. We can see $g(\mathbf{a}^{(i)} + h\mathbf{p}^{(i)})$ as a function of h . Write this function as $\phi(h)$. We start at $h = 0$ (which is the original value $\mathbf{a}^{(i)}$, so $\phi(0) = g(\mathbf{a}^{(i)})$), and

move in the direction of increasing h to find a small value of $\phi(h)$ that is less than $\phi(0)$. The descent direction was chosen so that for small $h > 0$, $\phi(h) < \phi(0)$; one way to tell we are at a minimum is we cannot choose a descent direction. Searching for a good value of h is known as **line search**. Typically, this search involves a sequence of estimated values of h , which we write h_i . One algorithm is to start with (say) $h_0 = 1$; if $\phi(h_i)$ is not small enough (and there are other tests we may need to apply — this is a summary!), we compute $h_{(i+1)} = (1/2)h_i$. This stops when some h_i passes a test, or when it is so small that the step is pointless.

14.2.2 Gradient Descent

One method to choose a descent direction is **gradient descent**, which uses the negative gradient of the function. Recall our notation that

$$\mathbf{a} = \begin{pmatrix} a_1 \\ a_2 \\ \dots \\ a_d \end{pmatrix}$$

and that

$$\nabla g = \begin{pmatrix} \frac{\partial g}{\partial a_1} \\ \frac{\partial g}{\partial a_2} \\ \dots \\ \frac{\partial g}{\partial a_d} \end{pmatrix}.$$

We can write a Taylor series expansion for the function $g(\mathbf{a}^{(i)} + h\mathbf{p}^{(i)})$. We have that

$$g(\mathbf{a}^{(i)} + h\mathbf{p}^{(i)}) = g(\mathbf{a}^{(i)}) + h(\nabla g)^T \mathbf{p}^{(i)} + O(h^2)$$

This means that we can expect that if

$$\mathbf{p}^{(i)} = -\nabla g(\mathbf{a}^{(i)}),$$

we expect that, at least for small values of h , $g(\mathbf{a}^{(i)} + h\mathbf{p}^{(i)})$ will be less than $g(\mathbf{a}^{(i)})$.

This works (as long as g is differentiable, and quite often when it isn't) because g must go down for at least small steps in this direction. There are two ways to evaluate a gradient. You can require that the software estimate a numerical derivative for you, which usually slows things down somewhat, or you can supply a gradient value. Usually this gradient value must be computed by the same function that computes the objective function value.

One tip: in my experience, about 99% of problems with numerical optimization codes occur because the user didn't check that the gradient their function computed is right. Most codes will compute a numerical gradient for you, then check that against your gradient; if they're sufficiently different, the code will complain. You don't want to do this at runtime, because it slows things up, but it's an excellent idea to check.

14.2.3 Stochastic Gradient Descent

Assume we wish to minimize some function $g(\mathbf{a}) = g_0(\mathbf{a}) + (1/N) \sum_{i=1}^N g_i(\mathbf{a})$, as a function of \mathbf{a} . Gradient descent would require us to form

$$-\nabla g(\mathbf{a}) = -\left(\nabla g_0(\mathbf{a}) + (1/N) \sum_{i=1}^N \nabla g_i(\mathbf{a})\right)$$

and then take a small step in this direction. But if N is large, this is unattractive, as we might have to sum a lot of terms. This happens a lot in building classifiers, where you might quite reasonably expect to deal with millions of examples. Touching each example at each step really is impractical.

Instead, assume that, at each step, we choose a number k in the range $1 \dots N$ uniformly and at random, and form

$$\mathbf{p}_k = -(\nabla g_0(\mathbf{a}) + \nabla g_k(\mathbf{a}))$$

and then take a small step along \mathbf{p}_k . Our new point becomes

$$\mathbf{a}^{(i+1)} = \mathbf{a}^{(i)} + \eta \mathbf{p}_k^{(i)},$$

where η is called the **steplength** (even though it very often isn't the length of the step we take!). It is easy to show that

$$\mathbb{E}[\mathbf{p}_k] = \nabla g(\mathbf{a})$$

(where the expectation is over the random choice of k). This implies that if we take many small steps along \mathbf{p}_k , they should average out to a step backwards along the gradient. This approach is known as **stochastic gradient descent** (because we're not going along the gradient, but along a random vector which is the gradient only in expectation). It isn't obvious that stochastic gradient descent is a good idea. Although each step is easy to take, we may need to take more steps. The question is then whether we gain in the increased speed of the step what we lose by having to take more steps. Not much is known theoretically, but in practice the approach is hugely successful for training classifiers.

Choosing a steplength η takes some work. Line search won't work, because we don't want to evaluate the function g , because doing so involves looking at each of the g_i terms. Instead, one uses a steplength that is large at the start — so that it can explore large changes in the values of the classifier parameters — and small steps later — so that it settles down. One useful strategy is to divide training into **epochs**. Each epoch is a block of a fixed number of iterations. Each iteration is one of the steps given above, with fixed steplength. However, the steplength changes from epoch to epoch. In particular, in the r 'th epoch, the steplength is

$$\eta^{(r)} = \frac{a}{r+b}$$

where a and b are constants chosen by experiment with small subsets of the dataset.

One cannot really test whether stochastic gradient descent has converged to the right answer. A better approach is to plot the error as a function of epoch on a validation set. This should vary randomly, but generally go down as the epochs proceed.

14.2.4 Newton's Method

Another way to choose a descent direction is **Newton's method**. We will start with Newton's method in one dimension, then describe it in multiple dimensions. We seek a point a so that $\frac{dg}{da}(a)$ is zero, and we will check the second derivative to find out whether the point we have is a maximum or minimum. Now assume our estimate $a^{(i)}$ is close to the right point. We would like to choose $p^{(i)}$ so that

$$\frac{dg}{da}(a^{(i)} + p^{(i)}) = 0.$$

Now we can expand g in a Taylor series to second order. We have that

$$g(a^{(i)} + p^{(i)}) = g(a^{(i)}) + \left(\frac{dg}{da}(a^{(i)})\right)p^{(i)} + \frac{1}{2}\frac{d^2g}{da^2}(a^{(i)})(p^{(i)})^2 + O((p^{(i)})^3).$$

so that

$$\frac{dg}{da}(a^{(i)} + p^{(i)}) \approx \left(\frac{dg}{da}(a^{(i)})\right) + \frac{d^2g}{da^2}(a^{(i)})(p^{(i)}).$$

Now we want $p^{(i)}$ to have the property that $\frac{dg}{da}(a^{(i)} + p^{(i)}) = 0$. We can write

$$\left(\frac{dg}{da}(a^{(i)})\right) + \frac{d^2g}{da^2}(a^{(i)})(p^{(i)}) = 0$$

so that

$$p^{(i)} = -\frac{\left(\frac{dg}{da}(a^{(i)})\right)}{\frac{d^2g}{da^2}(a^{(i)})}.$$

The same argument works in multiple dimensions, but we need to use vector calculus now. The point we seek has the property that $\nabla g(\mathbf{a}) = \mathbf{0}$. Now assume that our estimate $\mathbf{a}^{(i)}$ is rather close to the right point. We can then write a Taylor series. Write $\mathcal{H}(\mathbf{a}^{(i)})$ for the matrix of second partial derivatives of g , evaluated at $\mathbf{a}^{(i)}$. This is usually called the **Hessian**. The Taylor series for $g(\mathbf{a}^{(i)} + \mathbf{p})$ is then

$$g(\mathbf{a}^{(i)} + \mathbf{p}) \approx g(\mathbf{a}^{(i)}) + \nabla g(\mathbf{a}^{(i)})\mathbf{p} + \frac{1}{2}\mathbf{p}^T\mathcal{H}\mathbf{p}$$

and so we have

$$\nabla g(\mathbf{a}^{(i)} + \mathbf{p}) \approx \nabla g(\mathbf{a}^{(i)}) + \mathcal{H}\mathbf{p}.$$

Now we want $\mathbf{a}^{(i+1)} = \mathbf{a}^{(i)} + \mathbf{p}$ to be the right answer. So we want

$$\nabla g(\mathbf{a}^{(i)}) + \mathcal{H}\mathbf{p} = \mathbf{0}$$

or

$$\mathcal{H}\mathbf{p} = -\nabla g(\mathbf{a}^{(i)}).$$

In the ideal case, this update places $\mathbf{a}^{(i+1)}$ on the solution. This doesn't happen all that often (though it can happen in important cases), and we may need to iterate. When we are close to the right point, Newton's method can converge extremely fast.

However, Newton's method can behave rather badly; it is particularly prone to do so when we start too far from the right position. It is helpful to see why there can be a problem. If we choose \mathbf{p} so that

$$\mathcal{H}\mathbf{p} = -\nabla g(\mathbf{a}^{(i)}),$$

we can look at how $g(\mathbf{a}^{(i)} + h\mathbf{p})$ behaves *as a function of h* . We have

$$\begin{aligned} g(\mathbf{a}^{(i)} + h\mathbf{p}) &\approx g(\mathbf{a}^{(i)}) + h\nabla g(\mathbf{a}^{(i)})\mathbf{p} + O(h^2) \\ &= g(\mathbf{a}^{(i)}) - h(\nabla g(\mathbf{a}^{(i)})^T \mathcal{H}^{-1} \nabla g(\mathbf{a}^{(i)})^T) + O(h^2) \end{aligned}$$

Now imagine that \mathcal{H} is not positive definite; then we could have that

$$\nabla g(\mathbf{a}^{(i)})^T \mathcal{H}^{-1} \nabla g(\mathbf{a}^{(i)})^T$$

is negative; in turn, this would mean that for small values of h , $g(\mathbf{a}^{(i)} + h\mathbf{p})$ will be larger than $g(\mathbf{a}^{(i)})$, so we would be going in the wrong direction. However, if \mathcal{H} is positive definite, there are no difficulties. This is worth knowing both because there is a fix that typical codes apply (they use a matrix that is close to \mathcal{H} *and* positive definite, which they construct using a variety of clever methods) and because problems where \mathcal{H} is positive definite everywhere must display specially good behavior.

Typical codes estimate a \mathbf{p} using a modified form of Newton's method, then apply the line search above. To use Newton's method, you must either allow the code to compute a numerical Hessian (which can be rather slow), or provide the Hessian yourself.

Another tip: in my experience, about 99% of problems with codes that use Newton's method occur because the user didn't check that both the gradient *and* the Hessian their function computed is right. Again, most codes will check a numerical estimate against your value if you want. It's an excellent idea to check.

14.3 LOGISTIC REGRESSION, MINIMIZING LOSS, AND REGULARIZATION

Assume we have a set of N example points \mathbf{x}_i that belong to two classes, which we indicate by 1 and -1 . These points come with their class labels, which we write as y_i ; thus, our dataset can be written as

$$\{(\mathbf{x}_1, y_1), \dots, (\mathbf{x}_N, y_N)\}.$$

We seek a rule that predicts the sign of y for any point \mathbf{x} ; this rule is our classifier.

14.3.1 Logistic Regression

Logistic regression is a classifier that models the class-conditional densities by requiring that

$$\log \frac{p(1|\mathbf{x}, \mathbf{a})}{p(-1|\mathbf{x})} = \mathbf{a}^T \mathbf{x}$$

where \mathbf{a} is a vector of parameters. Assume we know a good value of \mathbf{a} , and the loss for false positives and for false negatives has the same value. Write $\gamma = \mathbf{a}^T \mathbf{x}$. Our

classifier will be:

$$\text{choose } \begin{cases} 1 & \text{if } \gamma > 0 \\ -1 & \text{if } \gamma < 0 \\ \text{randomly} & \text{if } \gamma = 0. \end{cases}$$

The decision boundary here will be a hyperplane passing through the origin of the feature space. Notice that we can turn this into a general hyperplane in the original feature space by extending each example's feature vector by attaching a 1 as the last component of the feature vector \mathbf{x} . This trick simplifies notation, which is why we adopt it here. It is straightforward to estimate \mathbf{a} using maximum likelihood. Note that

$$p(1|\mathbf{x}, \mathbf{a}) = \frac{\exp \mathbf{a}^T \mathbf{x}}{1 + \exp \mathbf{a}^T \mathbf{x}}$$

and

$$p(-1|\mathbf{x}, \mathbf{a}) = \frac{1}{1 + \exp \mathbf{a}^T \mathbf{x}}.$$

We adopt a notational trick. The i 'th example has a label y_i , which is either 1 or -1 . This means that $(y_i + 1)/2$ can be thought of as a switch — the term is zero when $y_i = -1$, and one when $y_i = 1$. We can now write the log-likelihood of a dataset as

$$\begin{aligned} \sum_{i \in \text{data}} \log p(y_i | \mathbf{x}_i, \mathbf{a}) &= \sum_{i \in \text{positive examples}} \log p(1 | \mathbf{x}_i, \mathbf{a}) + \\ &\quad \sum_{i \in \text{negative examples}} \log p(-1 | \mathbf{x}_i, \mathbf{a}) \\ &= \sum_{i \in \text{positive examples}} \{ \mathbf{a}^T \mathbf{x}_i - \log(1 + \exp \mathbf{a}^T \mathbf{x}_i) \} + \\ &\quad \sum_{i \in \text{negative examples}} \{ -\log(1 + \exp \mathbf{a}^T \mathbf{x}_i) \} \\ &= \sum_{i \in \text{examples}} \left\{ \mathbf{a}^T \mathbf{x}_i \frac{(y_i + 1)}{2} - \log(1 + \exp \mathbf{a}^T \mathbf{x}_i) \right\} \end{aligned}$$

Now we could estimate the correct set of parameters $\hat{\mathbf{a}}$ by solving for the maximum of the log-likelihood. In fact, it is convenient to search for the minimum of the negative log-likelihood (which is at the same point). You can think of this as an error function, because we seek the smallest value (and because it is, as we shall see below). We will divide this function by the number of examples. This doesn't change the location of the minimum, but will make for clean notation later. We must choose $\hat{\mathbf{a}}$ to minimize

$$- \left[\sum_{i \in \text{examples}} \frac{1}{N} \left\{ \left(\frac{1 + y_i}{2} \right) \mathbf{a}^T \mathbf{x}_i - \log(1 + \exp \mathbf{a}^T \mathbf{x}_i) \right\} \right].$$

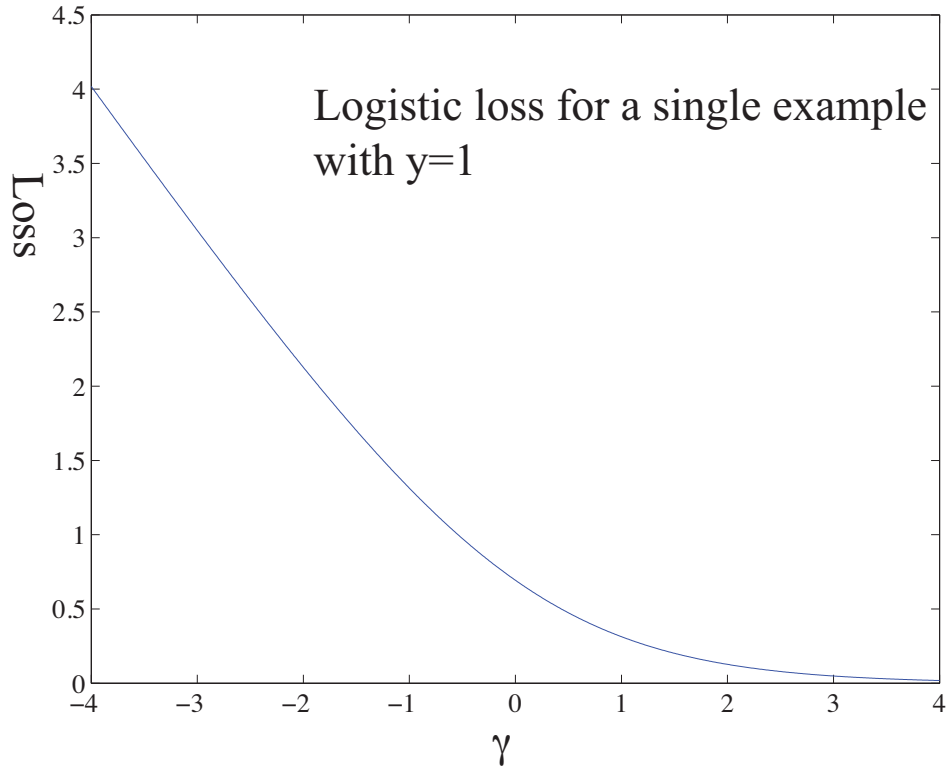


FIGURE 14.4: The logistic loss, plotted for the case $y_i = 1$. In the case of the logistic loss, the horizontal variable is the $\gamma_i = \mathbf{a} \cdot \mathbf{x}_i$ of the text. Notice that giving a strong negative response to this positive example causes a loss that grows linearly as the magnitude of the response grows. Notice also that giving an insufficiently positive response also causes a loss. Giving a strongly positive response is cheap or free.

You can think of the objective function

$$- \left[\sum_{i \in \text{examples}} \frac{1}{N} \left\{ \left(\frac{1 + y_i}{2} \right) \mathbf{a}^T \mathbf{x} - \log (1 + \exp \mathbf{a}^T \mathbf{x}) \right\} \right].$$

as the average, over the examples, of a loss function evaluated at each example. It is easier to see this if we rewrite it. Write $\gamma_i = \mathbf{a}^T \mathbf{x}_i$. Then the term for the i th example is

$$L(y_i, \gamma_i) = - \left\{ \frac{1}{2} (1 + y_i) \gamma_i - \log (1 + \exp \gamma_i) \right\}.$$

Now if $y_i = -1$, we have

$$\begin{aligned} L(y_i, \gamma_i) &= - \{ - \log (1 + \exp \gamma_i) \} \\ &= \{ \log (1 + \exp (-\gamma_i)) \} \end{aligned}$$

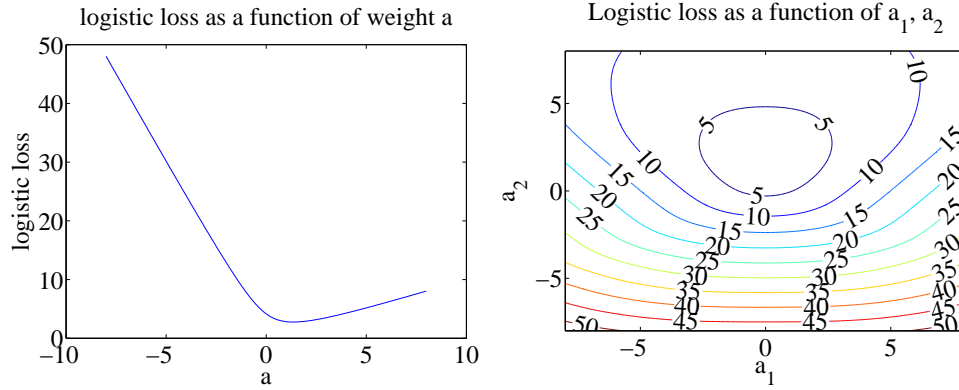


FIGURE 14.5: The logistic loss plotted for the two cases of example 1.

and if $y_i = 1$, we have

$$\begin{aligned}
 L(y_i, \gamma_i) &= -\{\gamma_i - \log(1 + \exp \gamma_i)\} \\
 &= \log(1 + \exp(y_i \gamma_i)) - \gamma_i \\
 &= \log\left(\frac{\exp(-y_i \gamma_i) + 1}{\exp(-y_i \gamma_i)}\right) - \gamma_i \\
 &= \log(\exp(-y_i \gamma_i) + 1) - \log \exp(-y_i \gamma_i) - \gamma_i \\
 &= \{\log(1 + \exp(-y_i \gamma_i))\}
 \end{aligned}$$

This means that we can write the objective function as

$$\frac{1}{N} \sum_{i=1}^N L(y_i, \gamma_i) = \frac{1}{N} \sum_{i=1}^N \log(1 + \exp(-y_i \gamma_i))$$

The function $L(1, \gamma)$ is plotted in Figure 14.4. This loss is sometimes known as the **logistic loss**. This loss very strongly penalizes a large positive γ_i if y_i is negative (and vice versa). However, there is no significant advantage to having a large positive γ_i if y_i is positive. This means that the significant components of the loss function will be due to examples that the classifier gets wrong, but also due to examples that have γ_i near zero (i.e., the example is close to the decision boundary). Maximizing the likelihood is equivalent to minimizing this risk as a function of \mathbf{a} .

Worked example 14.1 *Logistic loss*

There are six points in one dimension, $(-2, -1, -0.5, 0.5, 1, 2)$. The labels are $(-1, -1, 1, -1, 1, 1)$. Plot the logistic loss for this set of datapoints for two cases. In the first case, the feature vector is just the value of the data item. In the second case, you append a 1 (so that for the first point, you would have a feature vector $[-2, 1]$, and so on).

Solution: In the first case, a is one dimensional. Write x_i for the i 'th element of this dataset, and y_i for its label. The total loss is

$$\sum_i \log 1 + \exp(-y_i a x_i)$$

which is a function of a . The classifier associated with each value of a is

$$\text{choose} \begin{cases} 1 & \text{if } ax > 0 \\ -1 & \text{if } ax < 0 \\ \text{randomly} & \text{if } ax = 0. \end{cases}$$

This explains the shape of the loss I have plotted in Figure 14.5. If a is negative, then all but two points have the wrong answer. The loss strongly penalizes large magnitude values of a when a is negative, because in these cases ax has the wrong sign, and is large. If a is positive, then only two points have the wrong answer. As a gets more positive, ax has the wrong sign for these points, and is large; but the cost grows slowly because the other four points are right. In the second case, \mathbf{a} is two dimensional.

$$\sum_i \log 1 + \exp(-y_i \mathbf{a}^T \mathbf{x}_i).$$

This is a function of \mathbf{a} . Write $\mathbf{a} = (a_1, a_2)$; then we have

$$\sum_i \log 1 + \exp(-y_i (a_1 x_i + a_2)).$$

The classifier associated with each value of \mathbf{a} is

$$\text{choose} \begin{cases} 1 & \text{if } a_1 x + a_2 > 0 \\ -1 & \text{if } a_1 x + a_2 < 0 \\ \text{randomly} & \text{if } a_1 x + a_2 = 0. \end{cases}$$

I have plotted the loss as a contour plot (look at `contour`; labels with `clabel`) in Figure 14.5. Each curve is a contour of a fixed height. You should notice how, in each case, there is only a single minimum. The logistic loss is convex.

14.3.2 Regularization

Now assume we want to use Newton's method to minimize the loss for a logistic regression method. The Hessian will be

$$\mathcal{H} = \sum_{i \in \text{examples}} \frac{\exp \gamma_i}{(1 + \exp \gamma_i)^2} \mathbf{x}_i \mathbf{x}_i^T.$$

The Hessian will always be positive semidefinite. It is a weighted sum of terms $\mathbf{x}_i \mathbf{x}_i^T$ which are positive semidefinite, because

$$\mathbf{v}^T (\mathbf{x}_i \mathbf{x}_i^T) \mathbf{v} = (\mathbf{v}^T \mathbf{x}_i) (\mathbf{x}_i^T \mathbf{v}) = (\mathbf{v}^T \mathbf{x}_i)^2 \geq 0.$$

The weights are positive, so the Hessian must be positive semidefinite.

Notice that data points where γ_i has a large absolute value make little contribution to the Hessian—it is affected mainly by points where γ_i is small, that is, points near the boundary. For these points, the Hessian looks like a weighted covariance matrix. If there are enough data points where the γ_i are small, then we expect the Hessian to be positive definite, because we expect that the corresponding data points will span the feature space. But if we have features that are strongly correlated, the Hessian may have one or more small eigenvalues. The easiest way to see this is to use the analogy with a covariance matrix. If the γ_i are small, the weights will be about even and don't matter that much. So the Hessian is basically a covariance. But, because the features are correlated, this covariance matrix will have some small eigenvalues.

These small eigenvalues present a serious nuisance. We would typically maximize using Newton's method, which involves updating an estimate $\mathbf{a}^{(i)}$ by computing $\mathbf{a}^{(i+1)} = \mathbf{a}^{(i)} + \delta \mathbf{a}$. Recall we get the step $\delta \mathbf{a}$ by solving $\mathcal{H}(\delta \mathbf{a}) = -\nabla f$. When \mathcal{H} has small eigenvalues, many different $\delta \mathbf{a}$ will produce acceptable solutions, because there are vectors \mathbf{u} such that $\mathcal{H}\mathbf{u} \approx 0$. This means that a wide range of different $\mathbf{a}^{(i+1)}$ have essentially the same value of loss. In turn, many choices of \mathbf{a} will give about the same loss *on the training data*. The training data offers no reason to choose between these \mathbf{a} .

However, \mathbf{a} with very large norm may behave badly on future test data, because they will tend to produce large values of $\mathbf{a}^T \mathbf{x}$ for test data items \mathbf{x} . In turn, these can produce large losses, particularly if the sign is wrong. This suggests that we should use a value of \mathbf{a} that gives small training loss, and also has a small norm. In turn, this suggests we change the objective function by adding a term that discourages \mathbf{a} with large norm. This term is referred to as a **regularizer**, because it tends to discourage solutions that are large (and so have possible high loss on future test data) but are not strongly supported by the training data. The objective function becomes

Training Loss + Regularizer

which is

$$\text{Training Loss} + \frac{\lambda}{2} (\text{Norm of } \mathbf{a})$$

which is

$$\frac{1}{N} \sum_{i \in \text{examples}} \{ \log(1 + \exp -y_i \mathbf{a}^T \mathbf{x}_i) \} + \frac{\lambda}{2} \mathbf{a}^T \mathbf{a}$$

where $\lambda > 0$ is a constant chosen for good performance, and I divided by 2 to make gradients have a cleaner form. Too large a value of λ , and the classifier will behave poorly on training and test data; too small a value, and the classifier will behave poorly on test data.

Usually, the value of λ is set with a validation dataset. We train classifiers with different values of λ on a test dataset, then evaluate them on a validation set—data whose labels are known, but which is not used for training—and finally choose the λ that gets the best validation error.

Regularizing training loss using the norm is a general recipe, and can be applied to most of the classifiers we describe. For some classifiers, the reasons this approach works are more recondite than those sketched here, but the model here is informative. Norms other than L_2 —that is, $\|x\|_2^2 = \mathbf{x}^T \mathbf{x}$ —can be used successfully. The most commonly used alternative is L_1 —that is, $\|x\|_1 = \sum_i \text{abs}(x_i)$ —which leads to much more intricate minimization problems but strongly encourages zeros in the coefficients of the classifier, which is sometimes desirable.

14.4 THE LINEAR SUPPORT VECTOR MACHINE

Again, assume we have a set of N example points \mathbf{x}_i that belong to two classes, which we indicate by 1 and -1 . These points come with their class labels, which we write as y_i ; thus, our dataset can be written as

$$\{(\mathbf{x}_1, y_1), \dots, (\mathbf{x}_N, y_N)\}.$$

We seek a rule that predicts the sign of y for any point \mathbf{x} ; this rule is our classifier.

14.4.1 Classifying with a Hyperplane

We will use a linear rule, so that for a new data item \mathbf{x} , we will predict

$$\text{sign}((\mathbf{w} \cdot \mathbf{x} + b)).$$

You should think of \mathbf{w} and b as representing a hyperplane which separates the positive data from the negative data. On our training data, we would like to have

$$y_i (\mathbf{w} \cdot \mathbf{x}_i + b) > 0$$

for every example point. Notice the devious use of the sign of y_i — the inequality means that, when y_i is negative (resp. positive), $(\mathbf{w} \cdot \mathbf{x}_i + b)$ is negative (resp. positive). For the moment, we assume that the data is such that we can satisfy all these inequalities. The inequalities identify a *family* of hyperplanes that separate the positive data from the negative data. Each of these hyperplanes must separate the convex hull of one set of examples from the convex hull of the other set of examples. The most conservative choice of hyperplane is the one that is farthest from both hulls. This is obtained by joining the closest points on the two hulls, and constructing a hyperplane perpendicular to this line and through its midpoint.

This hyperplane is as far as possible from each set, in the sense that it maximizes the minimum distance from example points to the hyperplane (Figure 14.6). We need an expression for this distance.

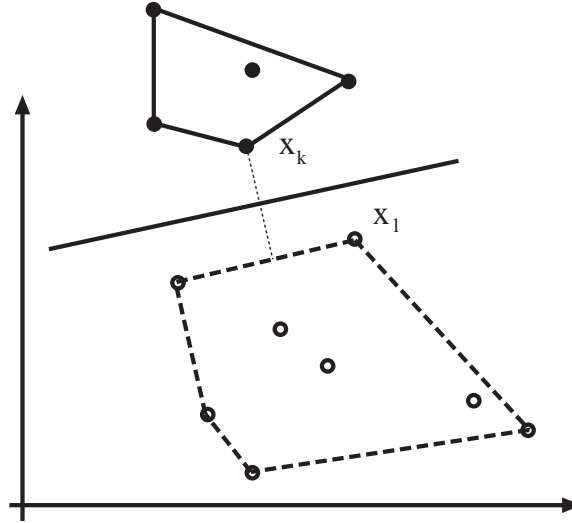


FIGURE 14.6: The hyperplane constructed by a support vector classifier for a plane dataset. The filled circles are data points corresponding to one class, and the empty circles are data points corresponding to the other. We have drawn in the convex hull of each dataset. The most conservative choice of hyperplane is one that maximizes the minimum distance from each hull to the hyperplane. A hyperplane with this property is obtained by constructing the shortest line segment between the hulls and then obtaining a hyperplane perpendicular to this line segment and through its midpoint. Only a subset of the data determines the hyperplane. Of particular interest are points on each convex hull that are associated with a minimum distance between the hulls. We use these points to find the hyperplane in the text.

Now we can choose the scale of \mathbf{w} and b because scaling the two together by a positive number doesn't affect the validity of the constraints $y_i(\mathbf{w} \cdot \mathbf{x}_i + b) > 0$. This means that we can choose \mathbf{w} and b such that for every data point we have

$$y_i(\mathbf{w} \cdot \mathbf{x}_i + b) \geq 1$$

and such that equality is achieved on at least one point on each side of the hyperplane. Now assume that \mathbf{x}_k achieves equality and $y_k = 1$, and \mathbf{x}_l achieves equality and $y_l = -1$. This means that \mathbf{x}_k is on one side of the hyperplane and \mathbf{x}_l is on the other. Furthermore, the distance from \mathbf{x}_l to the hyperplane is minimal (among the points on the same side as \mathbf{x}_l), as is the distance from \mathbf{x}_k to the hyperplane. Notice that there might be several points with these properties.

This means that $\mathbf{w} \cdot (\mathbf{x}_l - \mathbf{x}_k) = 2$, so that

$$\text{dist}(\mathbf{x}_k, \text{hyperplane}) + \text{dist}(\mathbf{x}_l, \text{hyperplane})$$

which is

$$\left(\frac{\mathbf{w}}{|\mathbf{w}|} \cdot \mathbf{x}_k + \frac{b}{|\mathbf{w}|}\right) - \left(\frac{\mathbf{w}}{|\mathbf{w}|} \cdot \mathbf{x}_1 + \frac{b}{|\mathbf{w}|}\right),$$

becomes

$$\frac{\mathbf{w}}{|\mathbf{w}|} \cdot (\mathbf{x}_1 - \mathbf{x}_2) = \frac{2}{|\mathbf{w}|}.$$

This means that maximizing the distance is the same as *minimizing* $(1/2)\mathbf{w} \cdot \mathbf{w}$.

In many cases, we cannot satisfy the inequalities. We introduce a set of **slack variables**, $\xi_i \geq 0$, which represent the amount by which the constraint is violated. We can now write our new constraints as

$$y_i (\mathbf{w} \cdot \mathbf{x}_1 + b) \geq 1 - \xi_i.$$

Equivalently, we could write

$$\xi_i = \max(0, 1 - y_i (\mathbf{w} \cdot \mathbf{x}_1 + b))$$

We would like to choose a \mathbf{w} such that

- $\mathbf{w}^T \mathbf{w}$ is small; and
- $(1/N) \sum_{i=1}^N \xi_i$ is small.

We do not know an appropriate weighting of these two terms. Write λ for some weight, which we will choose with the aid of a validation set. Then we could minimize

$$(1/N) \sum_{i=1}^N \max(0, 1 - y_i (\mathbf{w} \cdot \mathbf{x}_1 + b)) + \frac{\lambda}{2} \mathbf{w}^T \mathbf{w}.$$

14.4.2 The Hinge Loss

The expression

$$(1/N) \sum_{i=1}^N \max(0, 1 - y_i (\mathbf{w} \cdot \mathbf{x}_i + b)) + \frac{\lambda}{2} \mathbf{w}^T \mathbf{w}$$

fits into the rule of section 1, where we obtained a classifier by minimizing

$$\text{Training Loss} + \text{Regularizer}$$

The loss comparing the label value y_i and the prediction $y_i^{(p)} = (\mathbf{w} \cdot \mathbf{x}_i + b)$ can be written as

$$L_h(y_i, y_i^{(p)}) = \max(0, 1 - y_i y_i^{(p)}).$$

This loss is known as the **hinge loss**. It is always non-negative. For the moment, assume $y_i^{(p)} = 1$; then, any prediction by the classifier with value greater than one will incur no loss, and any smaller prediction will incur a cost that is linear in the prediction value (Figure 14.8). This means that minimizing the loss will encourage the classifier to (a) make strong positive (or negative) predictions for positive (or

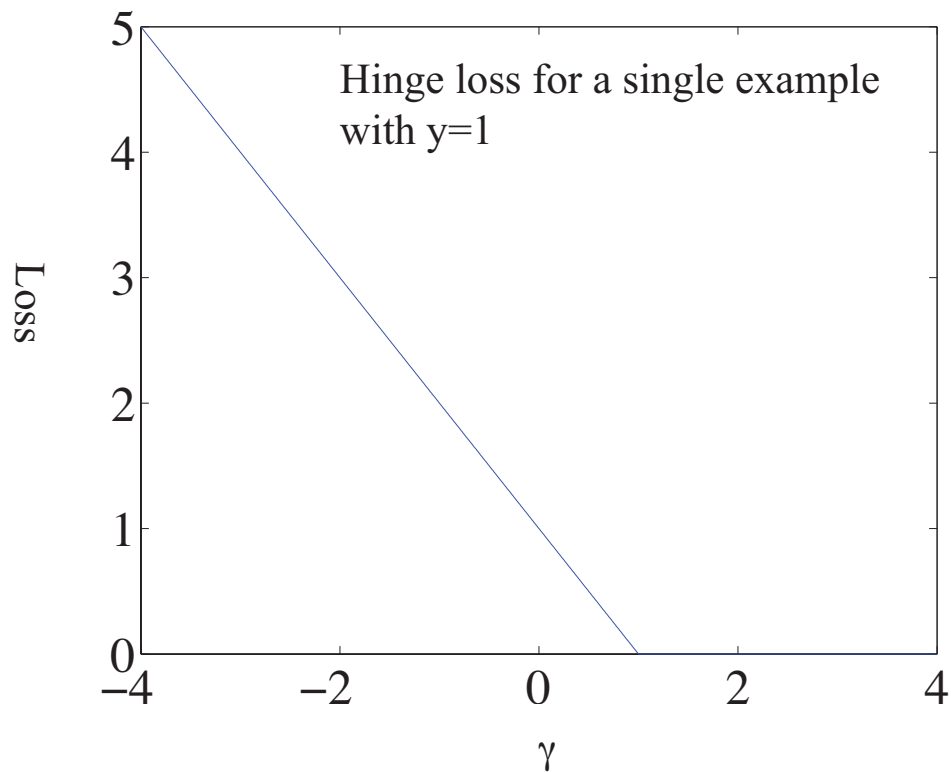


FIGURE 14.7: The hinge loss, plotted for the case $y_i = 1$. The horizontal variable is the $\gamma_i = \mathbf{a} \cdot \mathbf{x}_i$ of the text. Notice that giving a strong negative response to this positive example causes a loss that grows linearly as the magnitude of the response grows. Notice also that giving an insufficiently positive response also causes a loss. Giving a strongly positive response is or free. The loss should look a lot like the hinge loss to you.

negative) examples and (b) for examples it gets wrong, make the most positive (negative) prediction that it can.

Worked example 14.2 *Hinge loss*

There are six points in one dimension, $(-2, -1, -0.5, 0.5, 1, 2)$. The labels are $(-1, -1, 1, -1, 1, 1)$. Plot the hinge loss for this set of datapoints for two cases. In the first case, the feature vector is just the value of the data item. In the second case, you append a 1 (so that for the first point, you would have a feature vector $[-2, 1]$, and so on).

Solution: In the first case, a is one dimensional. Write x_i for the i 'th element of this dataset, and y_i for its label. The total loss is

$$\sum_i \max(0, 1 - y_i a x_i)$$

which is a function of a . I have plotted the loss in Figure 14.8. In the second case, \mathbf{a} is two dimensional.

$$\sum_i \max(0, 1 - y_i \mathbf{a}^T \mathbf{x}_i)$$

This is a function of \mathbf{a} . Write $\mathbf{a} = (a_1, a_2)$; then we have

$$\sum_i \max(0, 1 - y_i (a_1 x_i + a_2))$$

I have plotted the loss as a contour plot (look at `contour`; labels with `clabel`) in Figure 14.8. Each curve is a contour of a fixed height. You should notice how, in each case, there is only a single minimum. The hinge loss is convex. It's also piecewise linear — that is, it looks like a linear function at most points, but there are creases, which you can see as kinks in the contours of the figure.

14.4.3 Example: Training a Support Vector Machine with Stochastic Gradient Descent

We need to choose \mathbf{w} and b to minimize

$$C(\mathbf{w}, b) = (1/N) \sum_{i=1}^N \max(0, 1 - y_i (\mathbf{w} \cdot \mathbf{x}_i + b)) + \frac{\lambda}{2} \mathbf{w}^T \mathbf{w}.$$

There are several methods to do so. Section 1 describes some of the many available support vector machine training packages on the web; it is often, even usually, a good idea to use one of these. But it is worth understanding how such things work.

For a support vector machine, stochastic gradient descent is particularly easy. We have estimates $\mathbf{w}^{(n)}$ and $b^{(n)}$ of the classifier parameters, and we want to improve the estimates. We pick the k 'th example at random. We must now compute

$$\nabla \left(\max(0, 1 - y_k (\mathbf{w} \cdot \mathbf{x}_k + b)) + \frac{\lambda}{2} \mathbf{w}^T \mathbf{w} \right).$$

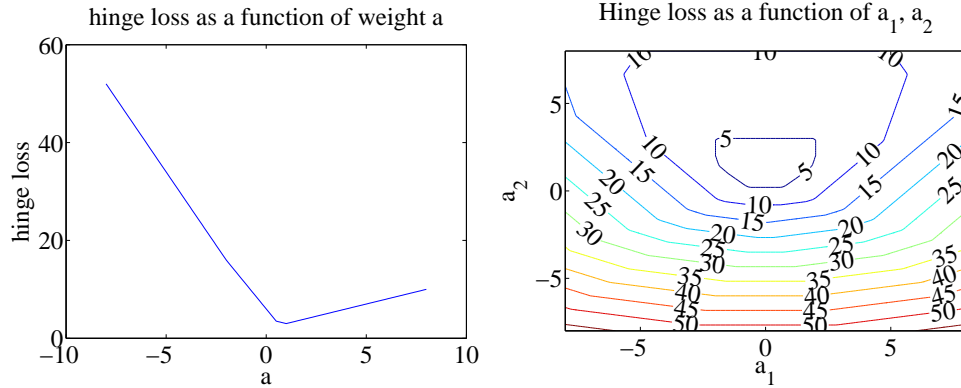


FIGURE 14.8: The hinge loss plotted for the two cases of example 2.

Assume that $y_k (\mathbf{w} \cdot \mathbf{x}_k + b) > 1$. In this case, the classifier predicts a score with the right sign, and a magnitude that is greater than one. Then the first term is zero, and the gradient of the second term is easy. Now if $y_k (\mathbf{w} \cdot \mathbf{x}_k + b) < 1$, we can ignore the max, and the first term is $1 - y_k (\mathbf{w} \cdot \mathbf{x}_k + b)$; the gradient is again easy. But what if $y_k (\mathbf{w} \cdot \mathbf{x}_k + b) = 1$? there are two distinct values we could choose for the gradient, because the max term isn't differentiable. It turns out not to matter which term we choose (Figure ??), so we can write the gradient as

$$p_k = \begin{cases} \begin{bmatrix} \lambda \mathbf{w} \\ 0 \end{bmatrix} & \text{if } y_k (\mathbf{w} \cdot \mathbf{x}_k + b) \geq 1 \\ \begin{bmatrix} \lambda \mathbf{w} - y_k \mathbf{x} \\ -y_k \end{bmatrix} & \text{otherwise} \end{cases}$$

We choose a steplength η , and update our estimates using this gradient. This yields:

$$\mathbf{w}^{(n+1)} = \mathbf{w}^{(n)} - \eta \begin{cases} \lambda \mathbf{w} & \text{if } y_k (\mathbf{w} \cdot \mathbf{x}_k + b) \geq 1 \\ \lambda \mathbf{w} - y_k \mathbf{x} & \text{otherwise} \end{cases}$$

and

$$b^{(n+1)} = b^{(n)} - \eta \begin{cases} 0 & \text{if } y_k (\mathbf{w} \cdot \mathbf{x}_k + b) \geq 1 \\ -y_k & \text{otherwise} \end{cases} .$$

To construct figures, I downloaded the dataset at <http://archive.ics.uci.edu/ml/datasets/Adult>. This dataset apparently contains 48,842 data items, but I worked with only the first 32,000. Each consists of a set of numeric and categorical features describing a person, together with whether their annual income is larger than or smaller than 50K\$. I ignored the categorical features to prepare these figures. This isn't wise if you want a good classifier, but it's fine for an example. I used these features to predict whether income is over or under 50K\$. I split the data into 5,000 test examples, and 27,000 training examples. It's important to do so at random. There are 6 numerical features. I subtracted the mean (which

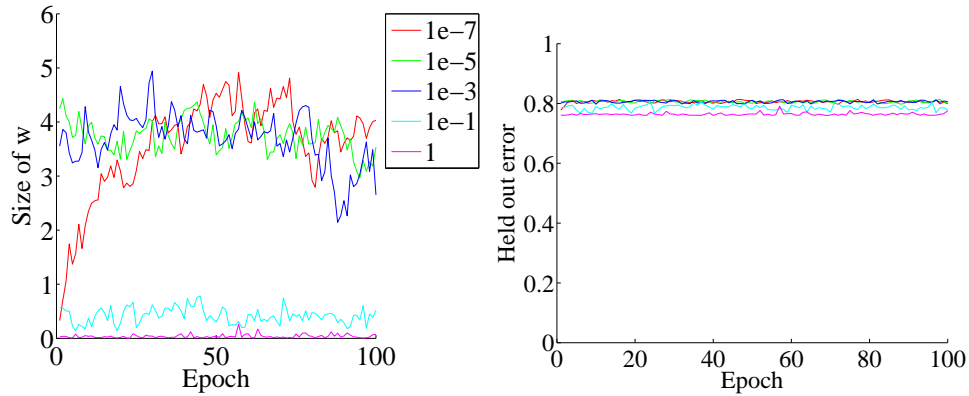


FIGURE 14.9: *On the left*, the magnitude of the weight vector \mathbf{w} at the end of each epoch for the first training regime described in the text. *On the right*, the accuracy on held out data at the end of each epoch. Notice how different choices of regularization parameter lead to different magnitudes of \mathbf{w} ; how the method isn't particularly sensitive to choice of regularization parameter (they change by factors of 100); how the accuracy settles down fairly quickly; and how overlarge values of the regularization parameter do lead to a loss of accuracy.

doesn't usually make much difference) and rescaled each so that the variance was 1 (which is often very important). I used two different training regimes.

In the first training regime, there were 100 epochs. In each epoch, I applied 426 steps. For each step, I selected one data item uniformly at random (sampling with replacement), then stepped down the gradient. This means the method sees a total of 42,600 data items. This means that there is a high probability it has touched each data item once (27,000 isn't enough, because we are sampling with replacement, so some items get seen more than once). I chose 5 different values for the regularization parameter and trained with a steplength of $1/(0.01 * e + 50)$, where e is the epoch. At the end of each epoch, I computed $\mathbf{w}^T \mathbf{w}$ and the accuracy (fraction of examples correctly classified) of the current classifier on the held out test examples. Figure 14.9 shows the results. You should notice that the accuracy changes slightly each epoch; that for larger regularizer values $\mathbf{w}^T \mathbf{w}$ is smaller; and that the accuracy settles down to about 0.8 very quickly.

In the second training regime, there were 100 epochs. In each epoch, I applied 50 steps. For each step, I selected one data item uniformly at random (sampling with replacement), then stepped down the gradient. This means the method sees a total of 5,000 data items, and about 3,216 unique data items — it hasn't seen the whole training set. I chose 5 different values for the regularization parameter and trained with a steplength of $1/(0.01 * e + 50)$, where e is the epoch. At the end of each epoch, I computed $\mathbf{w}^T \mathbf{w}$ and the accuracy (fraction of examples correctly classified) of the current classifier on the held out test examples. Figure 14.10 shows the results. You should notice that the accuracy changes slightly each epoch; that for larger regularizer values $\mathbf{w}^T \mathbf{w}$ is smaller; and that the accuracy settles down

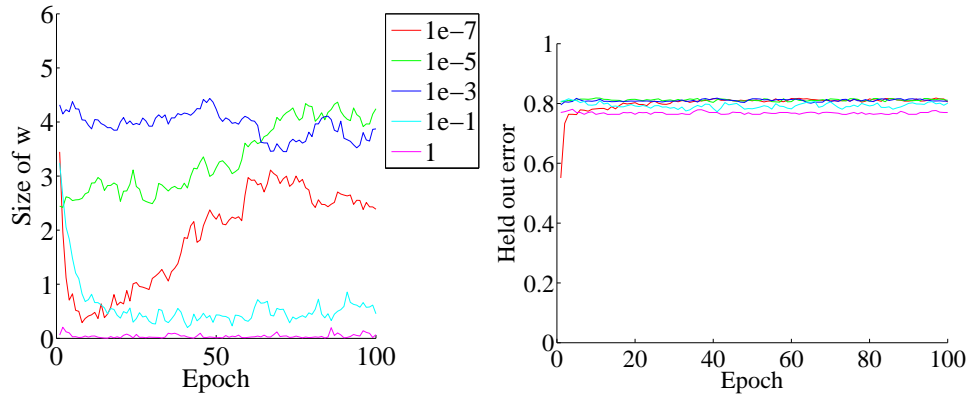


FIGURE 14.10: On the **left**, the magnitude of the weight vector \mathbf{w} at the end of each epoch for the second training regime described in the text. On the **right**, the accuracy on held out data at the end of each epoch. Notice how different choices of regularization parameter lead to different magnitudes of \mathbf{w} ; how the method isn't particularly sensitive to choice of regularization parameter (they change by factors of 100); how the accuracy settles down fairly quickly; and how overlarge values of the regularization parameter do lead to a loss of accuracy.

to about 0.8 very quickly; and that there isn't much difference between the two training regimes. All of these points are relatively typical of stochastic gradient descent with very large datasets.

14.5 PRACTICAL METHODS FOR BUILDING CLASSIFIERS

We have described several apparently very different classifiers here. But which classifier should one use for a particular application? Generally, this should be dealt with as a practical rather than a conceptual question: that is, one tries several, and uses the one that works best. With all that said, experience suggests that the first thing to try for most problems is a linear SVM or logistic regression, which tends to be much the same thing. Nearest neighbor strategies are always useful, and are consistently competitive with other approaches when there is lots of training data and one has some idea of appropriate relative scaling of the features. The main difficulty with nearest neighbors is actually finding the nearest neighbors of a query. Approximate methods are now very good, and are reviewed in Section 14.5.3. The attraction of these methods is that it is relatively easy to build multi-class classifiers, and to add new classes to a system of classifiers.

14.5.1 Manipulating Training Data to Improve Performance

Generally, more training data leads to a better classifier. However, training classifiers with large datasets can be difficult, and it can be hard to get enough training data. Typically, only a relatively small number of example items are really important in determining the behavior of a classifier (we see this phenomenon in greater detail in Section 14.4). The really important examples tend to be rare cases that

are quite hard to discriminate. This is because these cases affect the position of the decision boundary most significantly. We need a large dataset to ensure that these cases are present.

There are some useful tricks that help.

We train on a subset of the examples, run the resulting classifier on the rest of the examples, and then insert the false positives and false negatives into the training set to retrain the classifier. This is because the false positives and false negatives are the cases that give the most information about errors in the configuration of the decision boundaries. We may repeat this several times, and in the final stages, we may use the classifier to seek false positives. For example, we might collect pictures from the Web, classify them, and then look at the positives for errors. This strategy is sometimes called **bootstrapping** (the name is potentially confusing because there is an unrelated statistical procedure known as bootstrapping; nonetheless, we're stuck with it at this point).

There is an extremely important variant of this approach called **hard negative mining**. This applies to situations where we have a moderate supply of positive examples, but an immense number of negative examples. Such situations occur commonly when we use classifiers to detect objects (Section ??). The general procedure is to test every image window to tell whether it contains, say, a face. There are a lot of image windows, and it is quite easy to obtain a lot of images that are certain not to contain a face. In this case we can't use all the negative examples in training, but we need to search for negative examples that are most likely to improve the classifier's performance. We can do so by selecting a set of negative examples, training with these, and then searching the rest of the negative examples to find ones that generate false positives—these are hard negatives. We can iterate the procedure of training and searching for hard negatives; typically, we expand the pool of negative examples at each iteration.

14.5.2 Building Multi-Class Classifiers Out of Binary Classifiers

There are two standard methods to build multi-class classifiers out of binary classifiers. In the **all-vs-all** approach, we train a binary classifier for each pair of classes. To classify an example, we present it to each of these classifiers. Each classifier decides which of two classes the example belongs to, then records a vote for that class. The example gets the class label with the most votes. This approach is simple, but scales very badly with the number of classes.

In the **one-vs-all** approach, we build a binary classifier for each class. This classifier must distinguish its class from all the other classes. We then take the class with the largest classifier score. One possible concern with this method is that training algorithms usually do not compel classifiers to be good at ranking examples. We train classifiers so that they give positive scores for positive examples, and negative scores for negative examples, but we do nothing explicit to ensure that a more positive score means the example is more like the positive class. Another important concern is that the classifier scores must be calibrated to one another, so that when one classifier gives a larger positive score than another, we can be sure that the first classifier is more certain than the second. Some classifiers, such as logistic regression, report posterior probabilities, which require no calibration.

Others, such as the SVM, report numbers with no obvious semantics and need to be calibrated. The usual method to calibrate these numbers is an algorithm due to ?, which uses logistic regression to fit a simple probability model to SVM outputs. One-vs-all methods tend to be reliable and effective even when applied to uncalibrated classifier outputs, most likely because training algorithms do tend to encourage classifiers to rank examples correctly.

Neither strategy is particularly attractive when the number of classes is large, because the number of classifiers we must train scales poorly (linearly in one case, quadratically in the other) with the number of classes. If we were to allocate each class a distinct binary vector, we would need only $\log N$ bits in the vector for N classes. We could then train one classifier for each bit, and we should be able to classify into N classes with only $\log N$ classifiers. This strategy tends to founder on questions of which class should get which bit string, because this choice has significant effects on the ease of training the classifiers. Nonetheless, it gives an argument that suggests that we should not need as many as N classifiers to tell N classes apart.

14.5.3 Finding the Nearest Neighbors and Hashing

To build a nearest neighbors classifier, we need to find the members of a set of high dimensional vectors that are closest to some query vector. It turns out this is a general, and quite difficult, problem. A linear search through the dataset is fine for a small set of data items, but we will operate at scales where we need something more efficient. Surprisingly, exact solutions turn out to be only very slightly more efficient than a linear search through the dataset. Approximate solutions are much more efficient. They are approximate in the sense that, with high probability, they return a point that is almost as close to the query as the closest point. The main trick to obtaining a good approximate solution is to carve the space into cells, then look at items that lie in cells near the query vector; there are two methods that are worth discussing in detail here.

Locality Sensitive Hashing

In **locality sensitive hashing**, we build a set of hash tables containing the data items, using different hashing functions for each table. For a query item, we recover whatever is in each hash table at the location corresponding to the hash code computed for the query item. We search this set, keeping any data items from this set that are sufficiently close to the query. There are many choices of hash function; the most widely used in vision is **random projection**. Write \mathbf{v} for a vector, representing either a query or a data item. We now obtain a single bit of a hash code by choosing a random vector \mathbf{r} and then computing $\text{sign}(\mathbf{v} \cdot \mathbf{r})$. Computing a k -bit hash code involves choosing k such random vectors, then computing one bit for each. There is a set of k such random vectors associated with each hash table. Geometrically, choosing an \mathbf{r} corresponds to choosing a hyperplane in the data space, and the hashing bit corresponds to which side of the hyperplane \mathbf{v} lies on. A k -bit hash code identifies a cell in an arrangement of k hyperplanes in which \mathbf{v} lies. k will be small compared to the dimension, and so we are cutting the space into 2^k cells. This means that there will be relatively few data items that lie in the

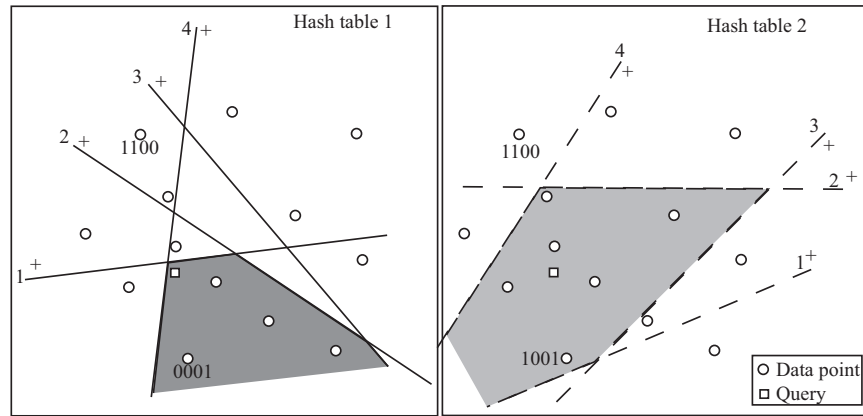


FIGURE 14.11: In locality sensitive hashing using a random projection hash function, the hash function is equivalent to a hyperplane in the data space. Items that lie on one side of the hyperplane corresponding to the n th bit have that bit set to one; otherwise, it is zero. These hyperplanes cut the space of data into a set of cells. All the data items in a cell get a binary hash code (shown for two points in each figure; we have marked the order of the bits by labeling the hyperplanes, and the $+$ s show which side of the hyperplane gets a one). To query, we find all data items in the same hash table entry as the query (the filled polygons in the figure), and then find the closest. However, the nearest neighbor might not be in this cell (for example, the case on the **left**). To reduce the probability of error from this cause, we use more than one hash table and search the union of the sets of points lying in the query cell. In the case illustrated, the nearest neighbor of the query lies in the query cell for the second hash table, on the **right**. The hash tables reduce the set of points we need to search, with high probability of finding a point that is almost as close as the nearest neighbor.

same cell as a query. Some nearby data items may not lie in the same cell, because they could be on the other side of a hyperplane, but these items should lie in the same cell in another hash table.

All these assertions can be made precise, resulting in a guarantee that: (a) a data item that is almost as close as the nearest neighbor will be found with high probability; and (b) all data items closer to the query than some threshold will be found with high probability, whereas data items that are significantly more distant will be found with low probability. Straightforward geometric intuition suggests that this approach will work best when the data items have zero mean, which is easy to arrange. Notice that using n k -bit hash tables is not the same as using one nk -bit hash table. In the first case, the list of points returned from a particular query is a union of the lists returned from each of the n hash tables. This means that points that are near the query but just happen to lie outside the query's cell

for one hash table, have a good chance of being found in another hash table. In the second case, the list we must handle is much shorter (because there are more cells), but there is a better chance of missing nearby points. The choice of n and k will depend on dimension and on the nature of the probabilistic guarantee one wants. There are a variety of other possible choices of hash function. Details of other choices, and precise statements of the relevant guarantees, can be found in (?).

KD-Trees for Approximate Nearest Neighbors

Random projection methods build a cell structure that is independent of the distribution of the data. This means trouble if data is heavily concentrated in some regions, because queries that land in a heavily populated cell of the hash table will need to search a long list. An alternative method is to use a **k-d tree** to build the cell structure. A k-d tree is built by recursively splitting cells. The root will be the whole space. To generate the children of a cell, select one dimension d , perhaps at random, and select some threshold value t_d . Write the d th component of \mathbf{v} as v_d . Now all data items in a cell with $v_d \leq t_d$ are placed in the left child, and all others in the right. We now apply this splitting procedure recursively to the root, until the children are sufficiently small. If we choose the threshold value appropriately (for example, the median of the data in the cell), we can ensure that cells are small in dense components of the space and large in sparse components.

The nearest neighbor to a query can then be found by walking the tree to find the cell containing the query point. We then check any data items in that cell. Write the distance from the query to the closest as d_c . We now backtrack, investigating cells that could contain points closer than d_c and updating d_c when we find a better point. We can prune any branch of the tree that represents a volume that is further from the query than d_c . This procedure works well for low dimensions, but becomes unattractive in high dimensions because we will need to explore too many cells (the number of neighbors of a cell goes up exponentially with dimension).

This difficulty can be avoided if an approximate nearest neighbor is sufficient. In the **best bin first** approach, we look at a fixed number N_c of cells, then report the best point found so far. Promising cells will tend to have some points that are close to the query, and we define the distance between a cell and the query to be the shortest distance from the query to any point on the cell's boundary. Whenever we investigate the child of a cell, we insert the other child into a priority queue, ordered by distance to the query. Once we have checked a cell, we retrieve the next cell from the priority queue. We do this until we have looked at N_c cells. We will look mainly at cells that are close to the query, and so the point we report is a good approximate nearest neighbor.

Good performance of a particular method depends somewhat on the dataset. For most applications, the choice of method can be made offline using the dataset, or a subset of it. ? describe a software package that can choose an approximate nearest neighbors method that is fastest for a particular dataset. Generally, they find that using multiple randomized k-d trees is usually the best; at the time of writing, software could be found at <http://www.cs.ubc.ca/~mariusm/index.php/FLANN/FLANN>.

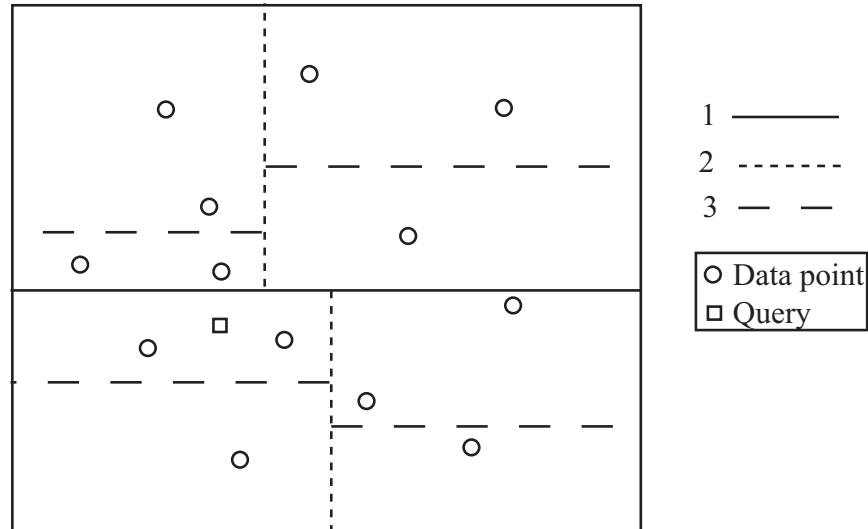


FIGURE 14.12: A k - d tree is built by recursively splitting cells along dimensions. The order in which cells are split for this tree is shown by the dashes on the lines. The nearest neighbor for the query point is found by (a) finding the closest item in the query point's cell, then (b) backtracking, and looking at cells that could contain closer items. Notice that in this example one will need to go right up to the root of the tree and down the other side to find the nearest neighbor. In high dimensions, this backtracking becomes intractable, but if an approximate nearest neighbor is sufficient, the amount of backtracking can be controlled successfully.

14.5.4 Software for SVM's

We obtain a support vector machine by solving one of the constrained optimization problems given above. These problems have quite special structure, and one would usually use one of the many packages available on the web for SVMs to solve them.

LIBSVM (which can be found using Google, or at <http://www.csie.ntu.edu.tw/~cjlin/libsvm/>) is a dual solver that is now widely used; it searches for nonzero Lagrange multipliers using a clever procedure known as SMO (sequential minimal optimization). A good primal solver is PEGASOS; source code can be found using Google, or at <http://www.cs.huji.ac.il/~shais/code/index.html>.

SVMLight (Google, or <http://svmlight.joachims.org/>) is a comprehensive SVM package with numerous features. It can produce sophisticated estimates of the error rate, learn to rank as well as to classify, and copes with hundreds of thousands of examples. Andrea Vedaldi, Manik Varma, Varun Gulshan, and Andrew Zisserman publish code for a multiple kernel learning-based image classifier at <http://www.robots.ox.ac.uk/~vgg/software/MKL/>. Manik Varma publishes code for general multiple-kernel learning at <http://research.microsoft.com/en-us/um/people/manik/code/GMKL/download.html>, and for multiple-kernel learning using SMO at <http://research.microsoft.com/en-us/um/people/manik/code/SMO-MKL/download.html>. Peter Gehler and Sebastian Nowozin publish code for

their recent multiple-kernel learning method at <http://www.vision.ee.ethz.ch/~pgehler/projects/iccv09/index.html>.

14.6 NOTES

We warn readers that a search over classifiers is not a particularly effective way to solve problems; instead, look to improved feature constructions. However, many application problems have special properties, and so there is an enormous number of different methods to build classifiers. We have described methods that reliably give good results. Classification is now a standard topic, with a variety of important textbooks. Accounts of most mainstream classifiers can be found in major recent texts. We recommend ?, ?, ?, and ?. An important theoretical account of when classification methods can be expected to work or to fail is in ?.

Listing 14.1: Matlab code used to train a histogram based classifier for the pizza data

```

[num, txt, raw]=xlsread('~/Current/Courses/Probcourse/SomeData/DataSets/cleanpiz
ndat=size(num, 1);
pdiams=num(:, 5);
labels=zeros(ndat, 1);
for i=1:ndat
    if strcmp(txt(i, 2), 'Dominos')==1
        labels(i)=1;
    else
        labels(i)=-1;
    end
end
% now we do a test-train split
ntrain=floor(0.8*250);
permutev=randperm(ndat);
%% how many boxes in the histogram
nboxes=20;
hn=(1-1e-9)*min(pdiams);
hx=(1+1e-9)*max(pdiams);
hs=(hx-hn)/nboxes;
histd=zeros(nboxes, 1);
histe=zeros(nboxes, 1);
%%
for i=1:ntrain
    pd=pdiams(permutev(i));
    lv=labels(permutev(i));
    hptr=floor((pd-hn)/hs+1);
    if lv>0
        histd(hptr)=histd(hptr)+1;
    else
        histe(hptr)=histe(hptr)+1;
    end
end
histe=histe+1e-9;
histd=histd+1e-9;
pxcd=histd/sum(histd);
pxce=histe/sum(histe);

```

Listing 14.2: Matlab code used to test a histogram based classifier for the pizza data

```

%%
% it's trained. Now test for various thresholds
thresholds=[0, 0.2, 0.4, 0.6, 0.8, 1, 1.2, 1.4, 1.6, 1.8, 2];
fps=zeros(11, 1);
fns=zeros(11, 1);
tps=zeros(11, 1);
tns=zeros(11, 1);
for th=1:11
    thresh=thresholds(th);
    for i=ntrain+1:ndat
        pd=pdiams(permutev(i));
        lv=labels(permutev(i));
        hptr=floor((pd-hn)/hs+1);
        testv=pxcd(hptr)/pxce(hptr);
        if testv>thresh
            prelabel=1;
        else
            prelabel=-1;
        end
        if prelabel>0&&lv>0
            tps(th)=tps(th)+1;
        elseif prelabel>0&&lv<0
            fps(th)=fps(th)+1;
        elseif prelabel<0&&lv>0
            fns(th)=fns(th)+1;
        else
            tns(th)=tns(th)+1;
        end
    end
end
tpr=tps/(ndat-ntrain);
fpr=fps/(ndat-ntrain);
figure(1); clf;
plot(fpr, tpr, 'r');
axis([0 1 0 1])
ylabel('True positive rate or detect rate');
xlabel('False positive rate');
hold on
plot([0, 1], [0, 1], '—r');
cd('~/Current/Courses/Probcourse/Classification/Figures');
print -depsc2 pizzaroc.eps

```

CHAPTER 15

Math Resources

15.1 USEFUL MATERIAL ABOUT MATRICES

Terminology:

- A matrix \mathcal{M} is **symmetric** if $\mathcal{M} = \mathcal{M}^T$. A symmetric matrix is necessarily square.
- We write \mathcal{I} for the identity matrix.
- A matrix is **diagonal** if the only non-zero elements appear on the diagonal. A diagonal matrix is necessarily symmetric.
- A symmetric matrix is **positive semidefinite** if, for any \mathbf{x} such that $\mathbf{x}^T \mathbf{x} > 0$ (i.e. this vector has at least one non-zero component), we have $\mathbf{x}^T \mathcal{M} \mathbf{x} \geq 0$.
- A symmetric matrix is **positive definite** if, for any \mathbf{x} such that $\mathbf{x}^T \mathbf{x} > 0$, we have $\mathbf{x}^T \mathcal{M} \mathbf{x} > 0$.
- A matrix \mathcal{R} is **orthonormal** if $\mathcal{R}^T \mathcal{R} = \mathcal{I} = \mathcal{I}^T = \mathcal{R} \mathcal{R}^T$. Orthonormal matrices are necessarily square.

Orthonormal matrices: You should think of orthonormal matrices as rotations, because they do not change lengths or angles. For \mathbf{x} a vector, \mathcal{R} an orthonormal matrix, and $\mathbf{u} = \mathcal{R}\mathbf{x}$, we have $\mathbf{u}^T \mathbf{u} = \mathbf{x}^T \mathcal{R}^T \mathcal{R} \mathbf{x} = \mathbf{x}^T \mathcal{I} \mathbf{x} = \mathbf{x}^T \mathbf{x}$. This means that \mathcal{R} doesn't change lengths. For \mathbf{y}, \mathbf{z} both unit vectors, we have that the cosine of the angle between them is $\mathbf{y}^T \mathbf{x}$; but, by the same argument as above, the inner product of $\mathcal{R}\mathbf{y}$ and $\mathcal{R}\mathbf{x}$ is the same as $\mathbf{y}^T \mathbf{x}$. This means that \mathcal{R} doesn't change angles, either.

Eigenvectors and Eigenvalues: Assume \mathcal{S} is a $d \times d$ symmetric matrix, \mathbf{v} is a $d \times 1$ vector, and λ is a scalar. If we have

$$\mathcal{S}\mathbf{v} = \lambda\mathbf{v}$$

then \mathbf{v} is referred to as an **eigenvector** of \mathcal{S} and λ is the corresponding **eigenvalue**. Matrices don't have to be symmetric to have eigenvectors and eigenvalues, but the symmetric case is the only one of interest to us.

In the case of a symmetric matrix, the eigenvalues are real numbers, and there are d distinct eigenvectors that are normal to one another, and can be scaled to have unit length. They can be stacked into a matrix $\mathcal{U} = [\mathbf{v}_1, \dots, \mathbf{v}_d]$. This matrix is orthonormal, meaning that $\mathcal{U}^T \mathcal{U} = \mathcal{I}$. This means that there is a diagonal matrix Λ such that

$$\mathcal{S}\mathcal{U} = \mathcal{U}\Lambda.$$

In fact, there is a large number of such matrices, because we can reorder the eigenvectors in the matrix \mathcal{U} , and the equation still holds with a new Λ , obtained by reordering the diagonal elements of the original Λ . There is no reason to keep track of this complexity. Instead, we adopt the convention that the elements of \mathcal{U} are always ordered so that the elements of Λ are sorted along the diagonal, with the largest value coming first.

Diagonalizing a symmetric matrix: This gives us a particularly important procedure. We can convert any symmetric matrix \mathcal{S} to a diagonal form by computing

$$\mathcal{U}^T \mathcal{S} \mathcal{U} = \Lambda.$$

This procedure is referred to as **diagonalizing** a matrix. Again, we assume that the elements of \mathcal{U} are always ordered so that the elements of Λ are sorted along the diagonal, with the largest value coming first. Diagonalization allows us to show that positive definiteness is equivalent to having all positive eigenvalues, and positive semidefiniteness is equivalent to having all non-negative eigenvalues.

Factoring a matrix: Assume that \mathcal{S} is symmetric and positive semidefinite. We have that

$$\mathcal{S} = \mathcal{U} \Lambda \mathcal{U}^T$$

and all the diagonal elements of Λ are non-negative. Now construct a diagonal matrix whose diagonal entries are the positive square roots of the diagonal elements of Λ ; call this matrix $\Lambda^{(1/2)}$. We have $\Lambda^{(1/2)} \Lambda^{(1/2)} = \Lambda$ and $(\Lambda^{(1/2)})^T = \Lambda^{(1/2)}$. Then we have that

$$\mathcal{S} = (\mathcal{U} \Lambda^{(1/2)}) (\Lambda^{(1/2)} \mathcal{U}^T) = (\mathcal{U} \Lambda^{(1/2)}) (\mathcal{U} \Lambda^{(1/2)})^T$$

so we can factor \mathcal{S} into the form $\mathcal{X} \mathcal{X}^T$ by computing the eigenvectors and eigenvalues.

15.1.1 Approximating A Symmetric Matrix

Assume we have a $k \times k$ symmetric matrix \mathcal{T} , and we wish to construct a matrix \mathcal{A} that approximates it. We require that (a) the rank of \mathcal{A} is precisely $r < k$ and (b) the approximation should minimize the **Frobenius norm**, that is,

$$\|(\mathcal{T} - \mathcal{A})\|_F^2 = \sum_{ij} (T_{ij} - A_{ij})^2.$$

It turns out that there is a straightforward construction that yields \mathcal{A} .

The first step is to notice that if \mathcal{U} is orthonormal and \mathcal{M} is any matrix, then

$$\|\mathcal{U} \mathcal{M}\|_F = \|\mathcal{M} \mathcal{U}\|_F = \|\mathcal{M}\|_F.$$

This is true because \mathcal{U} is a rotation (as is $\mathcal{U}^T = \mathcal{U}^{-1}$), and rotations do not change the length of vectors. So, for example, if we write \mathcal{M} as a table of row vectors $\mathcal{M} = [\mathbf{m}_1, \mathbf{m}_2, \dots, \mathbf{m}_k]$, then $\mathcal{U} \mathcal{M} = [\mathcal{U} \mathbf{m}_1, \mathcal{U} \mathbf{m}_2, \dots, \mathcal{U} \mathbf{m}_k]$. Now $\|\mathcal{M}\|_F^2 = \sum_{j=1}^k \|\mathbf{m}_j\|^2$, so $\|\mathcal{U} \mathcal{M}\|_F^2 = \sum_{i=1}^k \|\mathcal{U} \mathbf{m}_k\|^2$. But rotations do not change lengths, so $\|\mathcal{U} \mathbf{m}_k\|^2 = \|\mathbf{m}_k\|^2$, and so $\|\mathcal{U} \mathcal{M}\|_F = \|\mathcal{M}\|_F$. To see the result for the case of $\mathcal{M} \mathcal{U}$, just think of \mathcal{M} as a table of row vectors.

Notice that, if \mathcal{U} is the orthonormal matrix whose columns are eigenvectors of \mathcal{T} , then we have

$$\|(\mathcal{T} - \mathcal{A})\|_F^2 = \|\mathcal{U}^T(\mathcal{T} - \mathcal{A})\mathcal{U}\|_F^2.$$

Now write Λ_r for $\mathcal{U}^T \mathcal{A} \mathcal{U}$, and Λ for the diagonal matrix of eigenvalues of \mathcal{T} . Then we have

$$\|(\mathcal{T} - \mathcal{A})\|_F^2 = \|\Lambda - \Lambda_A\|_F^2,$$

an expression that is easy to solve for Λ_A . We know that Λ is diagonal, so the best Λ_A is diagonal, too. The rank of \mathcal{A} must be r , so the rank of Λ_A must be r as well. To get the best Λ_A , we keep the r largest diagonal values of Λ , and set the rest to zero; Λ_A has rank r because it has only r non-zero entries on the diagonal, and every other entry is zero.

Now to recover \mathcal{A} from Λ_A , we know that $\mathcal{U}^T \mathcal{U} = \mathcal{U} \mathcal{U}^T = \mathcal{I}$ (remember, \mathcal{I} is the identity). We have $\Lambda_A = \mathcal{U}^T \mathcal{A} \mathcal{U}$, so

$$\mathcal{A} = \mathcal{U} \Lambda_A \mathcal{U}^T.$$

We can clean up this representation in a useful way. Notice that only the first r columns of \mathcal{U} (and the corresponding rows of \mathcal{U}^T) contribute to \mathcal{A} . The remaining $k - r$ are each multiplied by one of the zeros on the diagonal of Λ_A . Remember that, by convention, Λ was sorted so that the diagonal values are in descending order (i.e. the largest value is in the top left corner). We now keep only the top left $r \times r$ block of Λ_A , which we write Λ_r . We then write \mathcal{U}_r for the $k \times r$ matrix consisting of the first r columns of \mathcal{U} . Then

$$\mathcal{A} = \mathcal{U}_r \Lambda_r \mathcal{U}_r^T$$

This is so useful a result, I have displayed it in a box; you should remember it.

Procedure: 15.1 *Approximating a symmetric matrix with a low rank matrix*

Assume we have a symmetric $k \times k$ matrix \mathcal{T} . We wish to approximate \mathcal{T} with a matrix \mathcal{A} that has rank $r < k$. Write \mathcal{U} for the matrix whose columns are eigenvectors of \mathcal{T} , and Λ for the diagonal matrix of eigenvalues of \mathcal{A} (so $\mathcal{A} \mathcal{U} = \mathcal{U} \Lambda$). Remember that, by convention, Λ was sorted so that the diagonal values are in descending order (i.e. the largest value is in the top left corner).

Now construct Λ_r from Λ by setting the $k - r$ smallest values of Λ to zero, and keeping only the top left $r \times r$ block. Construct \mathcal{U}_r , the $k \times r$ matrix consisting of the first r columns of \mathcal{U} . Then

$$\mathcal{A} = \mathcal{U}_r \Lambda_r \mathcal{U}_r^T$$

is the best possible rank r approximation to \mathcal{T} in the Frobenius norm.

Now if \mathcal{A} is positive semidefinite (i.e. if at least the r largest eigenvalues of \mathcal{T} are non-negative), then we can factor \mathcal{A} as in the previous section. This yields a procedure to approximate a symmetric matrix by factors. This is so useful a result, I have displayed it in a box; you should remember it.

Procedure: 15.2 *Approximating a symmetric matrix with low dimensional factors*

Assume we have a symmetric $k \times k$ matrix \mathcal{T} . We wish to approximate \mathcal{T} with a matrix \mathcal{A} that has rank $r < k$. We assume that at least the r largest eigenvalues of \mathcal{T} are non-negative. Write \mathcal{U} for the matrix whose columns are eigenvectors of \mathcal{T} , and Λ for the diagonal matrix of eigenvalues of \mathcal{A} (so $\mathcal{A}\mathcal{U} = \mathcal{U}\Lambda$). Remember that, by convention, Λ was sorted so that the diagonal values are in descending order (i.e. the largest value is in the top left corner).

Now construct Λ_r from Λ by setting the $k - r$ smallest values of Λ to zero and keeping only the top left $r \times r$ block. Construct $\Lambda_r^{(1/2)}$ by replacing each diagonal element of Λ with its positive square root. Construct \mathcal{U}_r , the $k \times r$ matrix consisting of the first r columns of \mathcal{U} . Then write $\mathcal{V} = (\mathcal{U}_r \Lambda_r^{(1/2)})$

$$\mathcal{A} = \mathcal{V}\mathcal{V}^T$$

is the best possible rank r approximation to \mathcal{T} in the Frobenius norm.